

Toward the Advanced LIGO optical configuration investigated in 40meter prototype

Aspen winter conference

Jan. 19, 2005

O. Miyakawa, Caltech and the 40m collaboration

LIGO- G050047-00-R

LIGO

Caltech 40 meter prototype interferometer

Objectives

- Develop lock acquisition procedure of detuned Resonant Sideband Extraction (RSE) interferometer, as close as possible to Advanced LIGO optical design
- Characterize noise mechanisms
- Verify optical spring and optical resonance effects
- Develop DC readout scheme
- Extrapolate to AdLIGO via simulation
- etc.





Advanced LIGO optical configuration



LIGO:Power recycled FPMI

PRM

Detuning

- » Optical noise is limited by Standard Quantum Limit (SQL)
- AdvLIGO:GW signal enhancement using Detuned RSE
 - » Two dips by optical spring, optical resonance



Power

GW signal

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Laser

FP cavity



Differences between AdvLIGO and 40m prototype

- 100 times shorter cavity length
- Arm cavity finesse at 40m chosen to be = to AdvLIGO
 - » Storage time is x100 shorter.
- Control RF sidebands are 33/166 MHz instead of 9/180 MHz
 - » Due to shorter PRC length.
- LIGO-I 10-watt laser, negligible thermal effects
 - » 180W laser will be used in AdvLIGO.
- Noisier seismic environment in town
 - » >1x10⁻⁶m at 1Hz
- Smaller stack, commercial active seismic isolation
 - » STACIS isolators in use on all test chambers, providing ~30 dB of isolation from 1-100 Hz.
- LIGO-I single pendulum suspensions are used
 - » AdvLIGO will use triple (MC, BS, PRM, SRM) and quad (ITMs, ETMs) suspensions.





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In-vacuum Faraday Isolator and In-vacuum Mode Matching Telescope





LIGO-I type single suspension

- Each optic has five OSEMs (magnet and coil assemblies), four on the back, one on the side Suspension Block Suspension Support Structure Suspension Wire Magnet/Standoff Assembly Stiffener Bar Guide Rod & Wire Standoff Head Holder Sensor/Actuator Head Safety Stop Housing Magnet The magnet occludes light LED
 - from the LED, giving position
 Current through the coil creates a magnetic field, allowing mirror control

LIGO

Coil

Standoff

Photodiode-



STACIS Active seismic isolation





Digital control system



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Signal extraction for AdvLIGO



- Arm cavity signals are extracted from beat between carrier and f_1 or f_2 .
- Central part (Michelson, PR, SR) signals are extracted from beat between f₁ and f₂, not including arm cavity information.



5 DOF for length control

Port Dem. **L**_ **I**_s L Freq. **ETMy** -3.8E-9 -1.2E-3 -1.3E-6 -2.3E-6 SP f₁ 1 -4.8E-9 AP f, 1.2E-8 1.3E-3 -1.7E-8 1 $f_1 \times f_2$ -1.7E-3 -3.0E-4 -3.2E-2 -1.0E-1 SP 1 -6.2E-4 1.5E-3 7.5E-1 7.1E-2 AP $f_1 \times f_2$ 1 3.6E-3 2.7E-3 4.6E-1 -2.3E-2 PO $f_1 \times f_2$ 1 ITMv **ETMx** PRM ITMx Laser sy BS **Common of arms** $: L_{+}=(L_{x}+L_{y})/2$ $: L_{=} L_{x} - L_{y}$ Differential of arms sx Power recycling cavity : $I_{+}=(I_{x}+I_{y})/2$ SRM **Michelson** $: I_{=} I_{x} - I_{y}$ Signal recycling cavity : $I_s = (I_{sx} + I_{sy}) / 2$ PO SP

Signal Extraction Matrix (in-lock)

Disturbance by sidebands of sidebands



- Sidebands of sidebands are produced by two series EOMs.
- Beats between carrier and $f_2 + f_1$ disturb central part.

| Port | Dem. Freq. | L ₊ | L_ | Ι ₊ | I_ | l _s |
|------|------------------|-----------------------|---------|----------------|---------|----------------|
| SP | f ₁ | 1 | -1.4E-8 | -1.2E-3 | -1.3E-6 | -6.2E-6 |
| AP | f ₂ | 1.2E-7 | 1 | 1.4E-5 | 1.3E-3 | 6.5E-6 |
| SP | $f_1 \times f_2$ | 7.4 | -3.4E-4 | 1 | -3.3E-2 | -1.1E-1 |
| AP | $f_1 \times f_2$ | -5.7E-4 | 32 | 7.1E-1 | 1 | 7.1E-2 |
| PO | $f_1 \times f_2$ | 3.3 | 1.7 | 1.9E-1 | -3.5E-2 | 1 |

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Mach-Zehnder interferometer on 40m PSL to eliminate sidebands of sidebands





MZ eliminates sidebands on sidebands

MCT light, series EOMs

parallel EOMs in MZ ifo No sidebands on sidebands!

