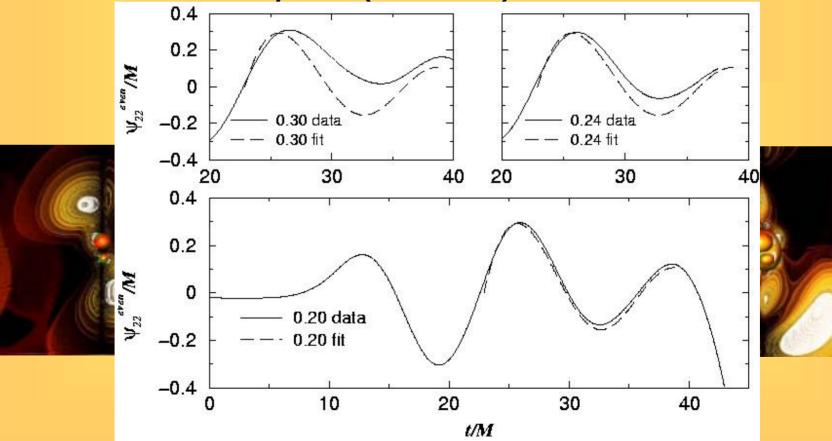
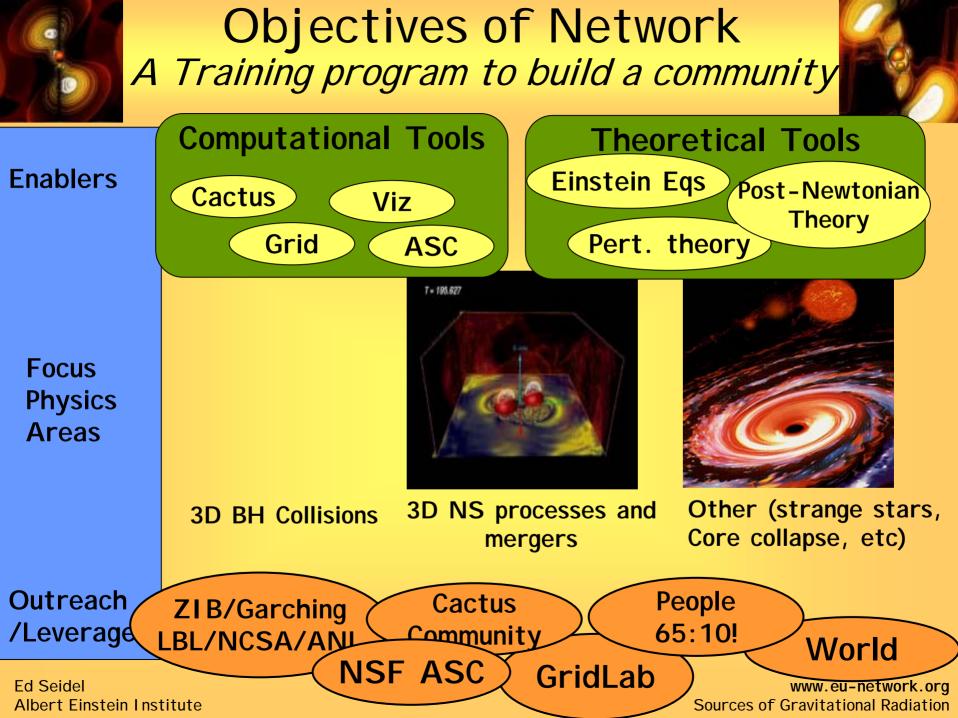
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EU Astrophysics Network Overview

Progress on 3D BH Collsions (me)
 3D Hydro (Luciano)



