ALLEGRO calibration for the LLO-ALLEGRO stochastic analysis

Martin McHugh Loyola University New Orleans

LIGO-G040138-00-Z

LSC meeting, LIGO Livingston Observatory 18 March, 2004

LLO - ALLEGRO correlation

- Good overlap (40 km separation) sensitive to higher frequency range than LLO-LHO
- Modulate rotate to align/misalign/anti-align antenna patterns
- Another independent detector



Outline of technique

- Correlate A1 with L1, taking same time stretch from each
 - Sampling rate very different (250Hz for A1, 2048Hz for LIGO) A1 data is heterodyned
- Match bins in frequency domain
- Resonant nature of the detector means long time constant for excitations in the raw data. Easier to calibrate all of the data and then work with *h*(*t*)







$$h^{H}[k] = \frac{z[k]}{T[k] \cdot F[k]}$$

Calibration

- determine *T*[*k*] *F*[*k*]
- Simple model for bar response (*R*)
 - track mechanical mode frequencies
 - track mode Q's (accounts for feedback -(H))
 - determine overall gain, T and F combined
 - Need to get phase shifts not included in mechanical response model (mainly electronic *F*)
 - also need overall phase of heterodyne reference with respect overall time standard (gps)
 - Tests indicate we can stitch together strain time series with overlapping stretches

LIGO-G040138-00-Z







Calibration pulses in raw and strain time series



Observatory 18 March, 2004



- ALLEGRO currently running same mechanical system as S2, mostly same electronics
- Allows us to perform some tests to verify model, measure phase shifts
- Have begun with noise injection measurement



LSC meeting, LIGO Livingston Observatory 18 March, 2004



|transfer function = response/excitation|