LASER INTERFEROMETRIC GRAVITATIONAL-WAVE OBSERVATORY

Recent Progress





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LIGO-G000017-00-D



OVERVIEW

- ☐ Facilities and construction complete
- Vacuum system complete
- □ Beam tube bakeout
 - LIGO Hanford Observatory: completed
 - LIGO Livingston Observatory: X-arm completed; Yarm underway
 - Seismic isolation stacks
 - LHO 2k and LLO complete
 - LHO 4k underway
- ☐ All core optics polished and coated
 - Metrology shows good results
- □ Suspensions
 - LHO 2k complete
 - **O** LLO ~ 50%



FACILITIES AND CONSTRUCTION

- All buildings and civil construction complete
- □ All vacuum equipment installed
- ☐ All beam tubes ready
- □ Beam tube bakeout
 - LHO: completed
 - LLO: X-arm completed; Y-arm underway
 - O H₂O partial pressures → < 10⁻⁹ torr
- Vacuum status
 - Maintaining < 3 x 10⁻⁸ torr even with chambers loaded
 - Partial pressure of H₂O carefully monitored/budgeted to avoid contamination of beam tube



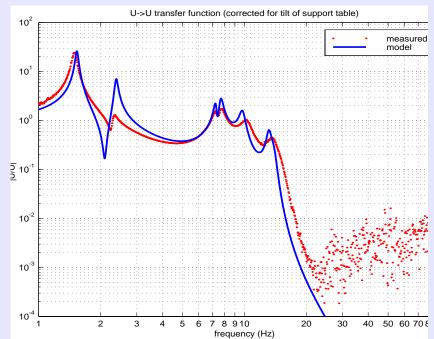
SEISMIC ISOLATION SYSTEM

□ Seismic isolation stacks

- O Stainless steel masses → ~ 600 kg each stage
- → Helicoil springs with lossy viscoelastic layer → Q ~ 40
- 3 stages $\Rightarrow 1/f^6$ for f > 10 Hz

□ Stack installation

- LHO 2k and LLO complete
- LHO 4k underway
- ☐ First-article testing of HAM stack



■ Water outgassing from viton seats



CORE OPTICS

- ☐ All 40 optics polished and coated
- □ Polished substrates
 - μ-roughness: < 10 ppm scatter</p>
 - Radius of curvature: δR/R < 5%
 (hand selected pairs → δR/R < 3%
- □ Coated optics
 - Point defects: < 2 ppm</p>
 - Coating absorption: < 1ppm</p>
 - Surface uniformity
 3.6 Å rms (polished) ⇒ 5.9 Å rms (coated)
- **□** Metrology
 - CSIRO and NIST (633 nm)
 - 1.064 mm Veeco (Wyko) interferometer
 - compares Core Optic and reference flat
 - 1 nm_{rms} surface variation accuracy
 - 3% radius of curvature accuracy
- □ Internal resonance Q factors



