



LIGO

Catching Gravitational Waves in Livingston, Louisiana

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Project Overview

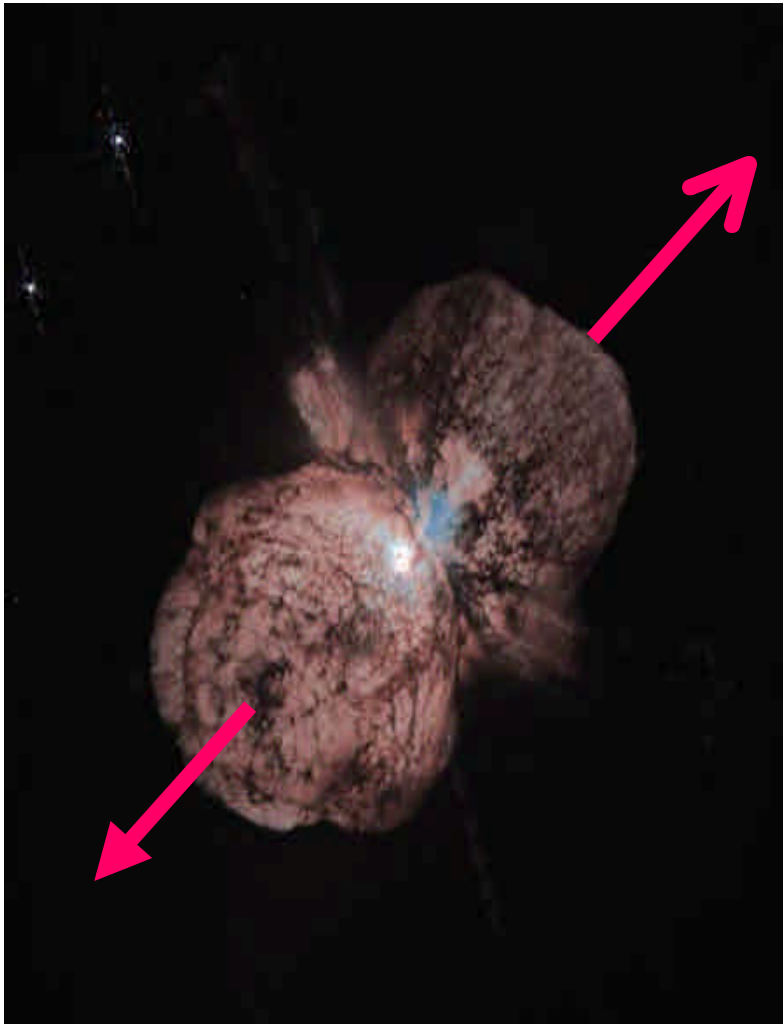
- LIGO: Laser Interferometer Gravitational Wave Observatory
- Operated by California Institute of Technology in partnership with Massachusetts Institute of Technology.
- Funded through cooperative agreement between National Science Foundation and Caltech
- LIGO is the result of more than 30 years of study and development of laser interferometer gravity wave antennae



Scientific Motivation

- Albert Einstein predicted the existence of gravitational waves - *ripples in the fabric of space and time* - as part of the general theory of relativity.
- LIGO is designed to detect these waves directly, opening up a new vantage point from which to study the universe.
- When operational, LIGO will allow scientists to understand the astrophysical sources of gravitational waves, including supernovae, binary neutron star systems, and the vibration of black holes.

Example: Supernova



Why does this picture look this way?

