

PulsarMon Update

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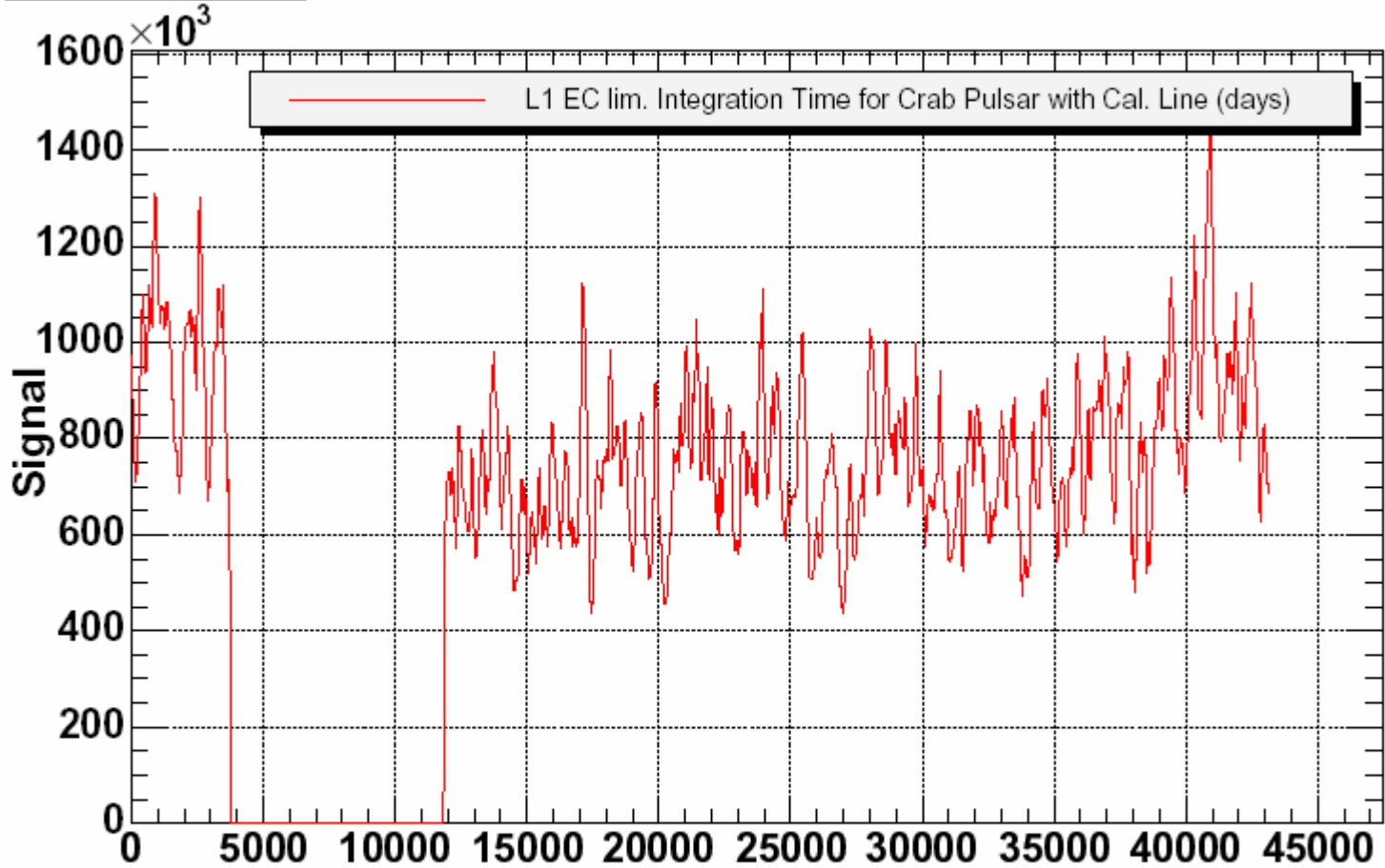
Detector Monitor for Pulsar Search

