



# GRB-triggered searches for gravitational waves from compact binary inspirals in LIGO and Virgo data during S5/VSR I

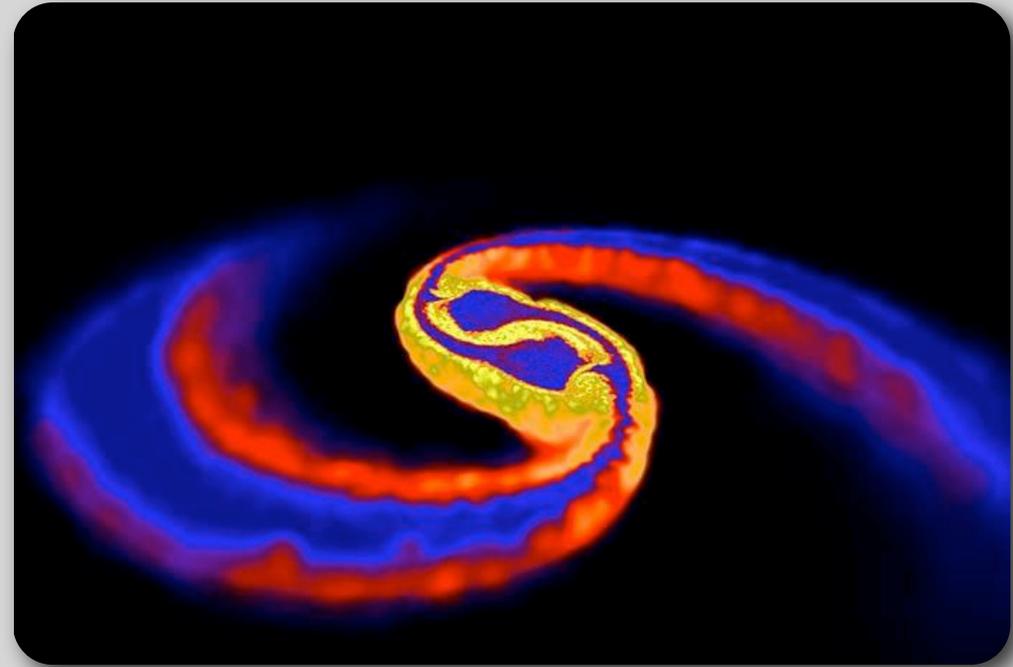
Nickolas Fotopoulos (UWM) for the [LIGO](#) Scientific Collaboration and the Virgo Collaboration

8th Edoardo Amaldi meeting, Columbia University, New York, NY | 2009.06.22

# Short GRBs: ideal targets for GW astronomy (I)

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- Most short GRBs are probably NSs disrupted by compact companions in the final stages of inspiral.
- A detection will constrain component masses and spins.\*
- A high-SNR detection will constrain NS equations of state.†
- Simultaneous EM/GW observations can measure absolute luminosity distance.‡



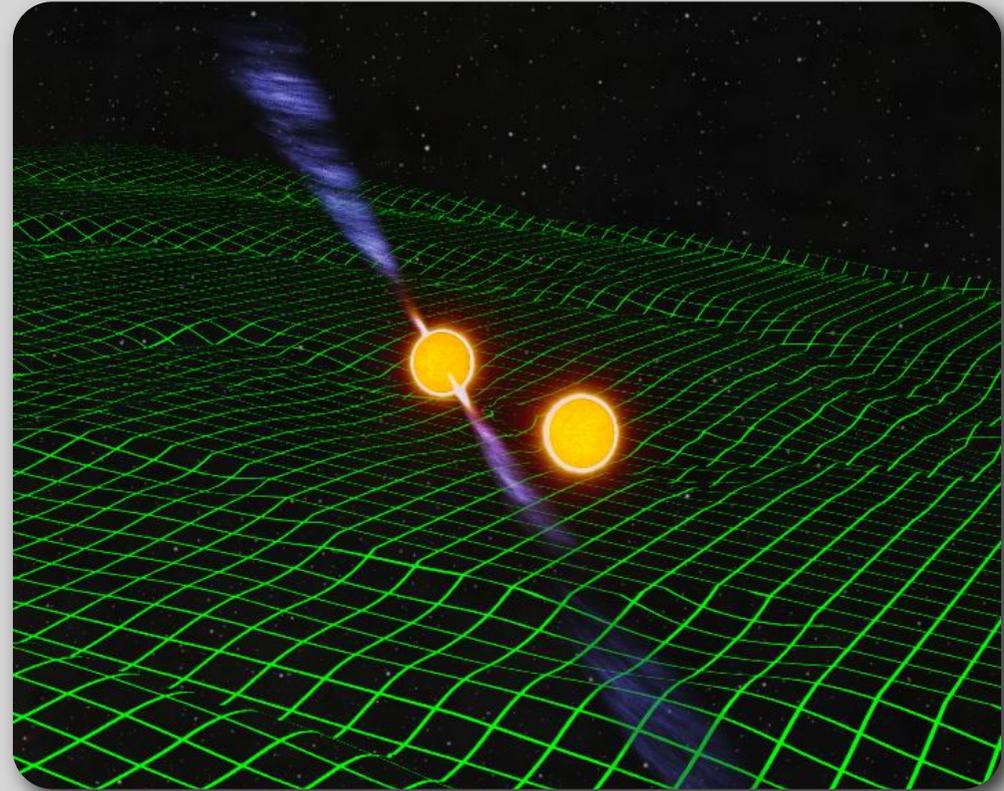
NS-NS merger simulation  
Price and Rosswog