Gravity plays a critical role in shaping the way the universe looks

Eta Carinae –
photo from Hubble Space Telescope

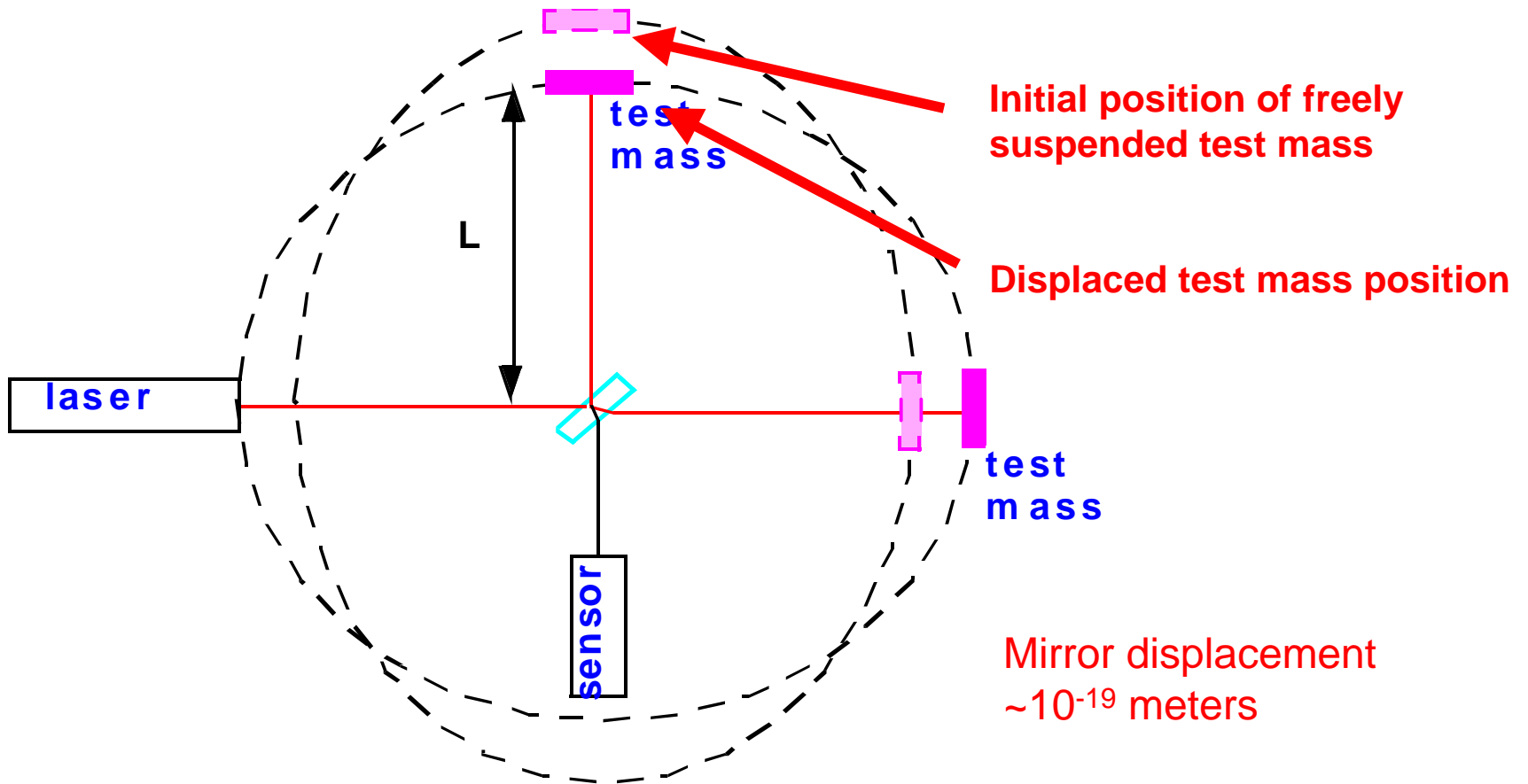


How LIGO works

- LIGO measures the tiny distortions of space-time caused by the passage of gravitational waves.
 - » Gravitational waves are waves in space and time
 - » Space is stretched and compressed at the frequency of the gravitational wave
- Distances between mirrors suspended in vacuum are precisely measured using a highly stabilized laser beam

