

AURIGA II run: stationary gaussian operation of a wideband acoustic gw detector

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- one sided $S_{hh}^{1/2} < 4 \cdot 10^{-21} \text{ Hz}^{-1/2}$ over 90 Hz band
- the 3 modes thermal at 4.5 K
- ~ 100% operation (except 3h/month > He transfer)
- veto time intervals under out-of-band triggers:
 δ -like events > stationary gaussian over 60% of time
GOAL Gaussian Operation At (thermal) Limit
- GOAL expected to increase after:
further improvement of suspensions
refinement of the analysis
- GOAL at ~ quantum limit (100mK, 30 h-bar, 20 MV/m)
one sided $S_{hh}^{1/2} < 4 \cdot 10^{-22} \text{ Hz}^{-1/2}$ over 90 Hz band



LIGO-G050072-00-Z



❖ new mechanical suspensions:
attenuation > 360 dB at 1 kHz
FEM modelled

❖ three resonant modes operation:
two mechanical modes
one electrical

❖ new data analysis and data acq.:
C++ object oriented code
frame data format
Monte Carlo software injections
improved noise matching algorithm
selectable templates

AURIGA run II: upgrades



❖ **three resonant modes operation:**

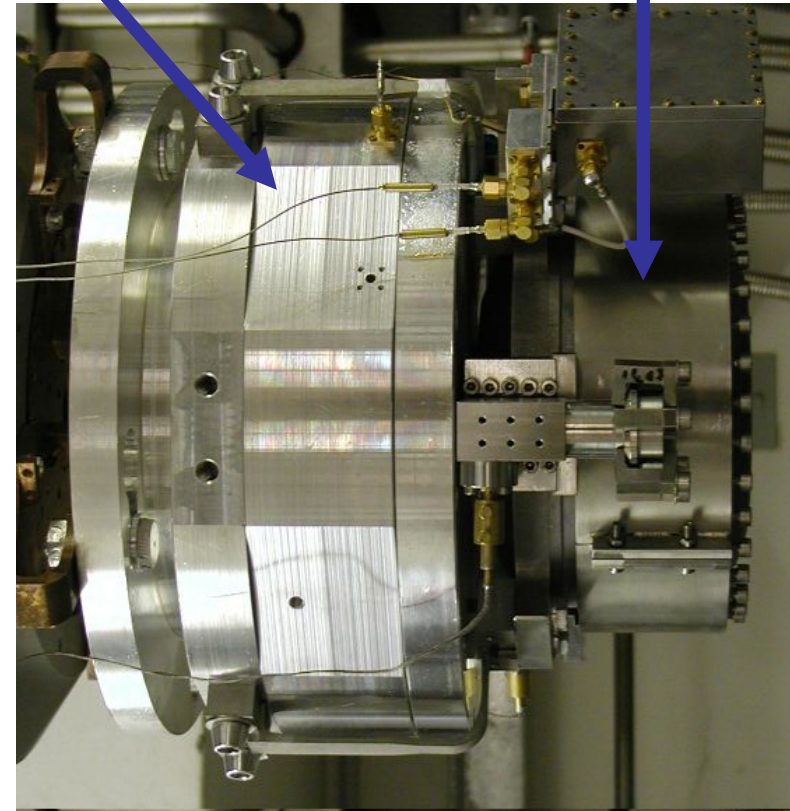
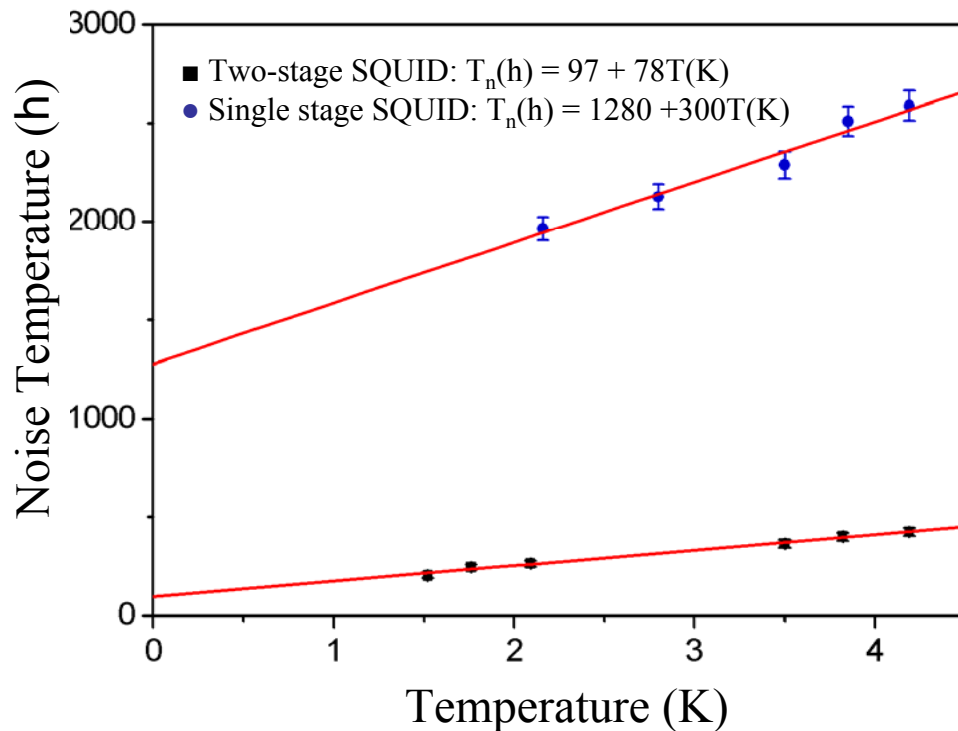
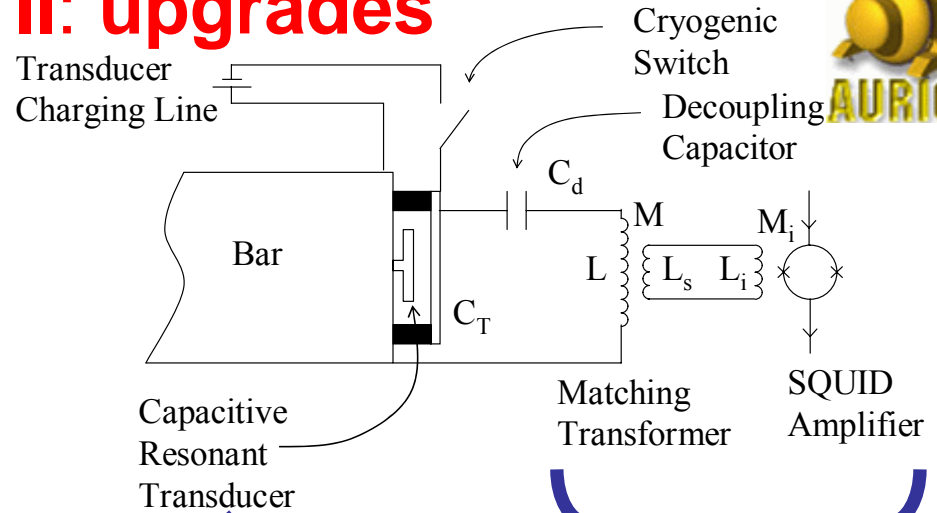
two mechanical modes
one electrical mode