Suspensions

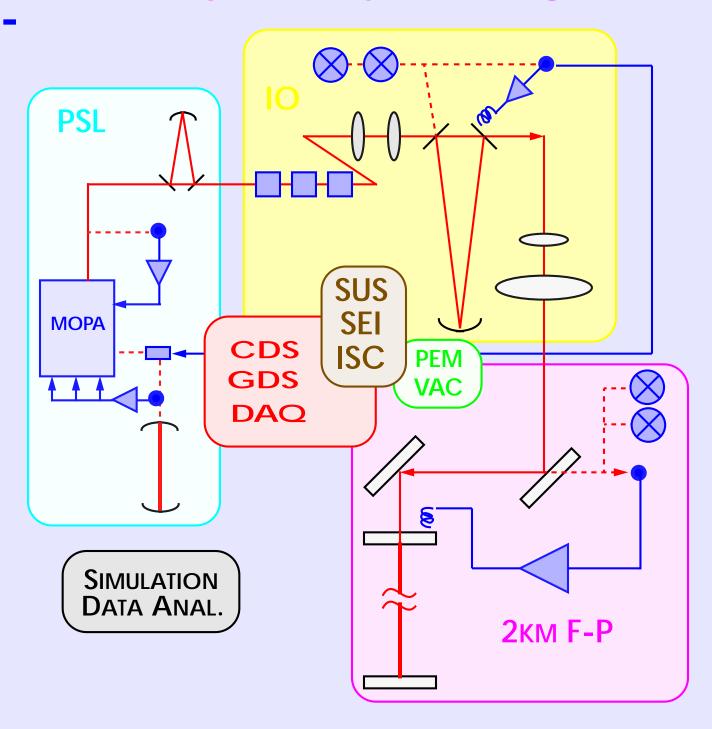
- □ Single wire-loop suspensions
- ☐ Four electromagnetic actuators
- □ Four 'shadow' sensors for local position sensing
- □ Installation
 - LHO 2k complete
 - **O** LLO ~ 50%

☐ Testing

- PD-LED package used for shadow sensors sensitive to 1.06 mm laser light
- Cross-coupling between principle degrees of freedom in sensors



COMMISSIONING ACTIVITIES AT LHO





PRE-STABILIZED LASER

□ Laser

- Master Oscillator Power Amplifier
- → 10 W in TEM₀₀ mode

□ PSL functions

- Frequency stabilization
- Intensity stabilization
- Pre-modecleaner

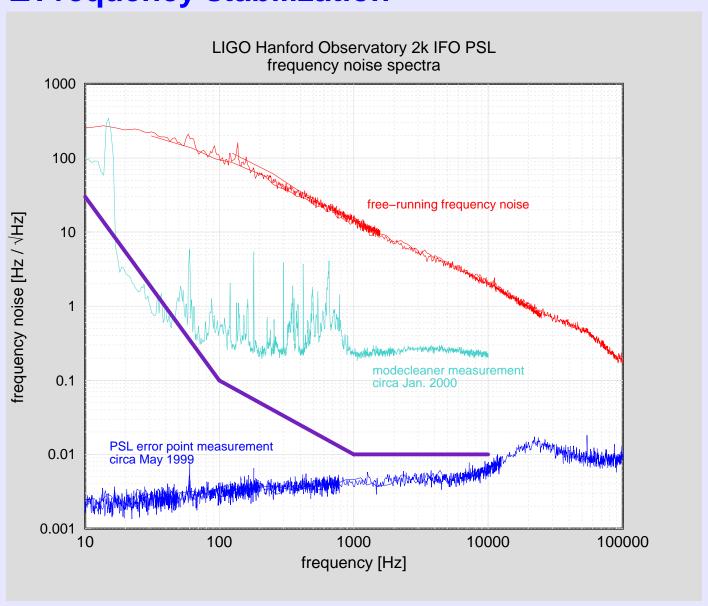
■ Nested loop strategy

- fixed-spacer reference cavity \Rightarrow 0.1 Hz/ $\sqrt{\text{Hz}}$
- 12 m suspended mode cleaner cavity \Rightarrow 10⁻⁴ Hz/ $\sqrt{\text{Hz}}$
- \bigcirc 4 km arm cavities $\Rightarrow 10^{-7} \text{ Hz}/\sqrt{\text{Hz}}$