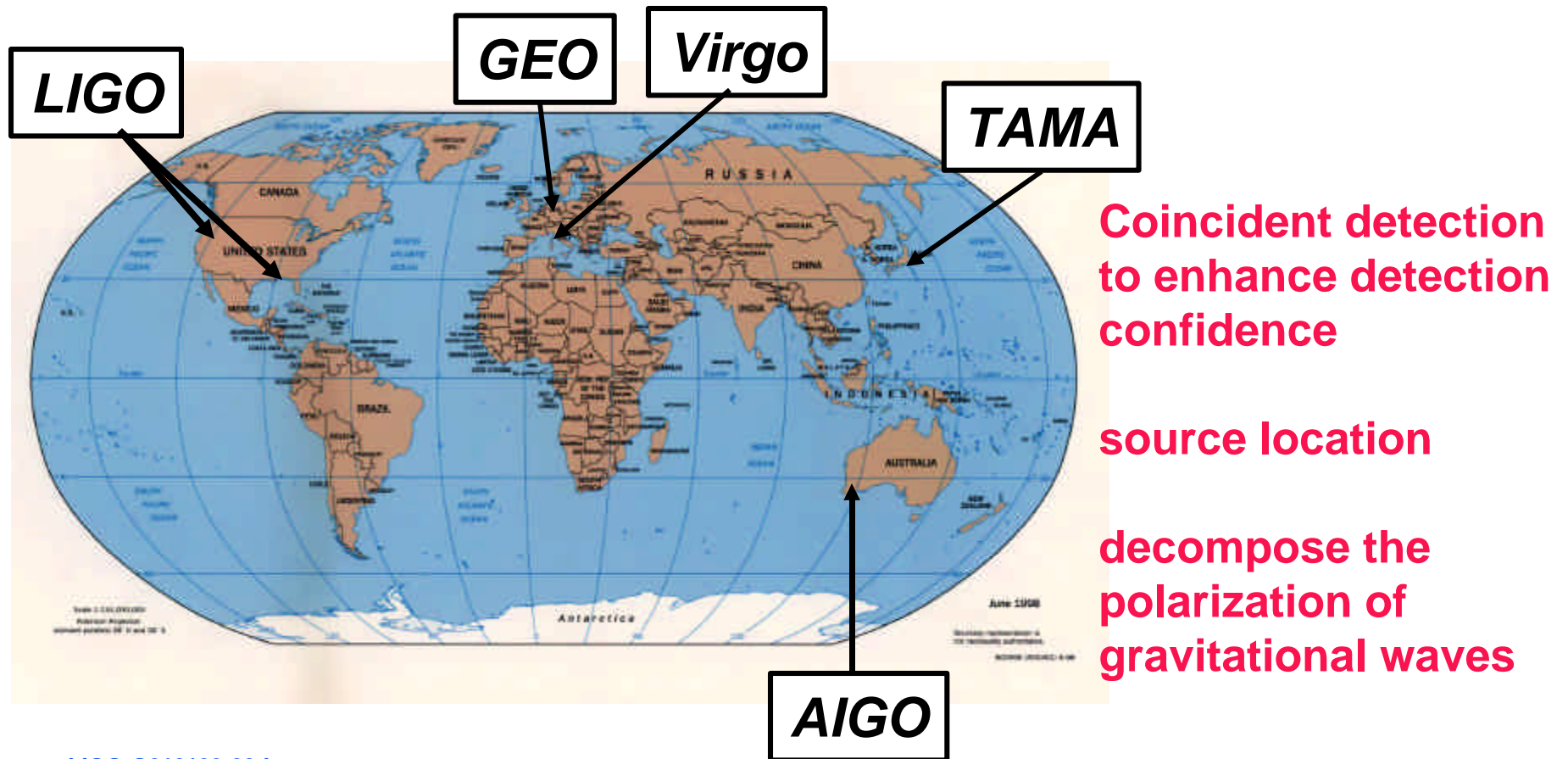


How LIGO Works



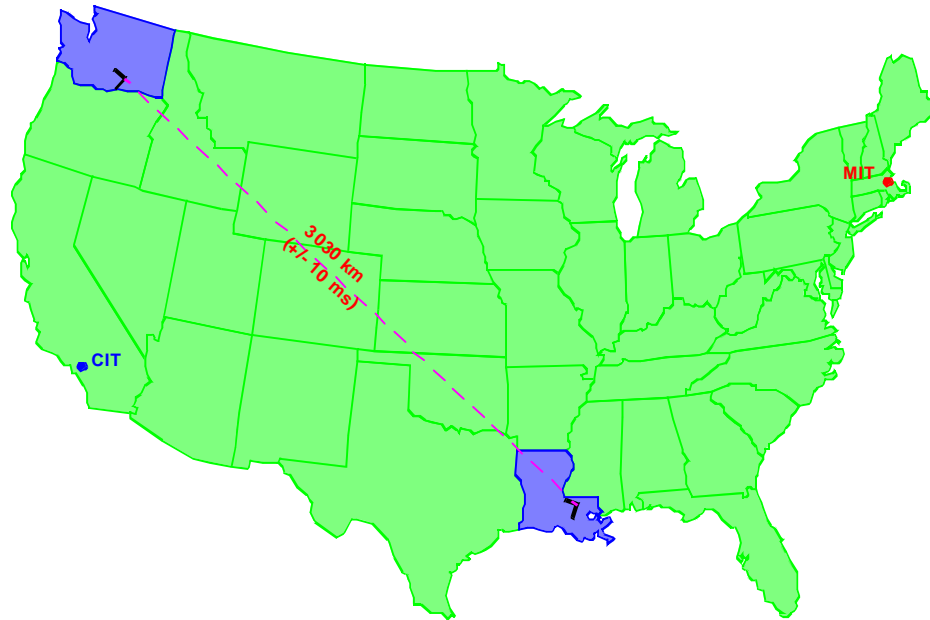


International network of gravitational wave interferometers now being readied for operation