\* Cutler and Flanagan, PRD 49, 2658 (1994);  
Finn and Chernoff, PRD 47, 2198 (1993);  
Poisson and Will, PRD 52, 848 (1995)  
† Flanagan and Hinderer, PRD 77, 021502 (2008);  
Read et al, arXiv:0901.3258  
‡ Nisanke et al, arXiv:0904.1017

# Short GRBs: ideal targets for GW astronomy (II)

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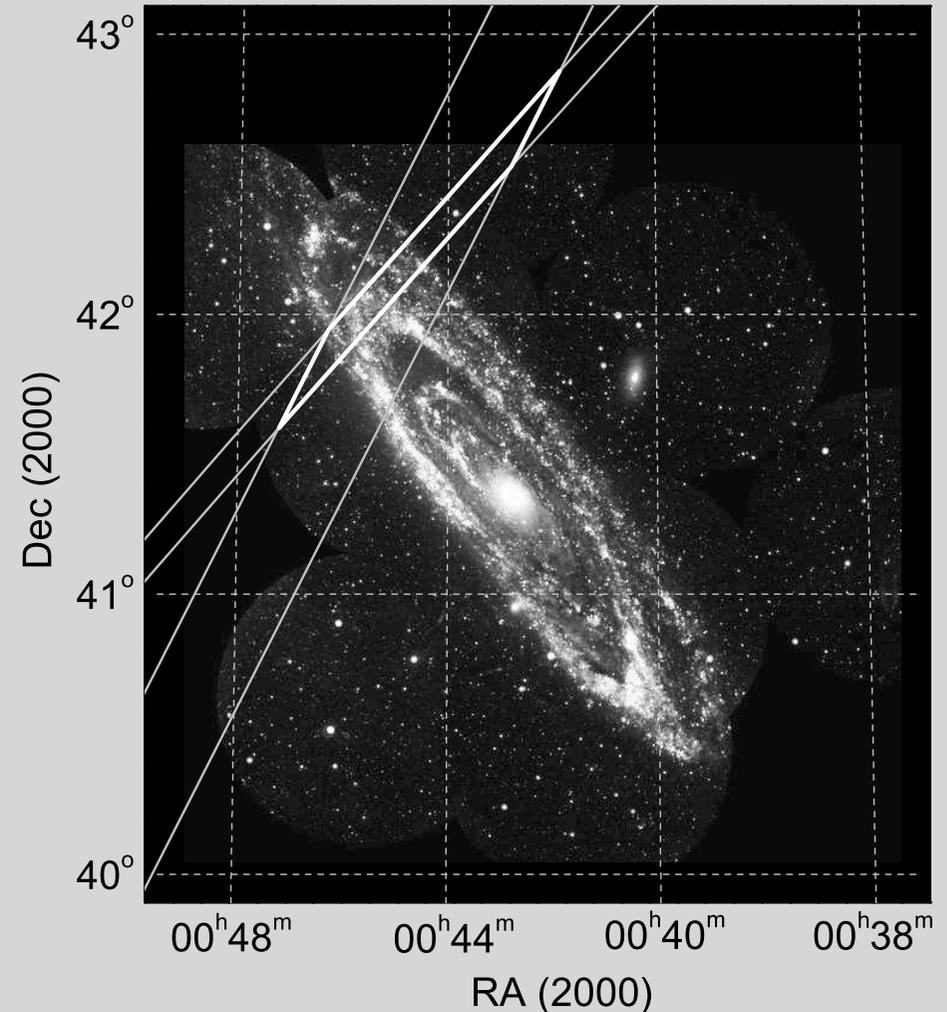
- A significant GW candidate with an EM counterpart is a far more compelling detection.
- A known time and sky location can be searched with significantly lowered thresholds.
- The GW emission during inspiral is well modeled. This enables matched filtering, which digs more deeply into the detector noise than unmodeled searches.



NS-NS inspiral depiction  
John Rowe Animation

# GRB 070201: not an inspiral in M31

- GRB 070201 occurred in the direction of M31, the Andromeda galaxy.
- M31 is  $\sim 770$  kpc away, well within LIGO's range.
- LIGO observations ruled out an inspiral progenitor in M31 at  $>99\%$  confidence.\* They allow a soft gamma repeater (SGR) progenitor.†
- The present search has lower thresholds and algorithmic improvements.



IPN  $3\sigma$  Sky Localization  
Mazets et al, ApJ 680, 545 (2008)

\* Abbott et al, ApJ 681, 1419 (2008)

† Ofek et al, ApJ 681, 1464 (2008);  
Mazets et al, ApJ 680, 545 (2008)

# S5: Nov 2005 – Nov 2007 | VSR I: May 2007 – Oct 2007

- 212 GRBs
- 33 short GRBs
- 22 short\* GRBs while two+ GW detectors were taking good data (duty cycle, data quality)

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051210	070429B
051211	070512
060121	070707
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060427B	070714B
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Already published: no  
inspiral in M3 I



**070201**



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Long duration but other suggestive features

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Analysis done differently;  
Three-detector analysis