(hard to directly compare because we can't turn the modulation depth up as high as we could before; but we can get up to $\Gamma = 0.25$ easily)





Important Milestones

September, 2003 Four TMs and BS: installed November 2003 **FP** Michelson locked February 2004 Power Recycling Mirror (PRM), Signal Extraction Mirror (SRM) installed June 2004 Mach-Zehnder installed August 2004 DRMI locked with carrier resonance October 2004 DRMI locked with sideband resonance November 2004 Off-resonant lock of arm cavities with DRMI







Lock Acquisition of Detuned RSE

1. lock central part



- Central part: not disturbed by carrier resonance on arm cavity (but disturbed by sidebands resonance)
- Lock acquisition
 - *I*: dither @ 1200 Hz
 - *I*₊ : <u>33MHz@SP</u> I : DDM@PO

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- After lock:
- \rightarrow DDM@AP
- $\rightarrow DDM@SP$
- $\rightarrow DDM@PO$

- Arm cavities: not disturbed by locked central part
- Lock each arm cavity independently
- Switch control servo to common/differential control





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Looking for good signal for lock acquisition

• Unfortunately, no way to lock central part directly using the original double demodulation ITMy PO **ITMx** BS Dither locking for *I* signal Laser PRM **Divide signal by inside power** SRM » Good cancellation of power recycling ^IAP $\Delta V_{l-} = \frac{d}{d l} \left(\frac{V_{\rm AP}}{V_{\rm PO}} \right)$ a few kHz LPF LPF $V_{\rm AP} (V_{\rm AP})'$ $V_{\rm PO}(V_{\rm PO})$ $=\frac{V_{\rm AP}' V_{\rm PO} - V_{\rm AP} V_{\rm PO}'}{V_{\rm PO}^2}$ Digital ΔV_{L} calculation

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I_ signal with dither





Lock I+ with DDM at SP

- With I- dither-locked, there's always a good I+ signal, for all values of Is.
- The locking point may not be at l+ = 0° !
- The PRM follows the swinging of the SRM; this signal keeps the combined cavity locked.
- Then, once Is is locked, we'll recover I+ = 0°.







$I_{\rm s}$ signal with $I_{\rm and}$ and $I_{\rm p}$ lock



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DRMI lock with

Unbalanced sideband by detuned cavity

August 19, 2004

DRMI locked with carrier resonance (like GEO configuration)

November 9, 2004

•DRMI locked with sideband resonance (Carrier is anti resonant preparing for RSE.)

November 16, 2004

- Switched to DDM control
- Can be locked with DDM directly
- Longest lock: 2.5 hours

Typical lock acquisition time ~10sec







Trial of Arm lock with DRMI





Off-resonant lock scheme for arm cavity



+ offset $\sqrt{\text{Transmitted power}}$ to avoid coupling of carrier in Michelson part when arm cavity is locked.



Off resonant Arm lock with DRMI

November 25, 2004

- Both arms locked with DRMI
- Off-resonant carrier on arm cavities
- Last < 1 min</p>
- Locked only 2 times

DRMI with single arm lock

- Not so difficult
- Last ~10 min
- Lock acquisition time ~1 min
- Reducing offset starts oscillation caused by optical lever servo, under investigation





Summary

- Optical configuration for AdvLIGO being developed at 40m prototype interferometer
- Stable operation of PSL and MC
- Locking of FPMI with digital LSC system (misaligned PRM, SRM), measurement of displacement noise
- Sidebands of sidebands: eliminated by M-Z interferometer
- Guided locking of DRMI using Dither-locking with carrier/sideband resonance
- Locking of DRMI with DDM with sideband resonance
- Off-resonant locking of both arms with DRMI (not perfect but very close to final configuration)

Hope we succeed in locking full RSE very soon!





40m

40m vs. Ad-LIGO

| Γ | Table 4: Length | n sensing s | signals. | \otimes means of | double den | nodulation | ι. |
|---|-----------------------|-------------|----------|--------------------|------------|------------|-----------|
| | Signal | L_+ | L_{-} | l_+ | l_{-} | l_s | |
| | SP, f_1 | 15.2 | 0.000 | -0.062 | 0.064 | -0.001 | |
| | AP, f_2 | 0 | 1.69 | 0 | 0.002 | 0 | |
| | SP, $f_2 - f_1$ | -0.0003 | 0.0001 | (0.214) | 0.029 | 0.039 | x6 |
| | AP, $f_2 \otimes f_1$ | 0 | 0 | 0.0025 | -0.0034 | -0.0004 | x1.5 |
| | PO, $f_2 - f_1$ | 0.005 | -0.004 | 1.000 | -0.277 | -2.980 | x3 |

Table 5: Length sensing signals for Advanced LIGO. \otimes means double demodulation. These numbers agree, up to an overall constant, with the table Peter Fritchel showed at the August 2000 LSC meeting (LIGO-G000225).

