



Current Status of the 40m Detuned RSE Prototype

Seiji Kawamura

2004 Aspen Winter Conference on Gravitational Waves
Gravitational Wave Advanced Detector Workshops

Feb. 15 - 21, 2004

40m Team:

A. Weinstein, O. Miyakawa, M. Smith, J. Heefner, B. Abbott,
S. Vass, R. Taylor, L. Goggin, F. Kawazoe, S. Sakata, and R. Ward



Objectives

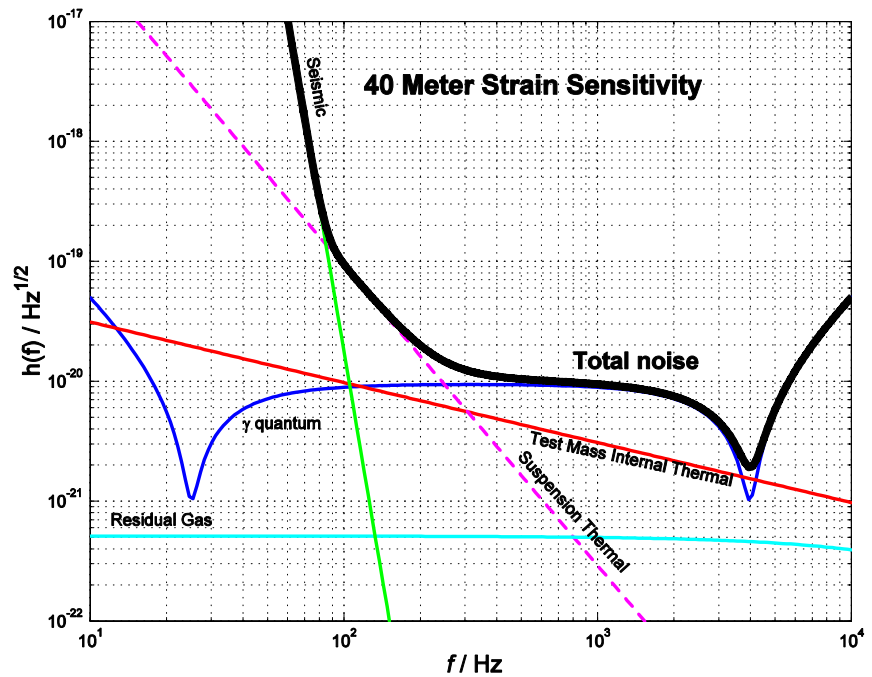
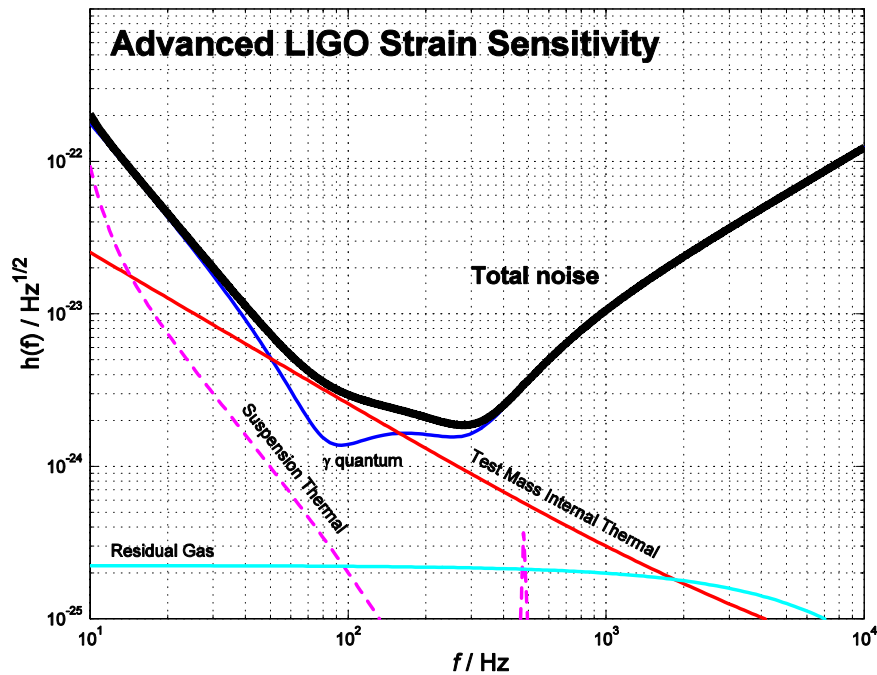
- Develop **lock acquisition procedure** of detuned Resonant Sideband Extraction (RSE),
- Characterize noise mechanism,
- Verify optical spring effect,
- Develop DC readout scheme,
- etc.

*for Advanced LIGO, LCGT,
and other future GW detectors*





Target Sensitivity of Advanced LIGO and 40m

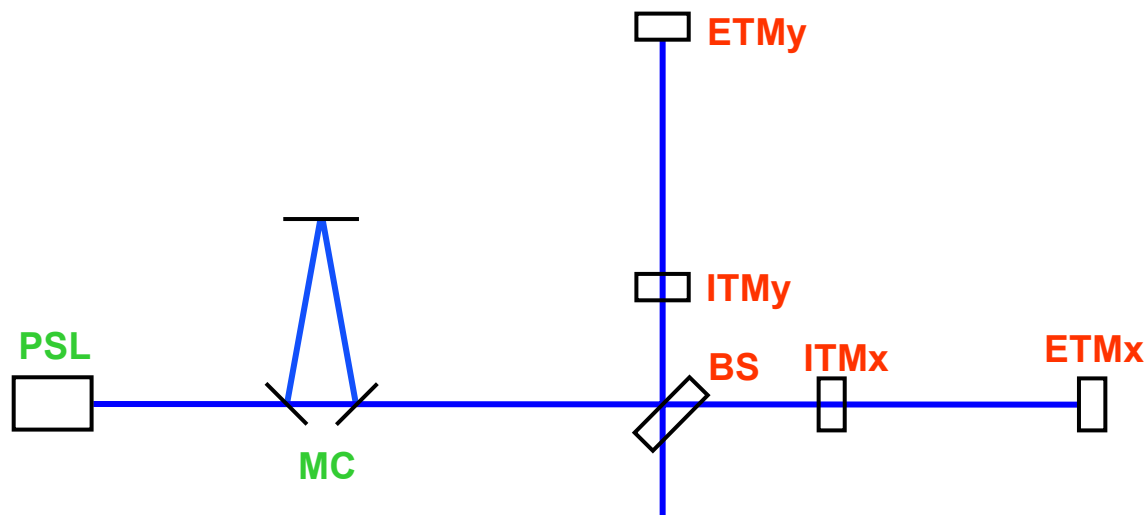




Important Achievement (1) Installation of FP Michelson

September, 2003

- Four TMs and BS: installed



Important Achievement (2)