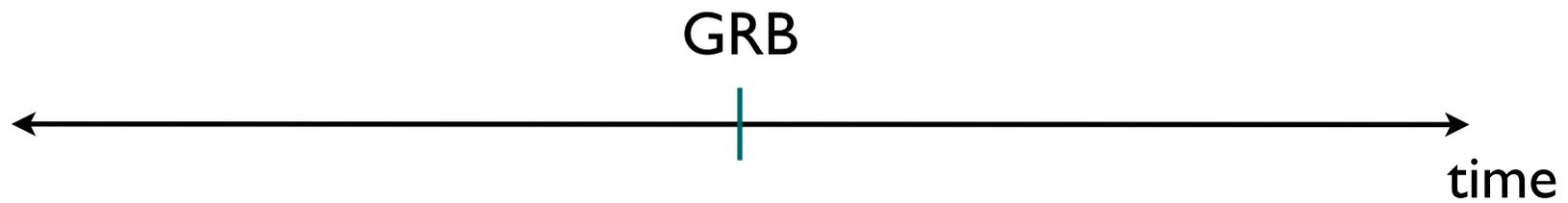


**070923**



# Experiment overview

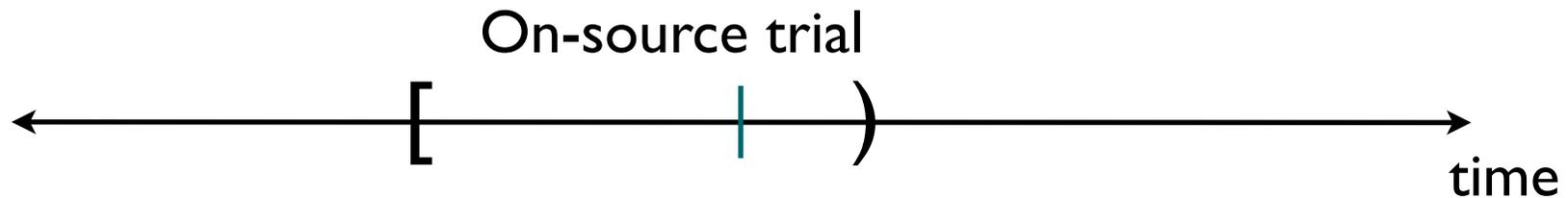
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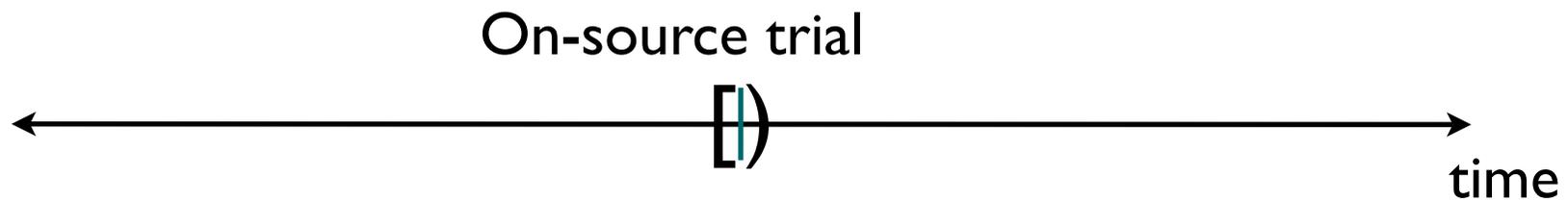
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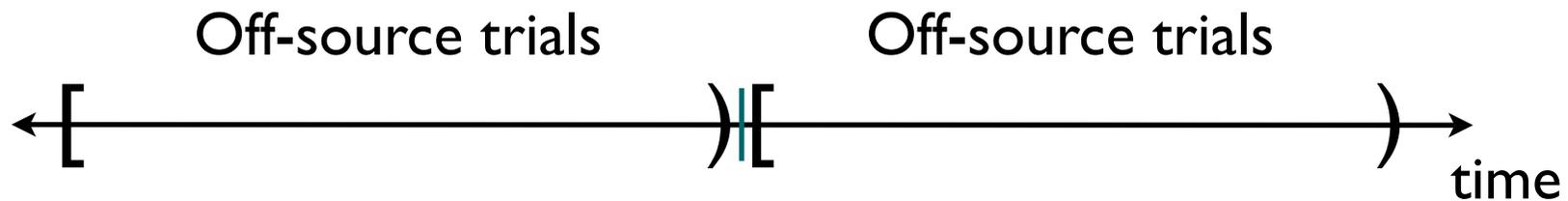
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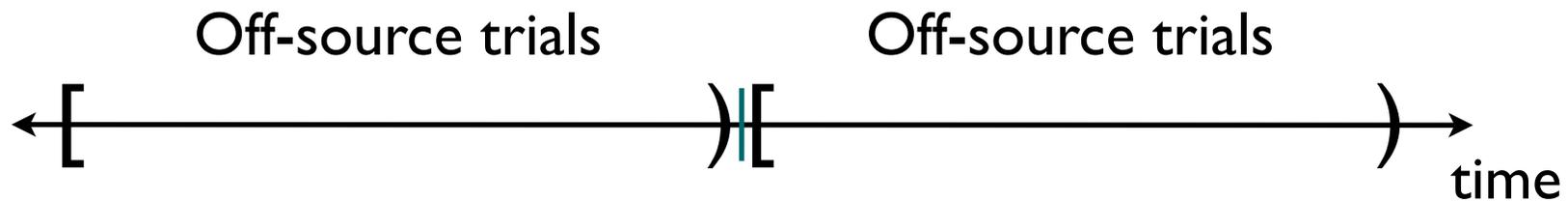
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- To estimate background, we analyze  $\sim 40$  minutes of nearby **off-source trials**.