❖ transducer bias field 8 MV/m

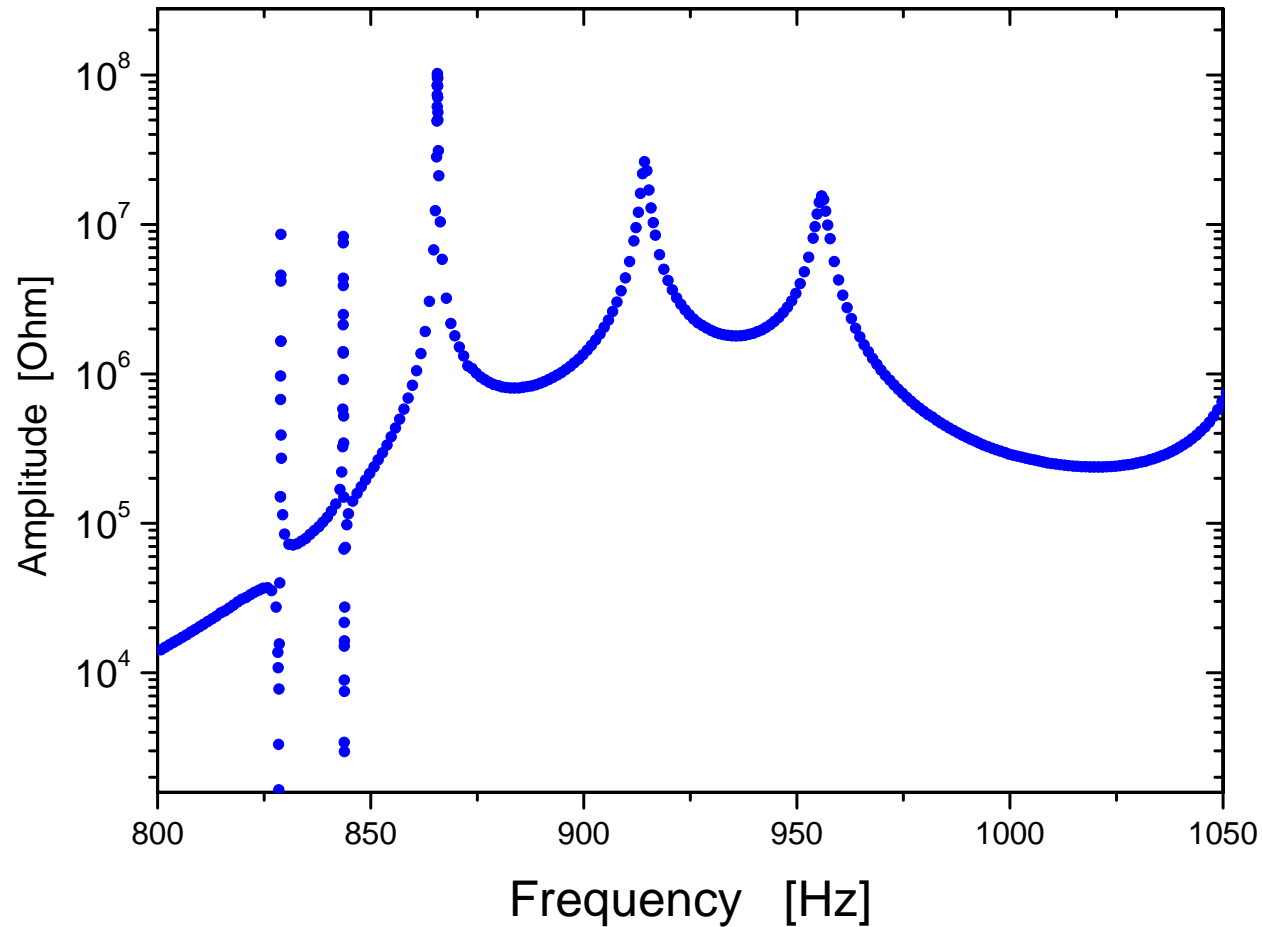
❖ **new SQUID amplifier :**

double stage SQUID

650 energy resolution at 4.5 K in the detector

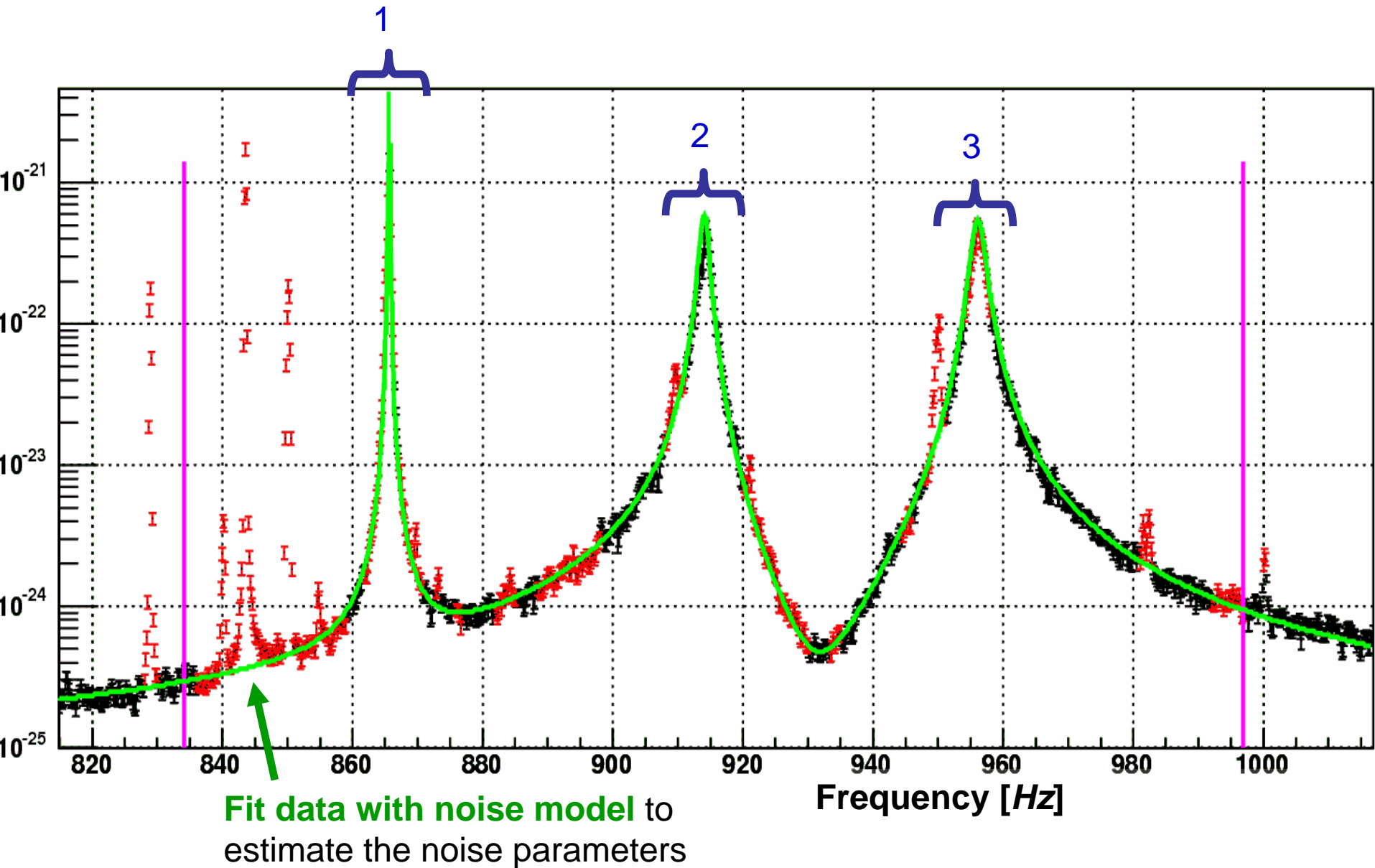


Measurement of the mechanical transfer function of the bar-transducer system



raw data PSD

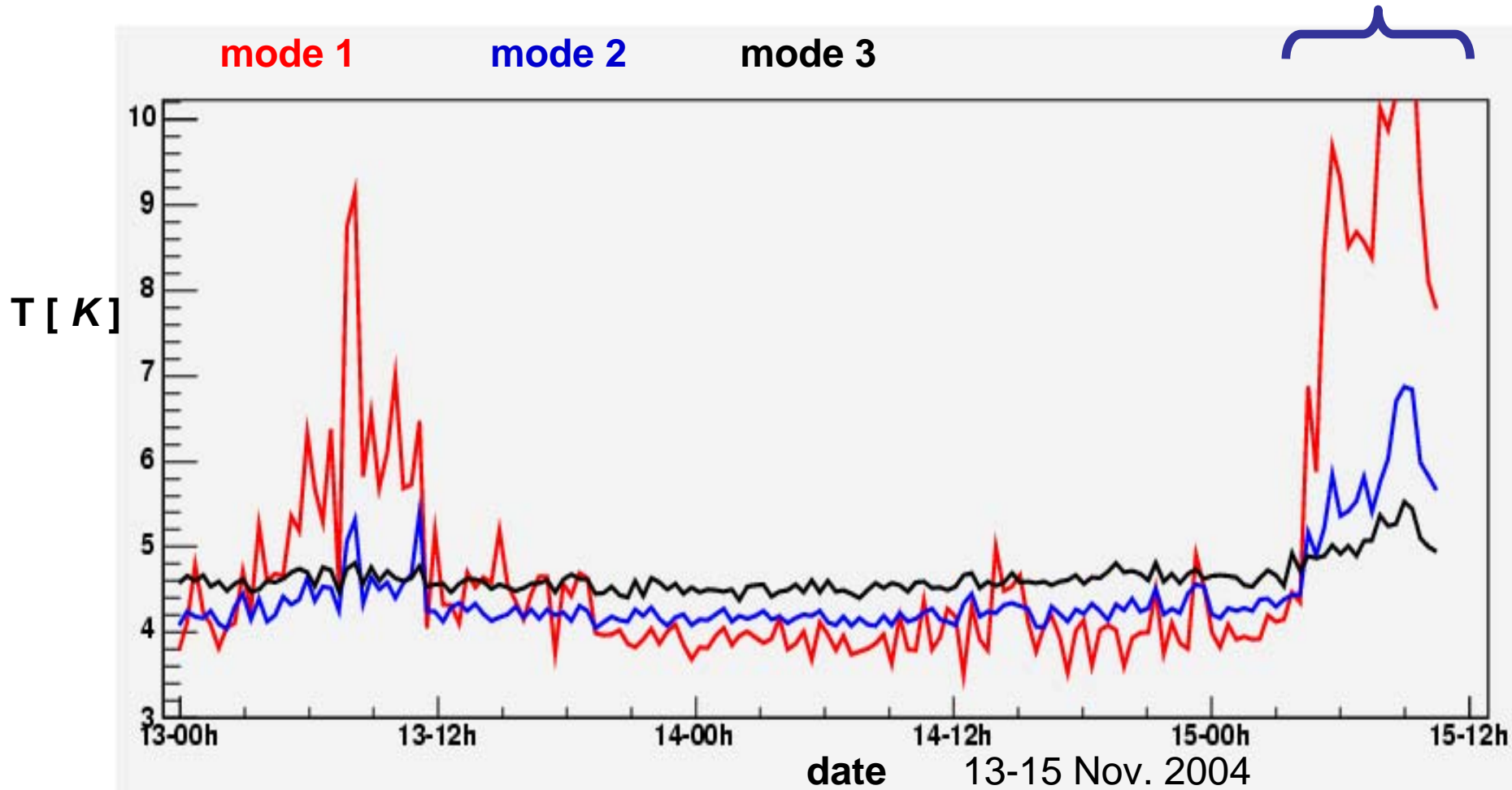
three resonant modes

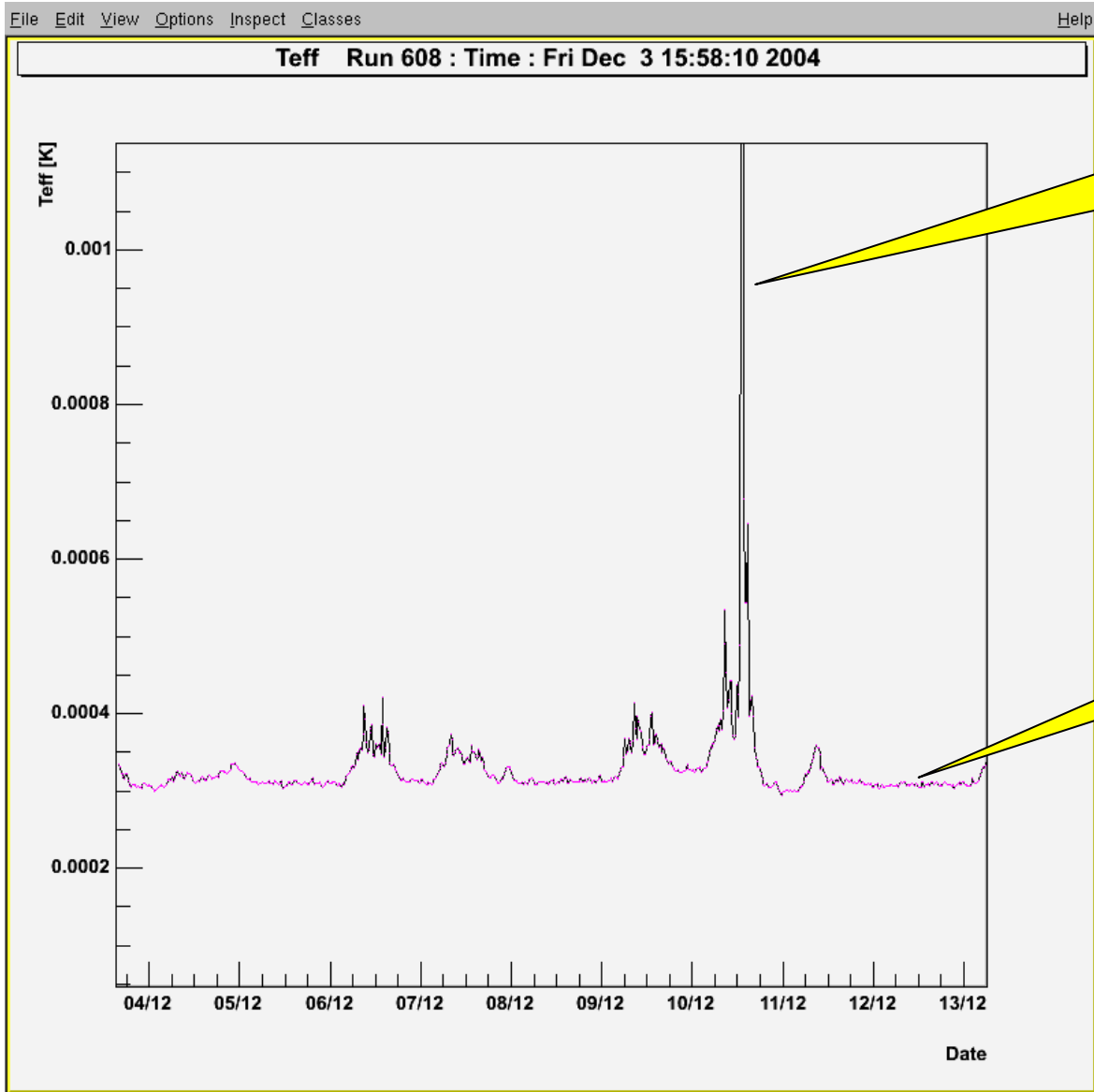


Temperatures of resonant modes

Detector operating at 4.5 K:
the 3 modes achieve the thermal noise level

