

**Optics in VIRGO** 

VIRGO Status

Injection Sytem in Virgo

Selected topics in Frequency stabilization

François BONDU bondu@obs-azur.fr VIRGO CNRS – Observatoire de la Côte d'Azur, Nice September 2002 LIGO-G020425-00-Z







Setup overview Sensitivity (m/√Hz) Reliability Details (servo loops, suspensions, etc.)



# **VIRGD**

#### Setup in June and July 2002







#### Digital filter used to lock PR

(design Jean-Pierre Coulon)

#### Bode diagram

#### Nichols diagram







#### Engineering Run #4, July 12-15 2002





Engineering Run #4, July 12-15 2002

Power in recycling cavity Vs time (3 days)



- 5 unwanted losses of lock (similar to previous runs)
- lock acquisition longer than before  $\Rightarrow$  duty cycle  $\sim 80\%$



VIRGO Central Interferometer Sensitivity Progress





VIRGO status / Summary

Commissioning of central interferometer is finished (injection system 10 W, full suspensions, data acquisition system, interferometer control...)

20 W laser to be installed on site end of sept. 2002

First long arm cavity to be commissioned this fall

End of commissioning foreseen for end 2003.



### Injection system

All functions demonstrated to work, almost within Virgo specifications

- laser beam
- suspensions inertial damping and local controls
- alignments
- filtering with a long Mode Cleaner cavity
- stabilizations (power, frequency, position)



--: digital control



10 W Laser bench se

#### 10 W Laser bench setup (Frédéric Cleva)



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### Injected Laser Performances



- Master Laser 700 mW, Slave Laser 11 W
- Injection bandwidth 100 kHz
- 11 W, mono-mode, single-frequency
- In operation since July 1999
- > 7000 hours of operation
- remote monitoring



#### Power Stabilization



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### Input Mode Cleaner





#### MC parameters Measurement

Finesse	1120 +/- 10	Cavity decay time	Expected: 1000; excess of 2380 ppm losses
Contrast in reflection	79 %		
Transmission	33 %	Ptransmitted/Pincident	
Coupling on 02 + 20 modes	7 %	Measurement of transmitted power	
Round trip length	285,46 +/- 0,3 m	Notch in TF at f=FSR	
Curvature	186,5 +/- 0,4 m	Notch in TF at 356 kHz	



#### Injection system

All functions demonstrated to work, almost within Virgo specifications, High reliability of the laser system

•More on laser frequency specs: see 3<sup>rd</sup> part

- To be fixed:
- Back scatter of light in mode cleaner
- transmission of the mode cleaner
- automatic lock acquisition
- mode cleaner length noise



Frequency stabilization

Design of feedback for the virgo interferometer

Design of efficient and stable servo loops

« optical transfer function »

CARM sensing for LHO-4K





# Design of efficient digital filters

Programm CROSSXP, by Jean-Pierre COULON:

X = filter order

Enter required phase and gain margin,

Enter band (wrt unity gain) over which high attenuation is required

=> Finds a « good filter » (depends on the starting seed)



# Example of filter (worked successfully)

Bode plot







Attenation at 10\*UG = 4000Gain at UG/10 = 8000



### Frequency stabilization spec

- Unity gain of « Common Mode » servo has to be less than Free Spectral Range of long Fabry Perot cavity
- (something « bad » happens at f = FSR ??)
- Specs for prestabilization stage
  - = specs for light entering ITF / common mode loop gain



### Optical transfer function

• Definition:

Optical transfer function

Photodiode demodulated current

Frequency noise

STF: MATLAB model of interferometer to study sensing function

« object oriented », computes compound optical objects reflectivities and/or transmissions, and then optical transfer function.



### Ex. Of a simple FP cavity



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### TF for frequency stabilization

VIRGO Transfer function between frequency noise and demodulated signal in phase on D2





#### Design of correction filter

Design of correction filter:

Zeros: f=2.3 kHz, Q=2 f=2.3 kHz, Q=2

Poles: f=10 Hz, Q=2 f=10 Hz, Q=2

#### **OPEN LOOP TRANSFER FUNCTION**





Overall loop: stable, gain margin 2.7, phase margin 70 degrees

New reference solution





### LIGO4K common mode sensing

#### TRANSFER FUNCTION, used « as is », UG = 6 kHz



10<sup>4</sup> 10<sup>3</sup>  $10^{2}$ 10 magnitude 0 🗙 5 kHz X 10 kHz 15 kHz 20 kHz 10 10<sup>-2</sup> 10<sup>-3</sup> 10<sup>-4</sup>\_\_\_\_\_ -900 -720 -540 -360 -180 0 Angle (degrees)

LIGO Sensing function for CARM

Bode plot

Nichols plot



Difficult to achieve high unity gain for common mode servo

=> specs for prestabilization stage (MC for Virgo) higher

... If modulation frequency (Virgo) is switched from 6 to 18 MHz, then bandwidths of few 100's kHz possible !! (LIGO4K : window at 40 MHz).



# Conclusions

- Phase of "central interferometer" is finished.
- VIRGO Injection system: some issues to be fixed. Change of topology w/ respect to frequency stabilization.
- Frequency stabilization: tough to make a high bandwdith « CARM »; specs for frequency prestabilization higher.
- Would be nice to check the sensing function of the interferometer...

# **((O))** VIRGD

#### **Co-workers**

Nice group:

Nary Man Alain Brillet François Bondu Eric Chassande-Mottin Hervé Trinquet (Ph.D.) Frédéric Cleva Magali Loupias Henrich Heitmann Jean Cachenaut Jean-Pierre Coulon Jean-Yves Vinet

And :

Pisa group (suspension, vacuum) Orsay group (Control-command) Annecy group (vacuum tanks, DAQ) Napoli group (environment monitoring) Roma group (marionettas) Lyon and Paris groups (mirrors)