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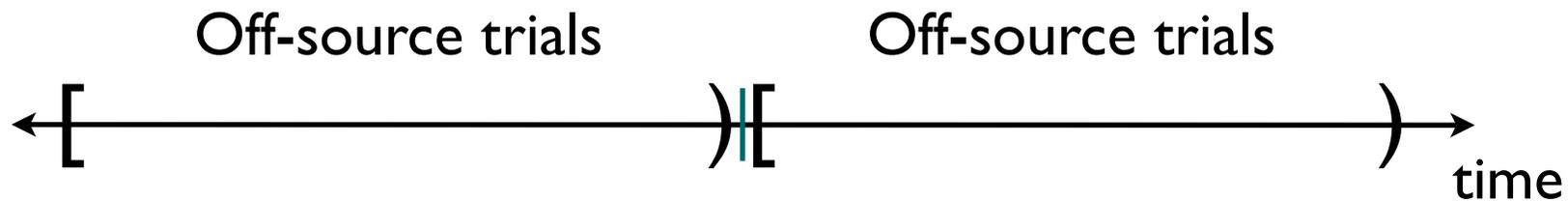
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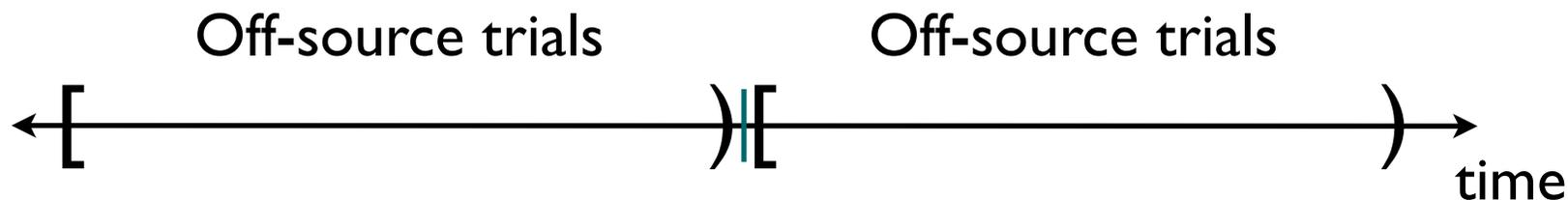


\* Abbott et al, PRD 73 (2006) 062001

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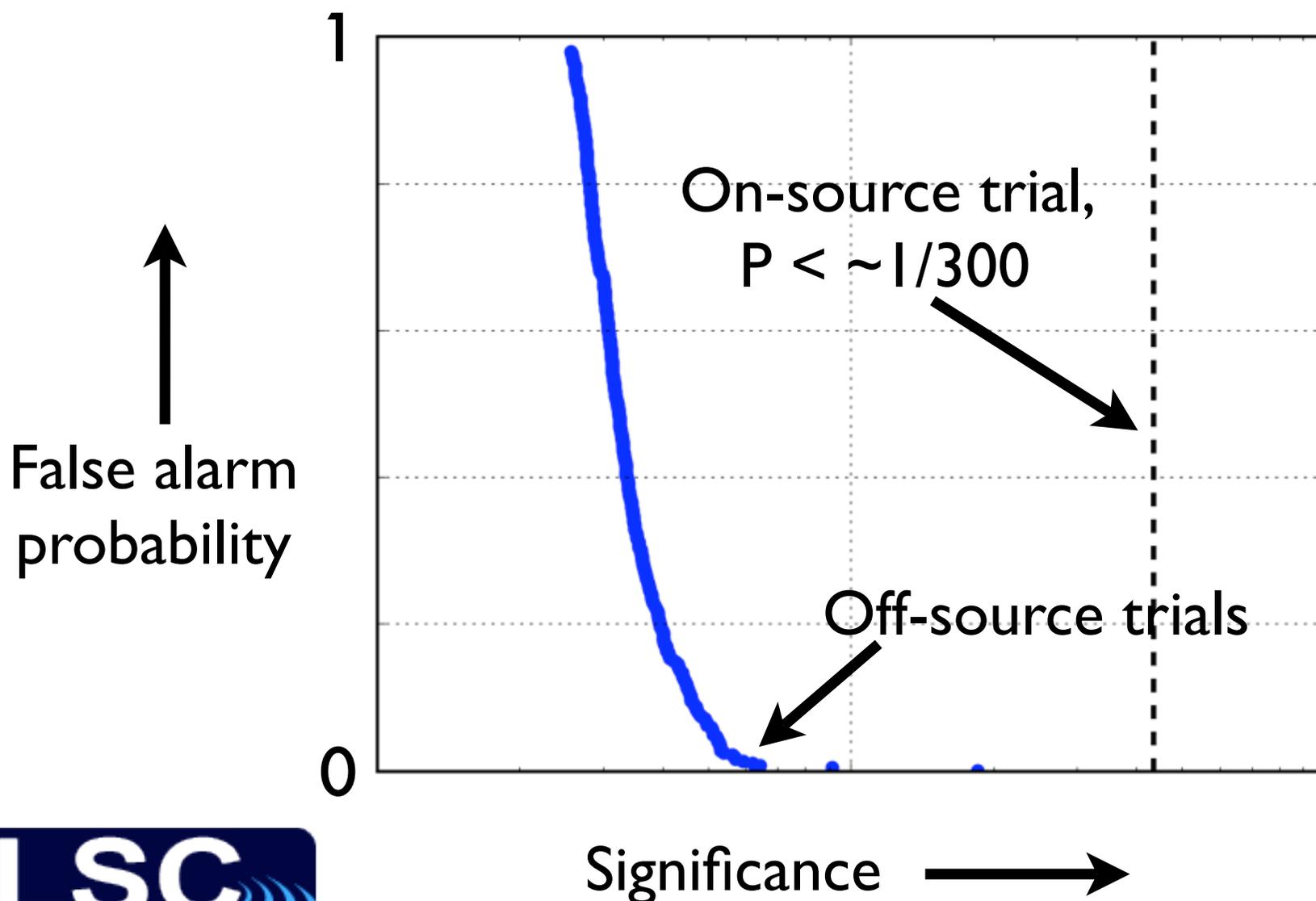