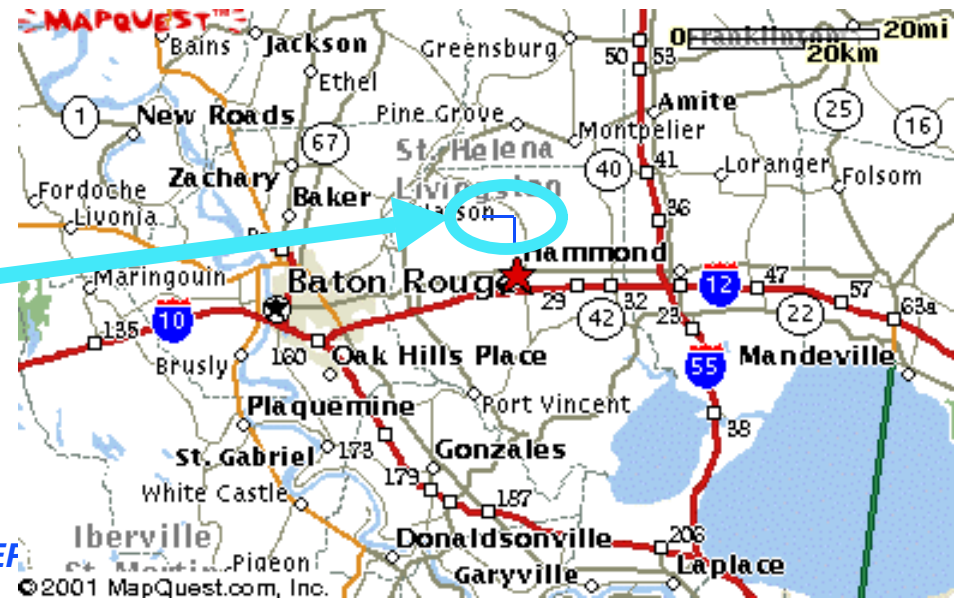


LIGO Locations



- Two well separated observatory sites to minimize correlated noise:
- Tri-cities area of Eastern Washington
 - Livingston, Louisiana

