

Proposal

Thermal and Thermoelastic Noise Research for Advanced LIGO Optics

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Outline

- Objective and Approach
- Output of the Prior NSF Support
 - Laser phase noise formulas for optical resonator and delay line
 - Coating noise estimation
- Proposed Work
 - Coating noise studies for mirrors of edge geometry
 - Thermo-elastic noise of coated mirrors
 - Thermal & thermo-elastic noises of realistic mirror designs
- Tasks & Time Lines
- Summary
- (Broader Impacts)

Objective and Approach

Objectives

- To develop laser-phase-noise formulas
 - Green's-function-based
 - Analytical and computational
 - two-point laser-phase-fluctuation correlations
 - complex test mass objects.
- To extend the noise estimation method to thermo-elastic noises.
- To estimate thermal and thermo-elastic noises of coated mirrors for advanced LIGO designs.
- To examine merits of interferometer design options for future LIGO.

Approaches

- Calculate phase noise via Green's function method
 - Elasticity
 - Thermo-elasticity & thermal diffusion
- Analytical mirror models
 - Half space
 - Quarter space
 - Thin coating layer
- Numerical calculations
 - Realistic mirror shapes
 - Coating loss
 - Delay-line vs. Fabry Perot

Output to Date

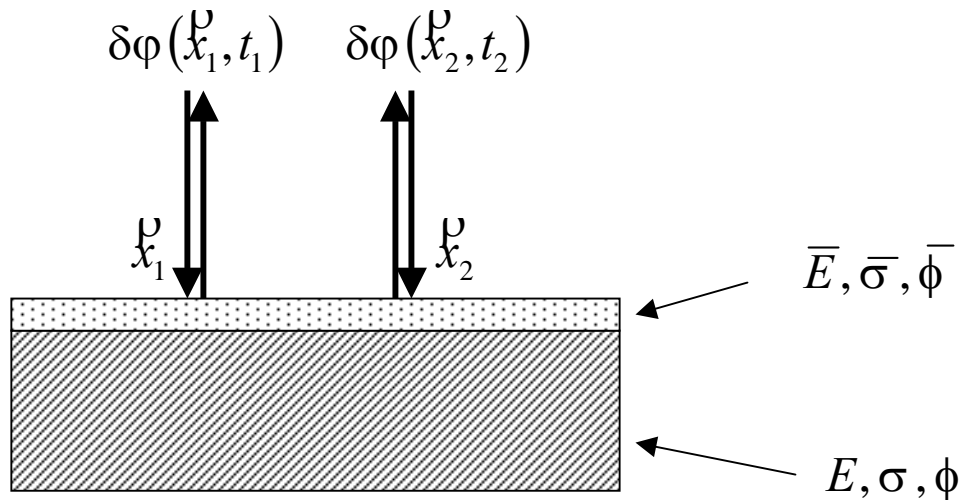
- Publications

- N. Nakagawa, Eric Gustafson, P. Beyersdorf and M. M. Fejer “Estimating the off resonance thermal noise in mirrors, Fabry-Perot interferometers and delay-lines: the half infinite mirror with uniform loss,” to appear in Phys. Rev. D
- N. Nakagawa, A. M. Gretarsson, E.K. Gustafson, and M. M. Fejer, “Thermal noise in half infinite mirrors with non-uniform loss: a slab of excess loss in a half infinite mirror,” submitted to Phys. Rev. D
- N. Nakagawa, E.K. Gustafson, and M. M. Fejer, “Thermal phase noise estimations for fabry-perot and delay-line interferometers using coated mirrors,” in preparation.
- D. Crooks, et al., “Excess mechanical loss associated with dielectric mirror coatings on test masses in interferometric gravitational wave detectors,” submitted to Classical and Quantum Gravity.
- Gregory M. Harry, et al., “Thermal noise in interferometric gravitational wave detectors due to dielectric optical coatings,” submitted to Classical and Quantum Gravity.

Phase-Noise Correlation

- Requirement
 - Compute laser phase noise correlation

$$\langle \delta\varphi(x_1^p, t_1) \delta\varphi(x_2^p, t_2) \rangle = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{-i\omega(t_1-t_2)} S_\varphi(\omega, x_1^p, x_2^p)$$



Ex. Coating noise model

Intrinsic Thermal Noise

- Phase Noise Formula

$$S_\phi(\omega, \rho_1, \rho_2) = 4k^2 \frac{2k_B T}{\omega} \iint dS' \int dS'' \psi_{00}^w(\rho' - \rho_1) \psi_{00}^w(\rho'' - \rho_2) \\ \times \int_V d^3x [\partial_i \chi_{nj}^\omega(x, \rho'; c')] c''_{ijkl}(x) [\partial_k \chi_{nl}^\omega(x, \rho''; c')]^*$$

$$c_{ijkl}^\omega = c'_{ijkl} - i c''_{ijkl} \approx [1 - i\phi(\omega)] c'_{ijkl}$$

$S_\phi(\omega, \rho_1, \rho_2)$	the two-point laser-beam phase-noise power-spectrum correlation
$\langle u_i(\rho^l) u_j(\rho^r) \rangle_\omega$	the displacement spectral correlation
ρ_1, ρ_2	The laser beam reflection points (the beam centers)
$\psi_{00}^w(\rho) \propto e^{-2 \rho ^2/w^2}$	the Gaussian laser-beam profile function

w, k	the laser beam spot size (amplitude radius), and wave number
χ_{ij}^ω	elastic Green's function
$c_{ijkl} [c'_{ijkl}, c''_{ijkl}]$	elastic constants [dispersive and absorptive parts]
$\phi(\omega)$	loss function
k_B, T	Boltzmann constant, and temperature

