
Search for gravitational-wave burst (GWB)
counterparts to gamma-ray bursts (GRB)
using data from the fifth LIGO science run
and the first Virgo science run

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For the LIGO Scientific Collaboration
and the Virgo Collaboration

LIGO-G0900392-v2

GRB-GWB triggered burst search with LIGO-Virgo

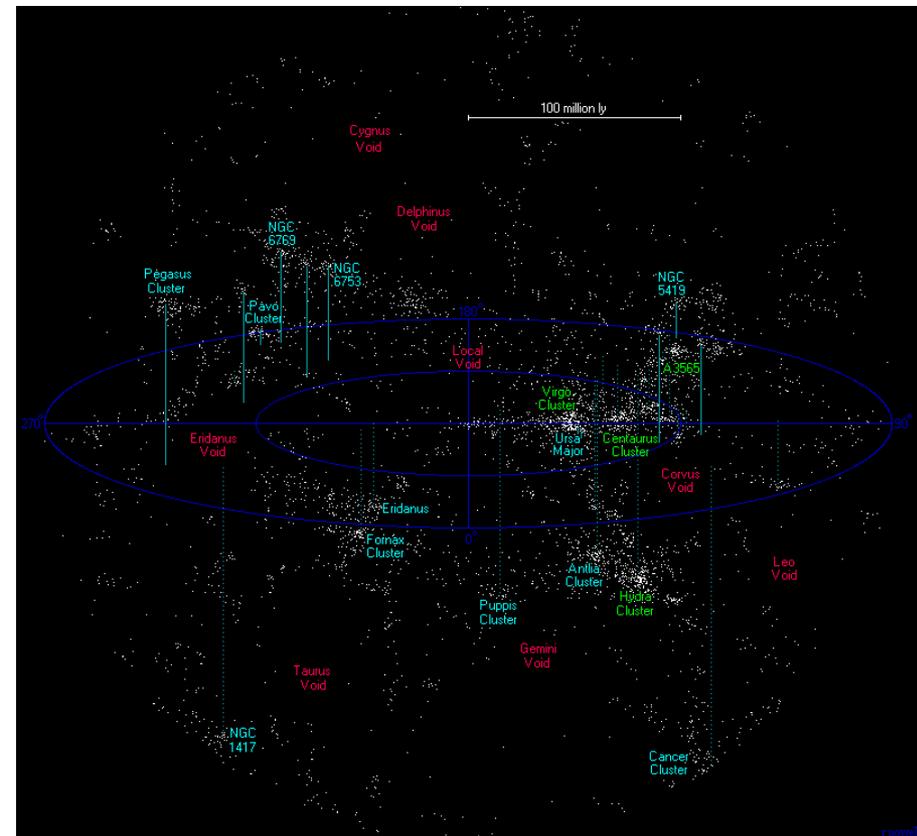


- ❖ GRBs are at cosmological distances ($z \sim 1$)
 - ❖ short GRBs nearer
- ❖ volume of sensitivity depends on strain amplitude sensitivity of interferometers

$$V \propto D_{\text{reach}}^3 \propto h_{\text{IFO}}^{-1/3}$$

- ❖ special GRBs, e.g. GRB 070201
- ❖ LIGO-Virgo S5/VS1 run
 - ❖ 2 years total run time; ~200 GRB triggers, mostly from Swift
- ❖ enhanced LIGO-Virgo
 - ❖ GRB triggers mostly from Fermi+Swift; factor of 3 increase in trigger rate

initial LIGO-Virgo: $D \sim 50$ Mpc
thousands of large galaxies,
hundreds of galaxy groups