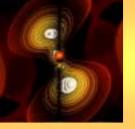


Our Team

7 Focus areas, links created, strengthened by Network

Post-Newtonian Schemes

allas **Potsdam CS** Efforts SOTON Worldwide Jena WashU **Friest** AUTH 12 10 1



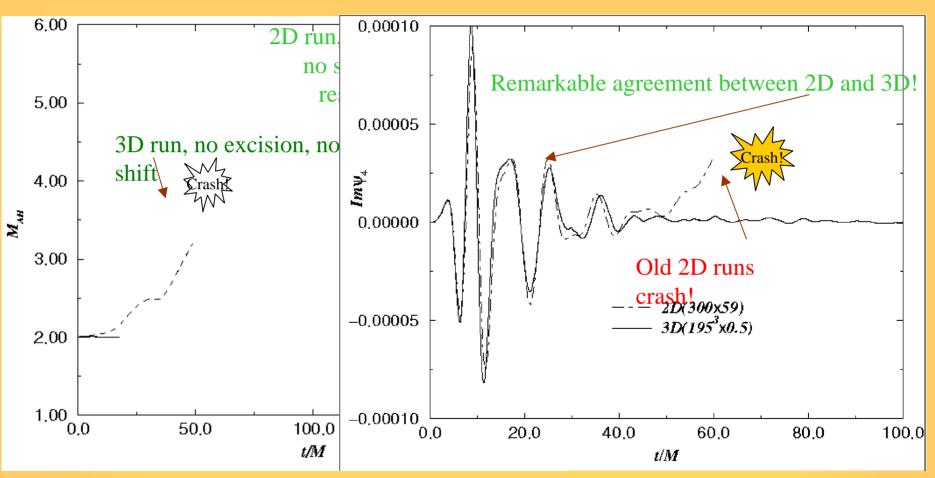
Steps to BBH Orbits (AEI/UT-Brownsville/Meudon)



- Step 0: Better Formulations (BSSN), gauges, excision
- Step 1: Do 3D single, rotating distorted BHs individually
 - *t* ~ *100s 1000s M* with convergence, solid masses, waveforms
 - Some instability may appear after t ~ 200 500M in full 3D
- Step 2: Do head -on BH collisions
 - 100s 1000s M with convergence, solid masses, waveforms
- Step 3: Go to co-rotating frame for orbits!
 - Can just about do this today...

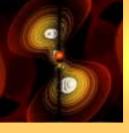
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Comparison of Schwarzschild Bighly Distorted Rotating BH

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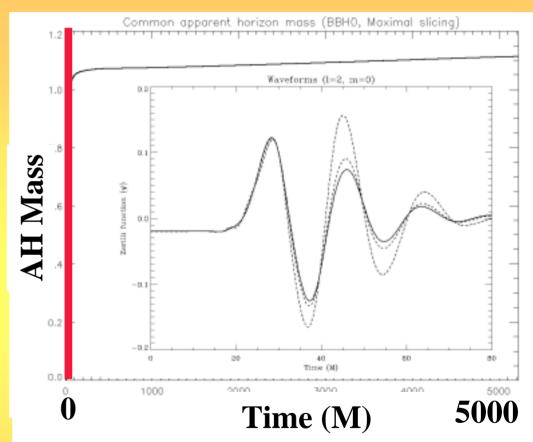


Step 2: 3D Head-On Collisions "Routine"



Brill-Lindquist

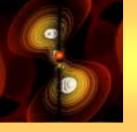
- Approx I SCO separation
- No momentum
- Similar to Misner μ=2.2
- New dynamic shift/lapse combo
- No Excision!!!



From Peter Diener...

www.eu-network.org Sources of Gravitational Radiation

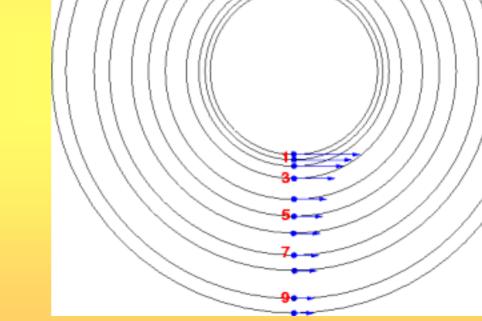
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Step 3: Co-rotating frame for Orbits

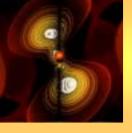


- New Techniques
 - Excision
 - Gauges, formulations
 - Co-rotating frame
- Baumgarte initial data
 - Estimated Innermost Stable Circular Orbit (ISCO)
 - Other circular orbits, too ("Pre-I SCO" sequence: PI -1-10)



From Denis Pollney...

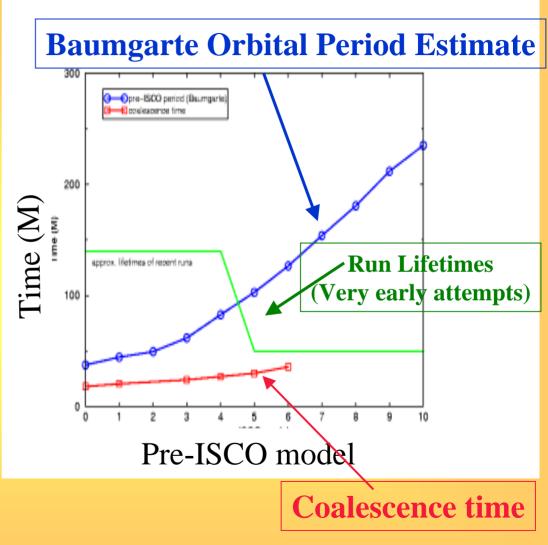
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Overview of Evolutions So Far