LIGO



LIGO-G010199-00-L
4/25/2001

Board of Regents/EF

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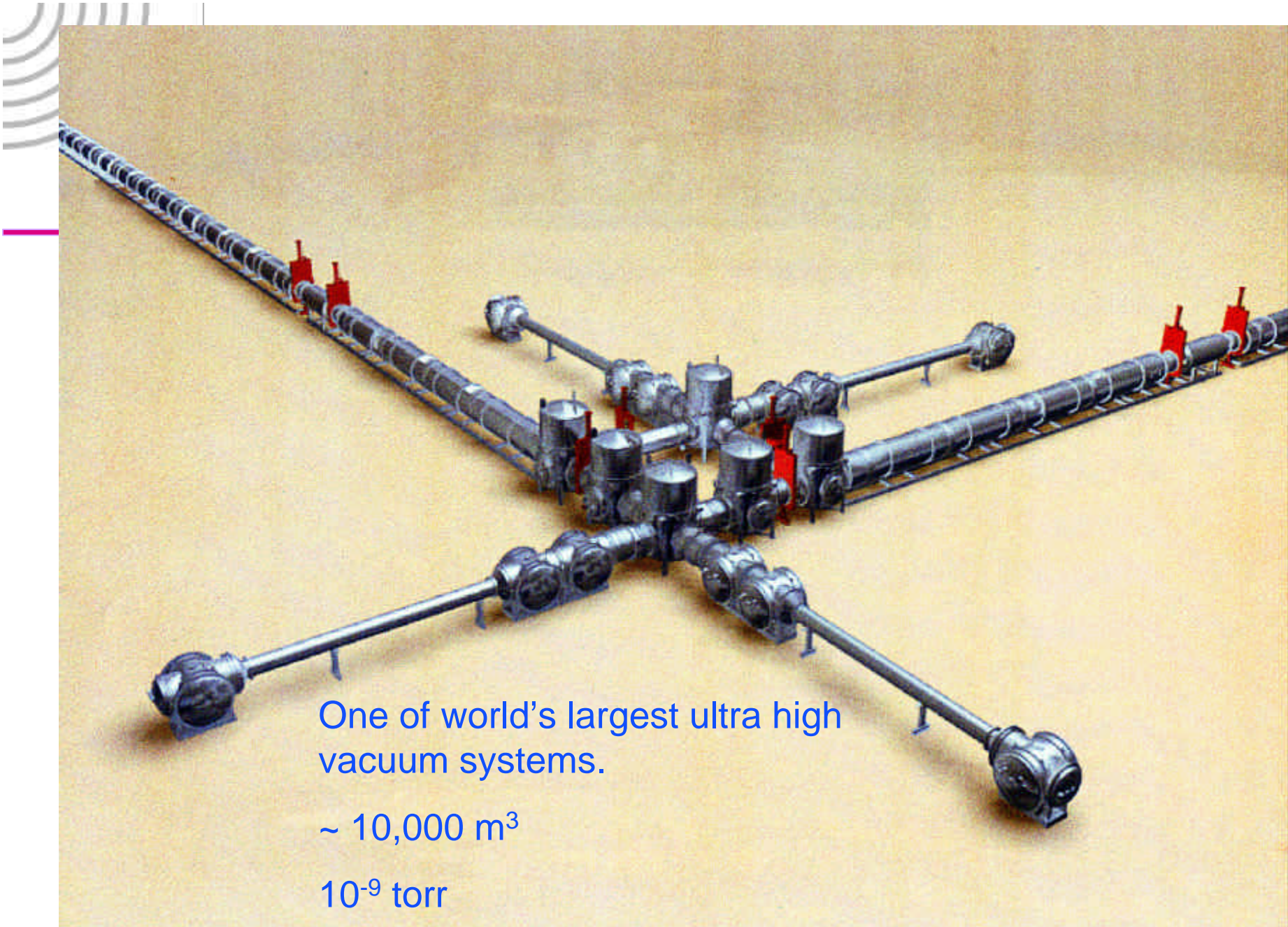


LIGO (Washington)



LIGO (Louisiana)





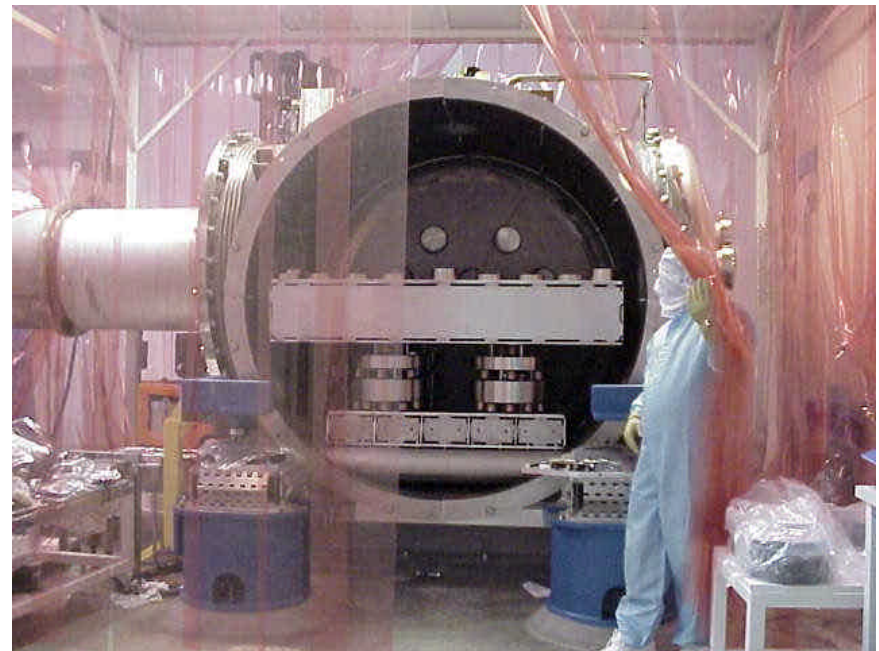
One of world's largest ultra high vacuum systems.