Source: www.atlasoftheuniverse.com

GRB-GWB triggered burst search



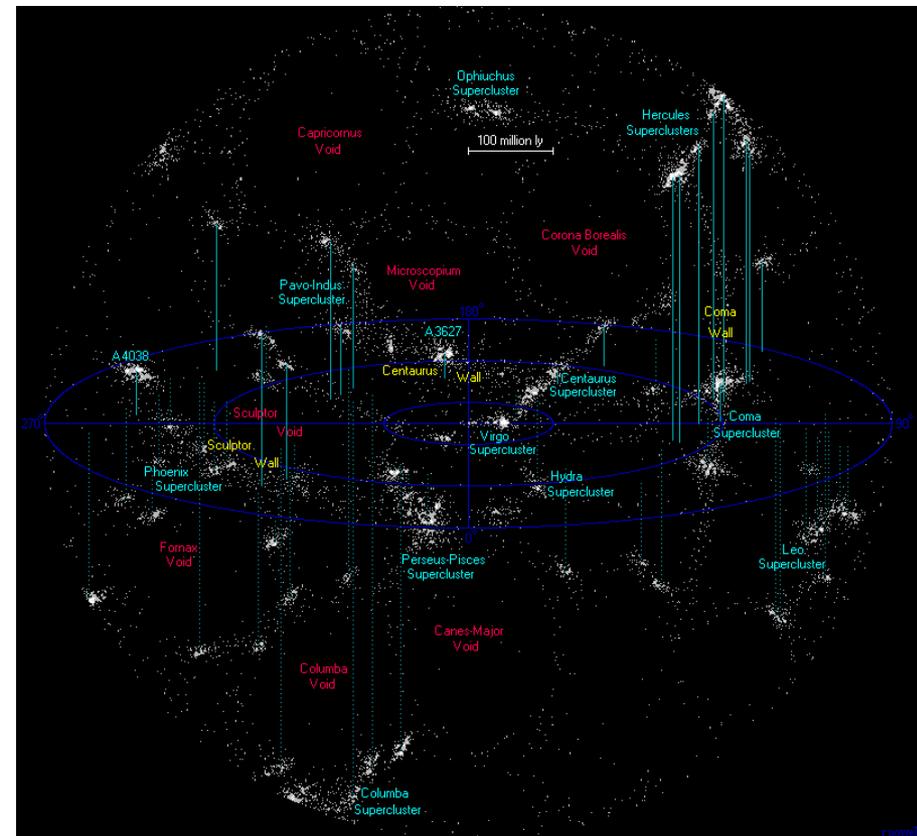
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enhanced LIGO-Virgo: $D \sim 100$ Mpc
several hundred thousand large galaxies,
dozens of galaxy clusters



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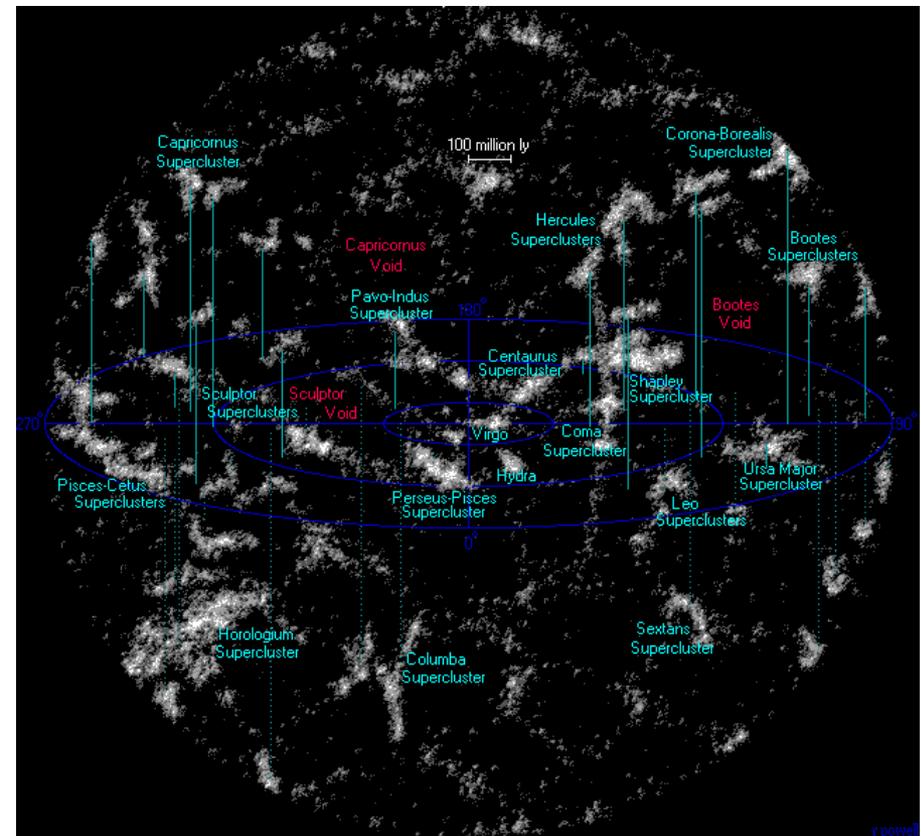