Monday morning

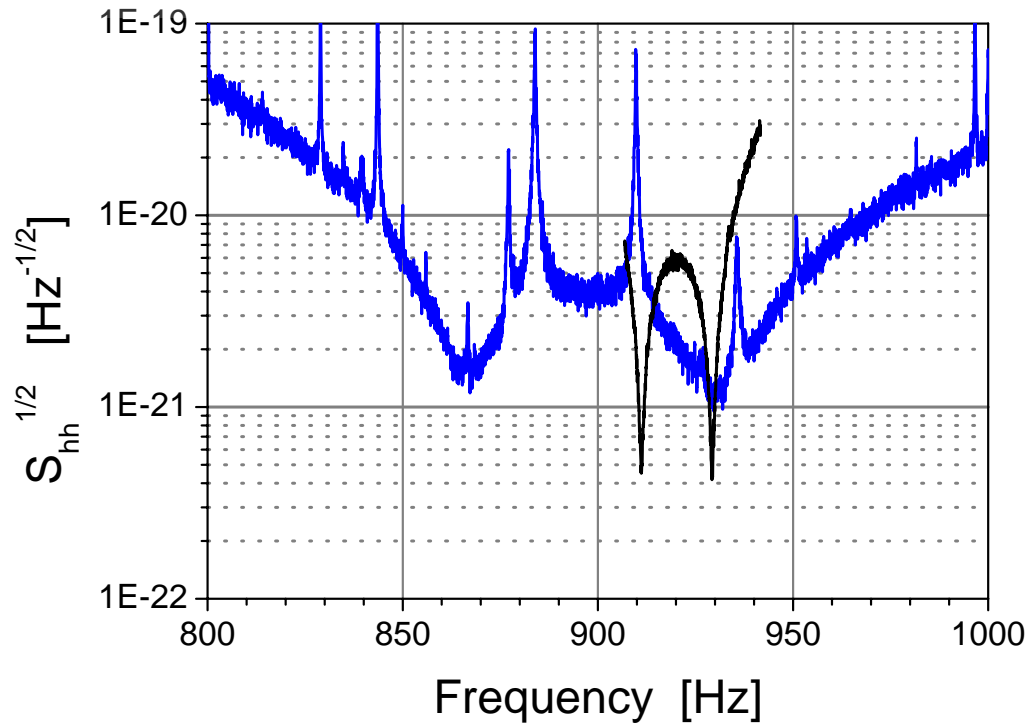




Liq. He
Transfer

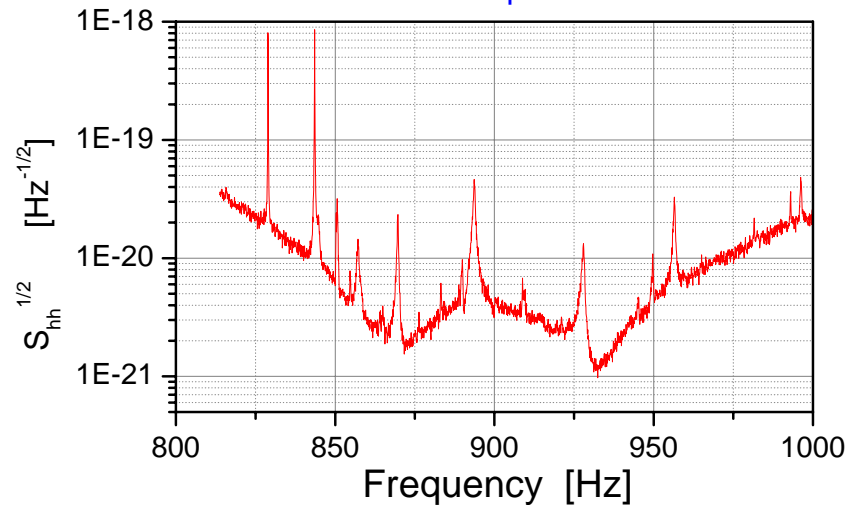
Teff = 300 μ k

initial operation at 4.5 K started
on Dec. 24th 2003

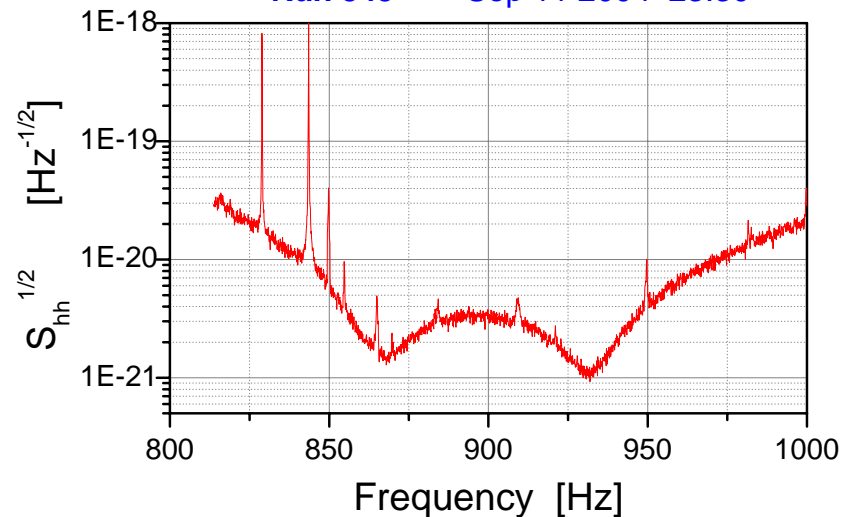


... months of diagnostic
measurements and noise hunting ...

Run 545 Sep 11 2004 19:00



Run 545 Sep 11 2004 23:30

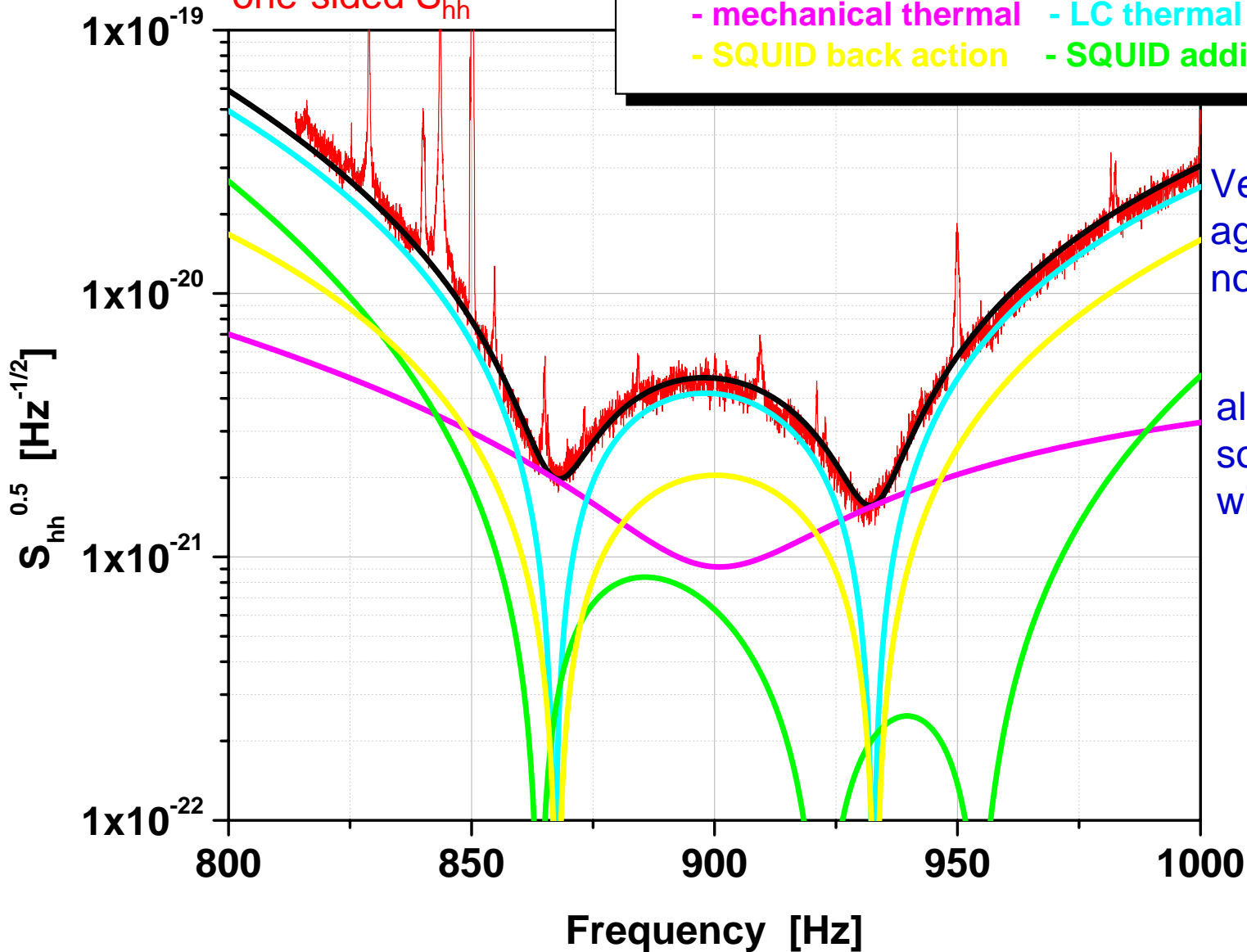


unmodeled spurious noise peaks within the
sensitivity bandwidth