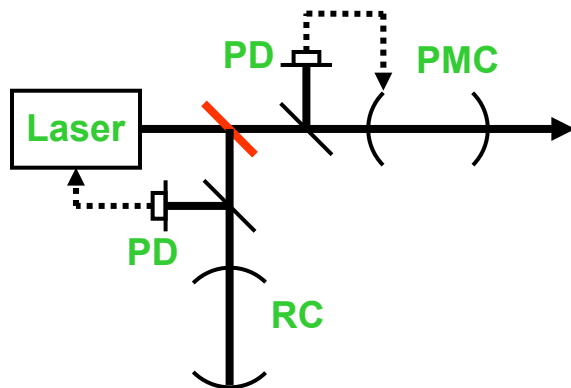
Modification of PSL

October 2003

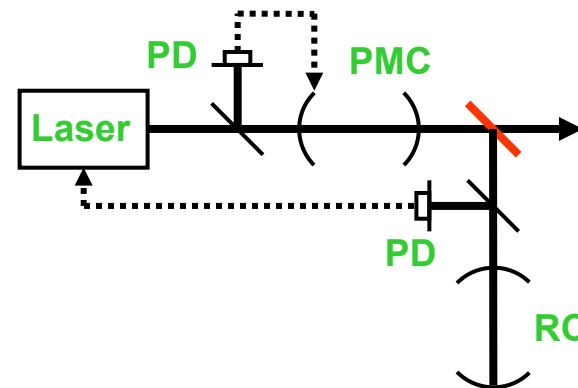
- Stabilize frequency of light with Reference Cavity (RF) **after** Pre-Mode Cleaner (PMC) instead of **before** PMC
- Noise sources associated with PMC and steering optics **after** PMC should be suppressed!

[S.Kawamura, "Configuration Study of Pre-Mode Cleaner and Reference Cavity in the 40m PSL System, LIGO-T030149-00-R (2003)]

[C. Mow-Lowry, R. Abbott, and B. Abbott, "Frequency Stability Servo Modifications Made at the 40 meter Laboratory", LIGO-T030205-00D (2003)]



Previous Configuration



New Configuration

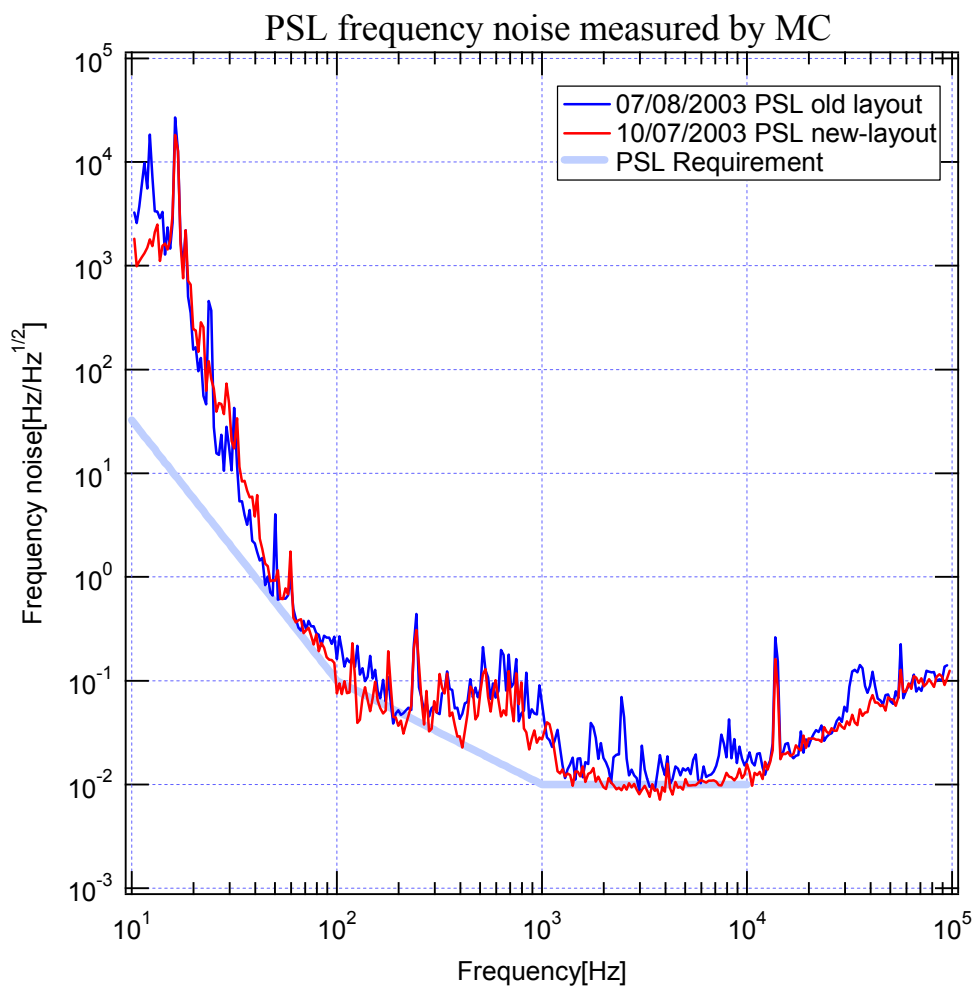


Important Achievement (3) Improvement of PSL Noise

October 2003

- PSL noise improved above 1 kHz
- New Calibration Method (AOM as a reference) used

[S.Kawamura and O. Miyakawa, "Convenient and Reliable Method for Measuring Frequency Noise of the Pre-Stabilized Laser", LIGO-T030239-00-R (2003)]

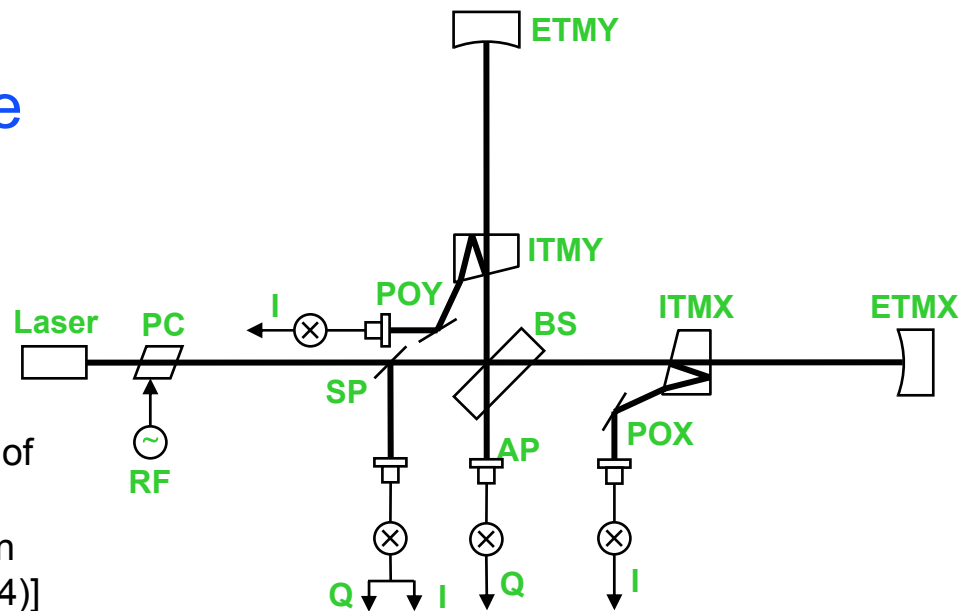


Important Achievement (4) Lock Acquisition of FP Michelson

November 2003

- FP Michelson locked
- Method similar to the one used for TAMA
- Interesting phenomena observed

[S. Kawamura and O. Miyakawa, "Polarity of Michelson Length Signal Obtained at the Symmetric Port In a Fabry-Perot Michelson Interferometer", LIGO-T030295-00-R (2004)]