PSL CHARACTERIZATION

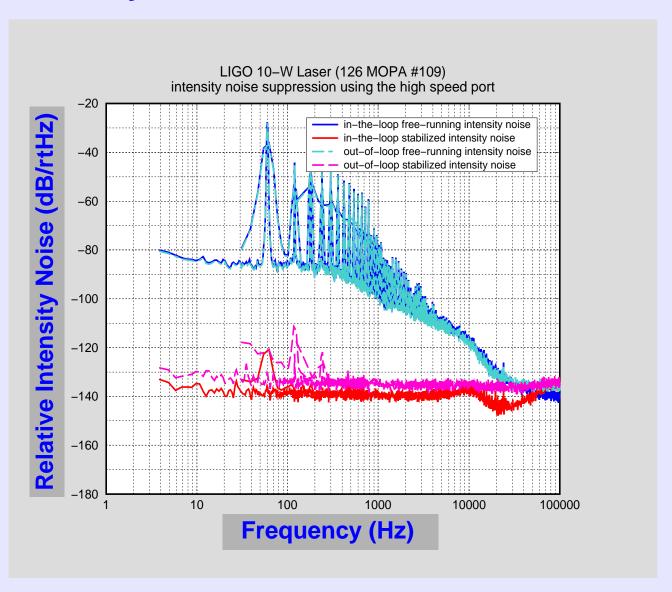
☐ Frequency stabilization





PSL CHARACTERIZATION (2)

□ Intensity stabilization





INPUT OPTICS: MODE CLEANER

- **□ EO Modulation**
- Mode-matching telescope
- **□** Suspended mode cleaner
 - 'Small' optics: 7.5 cm diameter
 - Triangular cavity
 - Finesse designed to be 1550
- □ Locks easily and stably for >24 hrs
- ☐ Length/frequency servo BW ~ 30 kHz
- □ Output beam angular stability
 - Long term drift ~ 4 μrad/hr
 - Fluctuations ~ 3 μrad rms

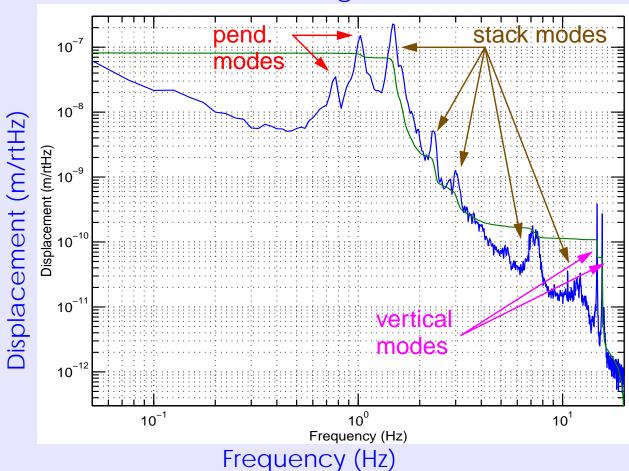


MODE CLEANER CHARACTERIZATION (2)

□ Length stability

- O Long term drift ~ 1 μm/hr
- Fluctuations ~ 0.1 μm rms (mostly at pendulum and stack eigenmodes)

Mode Cleaner Length Fluctuations





MODE CLEANER CHARACTERIZATION (3)

- Length measured by resonant RF sideband tuning: L = 15.239595 m ± 50 μm
- □ Cavity linewidth measured
 - ringdown/up and AM transfer function $\Rightarrow \Delta f = 7.31 \text{ kHz}$ ⇒ finesse, F = 1346
 - $T_{in} = T_{out} = 2255 \text{ ppm}, T_{curve} = 10 \text{ ppm}, F = 1346$ ⇒ total loss = 49 ppm/mirror
- ☐ Total transmission measured: 0.98 ± 4.7%
- ■Internal resonances of optics measured
 - **Q** = **0.75** million to **1.29** million for f_0 = 28.233 kHz mode
- **□** Ongoing work:
 - Operation at full power
 - Increase servo bandwidth
 - Implement intensity stabilization after MC