- not related to the dynamical linear response of the detector
- non gaussian statistics
- related to mechanical external disturbances

up-conversion of low frequency noise

AURIGA T=4.5K
one-sided S_{hh}



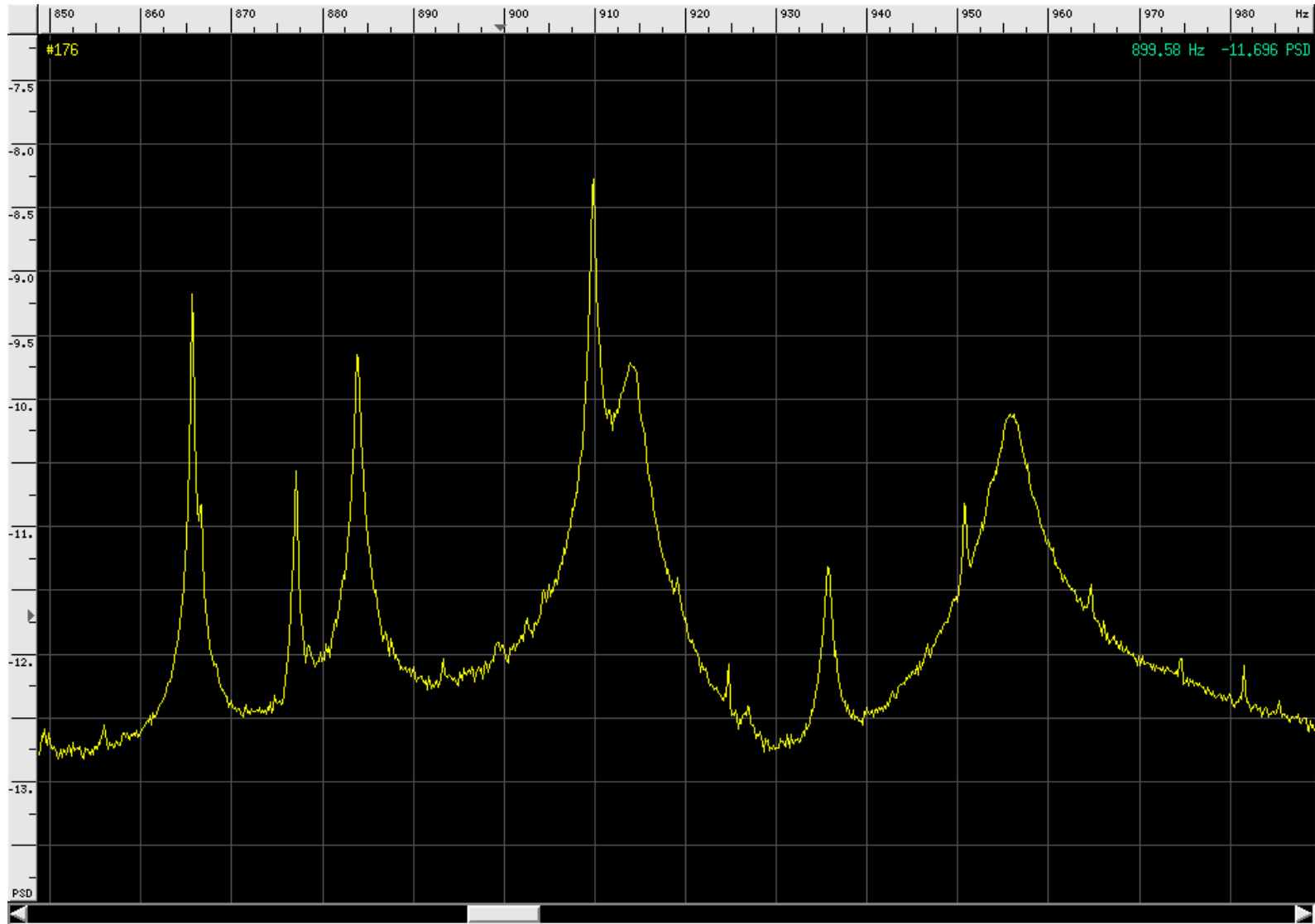
- noise prediction

- mechanical thermal
- LC thermal
- SQUID back action
- SQUID additive

Very good agreement with noise predictions

all these noise sources will scale with temperature

Typical FT output in the period Jan-March 04 with strong spurious lines of the “family” F1