- We **reuse** the hierarchical inspiral search pipeline used in previous LIGO analyses.\*
- We combine injection and off-source trials to form a **likelihood statistic**.

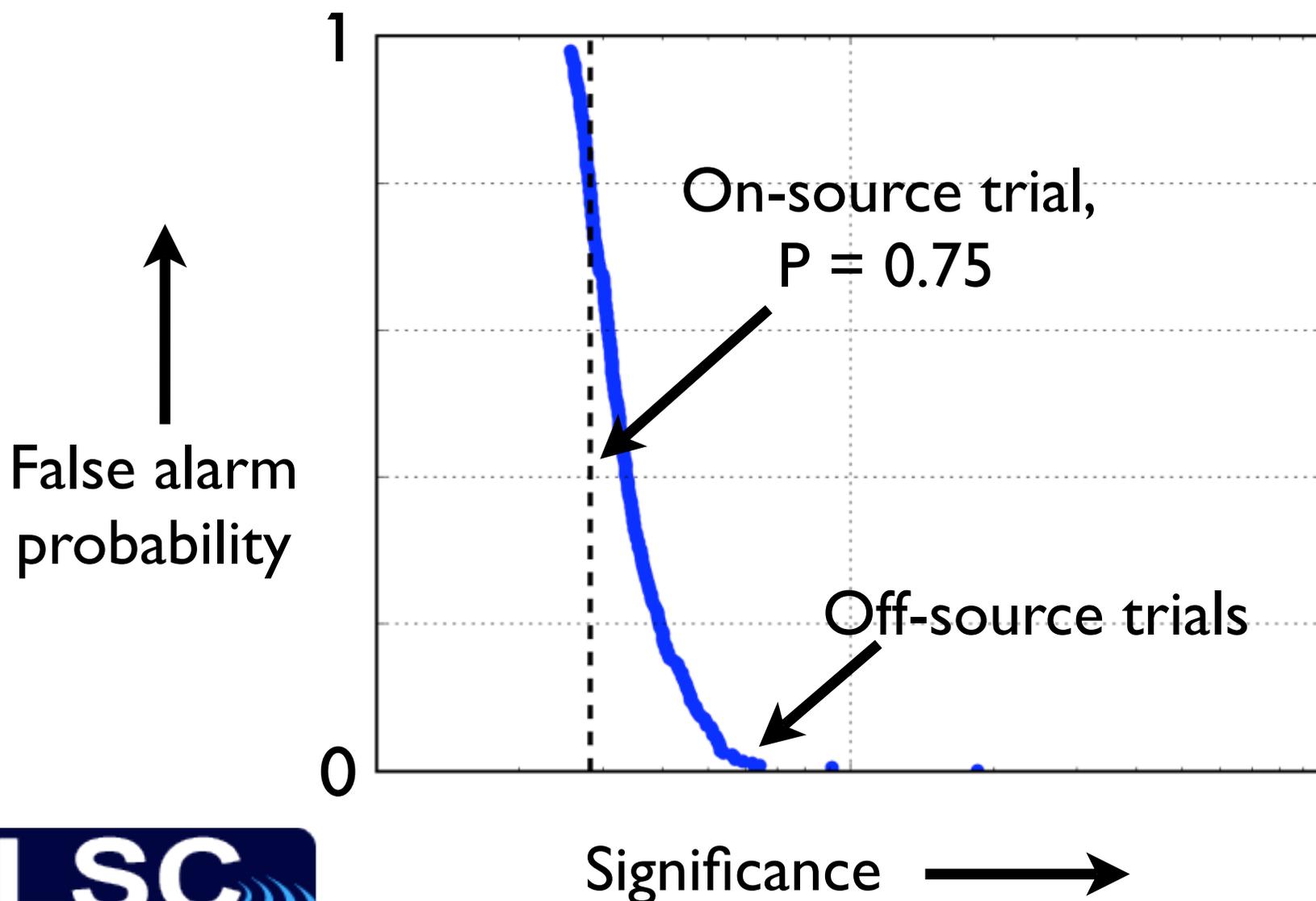


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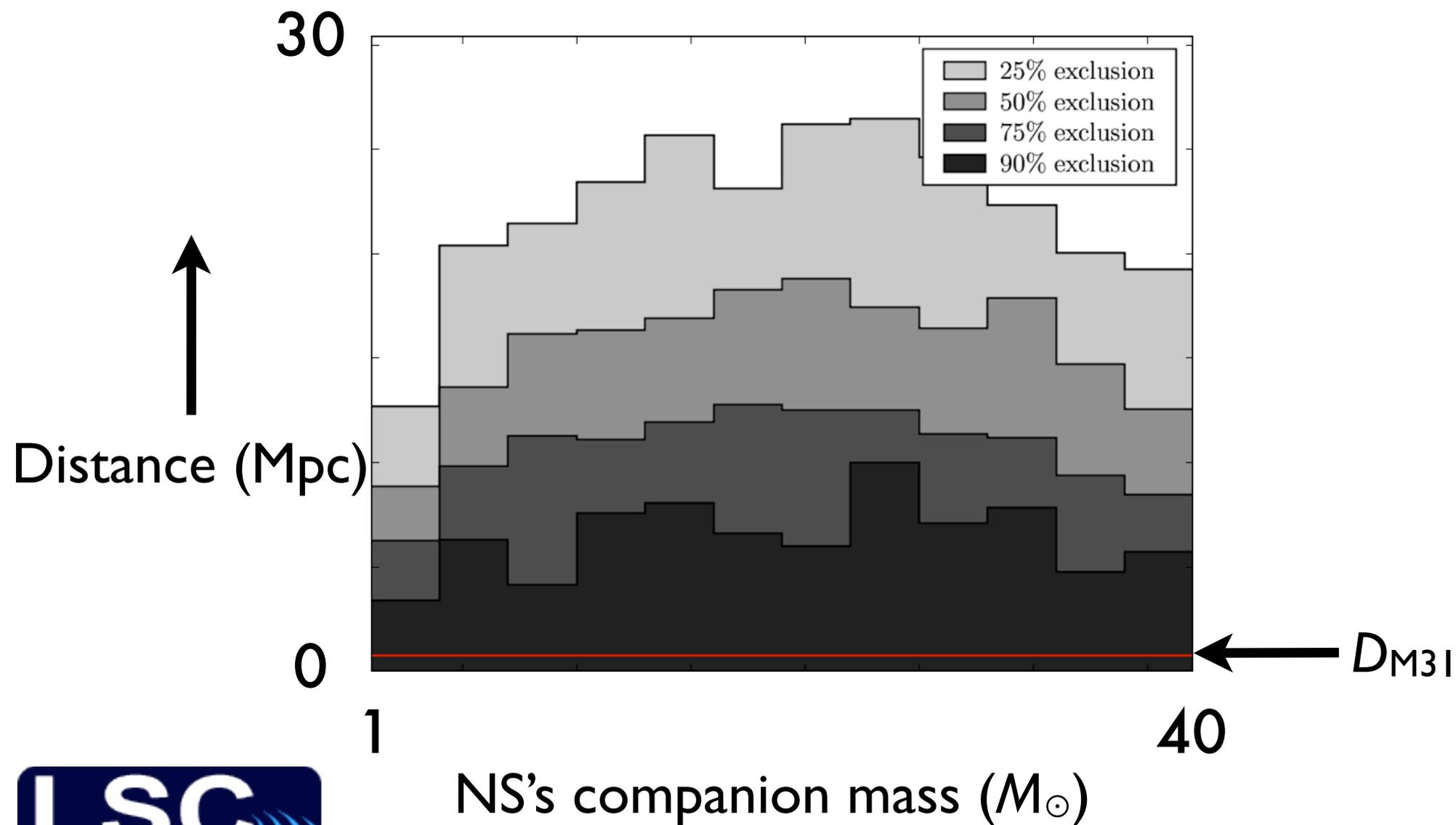
# What a **detection** might look like



