## Italian-American Research Collaboration Amundsen-Scott Station

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## The MOF at South Pole

## A collaboration between

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## Experiment Objectives

- Seismically map the sound speed of the sun's Cower atmosphere using high-frequency sound waves
- Improve the models for the sun's atmosphere
- Space Weather : early warning to explosive phenomena (flares, CME, ...)


## Experiment Description

The velocity and intensity signals are detected at two different heights in the solar atmosphere (two Magneto-Optical Filters tuned to the solar K and $\mathcal{N}$ a absorption lines)

The instrument has been given the name

$$
\mathcal{M} O \mathcal{T} \mathcal{H}
$$

Magneto Optical filters at Two Heights

## Experiment Description

- 2 MOFs tuned to the $\mathcal{K}(7699 \mathcal{A}-D 1)$ and $\mathcal{V a}$ (5890 $\mathcal{A}$ - (D2) lines
- Acquisition of separate and simultaneous images in the Red and Blue flanks of the above absorption fines
- 2 compact Telescopes ( 48 mm diam)
- $4 \times 12$-6it digital Cameras, running at 16 frames $/ \mathrm{s}$
- 10 s integration time for each recorded 16 bit image
- 1 TeraByte Raid Hard Disk recording System


## South Pole Remote site



## External Instrument Positioning



## The buried Control Room



## The ММОТभ



## Aligning and Focusing



## Hunting the Sun ...



## The MOTH゙ 's Optical Layout



## The System Layout



## First results




## Velocity Signals



## Potassium Log Power Spectrum - m corrected



## Sodium Log Power Spectrum - m corrected



## Na-K average Cross Spectrum



Time Lag Map@ 7 mHzz
(average: 27 seconds)


Time Lag Map@ $3 m \mathcal{H} z$ (average: 0 seconds)


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