the 3 modes

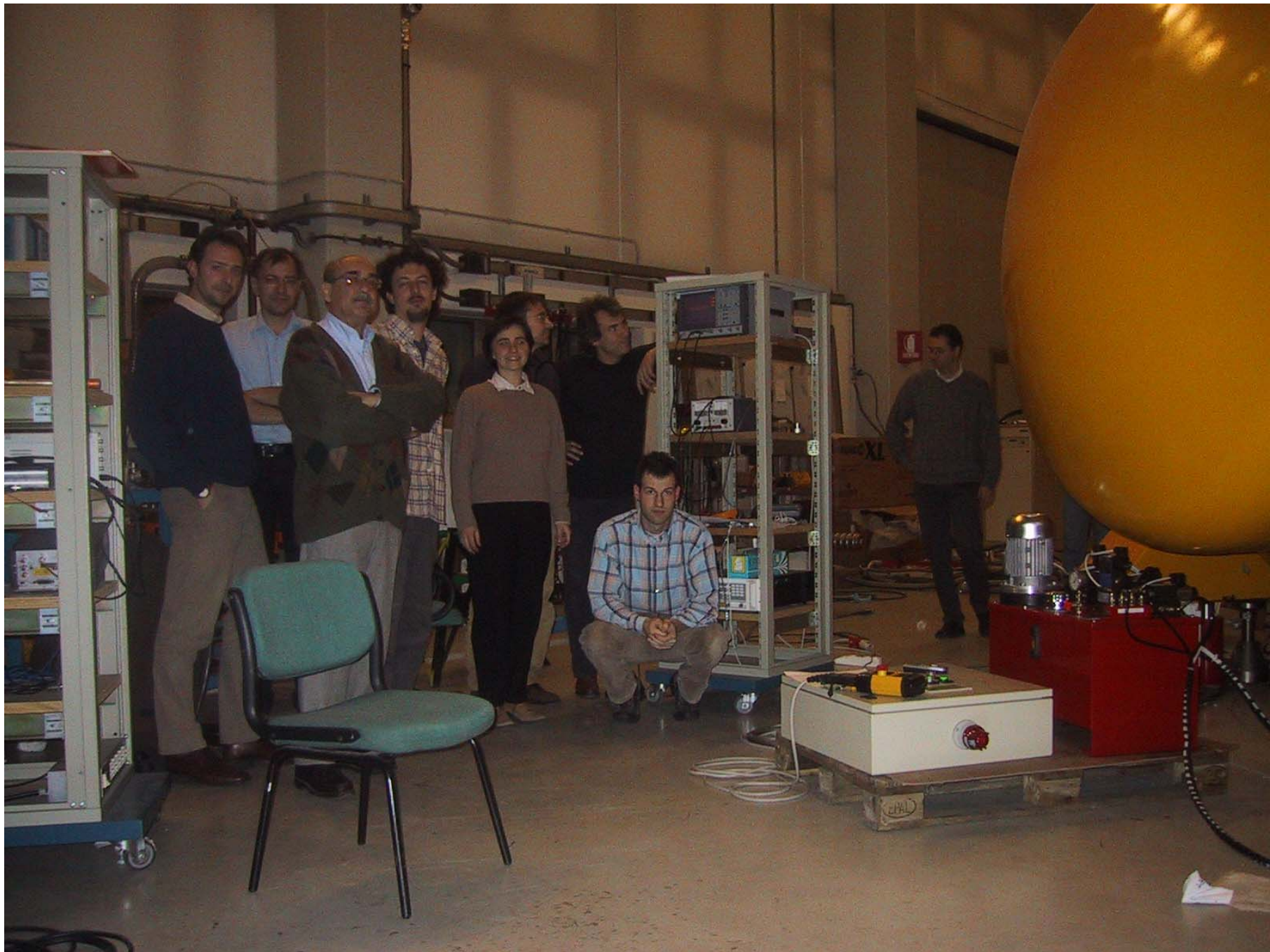
(with a few spuria going on/off and one all-band glitch)

“bar”

“transducer”

“electrical”

QuickTime™ and a
Graphics decompressor
are needed to see this picture.





