Coincident search for gravitational-wave and neutrino signals from core-collapse supernovae

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# Proposal for joint neutrino-GW search for nearby core-collapse supernovae

"Propose multi-stage collaborative work among gravitational-wave and stellar collapse neutrino communities that will put both in a better position to detect and extract the science of nearby core-collapse events in the immediate and near future."

#### Neutrino community

- S D'Antonio
- A Di Credico
- V Fafone
- ✤ W Fulgione
- K Scholberg

#### Theory/Phenomenology

- C D Ott
- G Pagliaroli
- F Vissani

#### LIGO-Virgo community

- L Cadonati
- E Coccia
- R Frey
- E Katsavounidis
- I Leonor
- C D Ott
- G Pagliaroli
- E Thrane

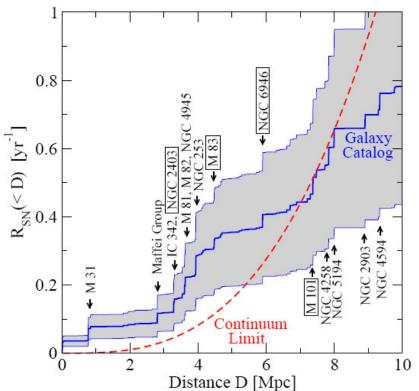
# Motivations for joint neutrino-GW nearby supernova search

- neutrino signal and gravitational-wave signal from core-collapse supernova are expected to be prompt and occur within a short time window of each other
  - onset of optical signal would be detected ~hours later
- both neutrino and GW signals would probe the innermost region of a SN core
- some supernovae might be optically dim but would still be seen in neutrinos and gravitational waves
- current generation of neutrino and GW detectors are expected to be sensitive to signals from Galactic/nearby supernova
- joint search will allow for shorter time windows and will tolerate higher single-detector false alarm rates, i.e. detection thresholds can be lowered
- use of worldwide network of neutrino and GW detectors will increase detection livetime

#### ➔ Increased sensitivity to supernova event

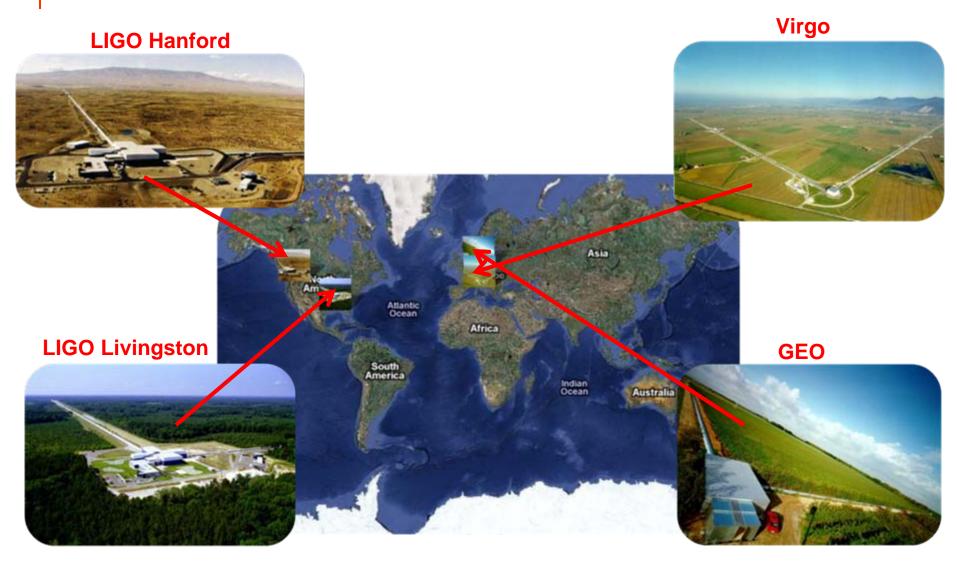
# Estimates of Galactic and nearby core-collapse supernova rate

- estimated Galactic rate is a few (~3) per century
- estimated rate in Local Group (out to ~1 Mpc) ~twice the Galactic rate
- observations indicate that the true SN core-collapse rates could be higher than these estimates
- existence of electromagnetically dark or obscured SN would also bring uncertainties to these rates