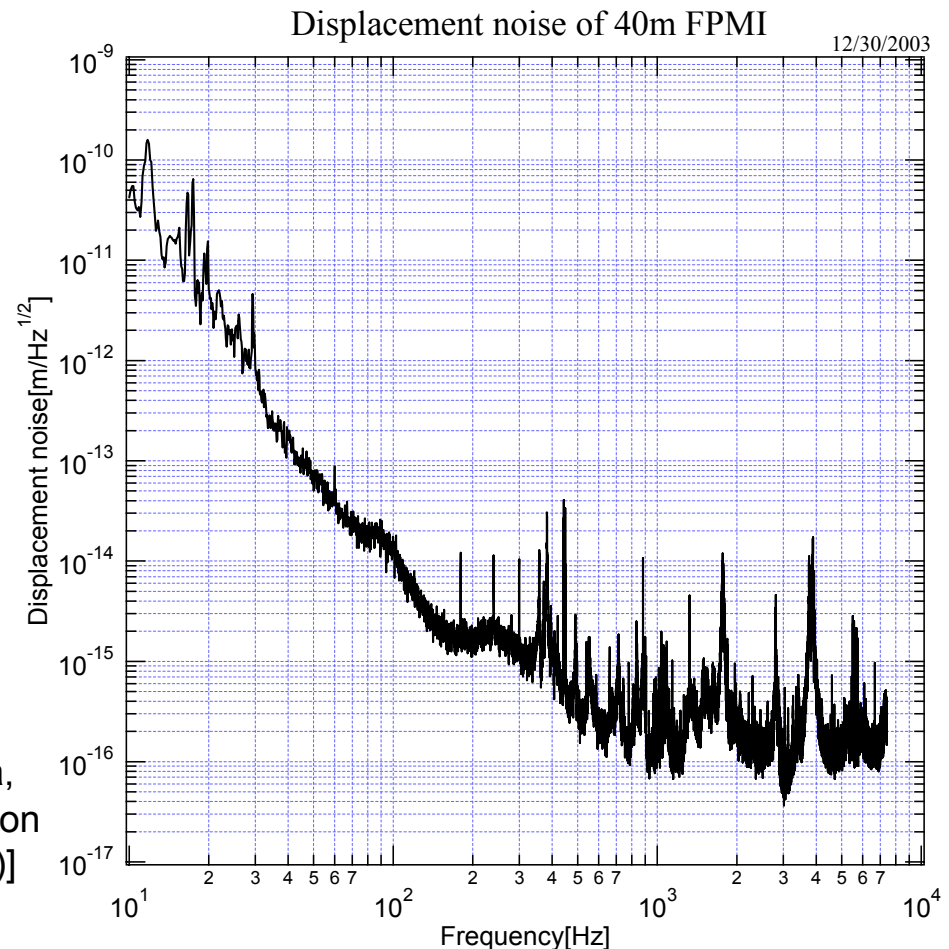


Important Achievement (5) Spectrum of FP Michelson

December 2003

- Displacement spectrum obtained
- With all the whitening/de-whitening filters ON
- Convenient calibration method (Michelson mid-fringe locking) used

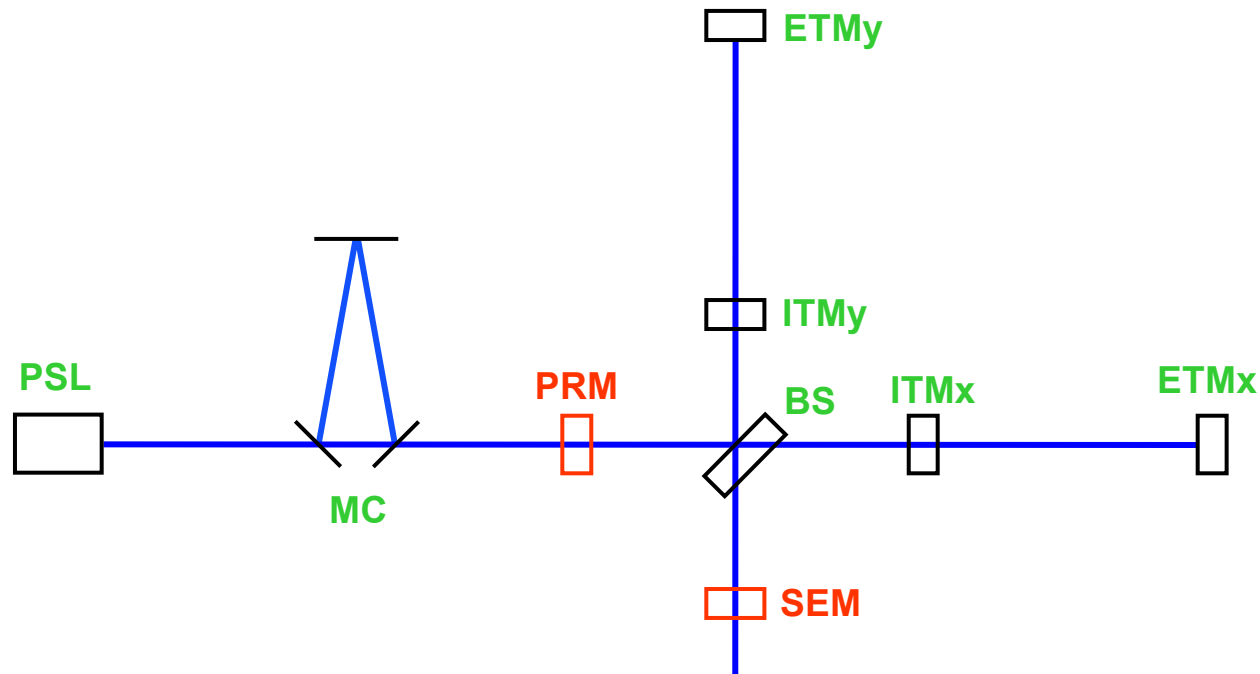
[S. Kawamura, O. Miyakawa, and S. Sakata, "Calibration of the 40m Fabry-Perot Michelson Interferometer", LIGO-T030287-00-R (2003)]



