

# **GW pulsars in binary systems Sco X-1**

#### **C** Messenger and A Vecchio



#### THE UNIVERSITY OF BIRMINGHAM

LSC General Meeting LIGO Livingston Observatory 17<sup>th</sup> – 20<sup>th</sup> March, 2003

LIGO-G030131-00-Z



## GW pulsars in binary systems

- Accretion always considered a viable mechanism to sustain GW emission (Wagoner, 1984)
- LMXB frequencies are clustered. It is quite possible that what's at work is loss of angular momentum through GWs
  - Bildsten (1998), and Ushomirsky et al. (2000)
  - Wagoner (2002)
  - LMXBs could be detected by advanced LIGO
- More in general, we aim at extending the range of application of present search codes



(from Cutler and Thorne, 2002)

![](_page_1_Picture_11.jpeg)

![](_page_2_Picture_1.jpeg)

# The data analysis problem

- We need to take into account the additional Doppler effect produced by the source motion, which actually dominates Earth orbit and rotation:
  - 3 parameters for circular orbit
  - 5 parameters for eccentric orbit
  - and relativistic corrections if the system is "hard" enough
- In the short term, we concentrate on Sco X-1, therefore <u>for the time</u> <u>being the search is simpler because:</u>
  - We know where the source is: <u>no search over position</u>
  - We integrate only over a short period (~ 1 week) so that the signal is monochromatic: <u>no search over spin-down parameters</u>
  - We target a source that is in circular orbit and whose period is fairly well known: no search over P for T < 1 month
    - The search problem for Sco X-1 in S1 data: search over 2 orbital parameters (discrete mesh) + frequency
- The search is intrinsically wide-band and over a parameter space

![](_page_2_Picture_14.jpeg)

![](_page_3_Picture_1.jpeg)

## Code development

- 1. LALDemodBinary and LALComputeSkyBinary
  - Completed and tested using Teviet's injection code
  - We loose < 1% of the SNR (not clear yet where it comes from)
  - Final testing and validation starting next week
- 2. LALBinaryMesh
  - Conceived to handle template placement for a metric not necessarily flat, and with  $N_D > 2$
  - At present we use it for a 2D flat grid (clearly, we could have used Tevient's code)
    - Thoroughly tested
    - Final test/validation in progress: detection of signals with the expected loss of SNR
  - Plan for the near future: extend the testing/validation to include spindown parameters

![](_page_3_Picture_13.jpeg)

![](_page_4_Picture_0.jpeg)

![](_page_4_Picture_1.jpeg)

![](_page_4_Figure_2.jpeg)

![](_page_4_Picture_4.jpeg)

![](_page_5_Picture_0.jpeg)

### Tests on Gaussian stationary noise

![](_page_5_Figure_3.jpeg)

![](_page_5_Picture_4.jpeg)

![](_page_6_Picture_0.jpeg)

#### Tests with/without signal

![](_page_6_Figure_3.jpeg)

![](_page_7_Picture_0.jpeg)

## Distribution in the "adjacent bin"

F Statistic Probability distribution

![](_page_7_Figure_3.jpeg)

We don't fully understand yet this behaviour

![](_page_7_Picture_5.jpeg)

![](_page_8_Picture_1.jpeg)

## Activity for the immediate future

- Complete testing and validation of the demodulation code and computation of F-statistic:
  - Investigate SNR loss
  - Check distribution with and without signal
    - Wide variety of orbital parameters
    - Observation time
    - Length of short FFT
- Full testing and validation of template placement code:
  - So far tested only on Sco X-1 parameters T < 1 day
  - Need to increase observation time and wider range of parameters
- Ephemeris
  - Phase of Sco X-1 on the orbit at a given epoch
- "Loop" over bandwidth

![](_page_8_Picture_16.jpeg)

![](_page_9_Picture_1.jpeg)

## Work plan

- Main goal:
  - Code for coherent search over large bandwidth (10-100 Hz) for monochromatic signals and orbital parameters for e = 0
    - Upper-limit on Sco X-1 using S1 data
    - Ready to analyse S2 data
- Include spin-down
- Extend to other LMXBs (?)
- ...

![](_page_9_Picture_10.jpeg)