Fluctuation-Dissipation Relation

- Surface Force density \rightarrow Strain

$$F\Psi_{00}^w(\hat{r}' - \hat{r}_1) \Rightarrow u_{ij}^\omega(\hat{x}; \hat{r}_1) = F \iint dS' [\partial_i \chi_{nj}^\omega(\hat{x}, \hat{r}'; c')] \Psi_{00}^w(\hat{r}' - \hat{r}_1)$$

$$S_\varphi(\omega, \hat{r}_1, \hat{r}_2) = 4k^2 \frac{2k_B T}{\omega} \frac{1}{F^2} \int_V d^3x [u_{ij}^\omega(\hat{x}; \hat{r}_1)] c''_{ijkl}(\hat{x}) [u_{kl}^\omega(\hat{x}; \hat{r}_2)]^*$$

- Landau-Lifshitz

$$\overline{E_{mech}} = \underbrace{-\left(\kappa/T_0\right) \int dV \left(\nabla T\right)^2}_{1 \ 4 \ 4 \ 2 \ 4 \ 4 \ 3} \text{ thermo-elastic} - \underbrace{\frac{\omega^2}{2} \int dV u_{ij}^\omega \eta_{ijkl} u_{kl}^{\omega*}}_{1 \ 4 \ 4 \ 2 \ 4 \ 4 \ 3} \text{ viscous}, \quad c''_{ijkl} = \underbrace{\omega \eta_{ijkl}}_{1 \ 2 \ 3} \text{ viscous} + \underbrace{\Lambda}_{structural}$$

$$S_\varphi(\omega, \hat{r}, \hat{r}) = 4k^2 \frac{4k_B T}{\omega^2} \frac{1}{F^2} \left(-\overline{E_{mech}} \right)$$

Various Optical Configurations

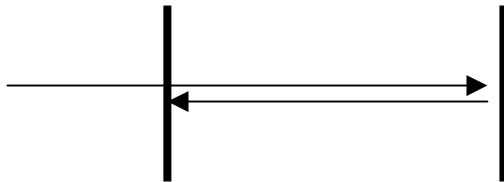
- Phase noise formulas

Computed explicitly for

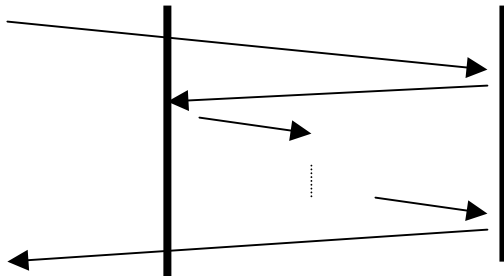
- Single-reflection mirror



- Fabry-Perot resonator



- Optical delay line



$$S_{\phi}^{Single}(\omega) = S_{\phi}(\omega, \rho, \rho)$$

$$S_{\phi}^{FP}(\omega) = \left[\frac{(1+r_I)^2}{1+r_I^2} \right] \left[1 - \frac{2r_I}{1+r_I^2} \cos 2\omega\tau \right]^{-1} \left[S_{\phi}^E(\omega) + r_I^2 S_{\phi}^I(\omega) \right]$$

$$S_{\phi}^{DL}(\omega) = \sum_{n=1}^N S_{\phi}^E(\omega, \mathbf{r}_n, \mathbf{r}_n) + 2 \sum_{n=2}^N \sum_{q=1}^{n-1} \cos[2(n-q)\tau\omega] \cdot S_{\phi}^E(\omega, \mathbf{r}_n, \mathbf{r}_q) \\ + \sum_{n=1}^{N-1} S_{\phi}^I(\omega, \mathbf{r}_n, \mathbf{r}_n) + 2 \sum_{n=2}^{N-1} \sum_{q=1}^{n-1} \cos[2(n-q)\tau\omega] \cdot S_{\phi}^I(\omega, \mathbf{r}_n, \mathbf{r}_q)$$

r_I	the input mirror reflection coefficient
τ	the transit time
$S_{\phi}^E(\omega), S_{\phi}^I(\omega)$	the single-reflection phase noises of the input and end-point mirrors.
ρ_n	the positions of the N-time reflections on the end-mirror surface
ρ_p	the positions of the (N-1)-time reflections on the input-mirror surface
E, σ	Young's modulus, and Poisson ratio

If Half space:

$$S_{\phi}(\omega, \rho_1, \rho_2) = \frac{8k_B T}{\sqrt{\pi}} \frac{\phi}{\omega} \frac{k^2}{w} \frac{1-\sigma^2}{E} e^{-(\rho_1-\rho_2)^2/2w^2} I_0\left(\frac{(\rho_1-\rho_2)^2}{2w^2}\right)$$

Fabry-Perot vs. Delay Line

- Fabry-Perot vs Delay lines
 - Analytical half-space mirror model
 - Fabry-Perot interferometer vs several delay lines
 - Storage time as proposed for LIGO II.
 - Delay line beam centers
 - evenly spaced
 - on a circle
- When the spots are not overlapping appreciably, the delay line is less noisy than the Fabry-Perot.
 - Noise levels are similar if
 - the spot circle radii comparable to the beam spot size
 - the spots are largely overlapping, and above several hundred Hertz.