F8 Veto

236.95 +/- 0.2 Hz

F8

F7

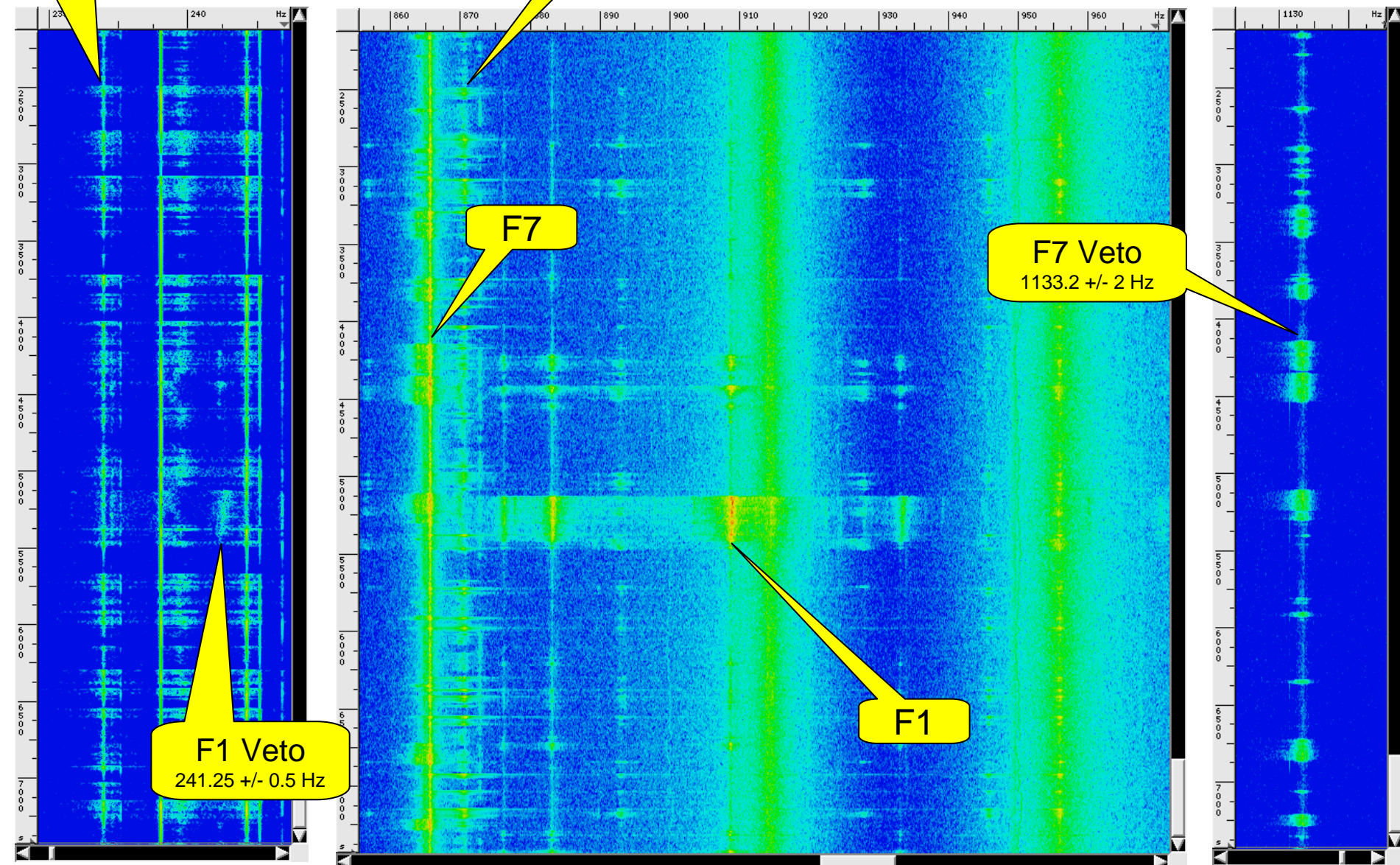
F7 Veto
1133.2 +/- 2 Hz

F1 Veto

241.25 +/- 0.5 Hz

F1

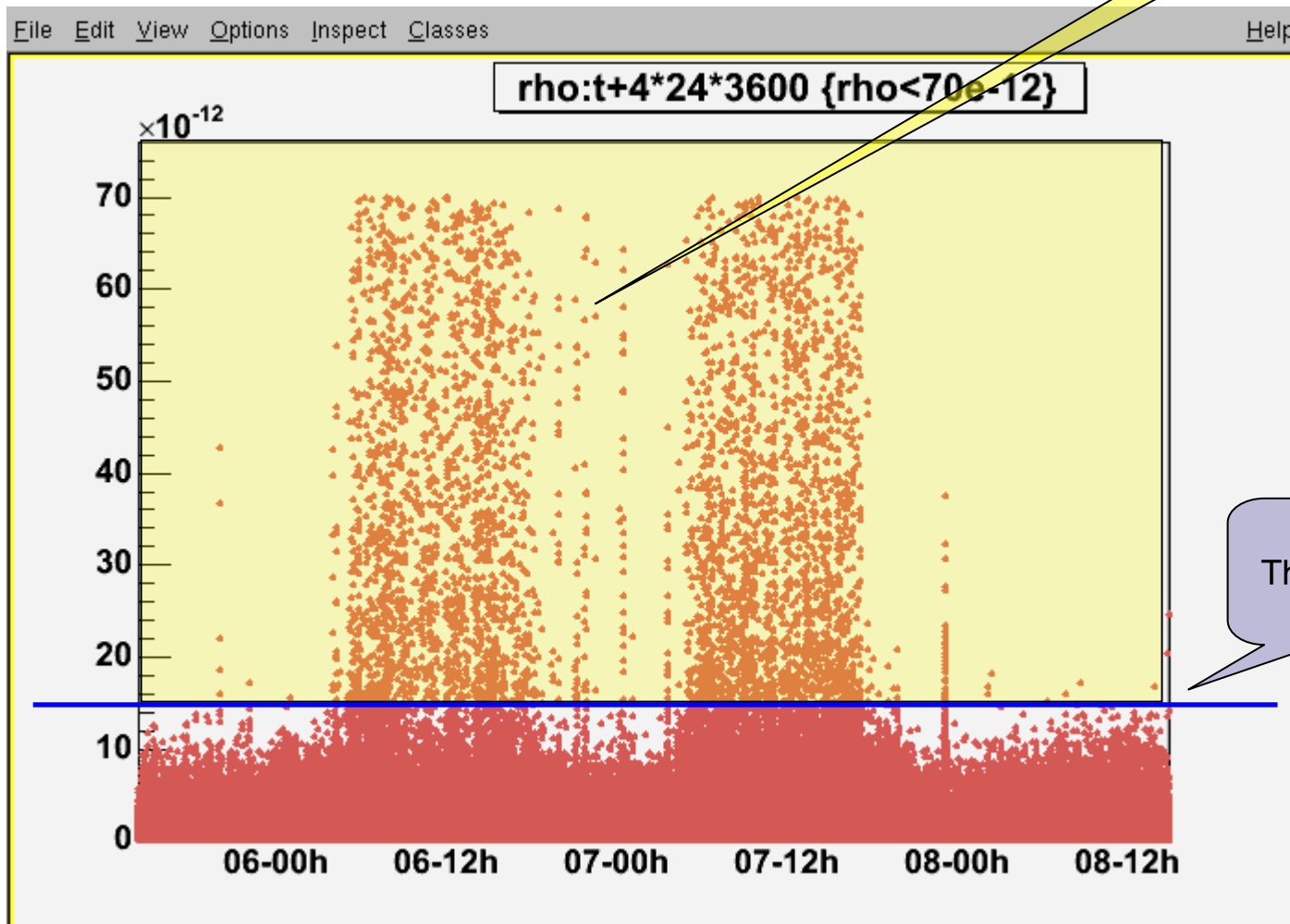
Veto



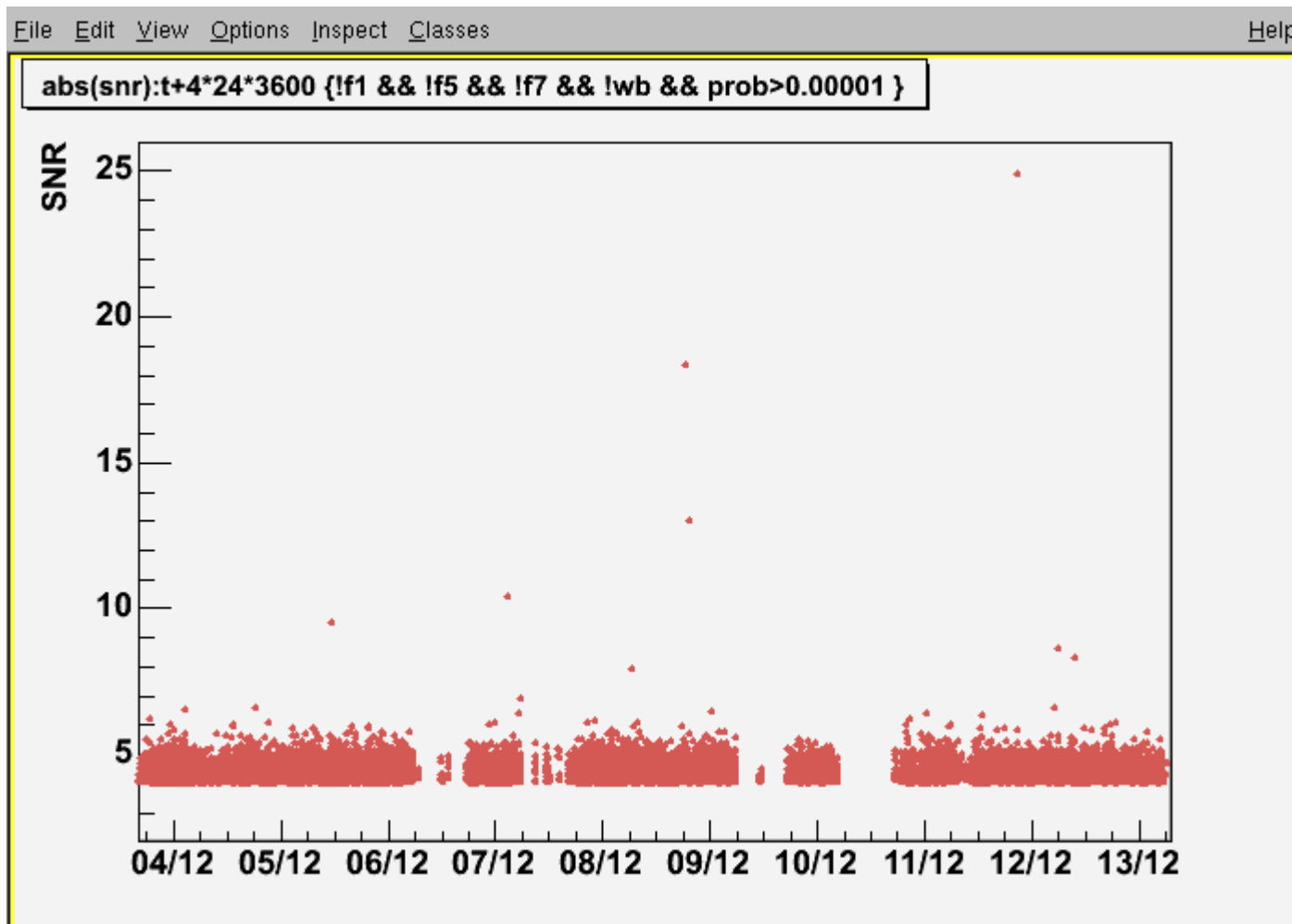