Ad-LIGO

| Signal | L_+ | L_{-} | l_+ | l_{-} | l_s |
|-----------------------|-------|---------|--------|---------|---------------|
| SP, f_1 | 1890 | 0.00 | -1.94 | 0.11 | 0.00 |
| AP, f_2 | 0 | -1500 | 0 | -1.88 | 0 |
| SP, $f_2 - f_1$ | -0.11 | -0.01 | 19.5 | -0.11 | 8.66 |
| AP, $f_2 \otimes f_1$ | 0.000 | 0.001 | -0.031 | 0.242 | 0.005 |
| PO, $f_2 - f_1$ | -0.42 | -0.01 | 8.84 | 5.81 | $\boxed{245}$ |

x2 x6 x17



SP33,DDM,+/-33M,+/-166M@SP





Original design (no offset)

- +33 : off-resonant -33 : off-resonant +166: resonant -166 : anti-resonant
- *I*₊ and *I*_s plot separated
- Difficult



LIGO- G05(



l₊ +0.56 degree

- +33 : resonant -33 : resonant +166: off-resonant -166 : anti-resonant
- *I*₊ and *I*_s plot overlapping
- DC line changed
- Difficult





Double demodulation signal of I_+

- What we expected
 - » Big offset when cavity is not locked
 - » No disturbance of carrier

- What we have seen
 - » No offset
 - » Big disturbance of carrier







Double Demodulation

- Double Demodulation used for *I*₊, *I*₋, and *I*_s
- Demodulation phases optimized to suppress
 DC and to maximize desired signal

[S.Kawamura, "Signal Extraction Matrix of the 40m Detuned RSE Prototype", LIGO-T040010-00-R (2004)]





Gain of dither locking signal

- I- dither locking signal gain depends strongly on Is
- But polarity of signal is always the same
- Can handle this with a limiter...
- I- dither locking signal doesn't depend on I+ at all!
- Signal is degraded by presence of RF sidebands... turn them down low (Γ<0.02) to acquire dither lock, then ramp them back up.





Optical configuration for Gravitational wave interferometer

 Gravitational wave detection using Michelson interferometer



 Signal and power enhancement using Fabry-Perot cavity in each arm







Once we acquire full lock

- Measure in-lock transfer functions.
- Verify optical resonance and optical spring
- Begin noise characterization
- Operate at a different SEC tune?
- Begin work towards DC readout
 - » output mode cleaner
 - » offset locking of arms
 - » Tuning of homodyne phase





Contents

- Over view of 40meter prototype
- Signal extraction for Advanced LIGO
- Lock acquisition of Dual Recycled Michelson (DRMI)
- Off-resonant lock of arm cavities with DRMI





Dual Recycling Summation cavity for End to End model

Calculation time is determined by shortest cavity (Michelson) length.

$$\tau = \frac{L_{\rm arm}}{c} >> \frac{l_{\rm Mi}}{c}$$

Calculating many time steps at once in Michelson part

$$E(t) = \mathbf{M} \cdot E(t - \tau)$$

= $\mathbf{M}^{2} \cdot E(t - 2\tau)$
M
= $\mathbf{M}^{N} \cdot E(t - N\tau)$
N = $\frac{L_{arm}}{l_{MI}}$

 $\mathbf{M} = \mathbf{M}_0 + \delta \mathbf{M}$

- $\mathbf{M} : \underset{\mathbf{M}_{0} \cdot \partial \mathbf{M} \neq \partial \mathbf{M} \cdot \mathbf{M}_{0}}{\text{matrix for DR summation cavity}}$
- $\mathbf{M} : \text{scalar for PR summation cavity} \\ \mathbf{M}_0 \cdot \partial \mathbf{M} = \partial \mathbf{M} \cdot \mathbf{M}_0$

- 400 times faster for LIGO
- 40 times faster for 40meter LIGO- G050047-00-R Aspen winter con



40m Team

On the payroll: Ben Abbott, Osamu Miyakawa, Bob Taylor, Steve Vass, Alan Weinstein

Grad students: Lisa Goggin, Rob Ward

LIGO engineering support: Jay Heefner, Rolf Bork, Alex Ivanov, Flavio Nocera, Michael Smith, Lisa Bogue, many others

Visitors: Seiji Kawamura, Fumiko Kawazoe, Shihori Sakata, Bryan Barr, Sascha Schediwy, Kentaro Somiya, Rana Adhikari (Hartmut Grote of GEO arrives Oct 1)



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First lock of Dual recycled Michelson





Control room



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