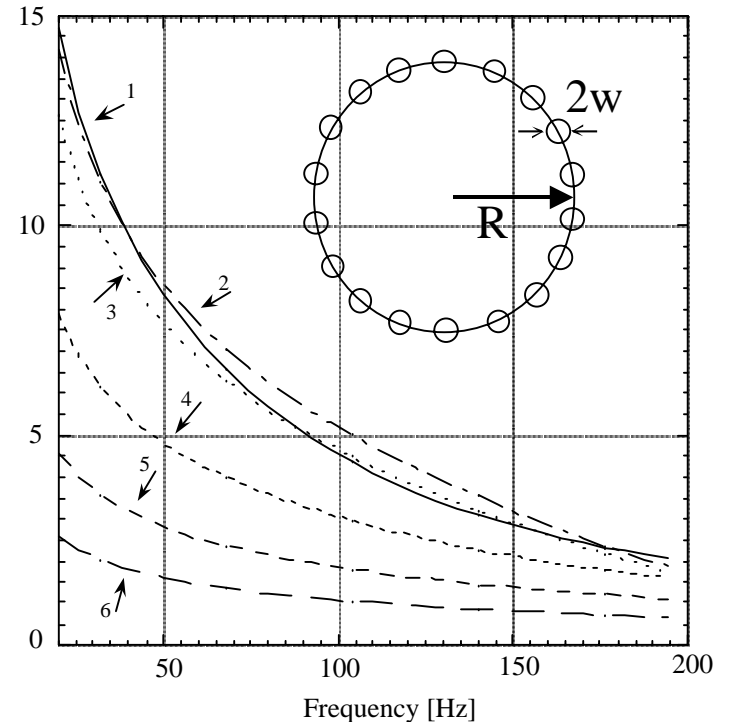
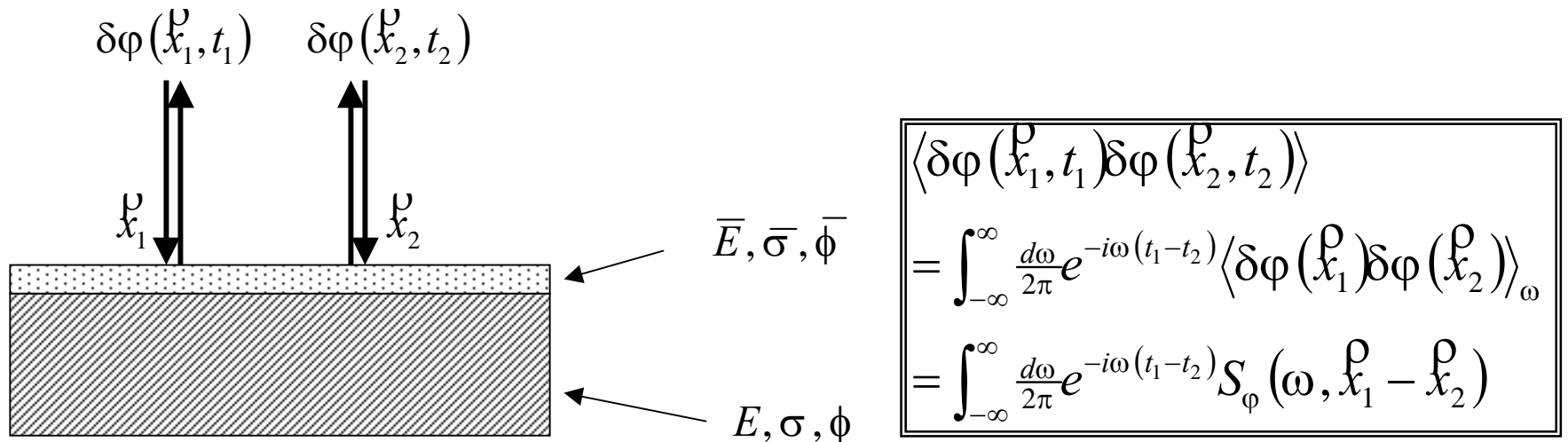


Figure 1: Comparison of the phase noise from a delay-line and a Fabry-Perot interferometer. The solid curve (1) is for a 4 km Fabry-Perot interferometer with an input mirror power reflectivity of $R_I=0.97$ and an end mirror power reflectivity of $R_E=1.00$. Curves 2, 3, 4, 5 and 6 correspond to 4 km delay lines all with 130 spots on the end mirror and laser beam spots of $1/e$ field radius w in a pattern with their centers on a circles of radius $R=w/3$ (2), $R=2w/3$ (3), $R=5w/2$ (4), $R=10w$ (5) and $R=20w$ (6) where $w=3.5$ cm, the spot size used for both mirrors of the Fabry-Perot interferometer. The mirror Q is assumed to be 3×10^8 and the material properties are those of Sapphire $E=71.8$ GPa and $s=0.16$ however we are treating sapphire as isotropic for the purpose of this illustration and assuming a single loss function.

Coating Noise; Problem Statement

- Coating noise model
 - Based on half-space mirror model
 - a lossy layer (thickness d) on a lossy host material
 - Requirement: Compute laser phase noise correlation
 - Approach: via analytical Green's function



Coating Noise

- Static Green's function for layer-on-substrate

$$\left[\chi_{z..}^{sos}(\tilde{p}, z) \right] = \frac{1}{p\bar{E}} \frac{1+\bar{\sigma}}{1-\bar{\sigma}} \left\{ \left[(2(1-\bar{\sigma}) - pz \cdot i\tau_1) G^{sos}(\tilde{p}, 0) + \frac{1}{2} pz \tau_3 \right] \cosh pz \right. \\ \left. + \left[(pz \tau_3 + (1-2\bar{\sigma})\tau_2) G^{sos}(\tilde{p}, 0) + \frac{3-4\bar{\sigma}}{2} - \frac{1}{2} pz(i\tau_1) \right] \sinh pz \right\}$$