$\sim 10,000 \text{ m}^3$

10^{-9} torr



Seismic isolation system





Suspended Optics

Mass	10.7 kg	<hr style="width: 50px; margin-left: 10px;"/>
Mirror diameter	25 cm	
Mirror material	Fused silica	
Mirror internal Q per mode	$>10^6$	
Mirror internal adsorption	~ 5 ppm	
Mirror loss	50 ppm	

Single Wire Loop Suspension

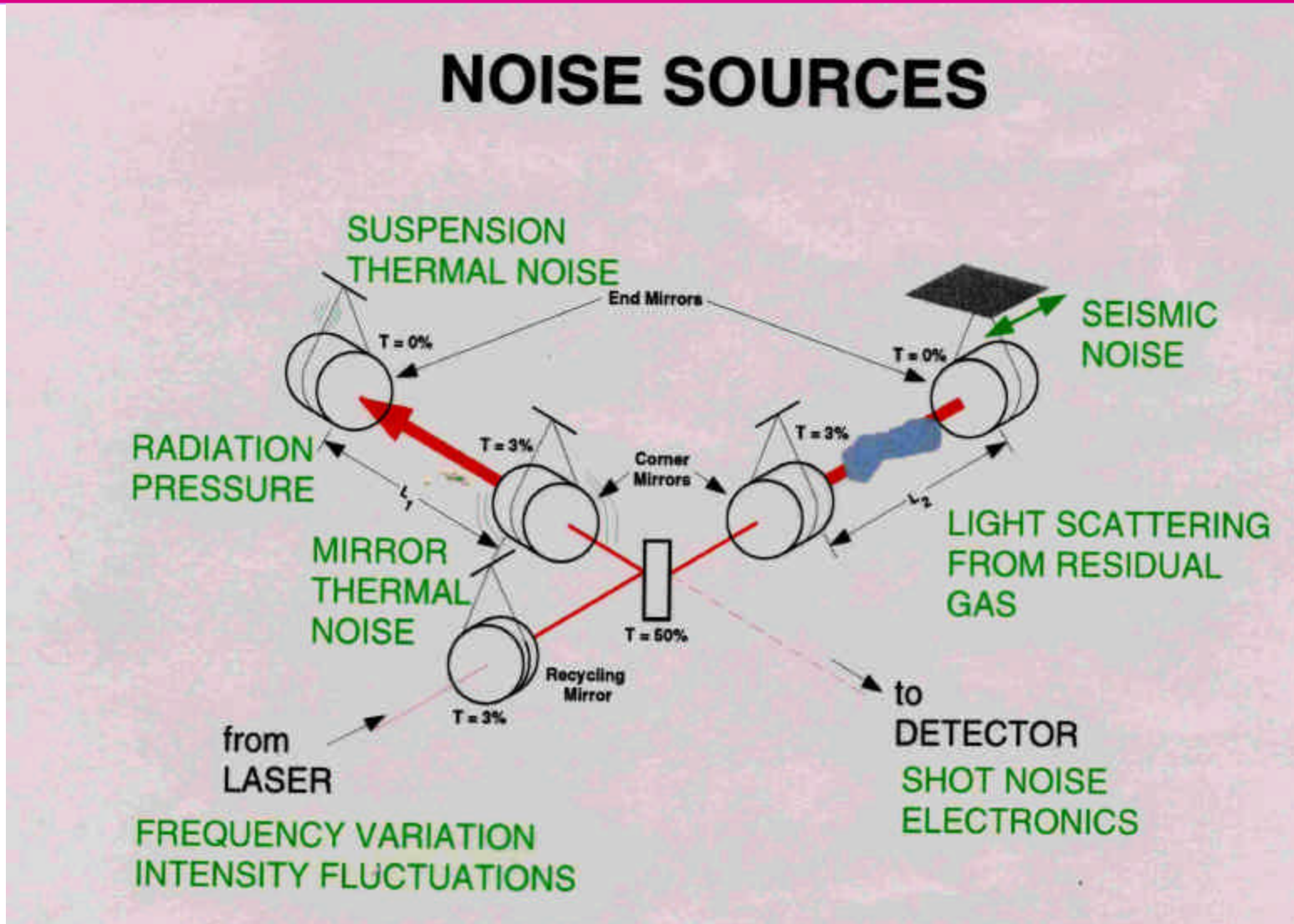
Pendulum material Steel wire, $Q=200,000$