Important Achievement (6) Installation of PRM and SEM

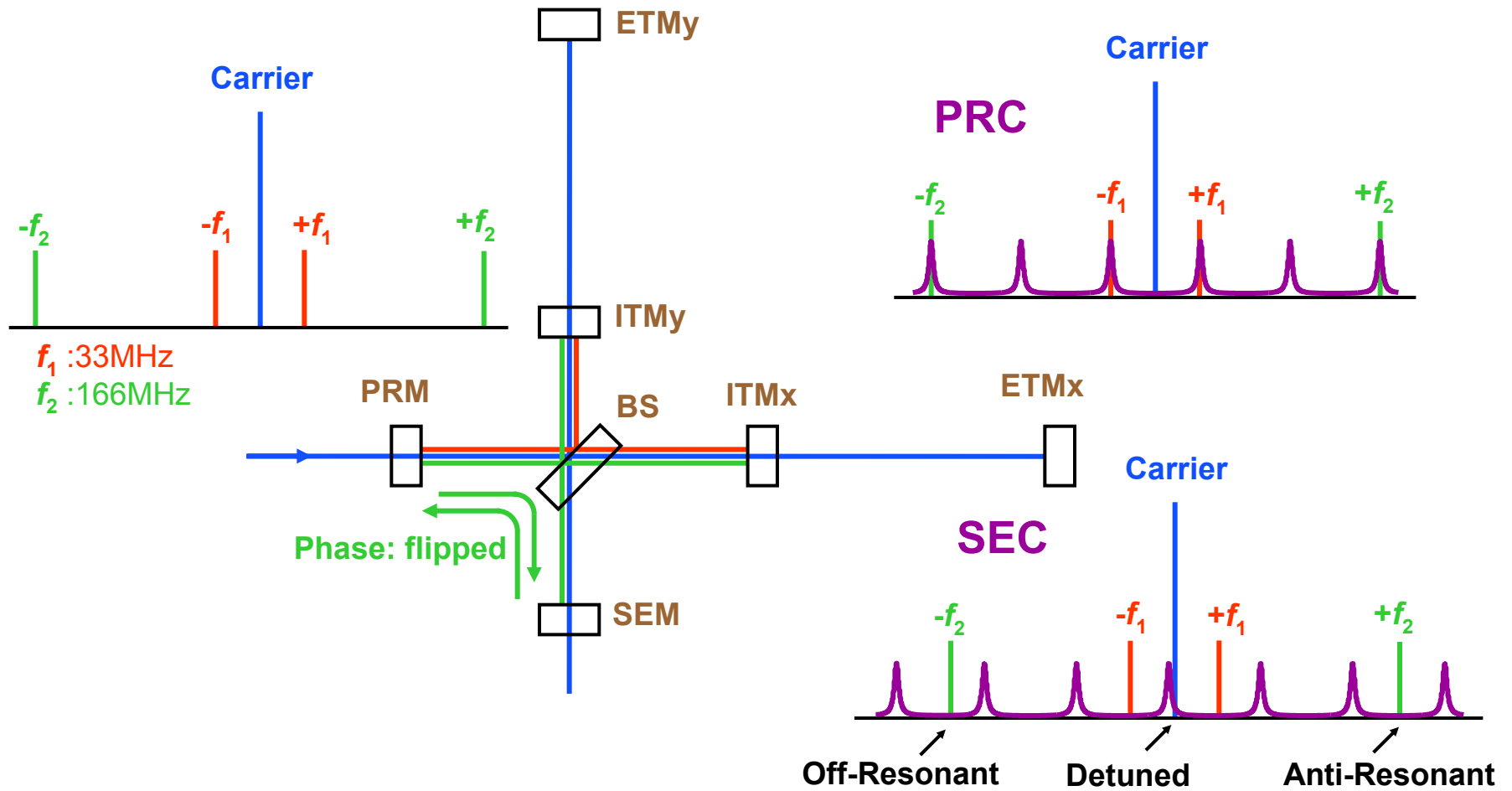
February 2004

- Power Recycling Mirror (PRM) and Signal Extraction Mirror (SEM) installed





Signal Extraction Method



Signal Extraction Matrix

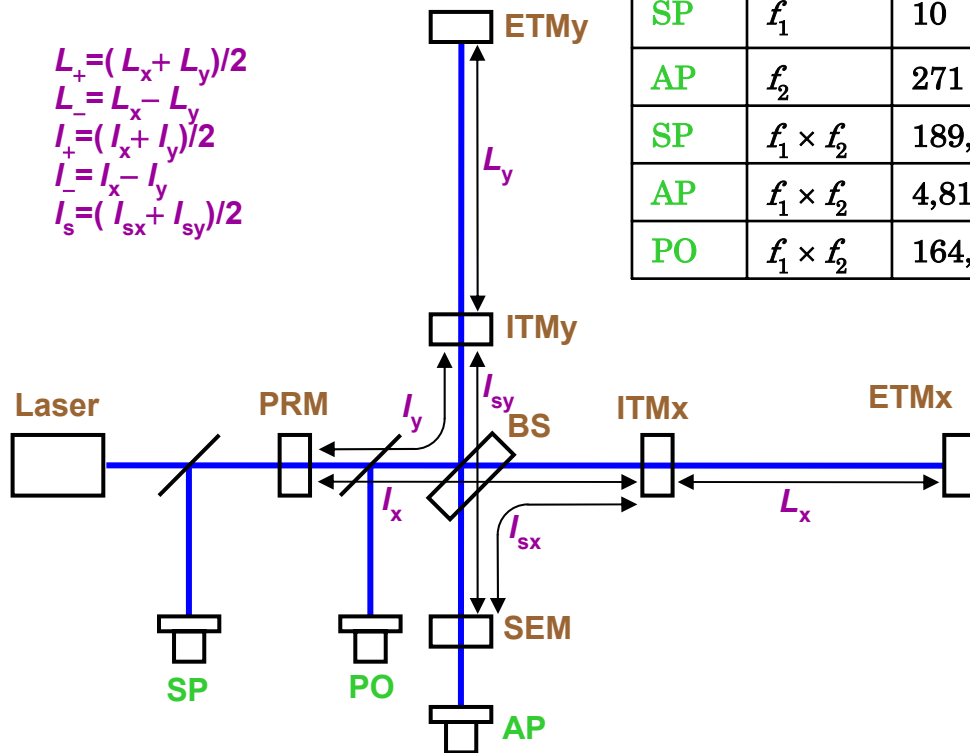
$$L_+ = (L_x + L_y)/2$$

$$L_- = L_x - L_y$$

$$I_+ = (I_x + I_y)/2$$

$$I_- = I_x - I_y$$

$$I_s = (I_{sx} + I_{sy})/2$$



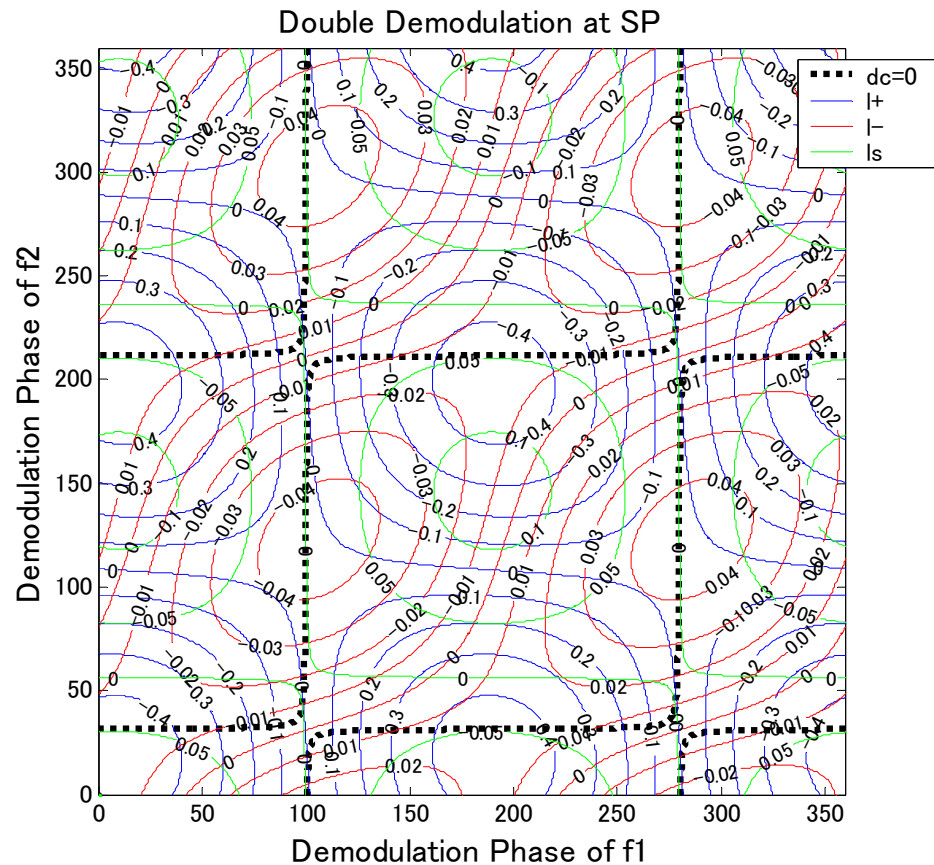
Port	Dem. Freq.	Dem. Phase	L_+	L_-	I_+	I_-	I_s
SP	f_1	10	1	-3.8E-9	-1.2E-3	-1.3E-6	-2.3E-6
AP	f_2	271	-4.8E-9	1	1.2E-8	1.3E-3	-1.7E-8
SP	$f_1 \times f_2$	189,32	-1.7E-3	-3.0E-4	1	-3.2E-2	-1.0E-1
AP	$f_1 \times f_2$	4,81	-6.2E-4	1.5E-3	7.5E-1	1	7.1E-2
PO	$f_1 \times f_2$	164,12	3.6E-3	2.7E-3	4.6E-1	-2.3E-2	1