$$\left[\Phi_{z..}^{sos}(\tilde{p}, z) \right] = \frac{1}{2(1-\bar{\sigma})} \left[2pz\tau_3 G^{sos}(\tilde{p}, 0) + 2(1-\bar{\sigma}) - pz \cdot i\tau_1 \right] \cosh pz \\ + \frac{1}{2(1-\bar{\sigma})} \left[2[1 - pz(i\tau_1)] G^{sos}(\tilde{p}, 0) + pz\tau_3 - (1-2\bar{\sigma})\tau_2 \right] \sinh pz$$

where

$$G^{sos}(\tilde{p}, 0) = \Delta^{-1} \left\{ \alpha(1-\sigma) \cosh 2pd + \frac{1}{4(1-\bar{\sigma})} \left\{ \frac{3-4\bar{\sigma}}{2} + \alpha(1-2\sigma)(1-2\bar{\sigma}) + \frac{\alpha^2}{2}(3-4\sigma) \right\} \sinh 2pd \right. \\ \left. + \left\{ \begin{array}{l} \frac{\alpha}{2} [(1-2\sigma) \cosh 2pd + \frac{1-\sigma}{1-\bar{\sigma}} (1-2\bar{\sigma}) \sinh 2pd] \\ + \frac{1}{8(1-\bar{\sigma})^2} [(1-2\bar{\sigma})(3-4\bar{\sigma}) - 2\alpha(1-2\sigma)(3-4\bar{\sigma}) + \alpha^2(1-2\bar{\sigma})(3-4\sigma)] \sinh^2 pd \\ - \frac{1}{8(1-\bar{\sigma})^2} (1-\alpha)[1+\alpha(3-4\sigma)](pd)^2 \end{array} \right\} \tau_2 \right. \\ \left. + \frac{1}{4(1-\bar{\sigma})} pd(1-\alpha)[1+\alpha(3-4\sigma)]\tau_3 \right\}$$

$$\Delta \equiv \left\{ \cosh pd - \frac{1}{2(1-\bar{\sigma})} [(1-2\bar{\sigma}) - \alpha(3-4\sigma)] \sinh pd \right\} \left\{ \cosh pd + \frac{1}{2(1-\bar{\sigma})} [(1-2\bar{\sigma}) + \alpha] \sinh pd \right\} \\ + \frac{1}{4(1-\bar{\sigma})^2} (1-\alpha)[1+\alpha(3-4\sigma)](pd)^2$$

$$\alpha \equiv \frac{1+\bar{\sigma}}{1-\bar{\sigma}} \frac{\bar{E}}{E}, \quad \tau_{1,2,3} = \text{Pauli matrices}$$

D. M. Burmister, J. Appl. Phys. 16, 89-94, 1945.

Coating Noise

- Intrinsic Thermal Phase-Noise Estimation

$$S_{\phi}^{coating}(\omega, \vec{r}) = 4k^2 \cdot \frac{2k_B T}{\omega} \frac{1}{\sqrt{\pi} w} \times \left\{ \begin{aligned} &\phi \cdot \frac{1-\sigma^2}{E} e^{-r^2/2w^2} I_0(r^2/2w^2) \\ &+ \left[\begin{aligned} &\bar{\phi} \frac{(1-2\bar{\sigma})(1+\bar{\sigma})}{1-\bar{\sigma}} \frac{1}{\bar{E}} + (\bar{\phi} - 2\phi) \frac{(1+\sigma)^2(1-2\sigma)^2}{1-\bar{\sigma}^2} \frac{\bar{E}}{E^2} \\ &- \phi \frac{2(1+\sigma)(1-2\sigma)\bar{\sigma}}{1-\bar{\sigma}} \frac{1}{E} \end{aligned} \right] \frac{d}{\sqrt{\pi} w} e^{-r^2/w^2} + O(d^2/w^2) \end{aligned} \right\}$$

$S_{\phi}^{coating}(\omega, \vec{r})$	The phase noise two-point correlation for a coated half-space mirror; double-sided
E, σ, ϕ	Young's modulus, Poisson ratio, and loss function of the substrate material
$\bar{E}, \bar{\sigma}, \bar{\phi}$	Those of the coating material
d	The coating thickness
$I_0(z)$	The 0-th order modified Bessel function of the first kind; $I_0(0)=1$

$\omega = 2\pi f$	Frequency
\vec{r}	a relative position vector between the two beam centers on the coating surface; $\vec{r} = 0$ for a single reflection.
k, w	The laser beam wave number, and spot size (amplitude radius)
k_B, T	The Boltzmann constant and the temperature