#### Ando, S. et al. 2005, PRL, 95, 171101

# Global network of GW detectors



# Some neutrino experiments with SN detection capability LVD **Borexino** Super-K Asia North America IceCube Atlantic Ocean Africa **KamLAND** South America Pacific Ocean Indian Ocean Australia

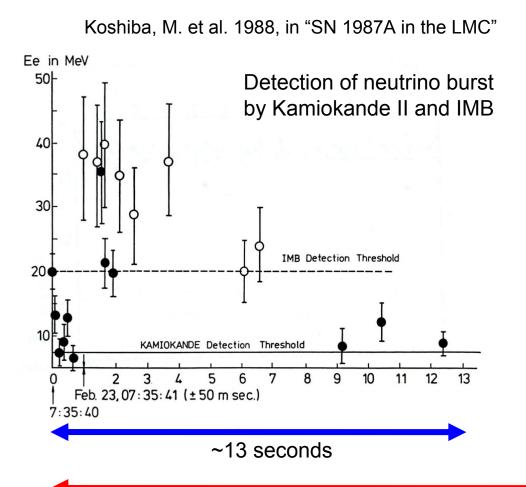
# Neutrino signal from supernovae: SN 1987A

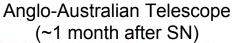
- in core-collapse supernova most (~99%) of the gravitational binding energy (~3E+53 ergs) is released in the form of neutrinos of all flavors
- neutrino energies are in the ~few tens of MeV range
- ~1E+58 neutrinos are emitted over a time scale of ~few tens of seconds
- neutrino burst from SN 1987A in LMC (~50 kpc) was detected by neutrino experiments

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#### Hubble Space Telescope Wide Field Planetary Camera 2

## Neutrino signal from supernovae: SN 1987A







#### first optical sighting occurred ~a few hours after time of neutrino burst

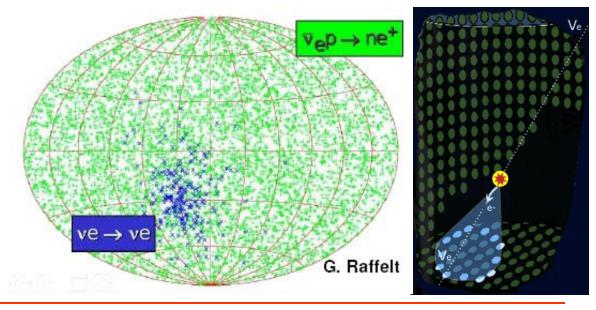
# Rough estimates of sensitivity of some neutrino detectors to core-collapse SN at 8.5 kpc

| Detector | Туре        | Mass (kton)  | Location   | Events at 8.5 kpc | Live period  |
|----------|-------------|--------------|------------|-------------------|--------------|
| Super-K  | $H_2O$      | 32           | Japan      | 8000              | 1996-present |
| SNO      | $D_2O$      | $1 (D_2 O)$  | Canada     | 400               | 1999-2006    |
|          |             | $1.4 (H_2O)$ |            | 450               |              |
| LVD      | $C_nH_{2n}$ | 1            | Italy      | 300               | 1992-present |
| KamLAND  | $C_nH_{2n}$ | 1            | Japan      | 300               | 2002-present |
| Borexino | $C_nH_{2n}$ | 0.3          | Italy      | 100               | 2005-present |
| AMANDA   | Long string | 0.4/PMT      | South Pole | N/A               | 1998-2009    |
| IceCube  | Long string | 0.4/PMT      | South Pole | N/A               | 2007-present |

these detect Cherenkov light from charged particles produced or scattered due to neutrino interactions in the medium