# What a **null result** might look like



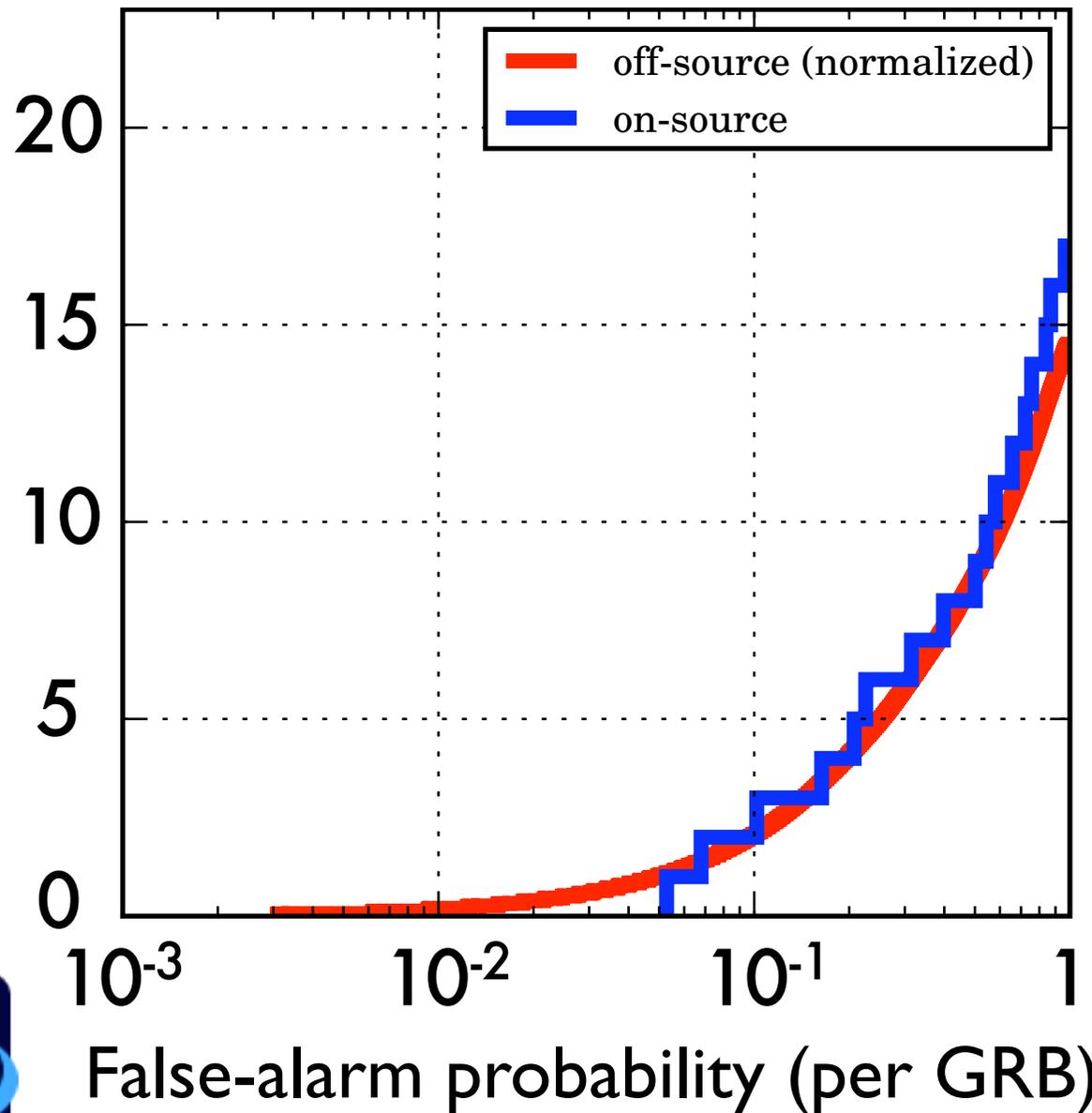
# Astrophysical exclusions from null results



(example recycled from previous GRB 070201 analysis) <sup>9</sup>

