LIGO Topics in simulation and DA for GW interferometers: the VIRGO case

- The VIRGO super-attenuator: a introduction
- Why a mechanical simulation
- How the simulation is set up
- The comparison with real data, and how these are analyzed
- Latest results on suspension performances
- How the simulation is helping for locking studies

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LIGO A VIRGO Super Attenuator

- An inverted pendulum for low frequency control
- 6 seismic filters (in all DOFs)
- 1 longitudinal-angular control stage (the marionetta)
- 1 longitudinal control stage (the reference mass)
- The system has a double role
 - » Filtering out the seismic noise
 - » Actuate on the mirror position



Why so many stages?



Attenuation goals



- Vertical seism is dominant
- On paper, above 3÷4 Hz the noise should become dominated by other sources
- The choice is to achieve this goal only by passive means.

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LIGO Price to pay: LF amplification x 10⁻⁶ 10 0.8 0.04 0.6 0.02 0.4 0 0.2 -0.02 0 -0.04-0.2 -0.4-0.06 -0.6 -0.08 -0.8-0.1 5 10 15 20 25 35 40 45 0 5 10 15 20 25 30 35 40 45 0 30 Left: ground motion Right: mirror motion → control requirements

The SA as a control device



Why a mechanical simulation?...



- In the plot, the vertical TFs of a seismic filter
 - » The two curves correspond to two options for the fundamental frequency.
- What happens when chaining several such elements?

R.Flaminio, S.Braccini

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What one would like to know

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Input: force z on Marionnette

- In the plot, some of the (simulated) TFs which are relevant in controlling mirror position
- The design of control filters depend on these TFs.
- A complex system: one needs
 - » A good model, and
 - » direct measurements

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Model scope and requirements

- Assess attenuation performance in the detection band [4Hz – 10 kHz]
 - » Help in deciding where improvement is needed
 - » Requires modeling of the internal modes of elastic elements
 - » Limit: only the ITF shall be able to fully validate the results!
- Predict the motion of test masses
 - » Due to noise inputs (seism, thermal noise ...)
 - » Under the influence of control forces
- Provide a time domain model
 - » Needed to integrate with optics and study lock-acquisition
 - » Simple: as few DOF as possible to keep simulation time within reasonable limits
 - » Neglect internal modes

Is this strategy consistent?



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Model construction

- Describe elastic elements
 - » Only those relevant in the frequency band of interest!
- Keep the model simple
 - » Possibly limit to a effective potential representation

$$U = \frac{1}{2} \int_0^L EI \left[x''(z,t) \right]^2 + T \left[x'(z,t) \right]^2 dz$$

» Neglect higher order modes (violins) for control studies, keep them otherwise.



Left: VIRGO blades for vertical attenuation

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LIGO Simulation parameters and tuning

- In VIRGO we chose to start from *physical* parameters
 - » Masses, characteristics of wires and blades, strength of magnets ...
 - » Some are better known, other are actually approximate
- An alternative approach would have been to see the mechanics as a black box
 - » One could define it as a MIMO model and then fit its parameters
 - » Advantage of generality, but total loss of contact with the instrument
- The tuning is a typical (hard) inverse problem
 - » Physical parameters as a basis
 - » Mode identification to select subsystems
 - » Parameter tuning using mode characteristics and TF measurements

Example: vertical performance



S.Braccini et al

- Impossible to measure the entire SA chain TF !
 - » Only the ITF shall have the sensitivity needed
 - Possible to measure stage by stage
- Compose the partial SA TFs to obtain the full one

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Seismic noise and sensitivity



- Vertical seismic noise dominates over horizontal, in the detection band
- The stage-by-stage measurement agrees with simulation → confidence that no effect has been forgotten
- Blade resonances come close to the sensitivity curve, in quiet conditions: for safety, in the Cascina suspensions they have been damped.

Passive isolation is not enough



Courtesy G.Losurdo

- At low frequency the SA is an "amplifier"
- An array of sensors picks up the motion, where is larger, and feedbacks it.
 - » Below 10÷20 mHz the system is "locked" to the ground
 - » In the [20mHz, 5Hz] it is locked to the inertial frame
- How this system performs?