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advanced LIGO-Virgo: $D \sim 500$ Mpc
millions of large galaxies,
hundreds of superclusters



Source: www.atlasoftheuniverse.com

GRB-GWB triggered burst search with LIGO-Virgo



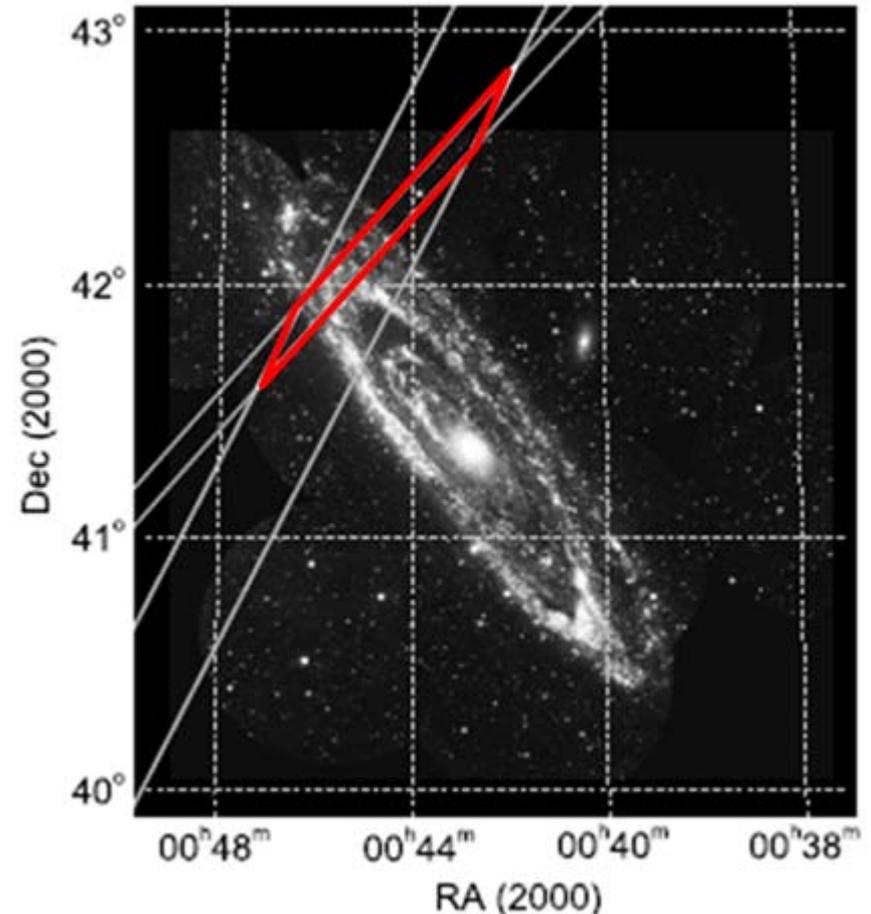
(arXiv:0712.1502)

GRB 070201 error box overlapped
spiral arm of Andromeda galaxy

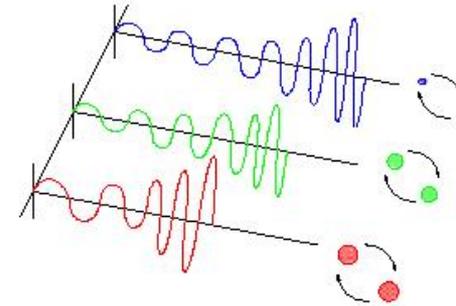
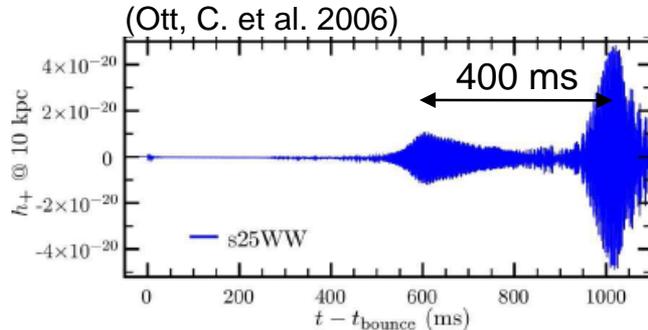
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Methods of searching for gravitational-wave (GW) counterparts to GRBs