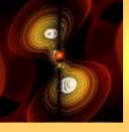


- Major problem is finding data sets that want to orbit!
- Baumgarte data
 - Seem to have momentum too low for circular orbits
 - Ω seems to be overestimated by factor of 2
- Evolutions solid (e.g., horizon mass accurate to 10% at t~100M)



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Example: Pre-ISCO 3



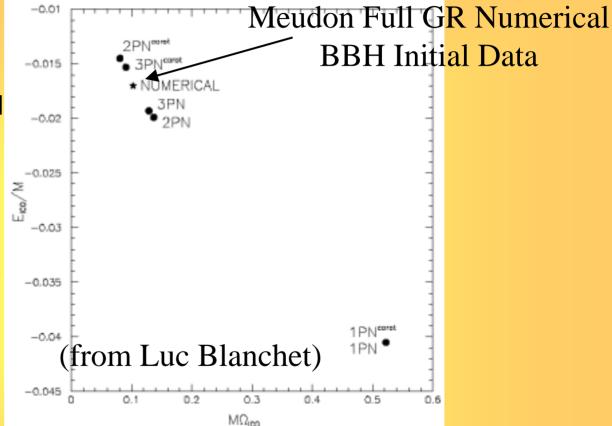
 $K_{zz} \text{ in x-y plane,} \qquad \text{AHs and } \Psi_4$ Showing excision (corotating) (Corotation taken out) $K_{zz} \text{ in x-y plane,} \qquad \text{AHs and } \Psi_4$ Showing excision (corotating) (Corotation taken out) $K_{zz} \text{ in x-y plane,} \qquad \text{AHs and } \Psi_4$ Sources of Gravitational Radiation

Increasing sophistication in IVP Grandclement (Meudon) BBH data



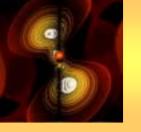
- Builds in approx helical Killing vector
- Has orbital period of

 2x that of potential method (also what we see in evolutions!)
- Post-Newtonian expansions seem to converge towards the numerical BBH data



Meudon Group: most astrophysically relevant binary BH data to date. Now in Cactus as EU Network Project

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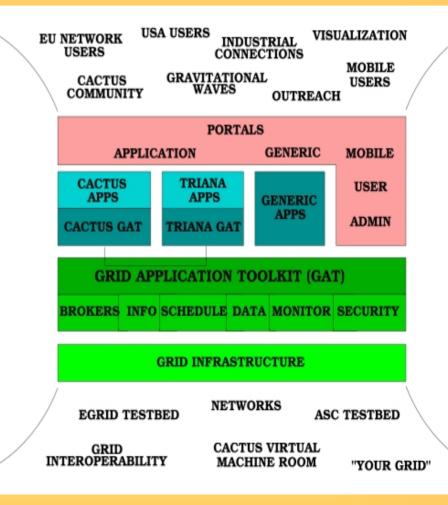
GridLab:

www.gridlab.org Enabling Dynamic Grid Applications

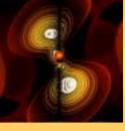


EU Project (€5M)

- AEI, ZIB, PSNC, Lecce, Athens, Cardiff, Amsterdam, SZTAKI, Brno, ISI, Argonne, Wisconsin, Sun, Compaq
- Grid Application Toolkit for application developers (API s/Tools)
- Develop new grid scenarios for 2 main apps:
 - Numerical relativity
 - Grav wave data analysis



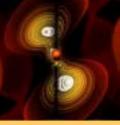
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Discovery Channel wants Pre-ISCO 3! This June...