- Figure of Merit (FOM) in terms of parameters relevant to Pulsar Search
- Main FOM is the Integration Time (IT) for the Crab ($f_s=58.6$)
- The complete list of FOMs is
- **Time Series:** IT for Crab, Ellipticity at 500 Hz, 1kpc distance
- **Power Spectra:** Noise Spectrum normalized to 1 year IT, Ellipticity for test pulsar 1 kpc vs. frequency
- **Scatter Plots:** hEC for known pulsars, IT for known pulsars, Ellipticity for known pulsars vs. frequency

Crab parameters: $f_s=59.6$ Hz $F_{\dot{t}}= 3.86e-10$ s⁻²,
 distance 2kpc

$$h_{EC} = \frac{5.7 \times 10^{-24}}{r/1kpc} \sqrt{\frac{f_s}{1kHz} \frac{\rho \&}{10^{-13} s/s} \frac{I}{10^{45} gcm^2}}$$

Time series



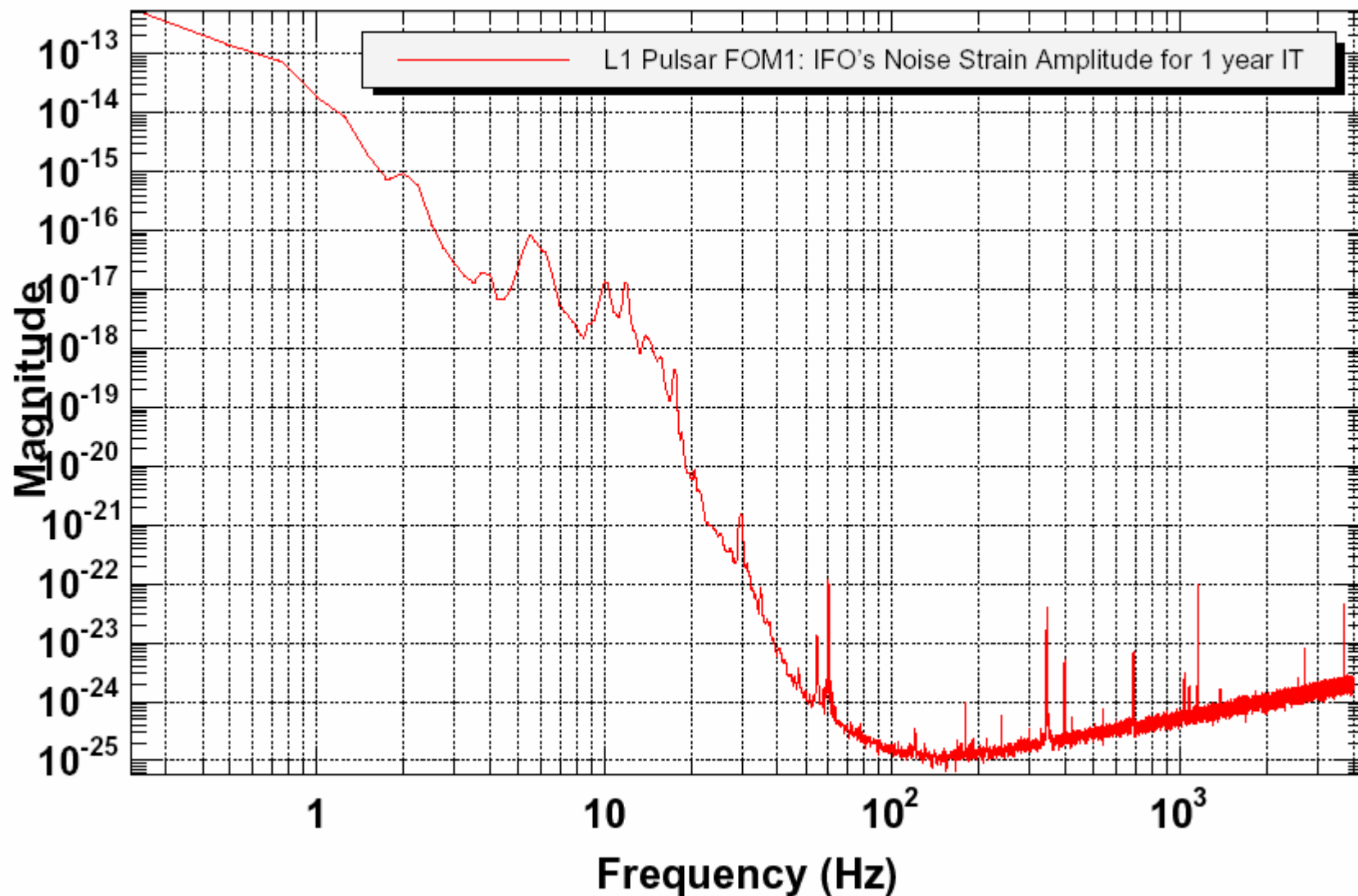
T0=06/03/2005 12:14:59

Time (s)
 Avg=1

$$IT = 11.5 \frac{S_n(f)}{h_{EC}^2}$$

Power spectrum