Actuator response



- On top of the IP the sensors allow to measure the response to control forces
- The simulation can be tuned to reproduce the main features.

Data: courtesy by A.Gennai

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Response to seism: open loop...



- Assuming a model for the noise, the IP motion can be estimated.
- The absolute scale is wrong, but the main features are reproduced

Data: courtesy by G.Losurdo

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...closed loop



- » Left: simulation and measurement on top of the IP
- » Right: the residual predicted RMS on the test mass, using as input the measured motion on top of the IP

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Next step: steer the mirror

- Force response from the reference mass is simple
 - » Just the response of a pendulum
- More complex is the response from the marionette
 - » This stage is *necessary* for yaw and pitch, and *desirable* for coarse longitudinal action
- Problem: not easy to tune the simulation parameters
 - » Poor inputs
 - » No permanent sensors to monitor the motion of the elements.

LIGO A system identification problem...



With A.Di Virgilio et al.

- Isolate the less known element: the steering filter
- Suspend and add a mock payload
- Actuate using the standard coils
- Read the motion in 3D using small LVDTs mounted on its surface
- Register inputs and outpus, compare with the model and tune its parameter

...with some difficulties



- » Left: a fitted linear model between inputs and outputs successfully fits the output spectrum → good data quality
- » Right: a model based on physics gives a less successful fit → extra DOF are excited, which the model ignores.

More accuracy is needed



- To be able to reach high gains, transfer functions like this are needed with good accuracy
- The pole/zero structure depends on the gain of the inertial damping!
- A (well tuned) model is indispensable to build an adaptive control loop

LIGO TF measurement: the "lavatrice"



- » A fiber Mach-Zender interferometer
- » Angular motion by beam translation, using a PSD, in open loop.
- » Longitudinal motions picked up by fringe interference, in closed loop, using a PZT to lock the interferometer
- » A single stage suspension is sufficient to reduce the seismic noise at a level which allows TF measurements.

Angular (yaw) TF on West Input



- The color of the data points is related to different runs, with different levels of input force.
- » The line is NOT the simulation, but a zero/pole fit to the data!!
- The measurement is taken with inertial damping on, to have some control on the mirror position

Exp. Data: L.Di Fiore

Still the yaw, against simulation



- Main features ok
- Low frequency resonances shifted: wrong momenta of inertia for the filters.
- After tuning, one expects good agreement ...

Exp. Data: L. Di Fiore

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Pitch motion



- Similar scenario: reasonable agreement, some tuning to be done.
- Note the zero/pole at 30 mHz in the experimental data: it is a mixing with longitudinal motion.

Exp. Data: L. Di Fiore

Longitudinal TF from steering filter



- Quite a large disagreement ... still some tuning work is required!
- Note however that this TF is less critical for the CITF: the noise level allows to lock from the reference mass

Exp. Data: L. Di Fiore

From the reference mass



- No surprises, and no need to compare with the simulation.
- Note however that the Q factor cannot be deduced from this curve: the measurement time was limited

Exp. Data: L. Di Fiore

Mirror longitudinal control



- » Should use both marionette and reference mass
- » Simple filter on the marionette, plus compensator for poles and zeros.
- » Unity gain set around 30 Hz
- Include saturation and digitization effects, electronic noise in DAC and ADC ...

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Central InTerFerometer locking



- While long arms are being build, VIRGO exercises locking in the central Michelson and the CITF
- Currently the MC and the Michelson are operated as separated ITF.
- Once both are locked, they shall be operated as a single instrument, until the arms are commissioned

Example: lock acquisition



- Simple Michelson (VIRGO CITF configuration)
- Longitudinal control using both marionetta and reference mass
- Initial experiments, still with a low gain
- "d1" is the dark fringe photodiode,
- Suspensions, optics, locking filters, all is simulated within the SIESTA package
- Soon in reality!

Courtesy of M.Barsuglia

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