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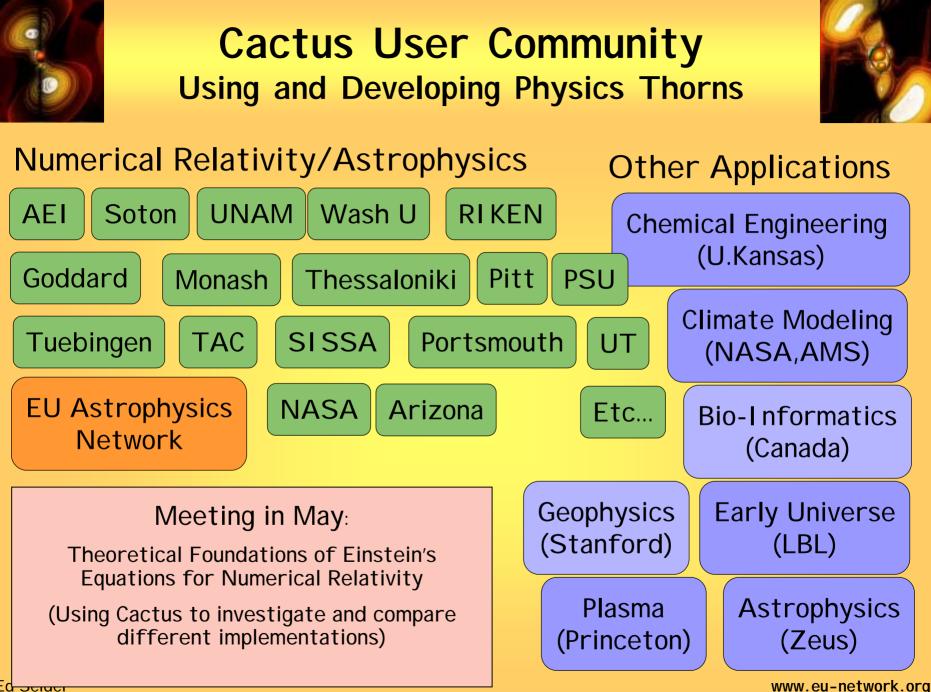


Building a Community (through) Code: Advantages



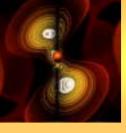
- Sharing of Expertise: No single group or community can do these problems
- Sharing of Code among projects
 - Infrastructure
 - BH routines apply directly to NS work, etc...
 - But free to keep routines in group as long as desired
- Better Code
 - Open source encourages people to be more careful in coding!
 - Encourages documentation
 - Encourages deeper thinking about how it interfaces to another code
- More trusted code, results
 - When code becomes open, and people can run it for themselves, they will begin to believe the results
- Improvements propagate quickly though community
- Not well accepted yet, but we are starting a good trend...

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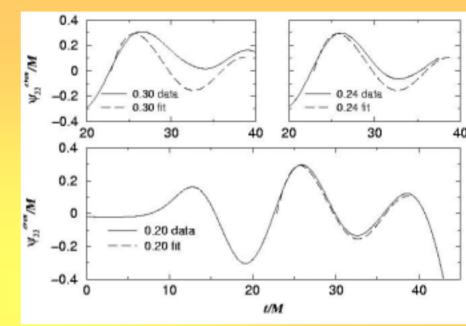
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www.eu-network.org Sources of Gravitational Radiation



Previous 3D BH Results new formulations, no shift, no excision

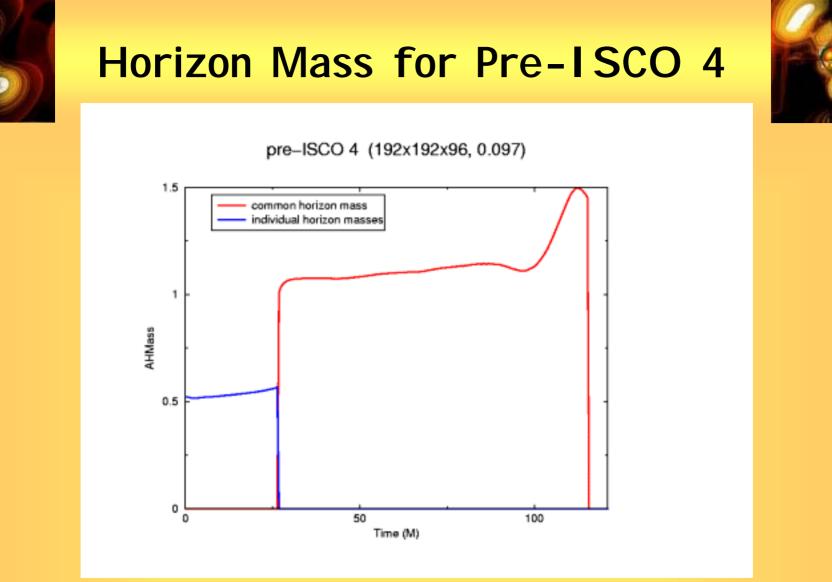




Grazing Collision 1999

- Big Advance: 3D, unequal mass, spin, J
- Physics!
- But Code crash by 40M

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Low resolution (192x192x96), but mass conserved to ~ 10% until ~100M
 Now exploring parameters to find orbital configurations

Now exploring parameters to find orbital configurations

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