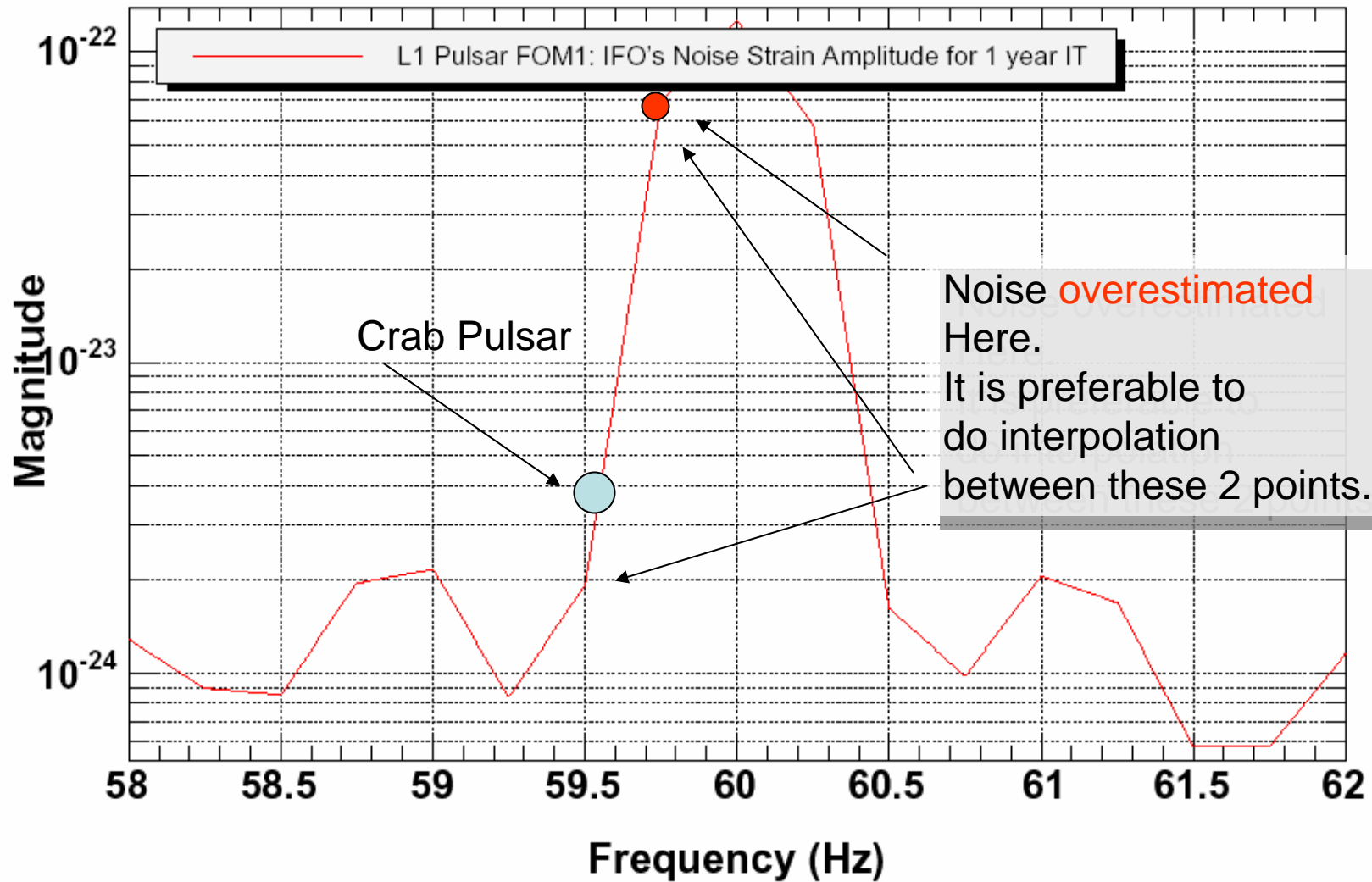
$$h(f) = \sqrt{S_n(f) / T}$$



T0=21/03/2005 01:01:45

Avg=15

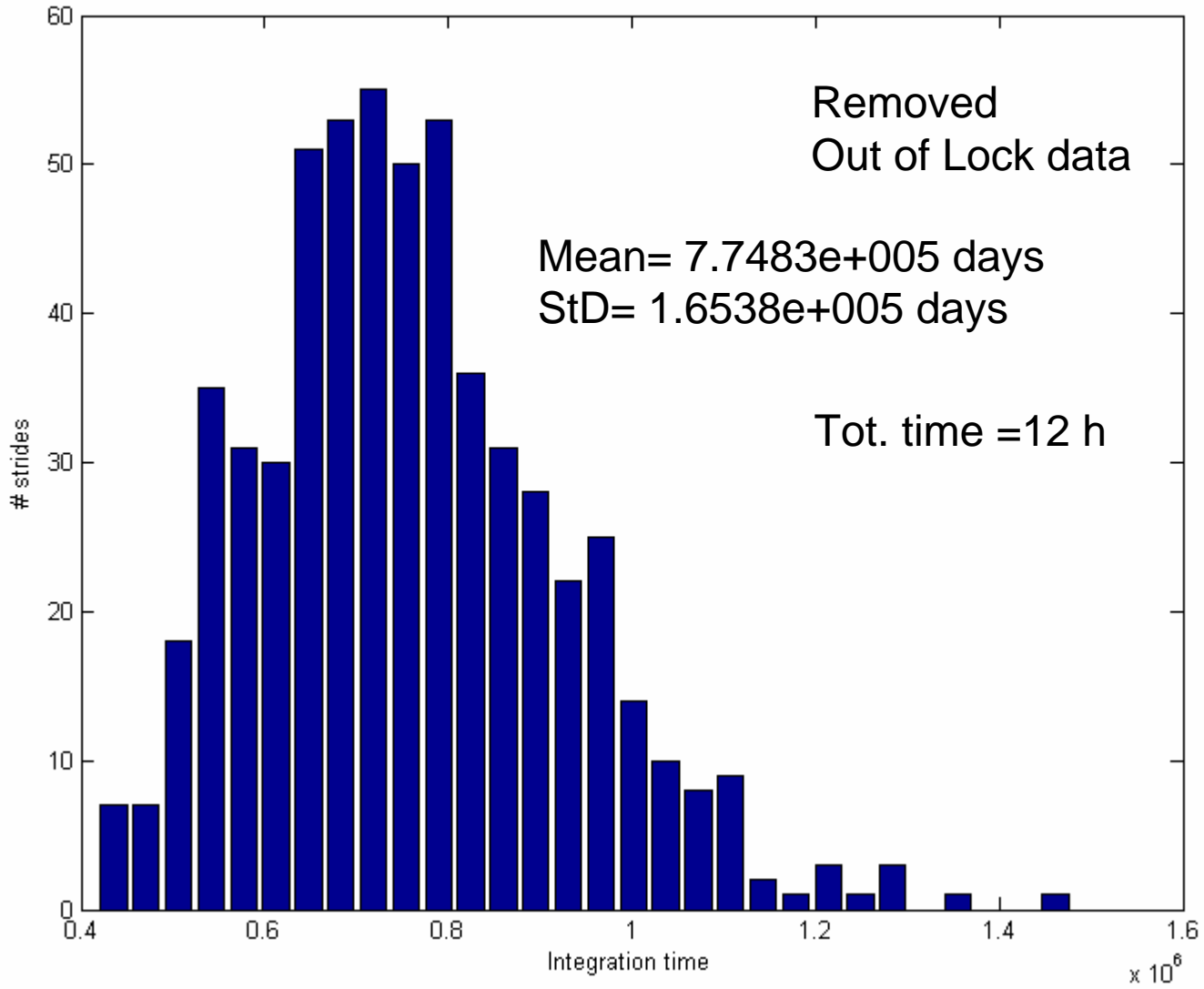
Power spectrum



T0=21/03/2005 01:04:45

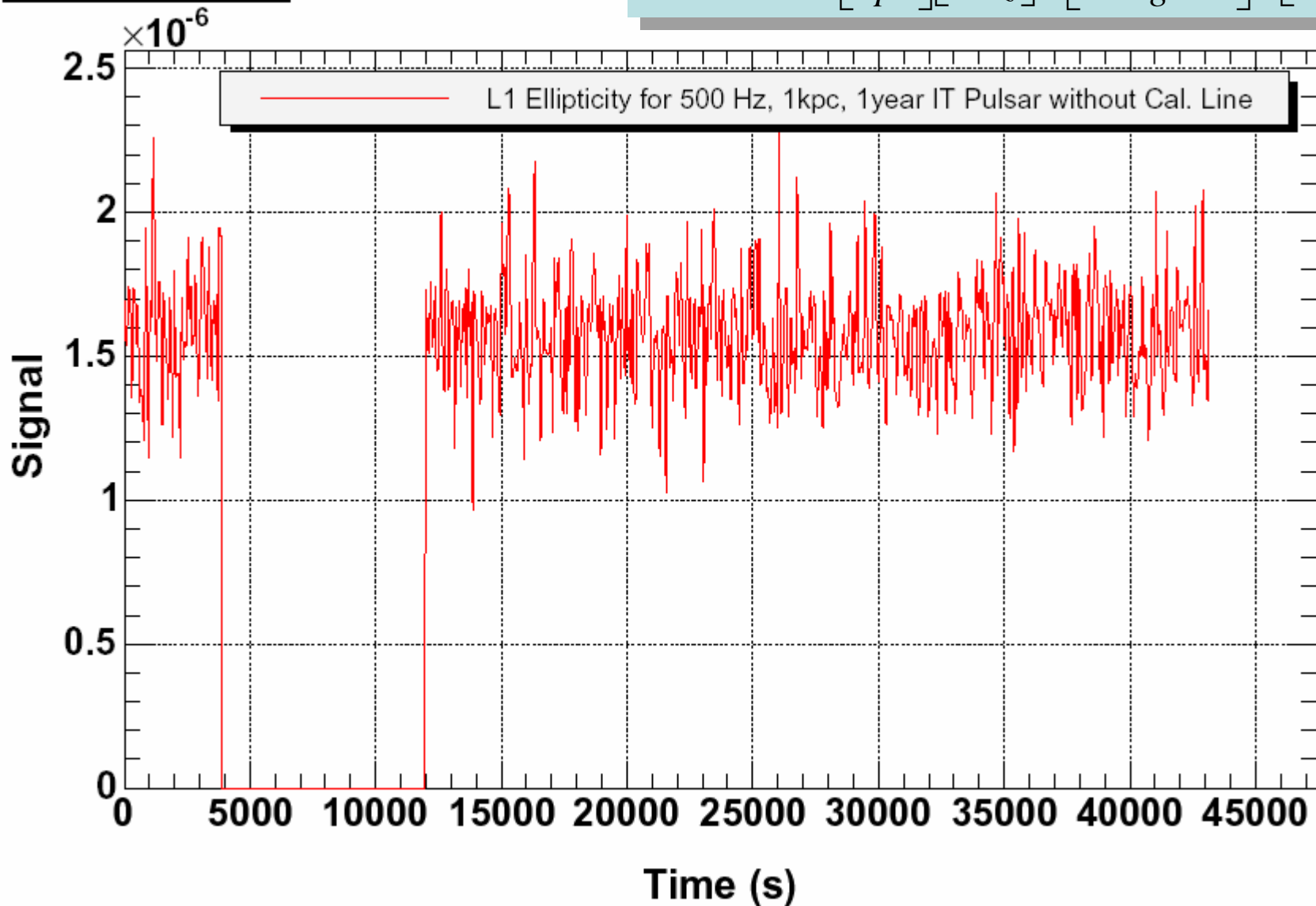
Avg=15

IT for Crab Statistics (03/06/2005)