Ada lck psp plugin @ 1133.2 +/- 2 Hz

F7

amplitude



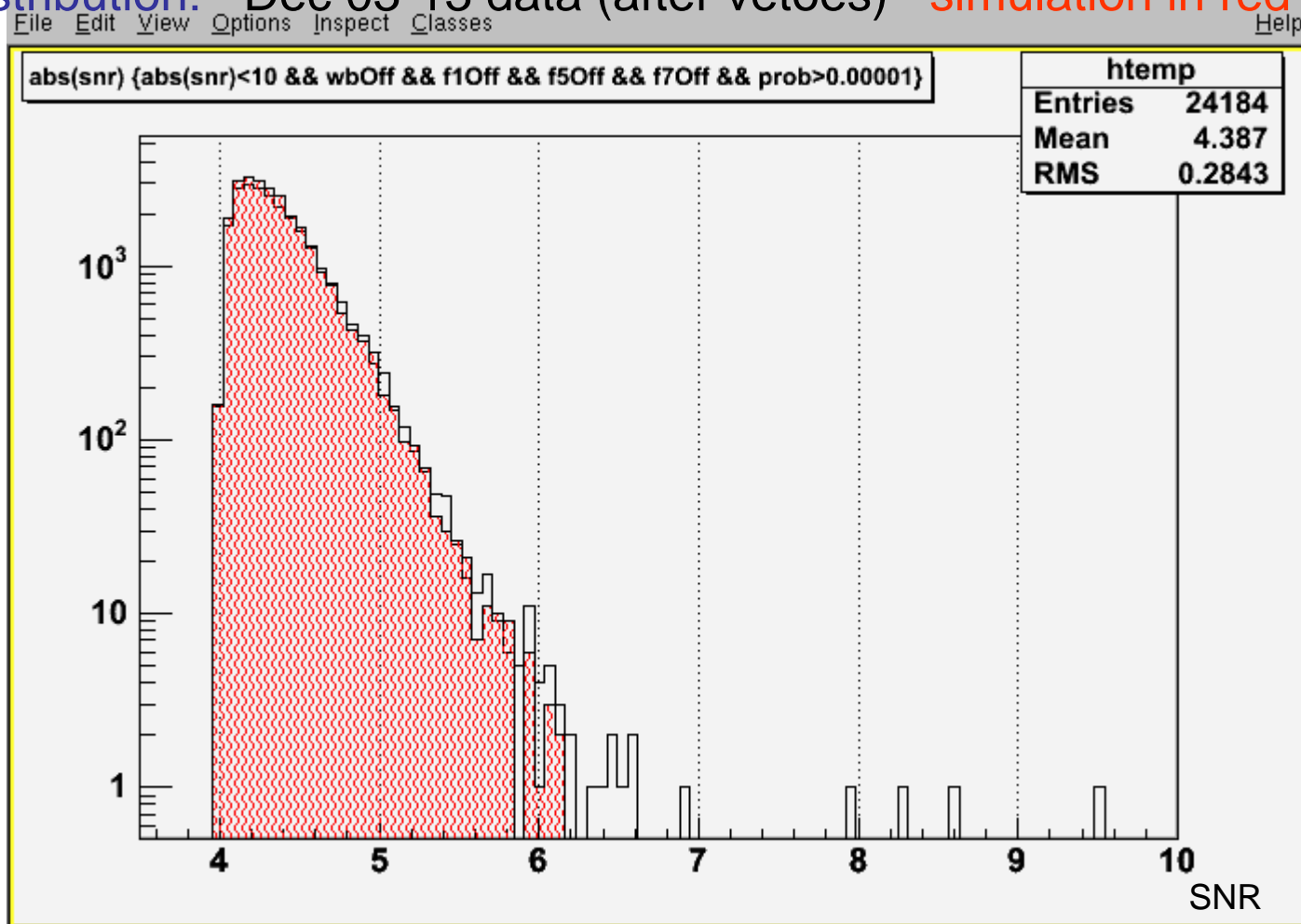
10 days of events after vetos and χ^2 test



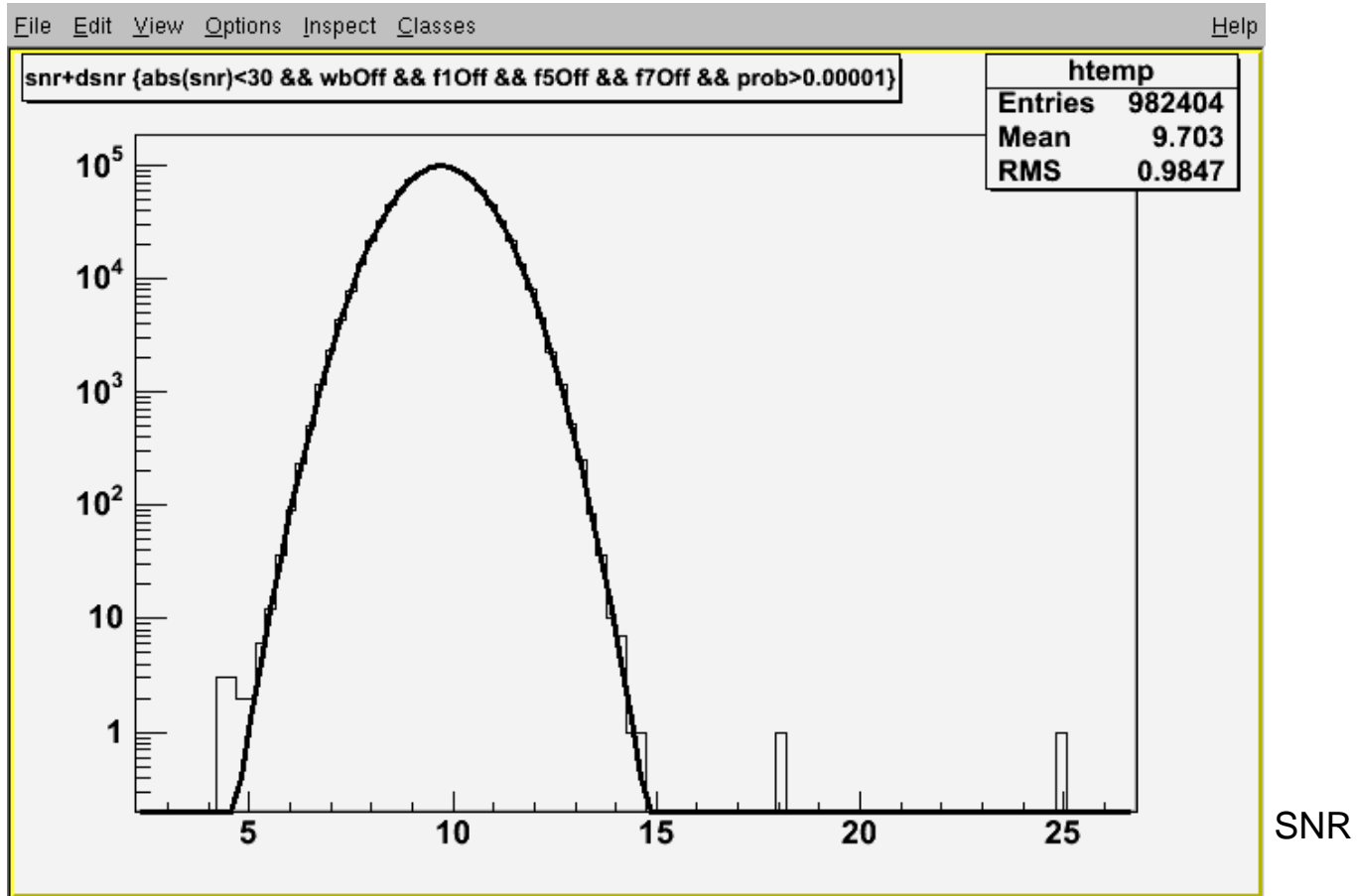
SNR = 5 corresponds to $h_{\text{burst}} = 7 \cdot 10^{-19}$

