

Planning for the S4 Run

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LIGO Scientific Collaboration Meeting
LIGO Hanford Observatory
August 16-19, 2004

Strawman Schedule

Hanford

- ~September 15
Launch new DMT monitors 24/7
- ~October 9
Mini-Engineering Run
- ~November 13-20
Engineering Run
- ~**Early January**
S4 Science Run

LIGO-G040335-00-Z

S4 Planning -- 2003.8.20

Livingston

- ~September 15
Launch new DMT monitors 24/7
- [~November 20
Mini-Engineering Run] maybe...
- ~Mid-December
Engineering Run
- ~**Early January**
S4 Science Run

Looking ahead to ~~S3~~ – DMT world S4

Old monitors:

- Need to be resurrected and retuned – main purpose of mini-E-runs in ~~September~~ **October/November**
DMT authors expected to participate
- Final tuning will likely be needed during ~~maxi-E-run~~ **engineering runs**
- Post-~~S3~~ **S4** goal: identify **additional** key monitors / FOM's for commissioning and make robust against IFO changes

New monitors:

- **Time is running short for completion!**
- **Some promises made after ~~S2~~ **S3** in danger of breakage...**
- **Please bear down to get things ready**

Looking ahead to S4 - Scimons

Making scimons more effective – discussions underway

- **Longer shift blocks with fewer different scimons**
(more cost-effective for groups & better training;
longer-term goal: more LSC students/postdocs at sites)
- **More focus on astrophysical figures of merit**
- **More focus on data quality flagging in the control room**
- **Groups should make requests for special consideration early – to avoid later use of scimon-swaps**

Looking ahead to S4 Analysis Feedback

- First line of defense against astrophysically crappy data:
SenseMonitor & other monitors of “expected” sensitivities
- Next line of defense:
DataQual, glitchmon, & other generic glitch finders
- Next line of defense:
BurstMon (& other monitors of realistic astrophysical sensitivity)
- Last line of defense
Quasi-online analysis jobs using actual inspiral template banks,
burst ETG’s, etc.

→ **DASWG** purview [volunteers welcome!

see Patrick Brady]

More on DMT issues

**Glad to see new astrophysical FOM and other DMT monitors
→ THANKS!**

But we need more monitors of known artifacts (see Fred's list)

**Embarrassing that we STILL don't have an airplane monitor!
(effects first seen in E1(!) engineering run – April 2000)**

**Upcoming detector investigation camp will have sessions
on DMT monitors, including how to write them, with
template examples**

Known Causes of Spurious Bursts

- Optical level laser AM burst
 - » Commonly due to diode aging
- Servo instability
 - » Commonly due to drifts in gain
- Saturate a coil driver
 - » Sum of fast and slow environmental factors
- Saturate a photodiode/mixer
 - » Might be obvious or not
- Clipping in a photodiode path
 - » introduces sensitivity to seismic or acoustic noise
- “Gimpy” cable
 - » often accompanied by level shift
- Dewar “pops”
- Dumb stuff
 - » Like touching up a picomotor

Can we do
better
monitoring
these
causes?

Most of these well suited to DMT
monitoring

• Varying levels of difficulty in
monitoring

• Mostly need people working on them!