Time series

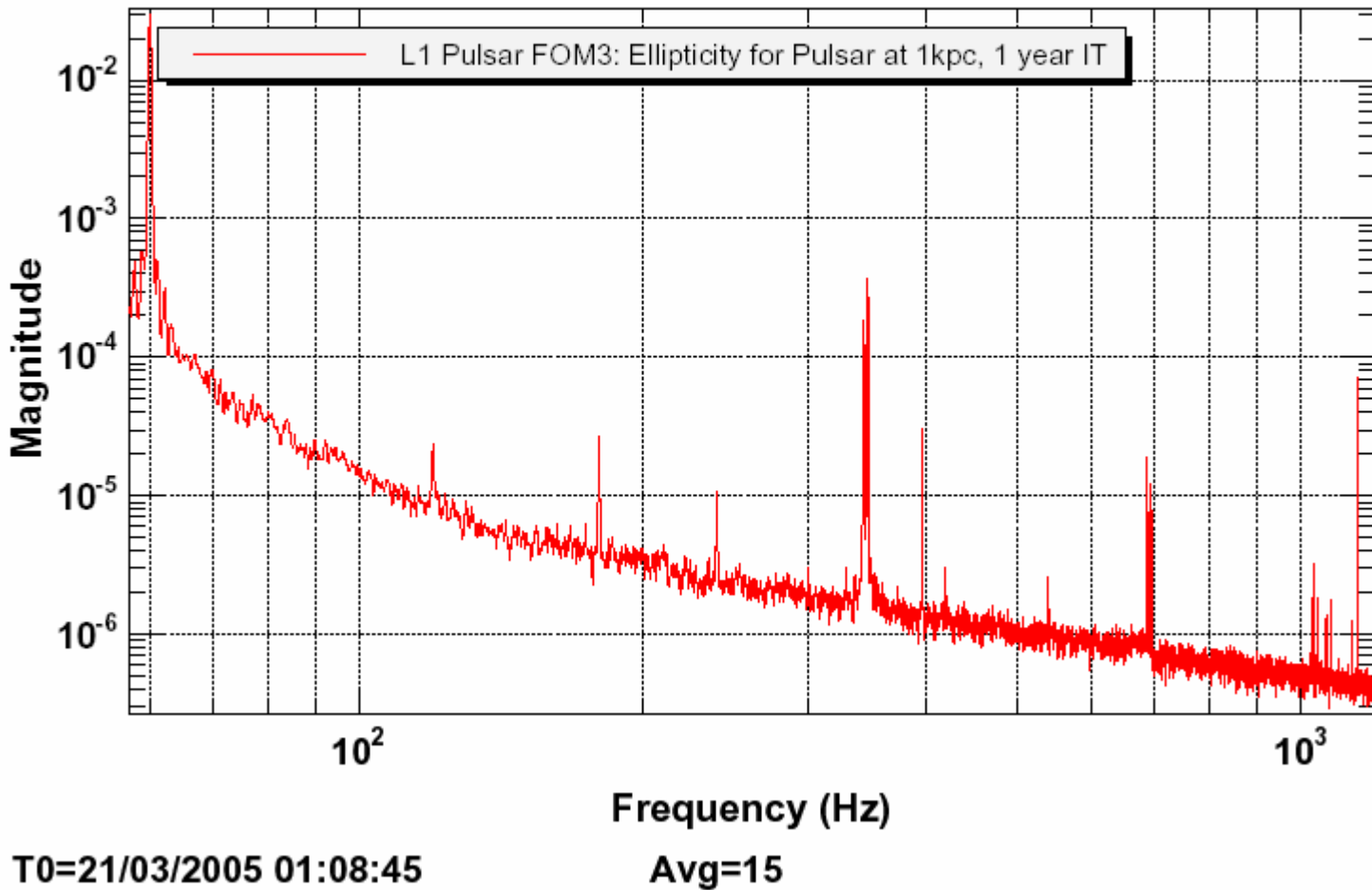
$$\varepsilon = 9.5 \times 10^{-6} \left[\frac{r}{\text{kpc}} \right] \left[\frac{f_s}{\text{kHz}} \right]^{-2} \left[\frac{I}{10^{45} \text{ gcm}^2} \right]^{-1} \left[\frac{h_0}{10^{-23}} \right]$$