Coating Noise

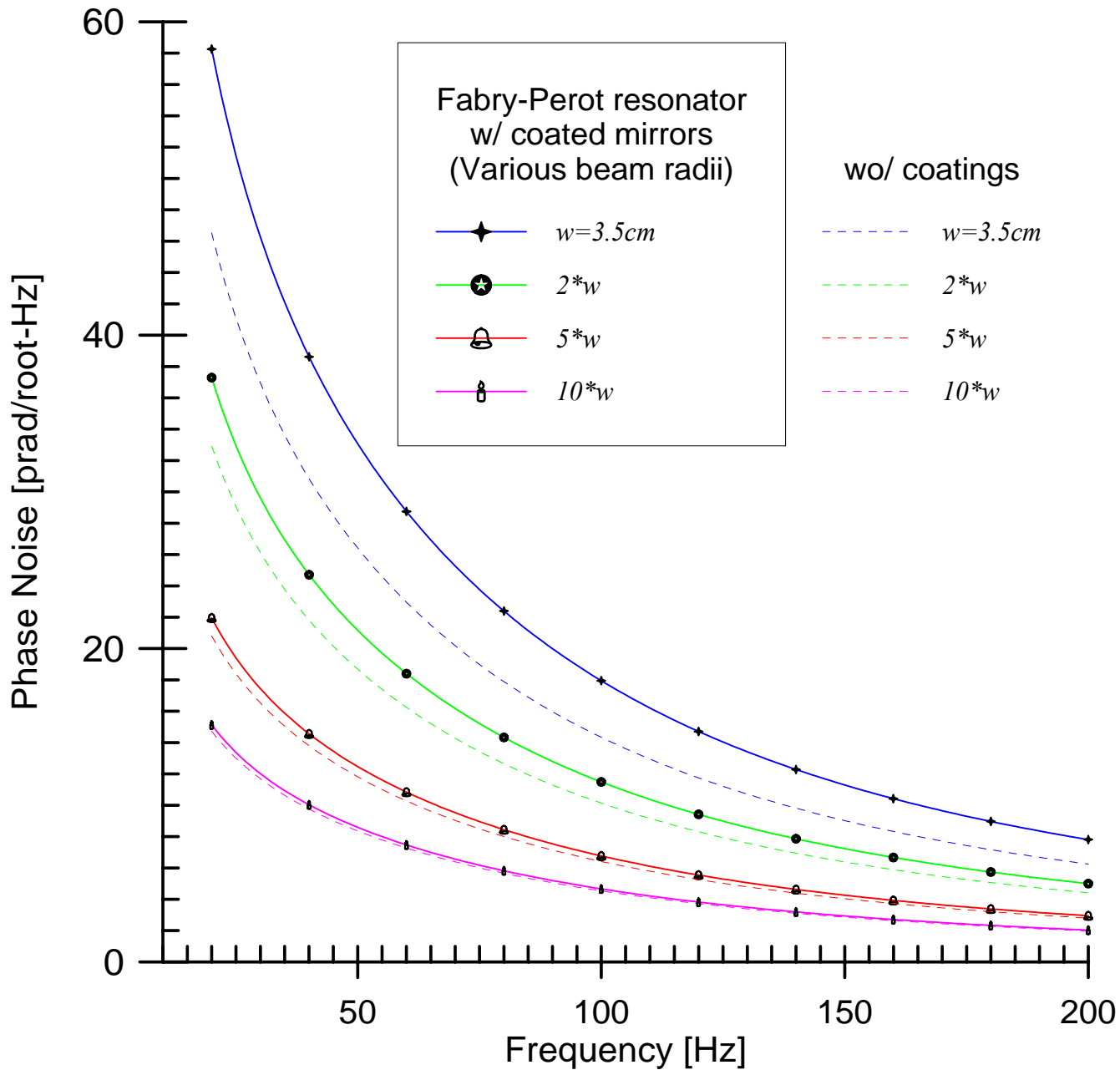
- Resonator

$$S_{\phi}^{coating}(\omega, 0) = 4k^2 \cdot \frac{2k_B T}{\omega} \frac{1}{\sqrt{\pi w}} \left\{ \phi \cdot \frac{1 - \sigma^2}{E} + \left[\phi \frac{(1 - 2\bar{\sigma})(1 + \bar{\sigma})}{1 - \bar{\sigma}} \frac{1}{E} + \Lambda \right] \frac{d}{\sqrt{\pi w}} \right\}$$

- Delay lines

$$S_{\phi}^{coating}(\omega, r) = 4k^2 \cdot \frac{2k_B T}{\omega} \frac{1}{\sqrt{\pi w}} \times \left\{ \phi \cdot \frac{1 - \sigma^2}{E} e^{-r^2/2w^2} I_0\left(\frac{r^2}{2w^2}\right) + \left[\phi \frac{(1 - 2\bar{\sigma})(1 + \bar{\sigma})}{1 - \bar{\sigma}} \frac{1}{E} + \Lambda \right] \frac{d}{\sqrt{\pi w}} e^{-r^2/w^2} \right\}$$