- Calculated by FINESSE
(A. Freise: <http://www.rzg.mpg.de/~adf/>)
- PO: light from BS to ITMy

Double Demodulation

- Double Demodulation used for I_+ , I_- , and I_s
- Demodulation phases optimized to **suppress DC** and to **maximize desired signal**

[S.Kawamura, "Signal Extraction Matrix of the 40m Detuned RSE Prototype", LIGO-T040010-00-R (2004)]





Length Tolerances

- Acceptable cavity length deviations from the ideal points:

6 cm for l_-

3 mm for l_+

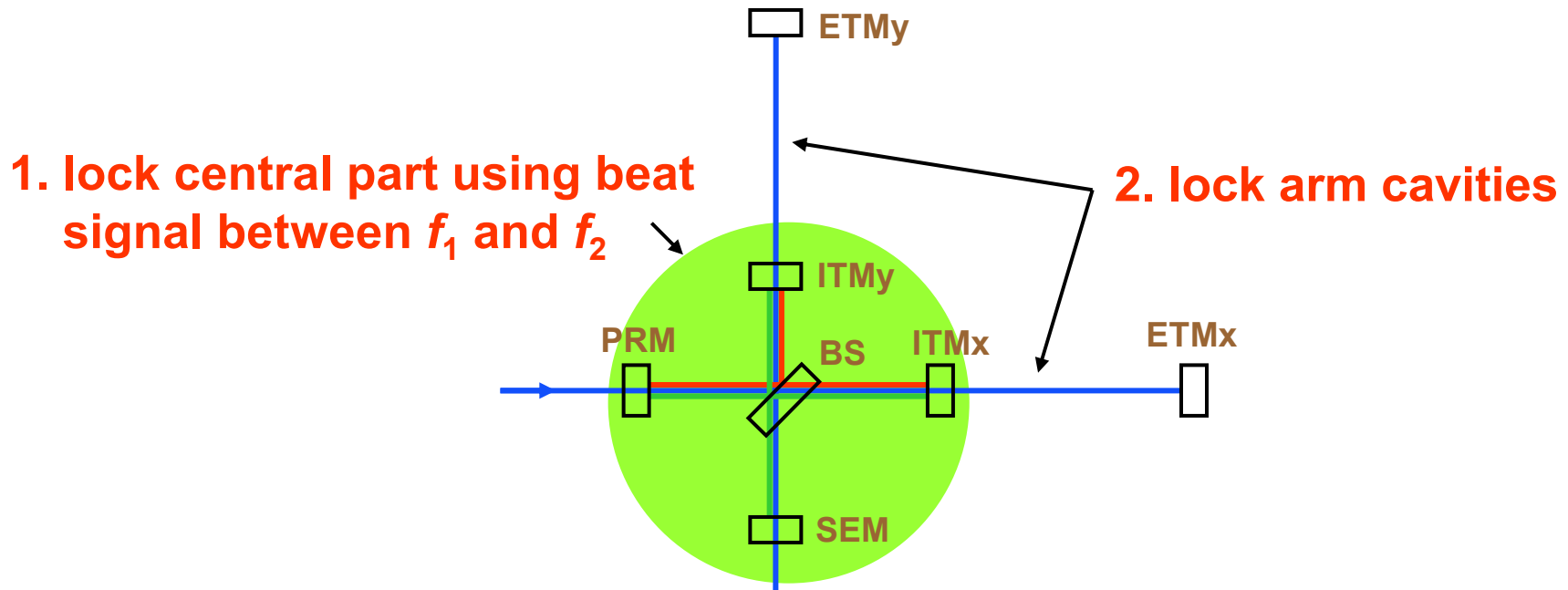
3 mm for l_s

Example: Signal matrix with l_+ deviation of 1 cm

Port	Dem. Freq.	Dem. Phase	L_+	L_-	l_+	l_-	l_s
SP	f_1	334	1	-7.6E-9	-1.2E-3	-4.1E-6	-2.3E-6
AP	f_2	230	-1.3E-9	1	3.0E-8	1.3E-3	-1.7E-8
SP	$f_1 \times f_2$	162,73	-6.5E-4	3.5E-4	1	-5.6E-2	1.3E-1
AP	$f_1 \times f_2$	173,218	1.5E-3	4.2E-4	-2.1	1	-2.4E-1
PO	$f_1 \times f_2$	329,153	1.1E-3	2.7E-3	2.6	-1.6E-1	1

S. Kawamura, "Signal Extraction Matrix of the 40m Detuned RSE Prototype", LIGO-T040010-00-R (2004)

Lock Acquisition of Detuned RSE



- Central part: not disturbed by lock status change of arm cavity
- Question: Not disturbed by flash of SBs or SBs of SBs in arms?
- Promising: TAMA using 3rd harmonic demodulation in PRFPMI

Lock Acquisition of Central Part

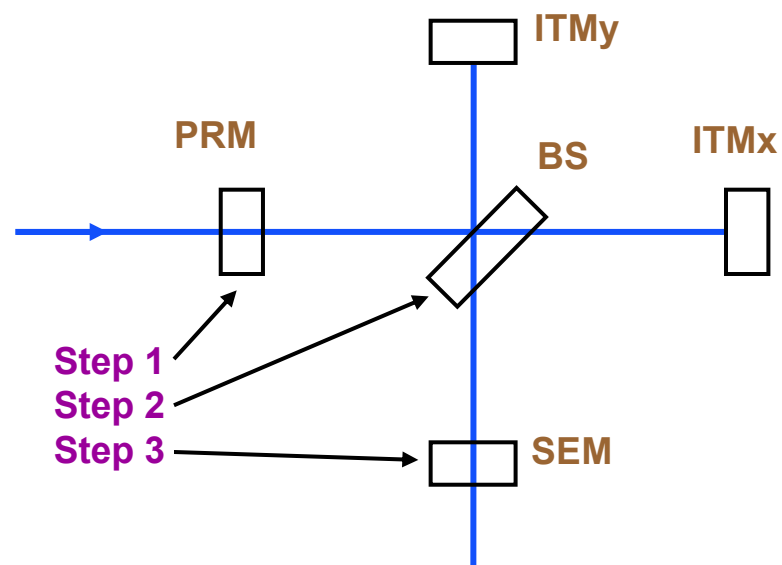
Ideal Procedure: Lock one by one

[for example]