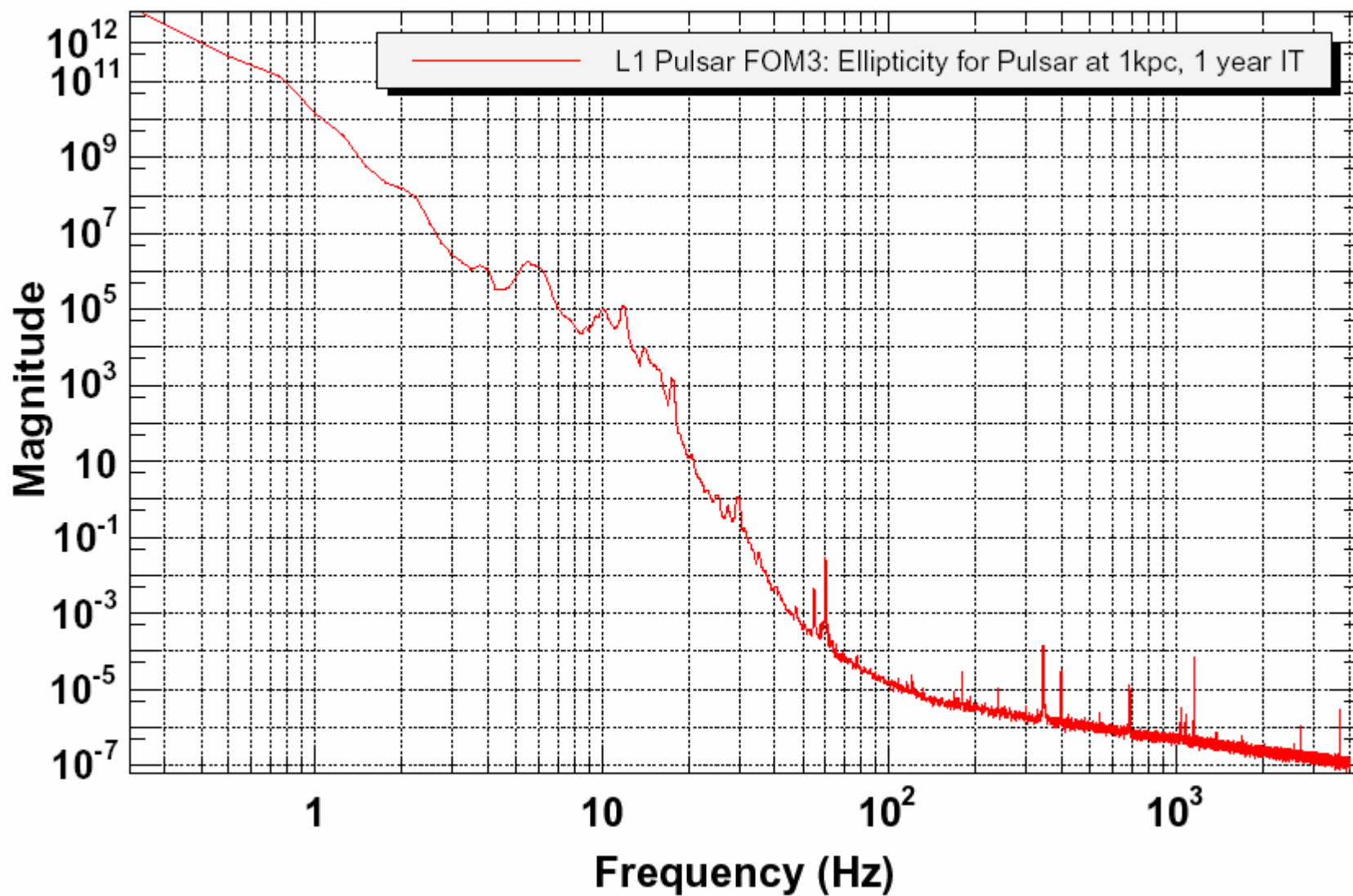
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Avg=1