$\xrightarrow{r \rightarrow \infty} \frac{1}{r}$

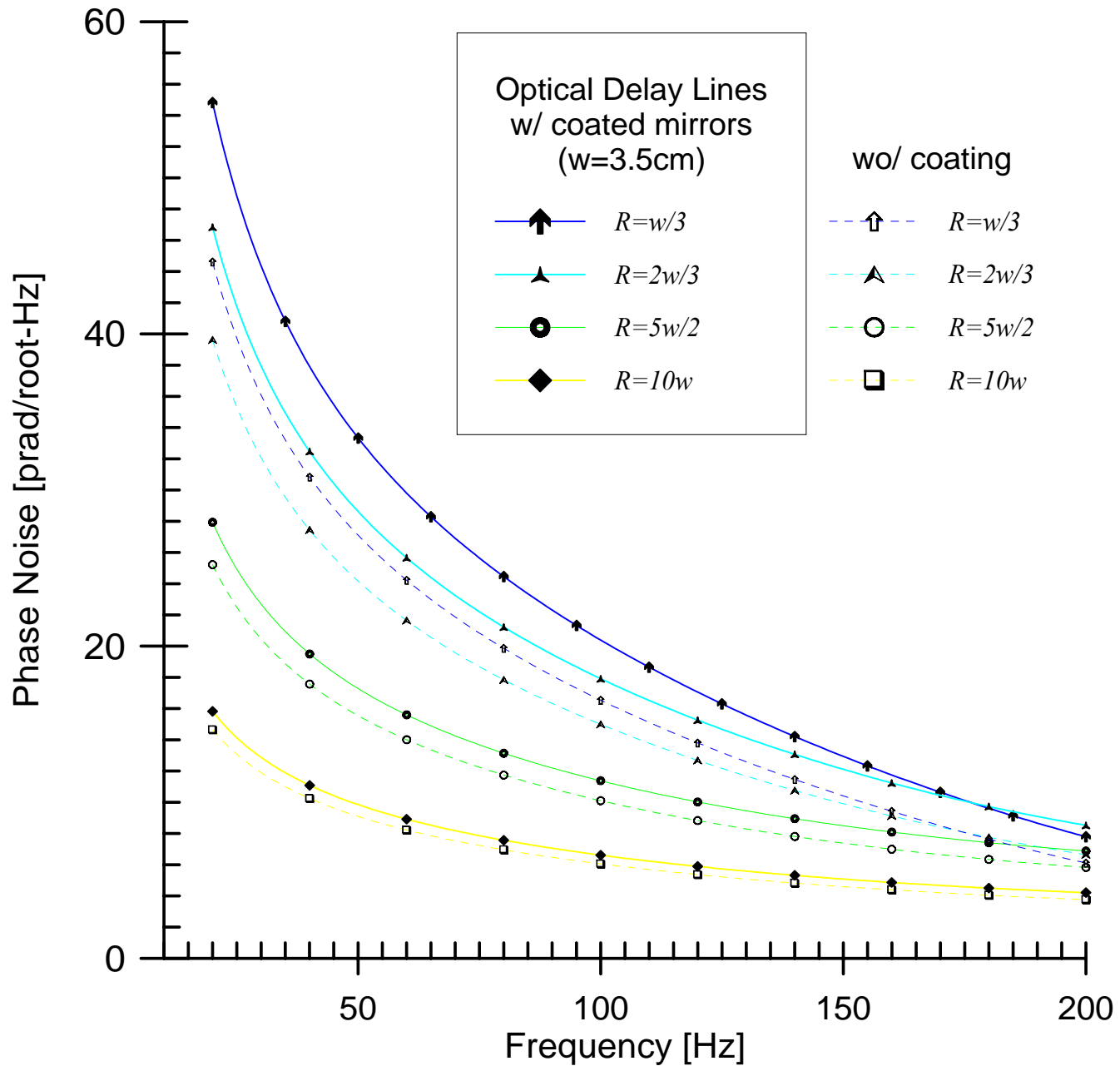


Substrate=Fused silica

$$\begin{cases} E = 72.6 \text{ GPa} \\ \sigma = 0.16 \\ Q = 3 \times 10^7 \end{cases}$$

Coating= average of
 Al_2O_3 and Ta_2O_5
(Crooks et al.)

$$\begin{cases} E = 260 \text{ GPa} \\ \sigma = 0.26 \\ Q = 1.6 \times 10^4 \end{cases}$$

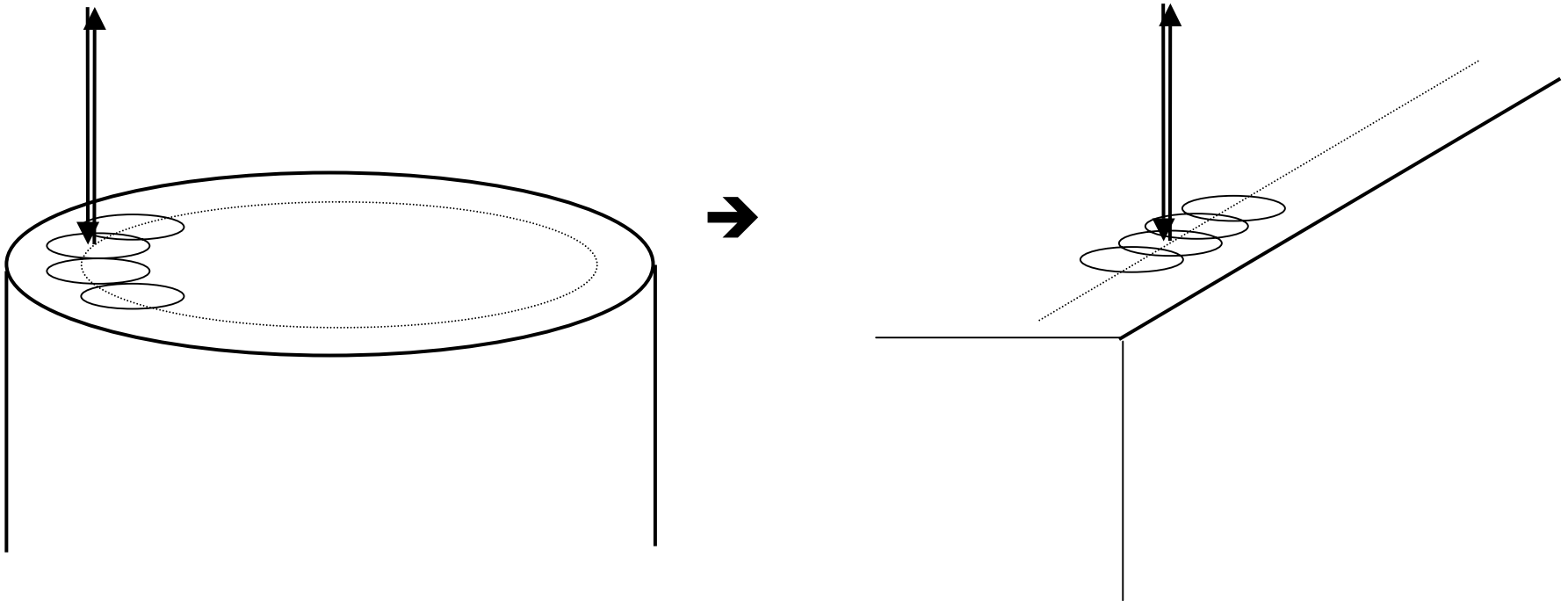


Proposed Work

- Edge effects on coated mirror noise
 - Analytical quarter-space model
- Extension to thermo-elastic noise
 - Coated mirrors
 - Delay-line interferometers
 - Mirrors at low temperatures
- Thermal & thermo-elastic noises
 - Mirrors of realistic shapes
 - Numerical Green's function calculation

