

Summary Requirements for Advanced LIGO ASC/LSC Photo-detectors

LIGO-T070110-00-C

R. Abbott, Caltech

17 May, 2007

1. **Overview and Background** – Requirements are being gathered to guide the design and prototype efforts for the Advanced LIGO ASC and LSC photo-detectors. This document serves as a repository for the growing body of information on this topic, and will be refined as new decisions are made.
2. **Requirements common to ASC and LSC photo-detectors**
 - 2.1. **Construction**
 - 2.1.1. Photodetector assemblies, including the diode elements and associated electronics will reside inside the vacuum envelope, and must be constructed in a manner compatible with an ultra-high vacuum environment.
 - 2.1.2. It is desirable to be able to open the hermetic package used to house the electronics
 - 2.1.3. No specific size constraint exists yet.
 - 2.2. **Thermal:** In-vacuum thermal management is required to be passive and must be part of the mechanical design associated with the photodetector element and amplifier electronics.
3. **LSC Photodetector Design Data**
 - 3.1. **Nomenclature:** To identify a specific photodetector operating frequency, a shorthand notation for the offset frequency from the optical carrier has been adopted. In general, there is a 1:5 relationship between the first and second RF modulation frequencies of the optical carrier, but other mixing products are present on the light.
 - 3.1.1. A letter prefix – I or Q – designates the in-phase or quadrature component of a specific RF sideband
 - 3.1.2. A numerical suffix – 1 or 2 – designates the first and second modulation sidebands applied by the IO modulators. Nominally, this refers to 9 MHz and 45 MHz. The exact frequencies will likely change slightly, but the frequency ratio of 1:5 will be preserved.
 - 3.1.3. A letter suffix – M or P – identifies the difference or sum frequency respectively of the first and second RF modulation sidebands.
 - 3.1.4. An abbreviation is used – consistent with previous LIGO naming conventions – to identify the specific optical port.
 - 3.1.5. Examples
 - 3.1.5.1. REFL I1 – Reflected Port, I-phase, 9 MHz sideband
 - 3.1.5.2. REFL IM – Reflected Port, I-phase, 36 MHz product (difference between 45 MHz and 9 MHz sidebands).

3.2. **Photodetector frequencies and locations** (displayed assuming 9 MHz and 45 MHz modulation)

Table 1

Modulation F1(MHz)	9	Modulation F2 (MHz)	45
Optical Port	Shorthand Identifier	Frequency (MHz)	Notes
REFL	I1	9	
REFL	IM	36	REFL IM&IP should be derived from the same photodiode
REFL	IP	54	REFL IM&IP should be derived from the same photodiode
REFL	I2	45	
POX	I1	9	
POX	Q2	45	
AS (Locking)	I2	45	

3.3. **Noise**

- 3.3.1. Minimum photocurrent for which a shot noise limited electronics performance applies - 30mA
- 3.3.2. Minimum offset frequency relative to RF center frequency for which a shot noise limited sensitivity requirement exists is 20Hz. This isn't really an impact on the detector design, but will influence demodulator design.
- 3.3.3. Shot noise at 30mA - 80pA/rtHz
- 3.3.4. Safety factor for electronics noise - ~4, more if possible
- 3.3.5. Electronics front end current noise requirement - 20pA/rtHz

3.4. **Dynamic Range**

- 3.4.1. Maximum sustained photo-current – 100mA

3.5. **Inputs and Outputs**

- 3.5.1. DC Power - +/-15VDC, +/-5% will be used for DC power.
 - 3.5.1.1. DC power filtration will be used in the vacuum system, but not active regulation
 - 3.5.1.2. Reverse polarity protection will be used to avoid damage is DC power polarity is reversed.
- 3.5.2. RF Outputs – One or more RF outputs will exist for each detector assembly. The exact number of outputs per detector will be the subject of a tradeoff study during the prototype process.
- 3.5.3. RF Inputs – One or more ASI type cancellation inputs will be required per detector. The characteristics of this input are **TBD**.
- 3.5.4. DC Output – A DC output will be supplied with internal whitening consistent with signal amplification to greater than 100nV/rtHz for noise immunity. Additional whitening will occur once the signal leaves the vacuum envelope.
- 3.5.5. Temperature Readback – A temperature sensor will be included in each in-vacuum detector assembly.
- 3.5.6. The need for external photodiode bias adjustment to compensate detector de-tuning resulting from variations in photocurrent depends on final circuit topology and the degree to which the design must be insensitive to this effect. The detectors must exhibit less than **TBD** degrees of phase variation over the range of photocurrent from **TBD** to 100mA. Provided the detector design intrinsically meets this requirement, there will be no provisions for external bias adjustment.