Search for GW burst signals

- ❖ model-independent
- ❖ cross-correlate different data streams; fully coherent network search
- ❖ target GW burst signals less than ~few seconds
- ❖ **used to search for GW counterpart to long and short GRBs**

Search for GW inspiral signals

- ❖ makes use of inspiral templates
- ❖ cross-correlate data stream with inspiral templates
- ❖ target GW inspiral signals from coalescing masses in the range $1 M_{\odot} < m_1 < 3 M_{\odot}$, $1 M_{\odot} < m_2 < 40 M_{\odot}$
- ❖ **used to search for GW counterpart to short GRBs**

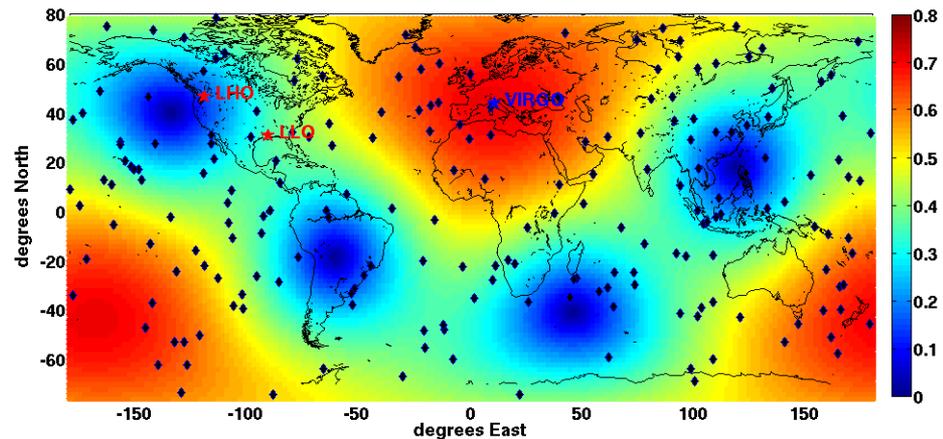
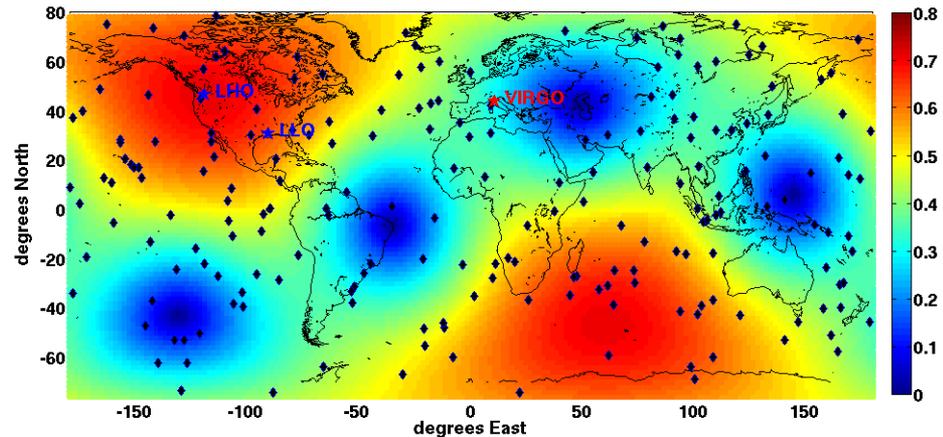
The GRB sample for the LIGO S5/VSR1 run