10 days of stationary gaussian operation:
< 2 outliers/day ~60% duty cycle

SNR distribution: Dec 03-13 data (after vetoes) simulation in red



Montecarlo injecting SNR=10 events: efficiency after vetoes

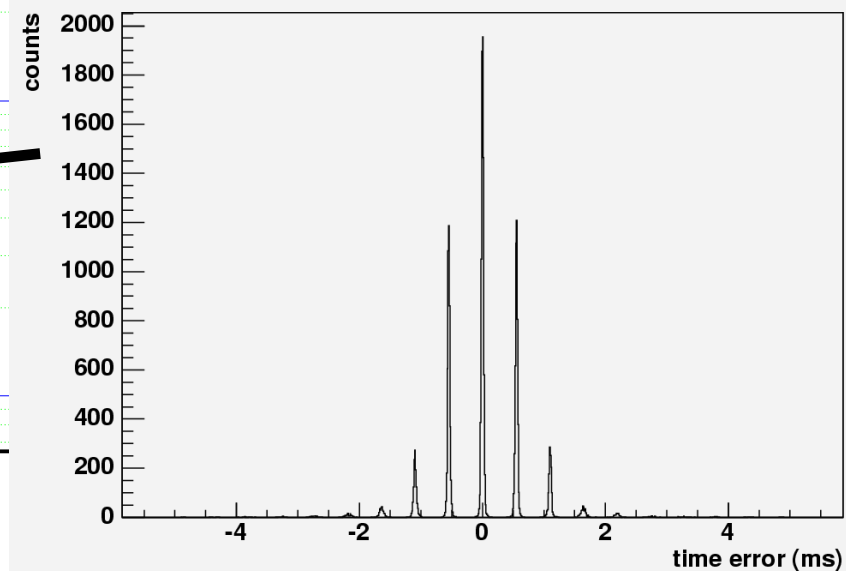
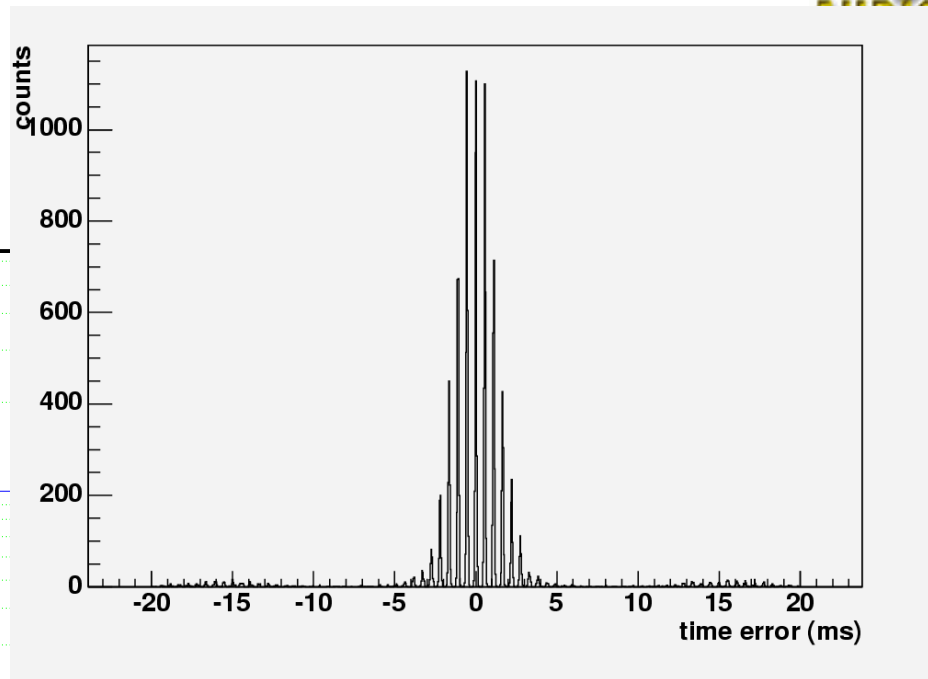
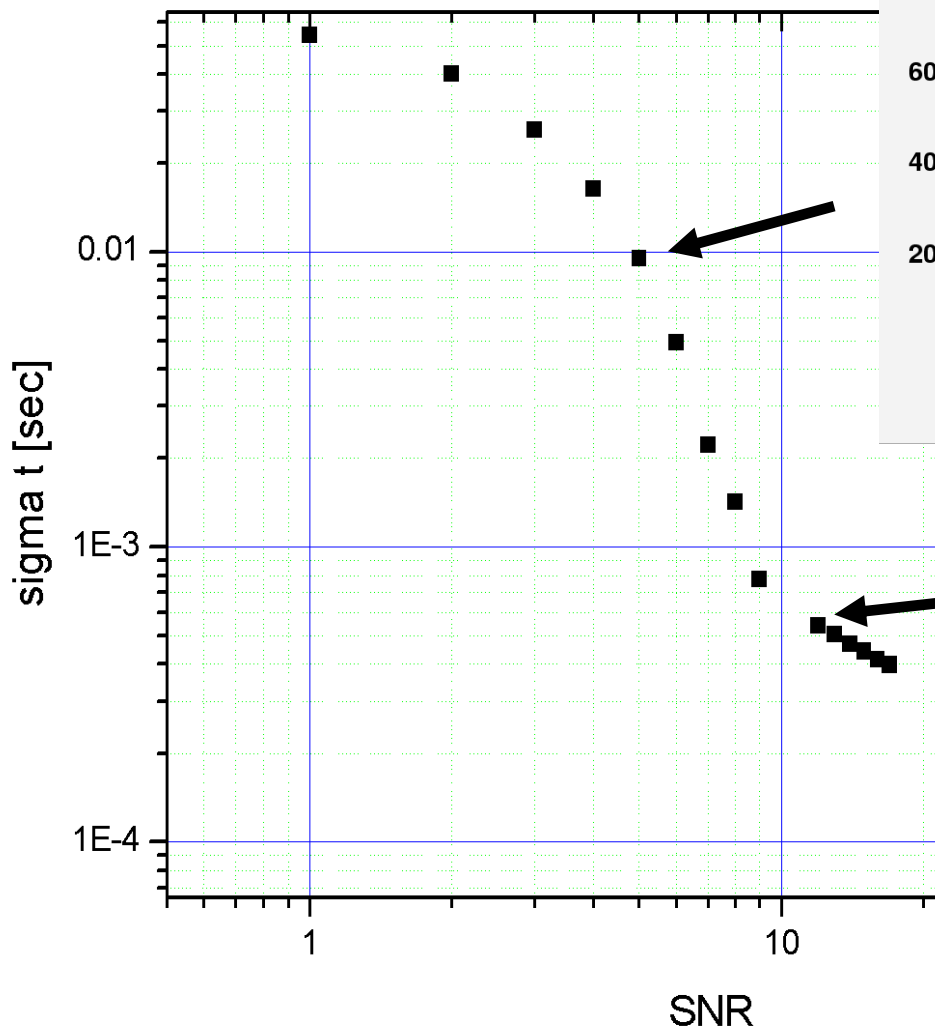


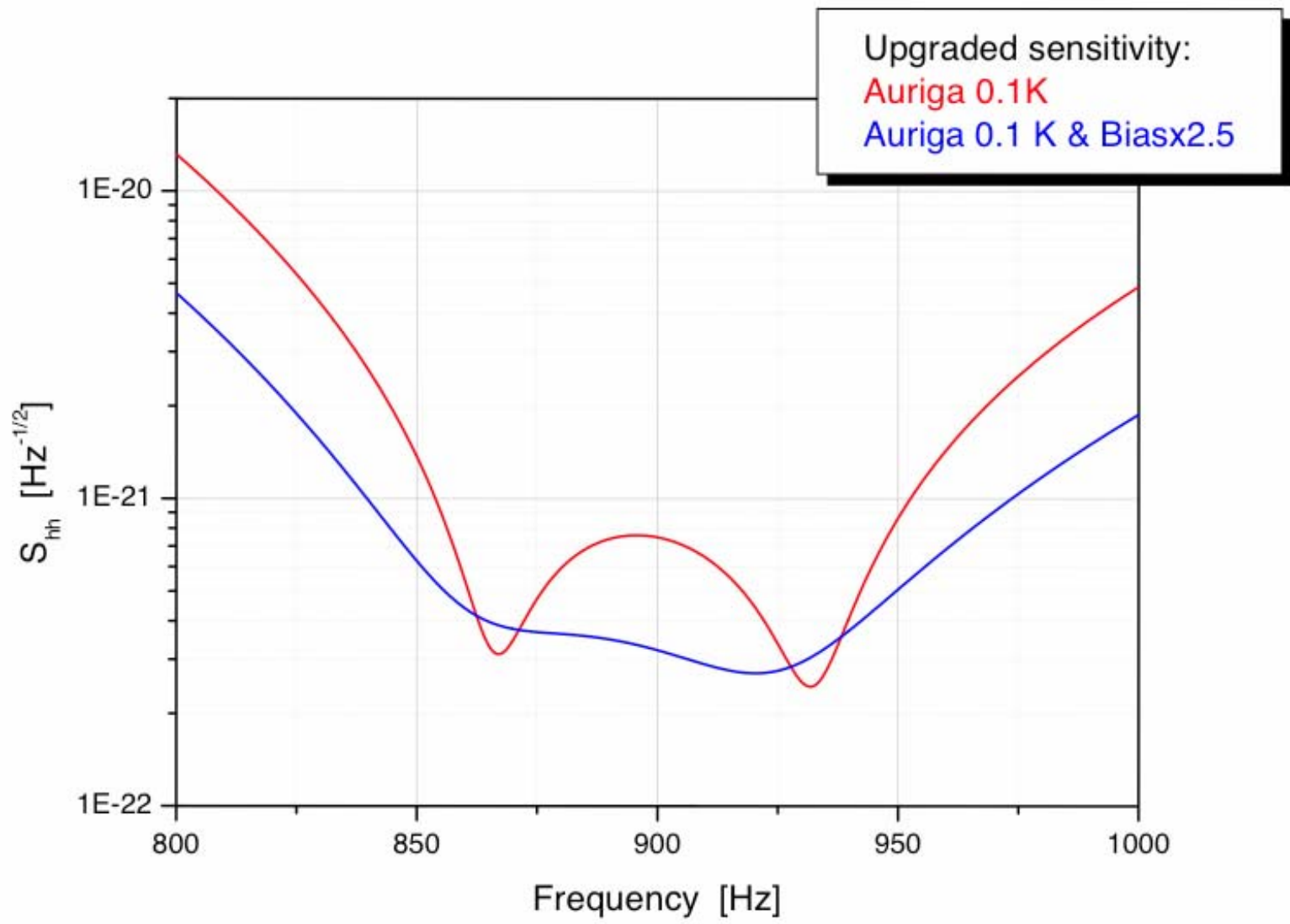
SNR distribution as predicted by gaussian statistics

mean $\text{SNR}_{\text{recovered}} = 9.7$

$N_{\text{recovered}} / N_{\text{injected}} = 0.59$ checks with fraction of time accepted after vetos

uncertainty on arrival time estimates for δ -like bursts





Upgraded sensitivity:
Auriga 0.1K
Auriga 0.1 K & Biasx2.5