# Preliminary results for 21 GRBs: no gravitational waves

Cumulative  
histogram



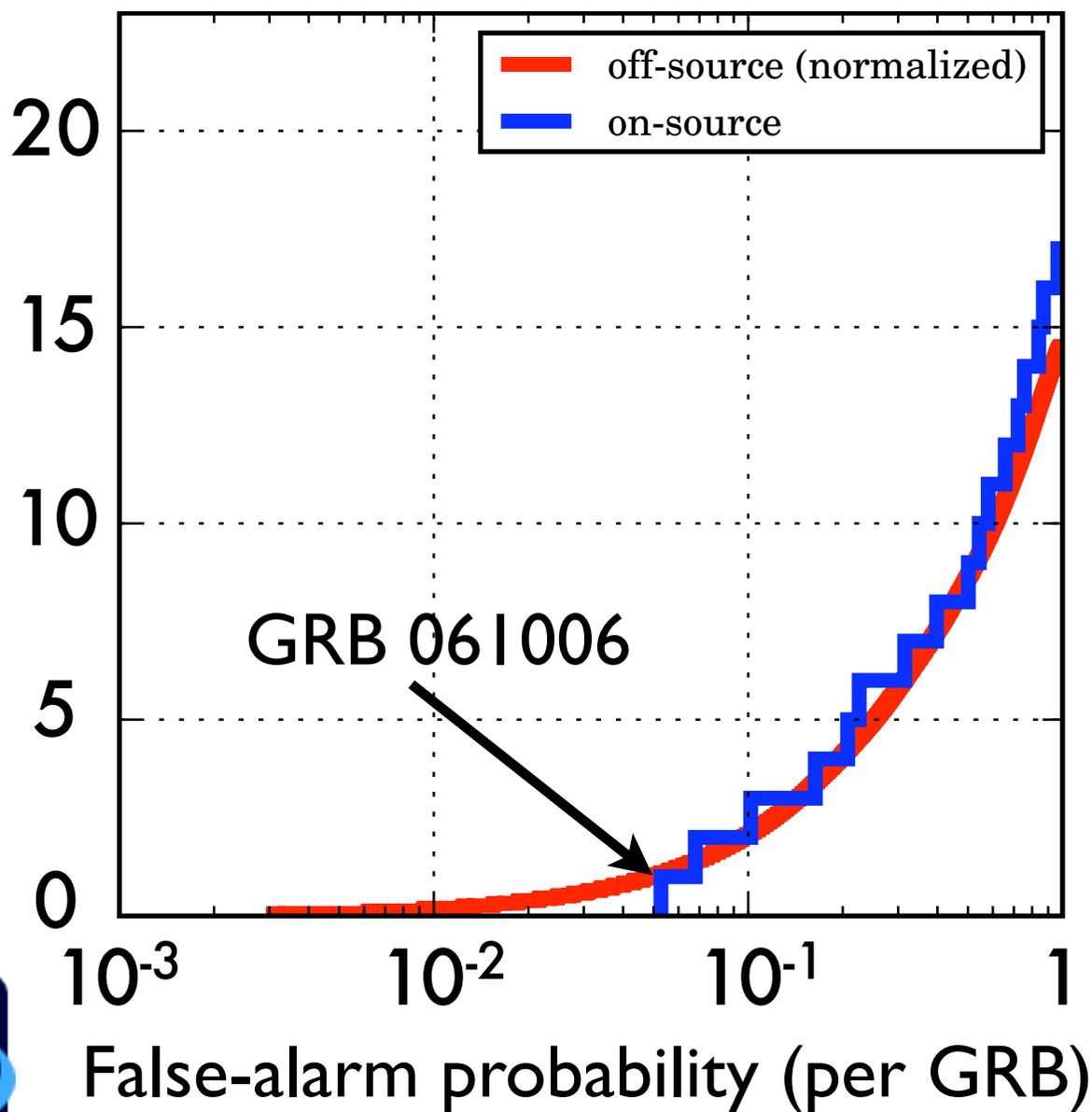
GRB 070923  
analyzed  
differently;  
result TBA