Step 1: Lock I_+ robustly

Step 2: Lock I_- robustly

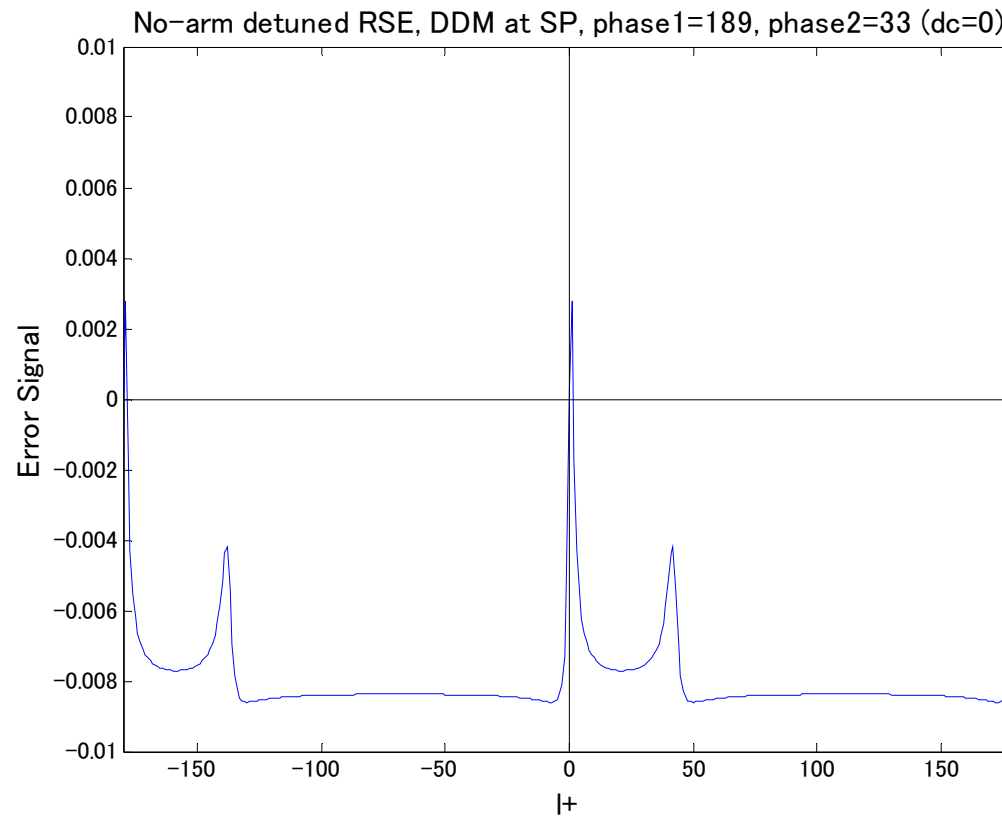
Step 3: Lock I_s



- Find primary signal not disturbed by the other two DOFs
- Find secondary signal not disturbed by the residual DOF



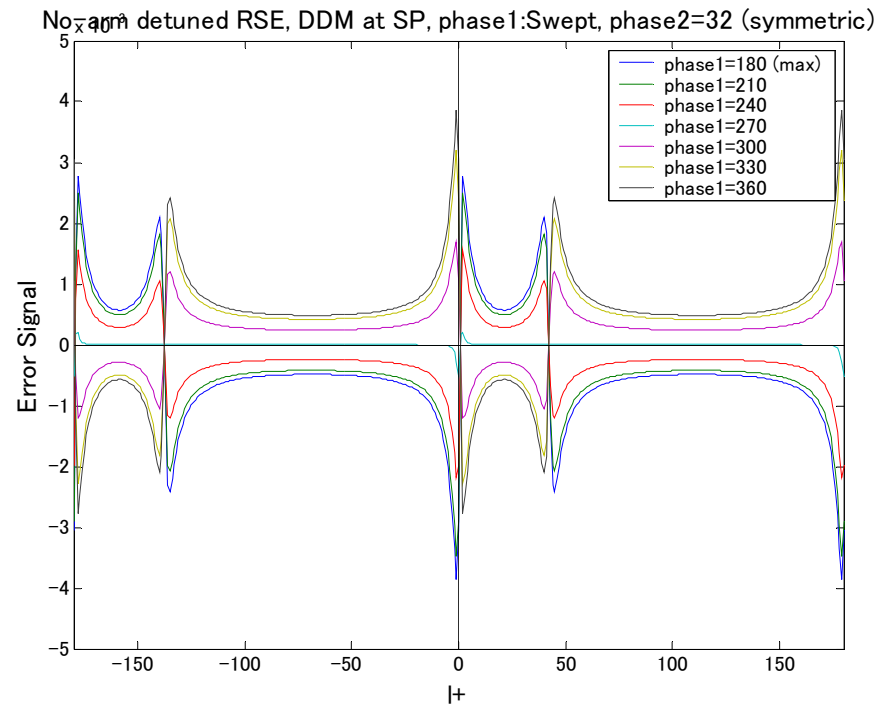
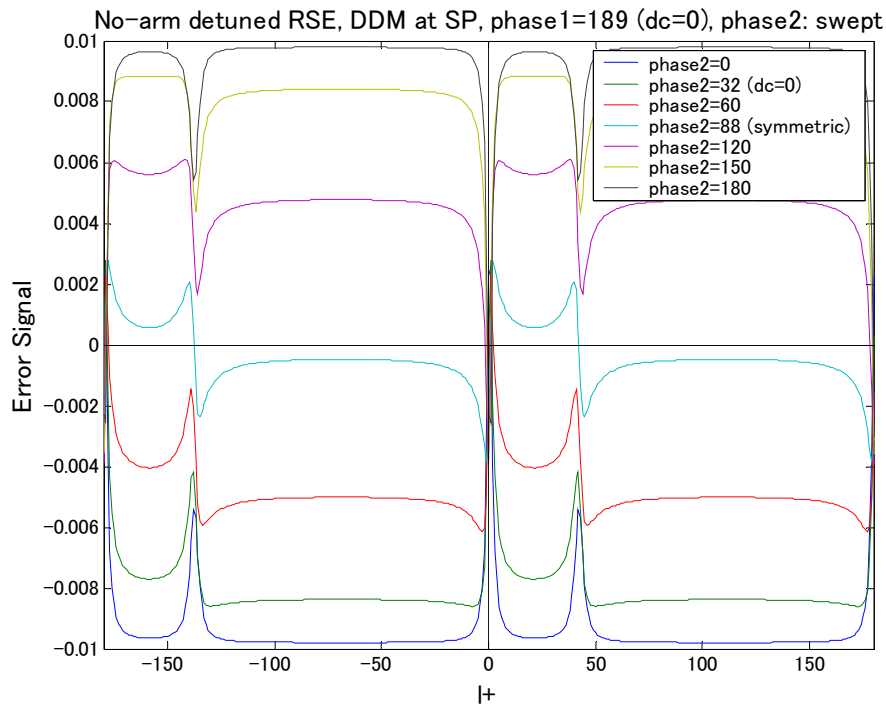
Quality of I+ Signal (dc=0, I+:Max)