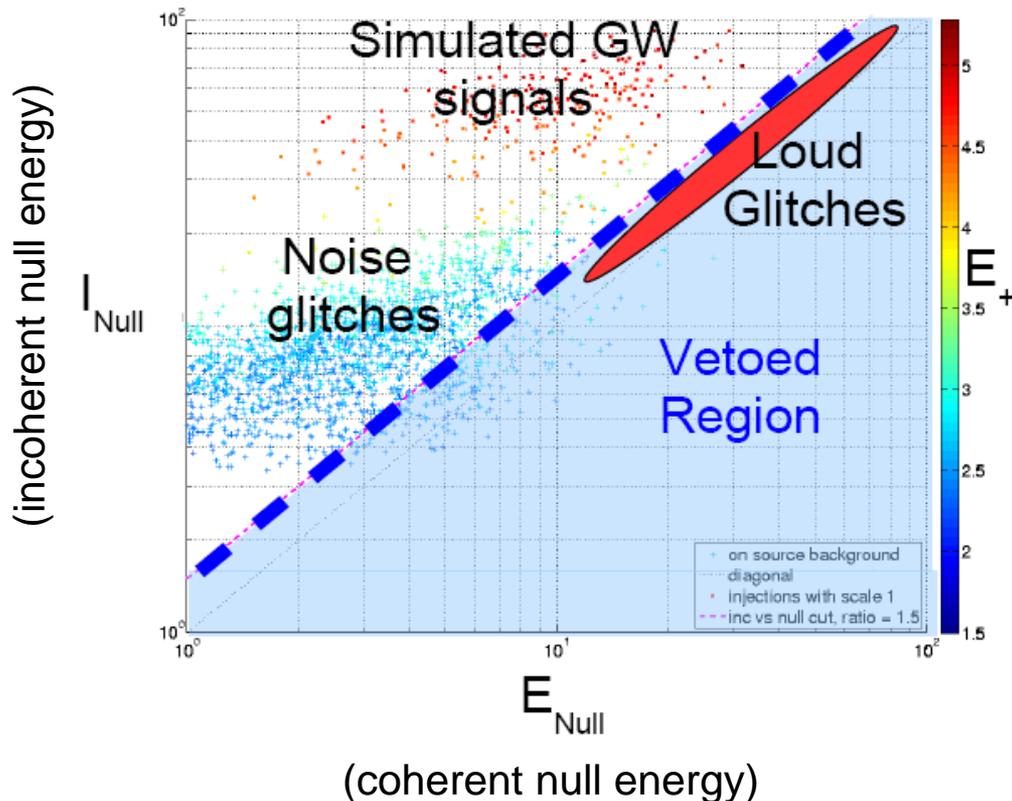
- ❖ **212 GRB triggers** from Nov. 4, 2005 to Oct. 1, 2007
 - ❖ **~70%** with double-IFO coincidence LIGO data
 - ❖ **~45%** with triple-IFO coincidence LIGO data
 - ❖ **~15%** short-duration GRBs
 - ❖ **~25%** with redshift
 - ❖ **~20%** fall in joint LIGO-Virgo times
 - ❖ all but a handful have accurate position information

burst analysis has been completed; paper due out soon

GRB triggers were mostly from Swift;
some were from IPN3, INTEGRAL, HETE-2



Current burst search algorithm: fully coherent network search



- ❖ takes into account different antenna responses of detectors to source at a given sky location
- ❖ coherently adds multiple data streams from a network of interferometers to maximize SNR
- ❖ reconstructs h_+ and h_x components of signal
- ❖ null stream can be used for consistency tests, i.e. better background rejection

S5/VSR1 GRB-GWB burst search sensitivity



- ❖ assuming energy emitted in GW

$$E_{\text{GW}} = 0.1 M_{\text{sun}}, \quad f_0 = 150 \text{ Hz}$$

$$E_{\text{GW}} \approx \frac{\pi^2 c^3}{G} D^2 f_0^2 h_{\text{rss}}^2$$

$$\Rightarrow D \sim 50 \left(\frac{E_{\text{GW}}}{0.1 M_{\text{sun}}} \right)^{1/2} \text{ Mpc}$$

- ❖ estimated GWB detection rate
(best estimates; unknown uncertainties; takes into account satellite detector efficiencies)

$$R_{\text{GW}} \sim \rho_{\text{GRB}} \left(\frac{4}{3} \pi D^3 \right)$$

long GRBs:

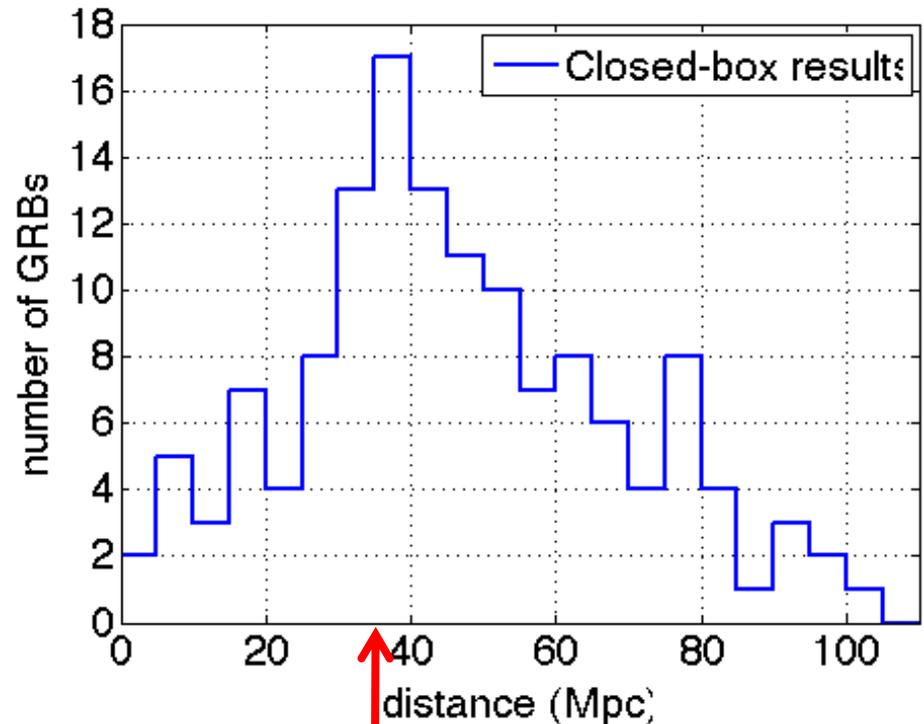
$$\rho_{\text{GRB}} \sim 0.5 \text{ Gpc}^{-3} \text{ yr}^{-1}$$

$$R_{\text{GW}} \sim 5 \times 10^{-5} \text{ yr}^{-1}$$

short GRBs:

$$\rho_{\text{GRB}} \sim [8 - 30] \text{ Gpc}^{-3} \text{ yr}^{-1}$$

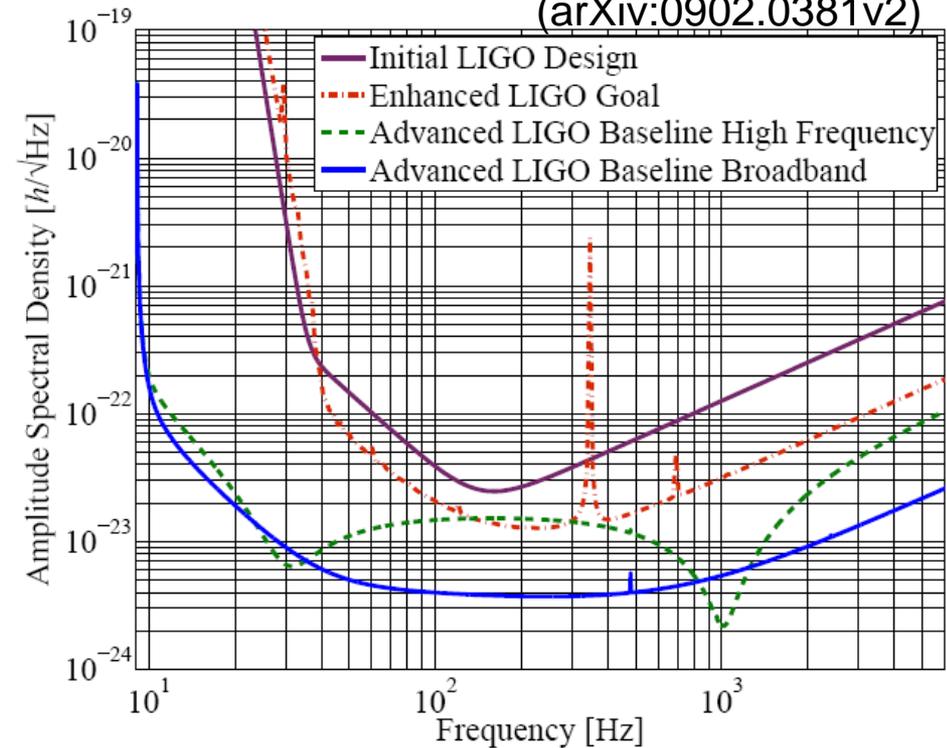
$$R_{\text{GW}} \sim [0.5 - 2] \times 10^{-3} \text{ yr}^{-1}$$