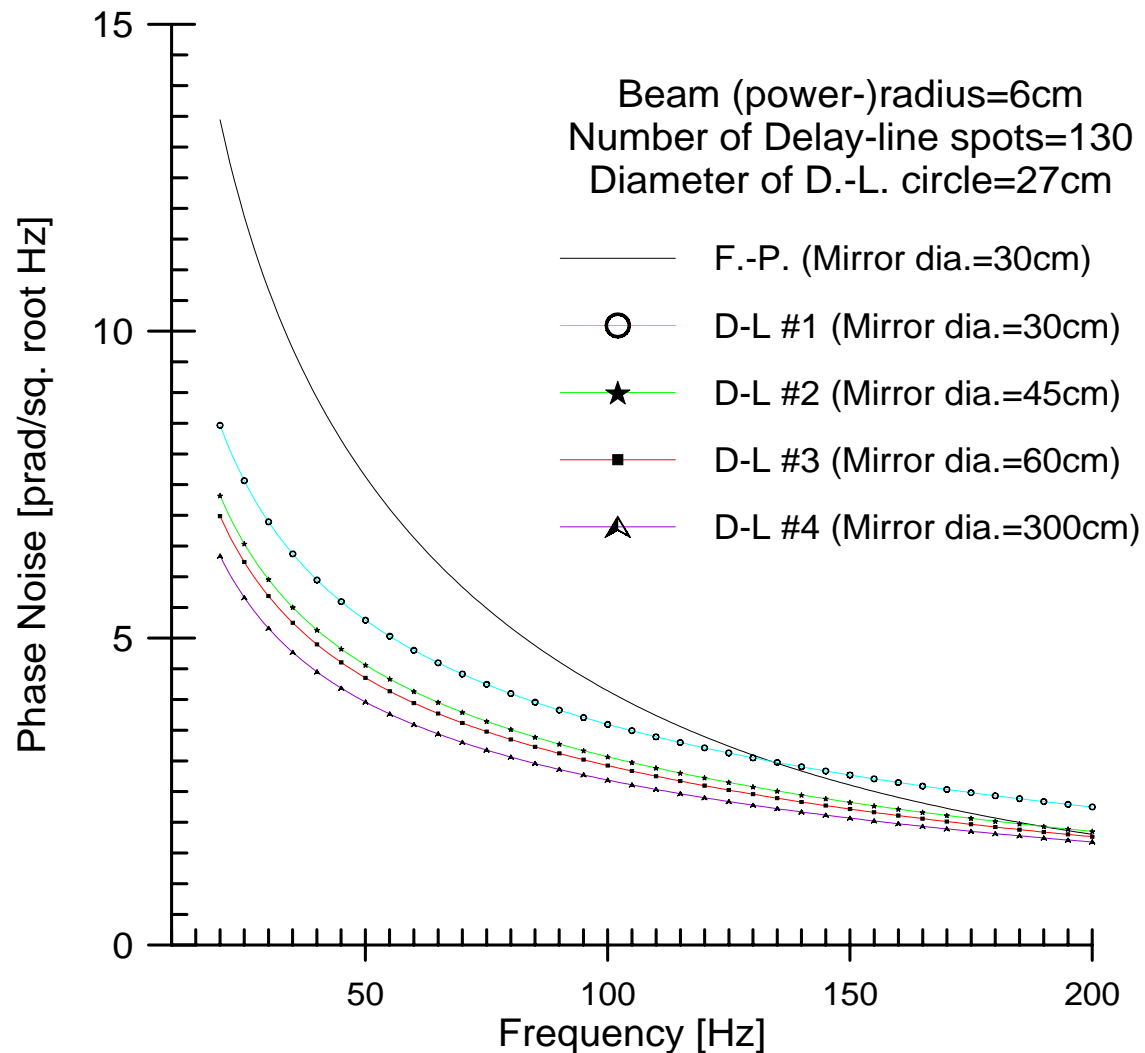
Quarter-Space Mirror Model

- Edge effects on noise calculation
 - Analytical model



Quarter-Space Mirror Model (con't)

- Finite-mirror size effect on thermal noise (scalar model)



Thermo-elasticity & Thermal diffusion

- From previous identification

$$S_\varphi(\omega, \rho, \rho) = 4k^2 \frac{4k_B T}{\omega^2} \frac{1}{F^2} \left(-\overline{\mathcal{E}_{mech}} \right) \quad \mathcal{E}_{mech} = -(\kappa/T_0) \int_V dV (\nabla T)^2 - \int_V dV u_{ij} c_{ijkl} u_{kl}$$

$$\overline{\mathcal{E}_{mech}} = -(\kappa/T_0) \int_V dV (\nabla T)^2 - \frac{\omega}{2} \int_V dV u_{ij}^\omega c_{ijkl}'' u_{kl}^{\omega*}$$