2 KM FABRY-PEROT CAVITY TEST: WHY?

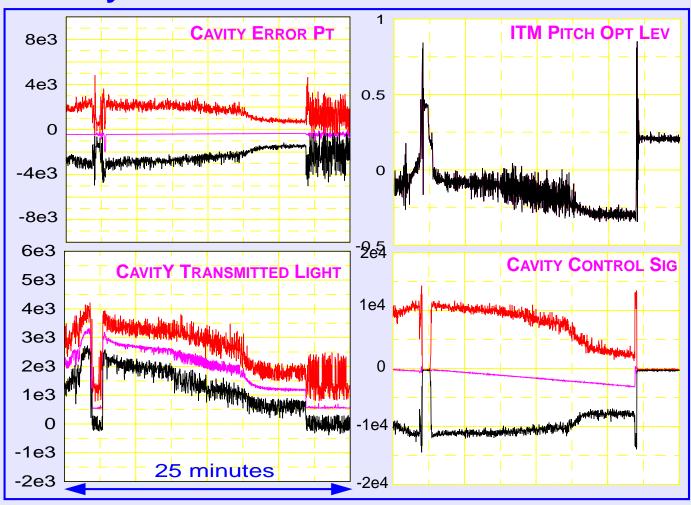
Exercise lock acquisition controls
 Measure frequency noise from IO mode cleaner
 Test nested loop frequency control
 Early look at optical properties (losses, radii of curvature) of core optics
 Mode matching
 Degree of excitation of suspension eigenmodes and test mass internal resonances

Measure environmental fluctuations



2 KM F-P TEST: LOCKING

- □ Gate valves between corner and mid stations first opened Dec '99
- □ Laser beam on suspension cage
- □ Cavity locks for ~15 minute stretches





2 KM CAVITY LOCKING

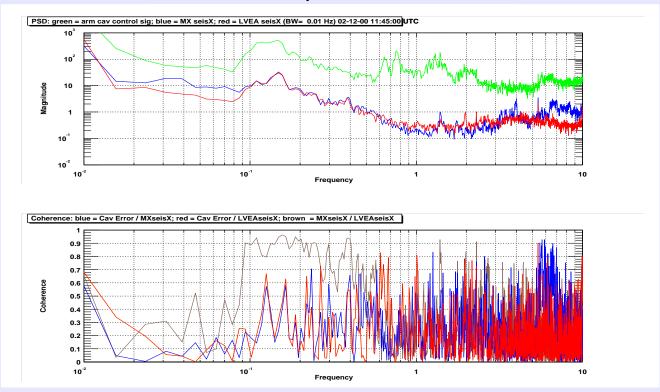
- □ Observe ~ 0.5 µm/min drift during 15 minute stretches
 - Arm cavity length drift consistent in magnitude and sign with drift in laser frequency due to reference cavity drift
 - Thermal control being implemented on ref cav
- ☐ Length-alignment coupling as length drifts
 - More coupling than can be accounted for by observed cross-coupling in actuators at DC
 - Possible saturations in coil drivers under investigation
 - Wavefront-sensing based alignment control next
- □ Excitation of 6.75 kHz 'butterfly' mode of TM
 - Sensitive to centering of laser beam on mirror
- □ Other internal resonances also excitable, e.g. 31.25 kHz !!
- □ Internal mode Q's for several core optics measured 200,000 to 10 million



2 KM F-P TEST: FIRST LOOK

□ Correlation studies: μ-seismic peak

Common-mode displacement of optics 2 km apart is ~
 2x smaller than each optic



- □ Cavity length: L = 2009.1096 ± 0.01 m (surveying → 2009.119 m)
- □ Cavity storage time \approx 460 μ sec \Rightarrow $F \approx$ 216
- □ Sensitivity dominated by frequency noise



Outlook for the Near Future

□ LHO 2km commissioning

- O Single arm test ⇒ Summer 2000
- Power-recycled Michelson ⇒ Winter 2000
- ⊃ PRM with FP arm cavities ⇒ Summer 2001

LLO

- Similar commissioning path
- ~3 months behind LHO schedule

□LHO 4km

- Early 2002
- Noise reduction and performance enhancement phase
- ☐ Science run begins 2002