GRB 980425/SN 1998bw
at 36 Mpc

Prospects for enhanced LIGO-Virgo and advanced LIGO-Virgo

- ❖ enhanced LIGO-Virgo (S6/VSR2 run)
 - ❖ factor of 2 increase in amplitude sensitivity vs. initial LIGO-Virgo
 - ❖ will commence summer of 2009
- ❖ advanced LIGO-Virgo
 - ❖ factor of 10 increase in sensitivity vs. initial LIGO-Virgo

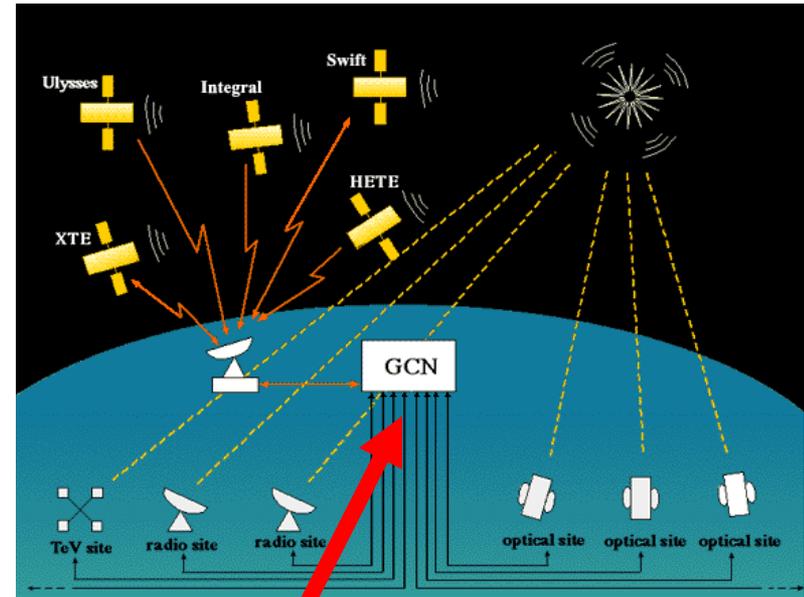


$$E_{\text{GW}} = 0.1 M_{\text{sun}}, f_0 = 150 \text{ Hz}$$

| Detector | Estimated distance reach | Estimated detection rate (yr^{-1}) | |
|-------------------------------|--------------------------|---|--------------------|
| | | Long GRBs | Short GRBs |
| Enhanced LIGO-Virgo (S6/VSR2) | ~ 100 Mpc | ~ 0.0021 | $\sim 0.03 - 0.13$ |
| Advanced LIGO-Virgo | ~ 500 Mpc | ~ 0.26 | $\sim 4 - 16$ |

S6/VSR2 low-latency GRB-GWB triggered burst search

- ❖ plan to have analysis results within a few hours of GRB trigger
- ❖ availability of results within ~hours means we can contribute in timely manner to discussion of interesting GRBs
- ❖ for interesting GRBs, disseminate results to science community within ~week
- ❖ expect GRB trigger rate of ~1 per day (high-confidence GRBs)



- ❖ fully coherent LIGO-Virgo analysis for searching for GW counterparts to GRBs is in place and working very well
- ❖ S5/VSR1 GRB-GWB search paper under review and will be out soon
- ❖ anticipating enhanced LIGO-Virgo S6/VSR2 run
 - ❖ factor of ~ 2 increase in amplitude sensitivity (~ 8 increase in volume sensitivity)
 - ❖ GRB trigger rate of ~ 1 per day from Fermi+Swift+other satellites
 - ❖ low-latency search (results \sim hours after GRB trigger)