- If the adiabatic condition holds

$$T - T_0 = -\frac{T_0 \rho \alpha}{C_p} (v_l^2 - v_t^2) u_{ll}$$

$$S_\varphi^{th-el}(\omega, \rho_1, \rho_2) = 4k^2 \frac{2k_B T}{\omega} \iint dS' \int dS'' \psi_{00}^w(\rho' - \rho_1) \psi_{00}^w(\rho'' - \rho_2) \\ \times \frac{\kappa}{\omega} T \left[\frac{\rho \alpha}{C_v} (v_l^2 - v_t^2) \right]^2 \int_V d^3 x \left[\partial_i \partial_l \chi_{nl}^\omega(x, \rho'; c') \right] \left[\partial_i \partial_k \chi_{nk}^\omega(x, \rho''; c') \right]^*$$

Numerical elasticity

- Noise study:
 - Cylindrical mirrors
 - Numerical Green's function computation
 - Betti-Rayleigh-Somigliana formula
 - Nodal discretization by the boundary element method.
 - Cylindrical mirror model
 - Single reflection
 - Fabry-Perot
 - Delay-line
 - Demonstrate the delay-line vs. resonator assertion
 - Effects of mirror aspect ratios.

$$-\rho\omega^2 u_i - \partial_j T_{ij} = 0$$

$$-\rho\omega^2 \Gamma_{ik} - \partial_j \Phi_{ijk} = \delta_{ik} \delta(\mathbf{x} - \mathbf{x}_P)$$

$$\Rightarrow \chi^V(\mathbf{x}_P) u_k(\mathbf{x}_P) = \int_S dS_j \left[-u_i \Phi_{ijk}^\omega + T_{ij} \Gamma_{ik}^\omega \right]$$

$u_i(\mathbf{x})$	Displacement
$T_{ij}(\mathbf{x})$	stress tensor ($T_{ij} \equiv c_{ijkl} \partial_k u_l$)
$\Gamma_{ij}^\omega(\mathbf{x})$	the fundamental solution (Green's function)
$\Phi_{ijk}^\omega(\mathbf{x})$	$\equiv c_{ijlm} \partial_l \Gamma_{mk}$
c_{ijkl}	elastic constants
ρ	Density
V, S	volume of a region and its boundary surface
$\chi^V(\mathbf{x})$	the characteristic function of V ($=1$ inside V , $=0$ outside)

Static Elasticity

- Needs to avoid rigid-body motions

- Mass and moment of inertia

$$M = \int_V dV \rho, \quad I_{ij} = \int_V dV \rho (x^2 \delta_{ij} - x_i x_j)$$

- Given surface forces f can be counter-balanced by volume forces

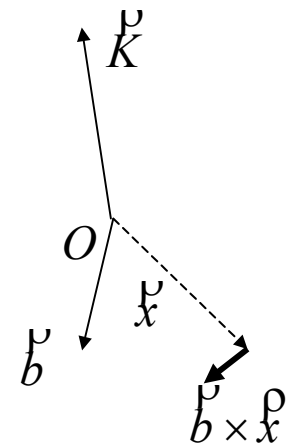
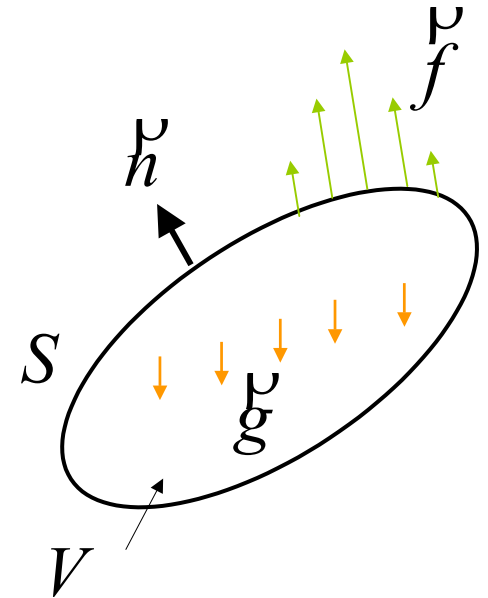
$$\boxed{\mathbf{g}(\mathbf{x}) = \mathbf{a} + \mathbf{b} \times \mathbf{x}, \quad \begin{cases} \mathbf{a} \equiv -\frac{1}{V} \mathbf{F} \\ \mathbf{b} \equiv -\frac{M}{V} \overline{\mathbf{I}^{-1}} \cdot \mathbf{K} \end{cases}}$$

where

$$\mathbf{F} \equiv \int_S dS \mathbf{f}(\mathbf{x}), \quad \mathbf{K} \equiv \int_S dS \mathbf{x} \times \mathbf{f}(\mathbf{x})$$

$$\mathbf{F}_g \equiv \int_V dV \mathbf{g}(\mathbf{x}), \quad \mathbf{K}_g \equiv \int_V dV \mathbf{x} \times \mathbf{g}(\mathbf{x})$$

$$\boxed{\mathbf{F} + \mathbf{F}_g = 0, \quad \mathbf{K} + \mathbf{K}_g = 0}$$



Summary

Accomplishments

- Obtained two-point phase noise correlation formulas
 - Resonator
 - Delay line
 - Delay lines can be quieter
- Given coating noise formulas
 - Relative magnitudes between coating and substrate noises.
 - Their behaviors against the beam size and other optical parameters

Proposed Activities

- Studies of intrinsic noises of coated mirrors
 - Thermal noise
 - Thermo-elastic noise
 - Relative significance of coating noise
 - Realistic mirror shapes

Broader Impacts

- Education
 - Student involvement through the collaboration with Stanford Group
- Contribution to LSC
 - The thermal noise estimation methodology
 - Impact on sorting out advanced optics designs
 - Impact on mirror material selection
- Impacts on other federally funded programs
 - Through advancement of computational physics methodology
 - NSF Industry/University Cooperative Research Center program
 - DOE/NERI project; “On-line NDE for advanced reactors”