See Movie

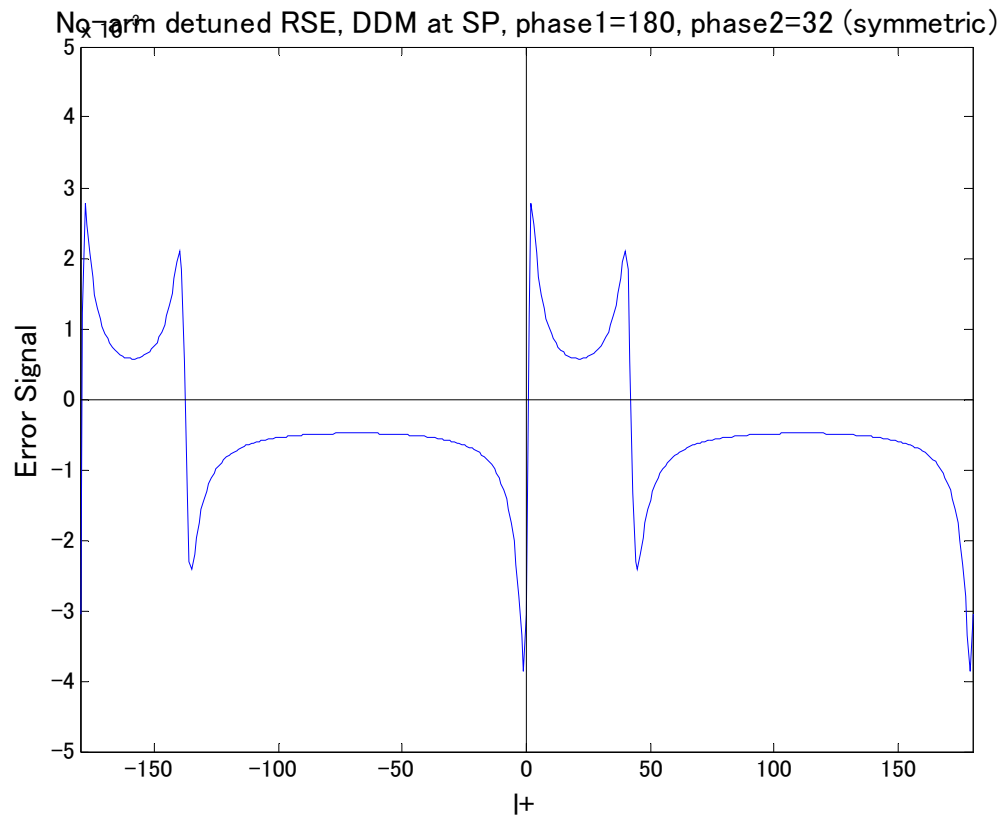


Dependence of I+ Signal on Demodulation Phases





Quality of I+ Signal (Symmetric, $dc \neq 0$)

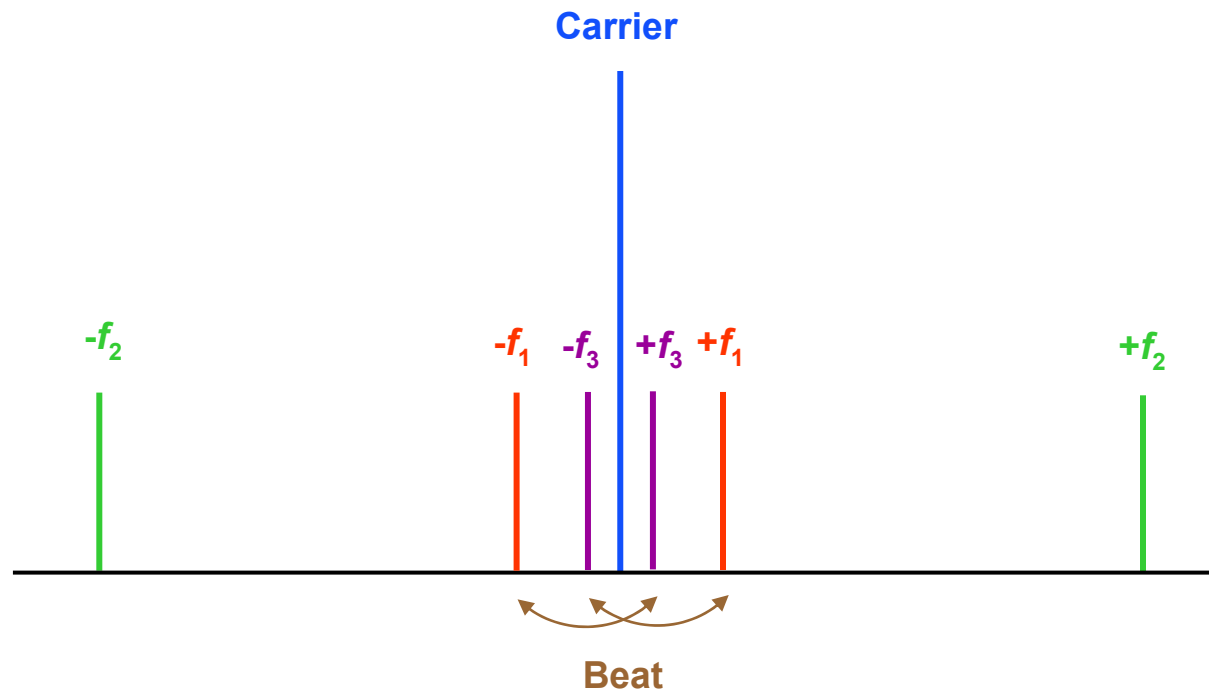


See Movie



Beat between f_1 and Non-Resonant SB f_3 (AM)