Four GRBs  
yielded no  
candidates



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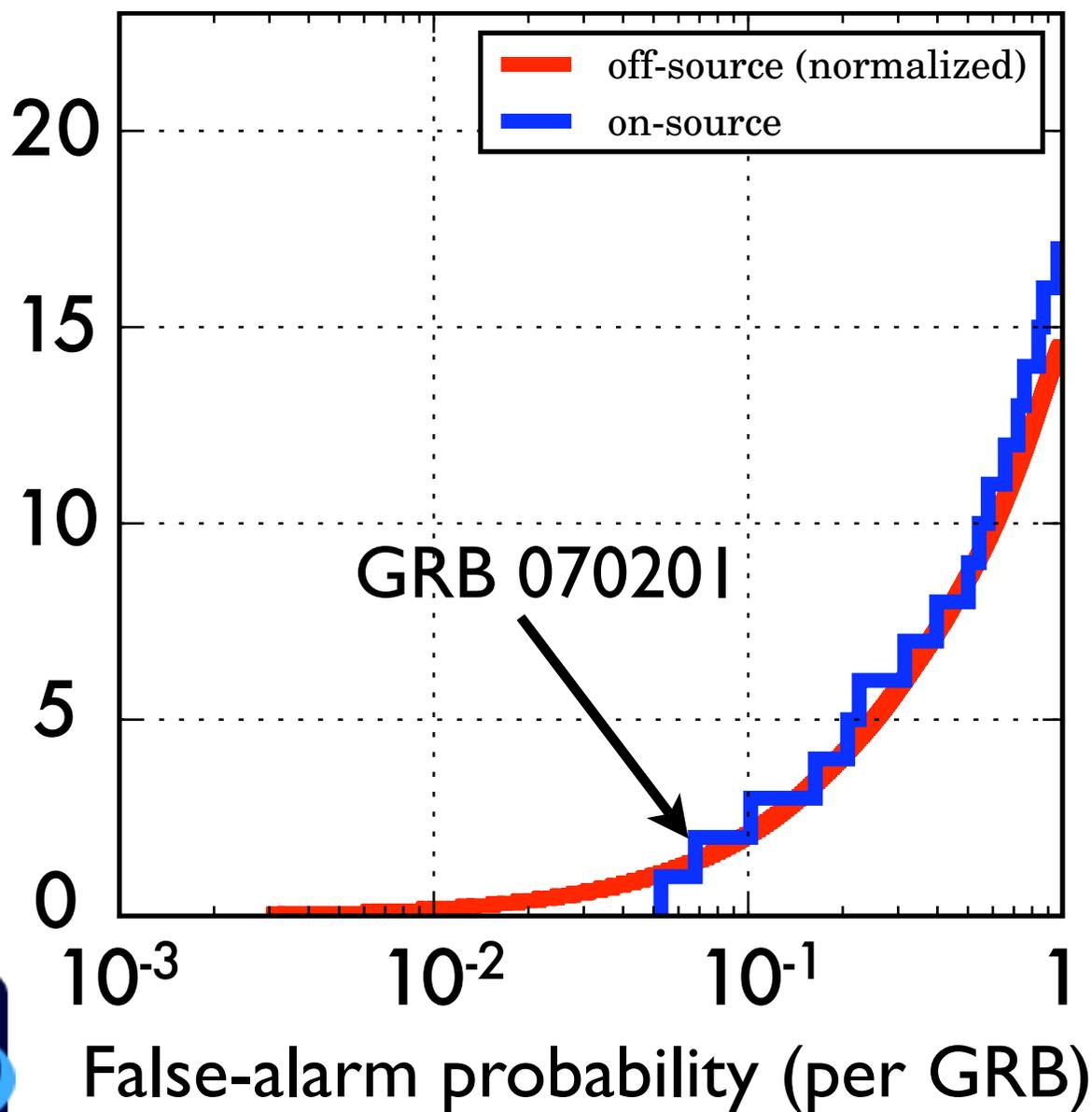
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# Where we are, where we will be

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- We discovered no GWs from **compact binary inspirals** in coincidence with **21 short GRBs** in S5/VSR1. GRB 070923, distance exclusions, and a population search are forthcoming.
- S6/VSR2 begins in a few weeks with enhanced detectors.
- Advanced detectors are due to come online around 2014.
- LIGO and Virgo are committed to **multi-messenger astronomy**. A coincident detection would provide enormous science.