Pendulum frequency 1 Hz



Limiting Factors in Operation

NOISE SOURCES





Design considerations to reduce noise

- Insensitivity to ground motion
 - » Seismic isolation
 - » Low vibration support equipment
 - » Active vibration isolation and servo control systems
- Brighter lasers
 - » Materials (mirrors, coatings, modulators) that can handle very high CW optical powers (up to about 1 MW at 1.06μ)
 - » 200 Watt single mode lasers at $\lambda = 1.06 \mu$
- Reduced thermal noise
 - » High Q suspensions
 - » High Q materials – sapphire, fused quartz, YAG, etc.



Collaborating Institutions

- Approximately 30 institutions world-wide are part of the LIGO science collaboration.
- LIGO encourages the participation of additional institutions in the science collaboration, particularly those in the regions close to each of the LIGO sites.
- Presently participating Louisiana institutions:
 - » LSU
 - » Southern Univ. of Baton Rouge
 - » Louisiana Tech Univ.
 - » Southeastern Louisiana Univ.
 - » Loyola University
- Others welcome!



Observatory Facilities and Staff

- Approximately 30 scientists and engineers resident at each LIGO site.
- Joint faculty appointments with U of Florida, SLU
- 5-10 scientific visitors from Caltech and MIT spend weeks to years
- Regular visitors from local universities
- Summer program for undergraduates
 - » This year about 20 students will work at Livingston during the summer
- Summer teacher program
 - » Strengthen science education in the region
- Laboratories and shops to conduct supporting research and development
- T1 internet (to be upgraded to OC3)



Status

- Construction complete
- All equipment installed
- Systems integration and test nearly complete
- Engineering studies continue through 2001
- Scientific data taking for 3-4 years begins in 2002
- Major upgrade planned for 2006
 - » 1000X increase in astrophysics event rate
 - » Advanced R&D and design engineering to be undertaken in parallel with scientific data taking



How Colleges and Universities Can Become Involved in LIGO

- *Join LIGO science collaboration:*

- » Contribute intellectually to the understanding of the interferometer, instrumental and environmental sources of noise, data analysis algorithms, etc.
- » Significant advantage in being located close to LIGO – data is dominated by instrumental signatures which need to be understood
- » Participate in the advanced R&D program for the LIGO upgrade:
 - Optical materials research
 - Coatings
 - Servo control system design
 - High power laser systems
 - Cryogenics
 - Earth science too! Characterization and adaptive reduction of seismic noise environment, earth tides



Educational Outreach

- Field trips by community and professional groups at both sites
- More than 3,000 visitors in last year at LLO (mostly school classes), 750 on one day during public open house
- Teacher open houses in summer and winter, more than 100 middle and high school science teachers in Livingston Parish have toured LIGO as part of teacher in-service
- LLO hosted more than 100 African-American high school science students participating in Southern University's Timbuktu Academy