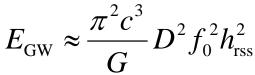
#### Pointing

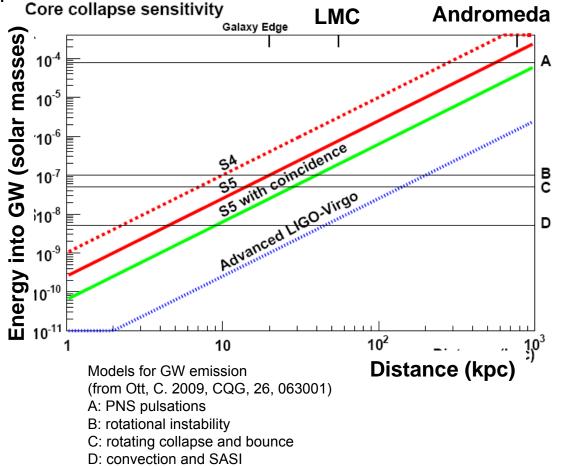
- if there are enough events, electrons from elastic scattering with neutrinos can be used to reconstruct direction
- events from dominant absorption interaction would be approximately isotropic
- with Super-K, expect a pointing accuracy of ~4 degrees for a SN at 8.5 kpc



# LIGO-Virgo sensitivity and expected improvement with joint neutrino search $\pi^2 c^3 p^2 c^2 t^2$

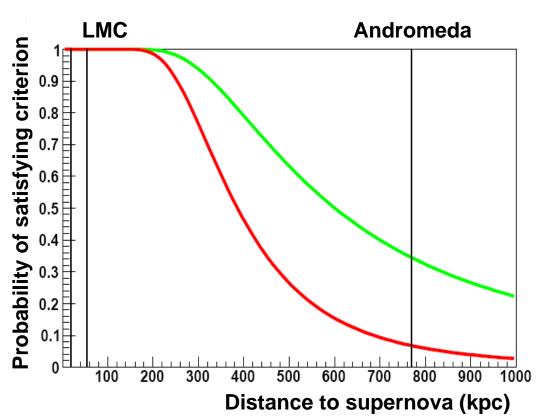
- in contrast to neutrino signal, energy emitted as GW radiation is expected to be small
- currently, there are large uncertainties in models of corecollapse SN, e.g. simulations have difficulty making a SN explode
- like neutrino signal, GW signal would probe the innermost region of SN core
- requiring coincidence of GW and neutrino signals to within a short time window of ~few seconds would allow lower detection thresholds
  - ➔ improvement in sensitivity





# Joint search could benefit neutrino search as well

- criterion for neutrino search can be relaxed
- example: for Super-K distant SN search, criterion is at least 2 neutrino events per 20 seconds and high energy threshold of 17 MeV
- if coincidence with GW signal is required, then criterion can be relaxed to a single neutrino event; odds will increase that distant core-collapse will satisfy this criterion
- energy threshold could also be lowered



Detection probability

# Supernova early warning system (SNEWS) http://snews.bnl.gov

- alert system which would send out notification of high-confidence SN to astronomical community a few minutes after detection of neutrino burst by multiple detectors
- LIGO-Virgo is signed up to get these alerts in the control rooms
- low-latency search for a GW signal coincident with a SNEWS trigger is planned for the LIGO-Virgo S6/VSR2 run
- the proposed joint GW-neutrino search will complement the existing infrastructure and procedures which are in place in the event of a SNEWS alert



## Data sets and status of proposal

- past runs--sufficient overlapping data exists
  - LIGO-Virgo S5/VSR1 run (Nov 2005 to Sep 2007;
    ~70% to ~80% duty cycle depending on interferometer)
  - SK-III run (Aug 2006 to ?; duty cycle ?)
  - LVD run 8 (Feb 2005 to May 2007; >99% duty cycle),
    LVD run 9 (June 2007 to Dec 2008; >99% duty cycle)
- future runs
  - LIGO-Virgo S6/VSR2 run
  - (I don't have info for the neutrino detectors)
- LIGO-Virgo collaborations have reviewed the proposal; awaiting final approval
- neutrino collaborations (currently Super-K, LVD, Borexino) are examining the proposal and discussions are ongoing
- Join us if you are interested in this proposal!

## Summary

- A proposal for a joint neutrino-GW search for nearby core-collapse supernovae has been presented
- The proposed search is scientifically well motivated, with experimental benefits to both neutrino and GW communities
- This is also a good avenue for establishing a robust working relationship between the two communities