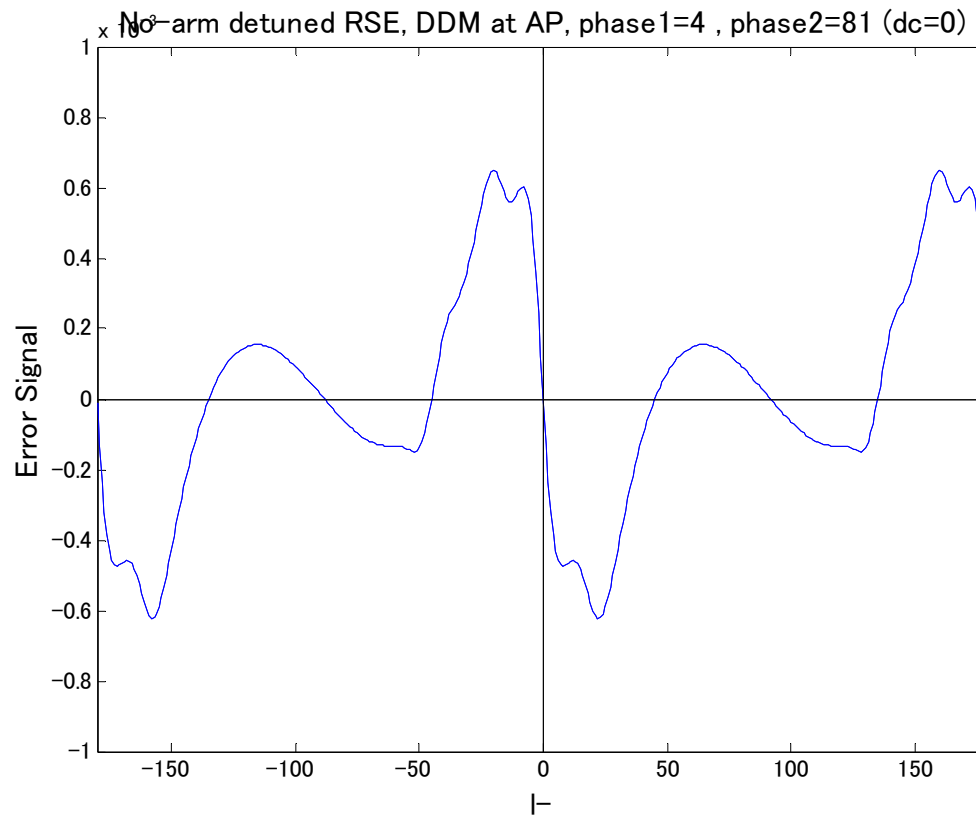
Hint: G. Heinzel at Garching 30m



See Movie



Quality of I- Signal

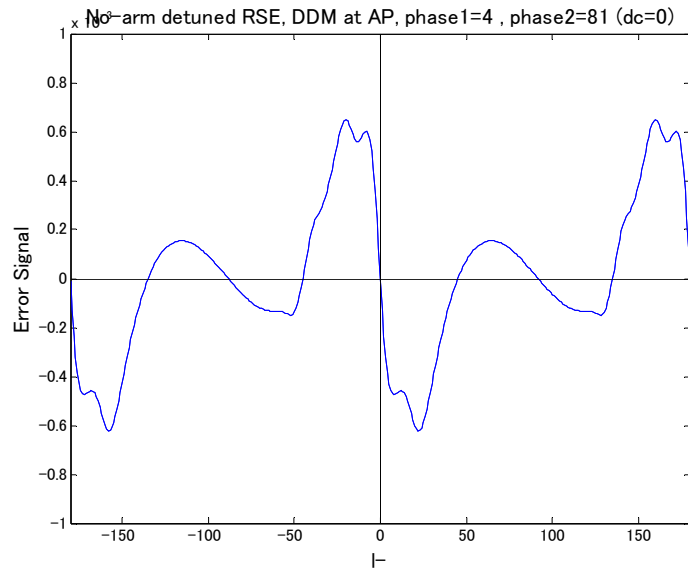


See Movie

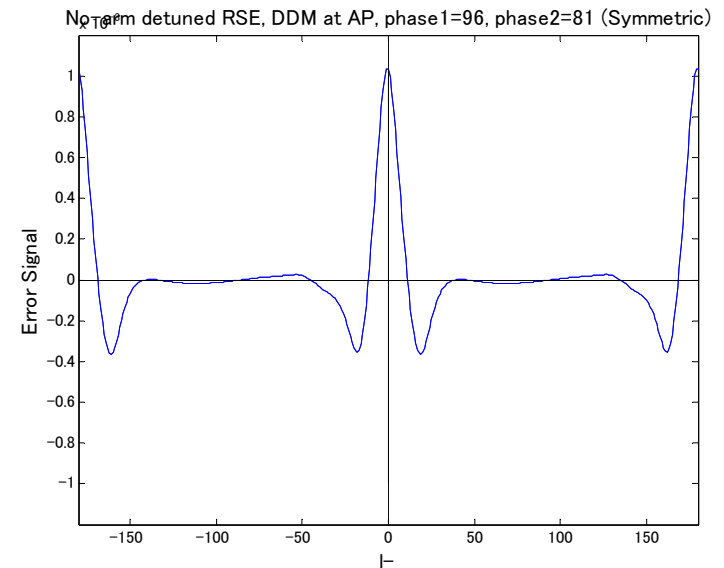


I- Signal Divided by AP signal with different Dem. Phase

Hint: H. Grote at GEO



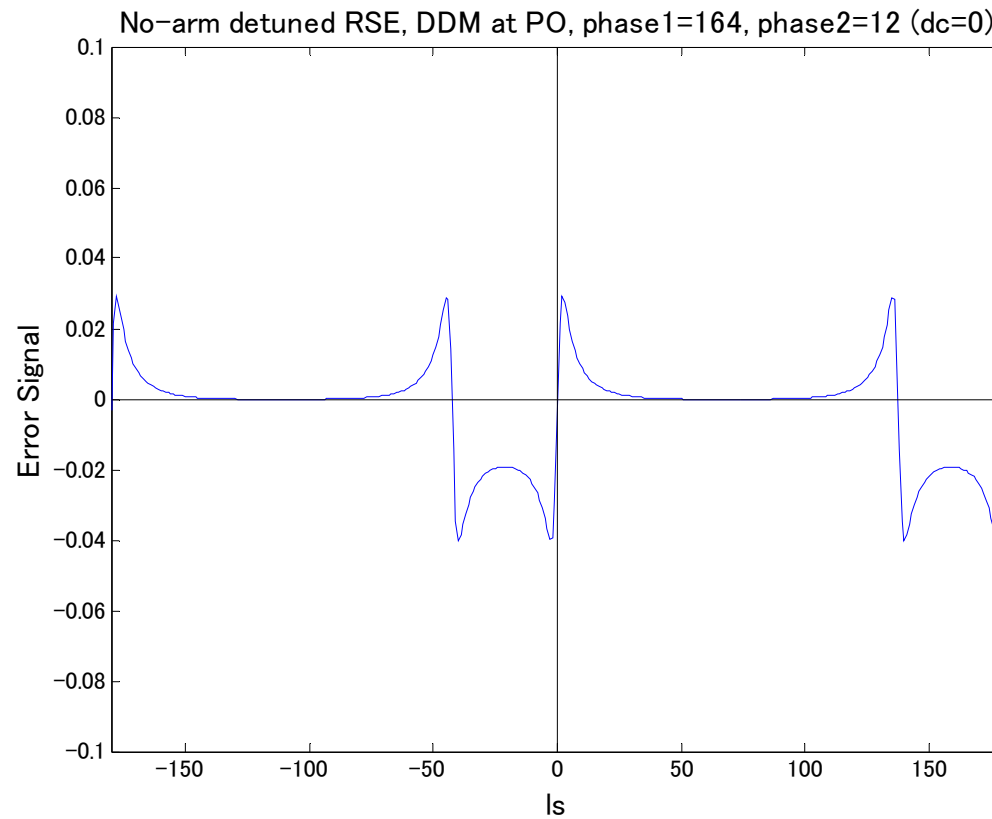
Divided by



See Movie



Quality of Is Signal



See Movie



Looking for Good Signal ...

Have tried the following, but not significant improvement...

- f1-dem, f2-dem
- For I+:
 - Single sideband f2, AM f2
 - NR pm, am, single sideband f3:
 - (f2-f3)-dem, (f1+f3)-dem
- For I-, Is:
 - Signal divided by the following:
 - 2×f1-dem, 2×f2-dem, DDM with different dem phases, (f2-2×f1)-dem



We Should Try More ...

- Analyze error signal with one DOF locked
- Try more various signals including mechanical dither signal
- Develop quick acquisition method
- Tighten suspended optics rigidly with respect to local frame, then manually bring three DOFs near lock point to acquire lock?



Summary

- Lots of achievements; experiment on 40m going very fast and smoothly
- Almost ready to try lock acquisition
- Lock acquisition: not so easy
- Need more investigation for lock acquisition

Hope we succeed in locking detuned RSE very soon!