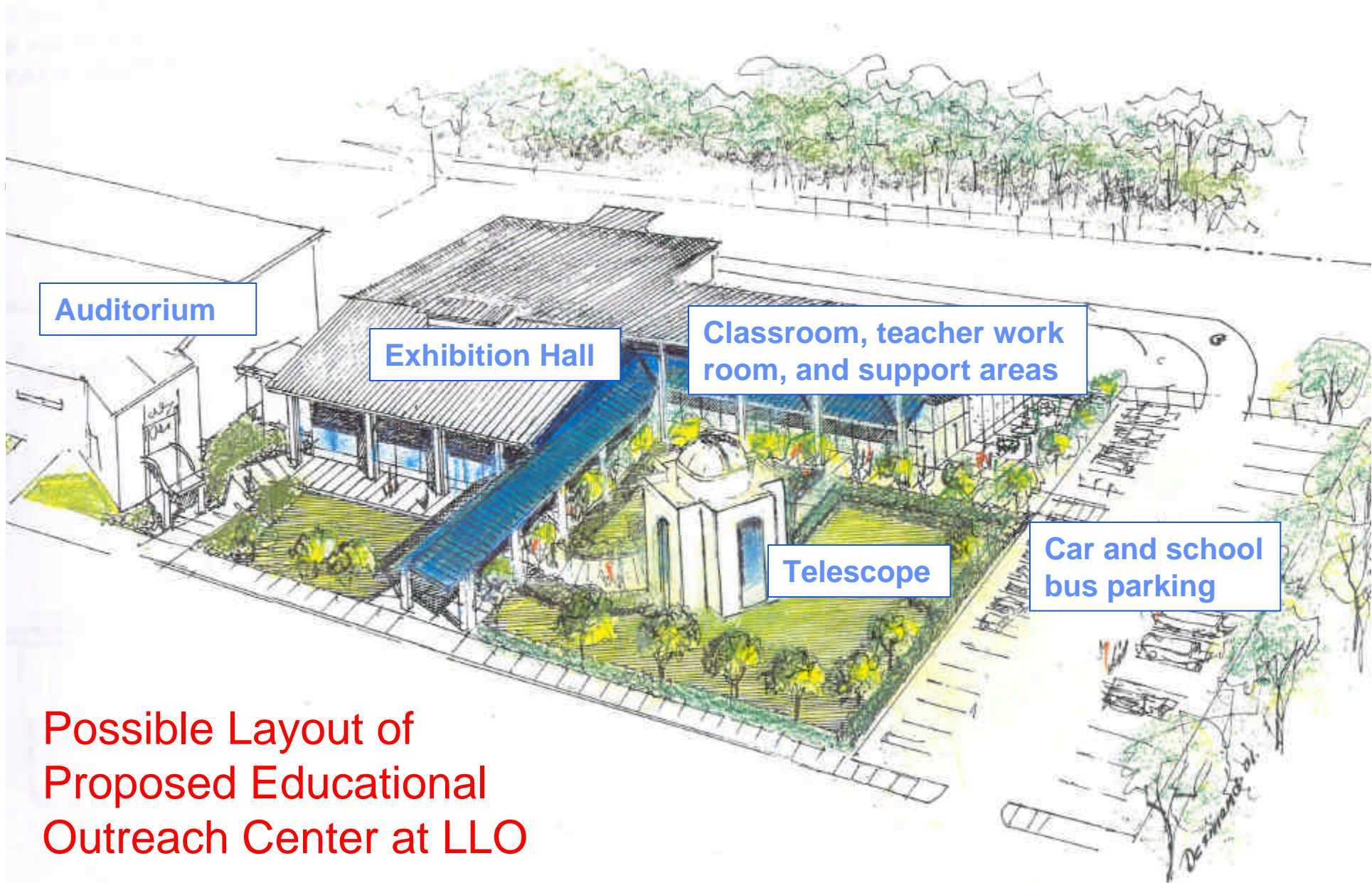
Become an Education Partner

- We want to be an important part of regional educational infrastructure
- Strengthen K-12 science education
 - » Pilot program in place at Northwestern State Univ. to develop LIGO related school curriculum materials
 - » Strong relationship with Southern University to promote community and K-12 outreach in traditionally underserved communities
 - » Additional opportunities to partner with school systems for in-class and site visits.
 - » Opportunity to develop proposals to NSF Division of Education and Human Resources
- Lots of room for others to be involved!



Outreach Center

- Congressman Richard Baker has enthusiastically supported establishment of an outreach center at LIGO Livingston Observatory along the lines of centers at Arecibo, Lowell Observatory, MacDonald Observatory.
- Center mission:
 - » Host site visitors with hands-on exhibits and science classes (similar to NSF-funded Arecibo and Lowell Observatory centers).
 - » Teacher in-service training and support for classroom enrichment (also like Arecibo and Lowell Centers).
 - » Host a modest school-to-work program for vocational training.
- We would like to partner with other education organizations to help create, operate, and utilize the proposed center.



**Possible Layout of
Proposed Educational
Outreach Center at LLO**



Summary

- LIGO welcomes opportunities for collaborative research and development with local institutions
- LIGO is committed to strengthening science education at the K-12 and university levels and wants to partner with universities, local school systems, and regional science centers to do this
- More information is available at the LIGO website:
www.ligo.caltech.edu