Power spectrum



Power spectrum



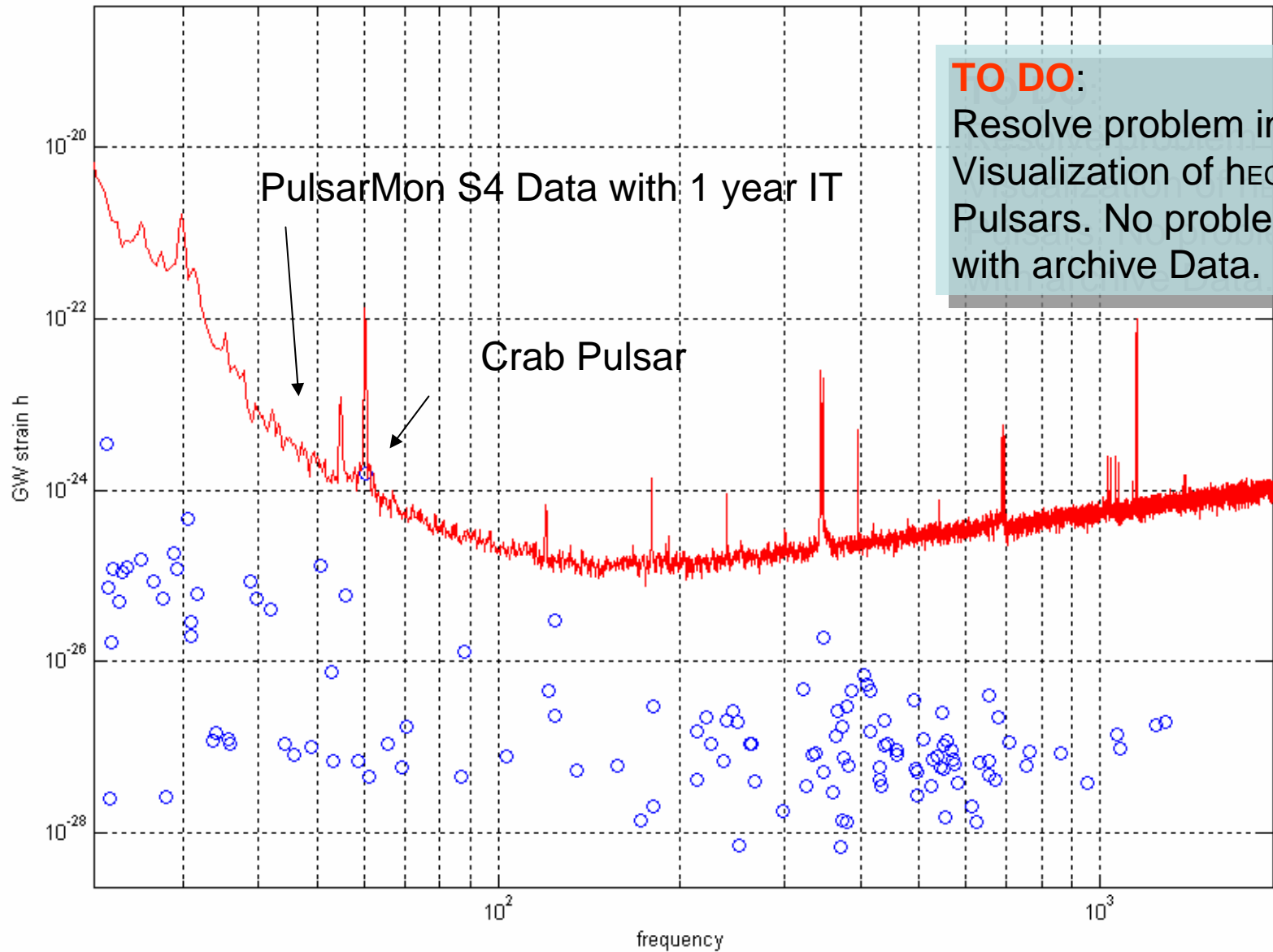
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Avg=15

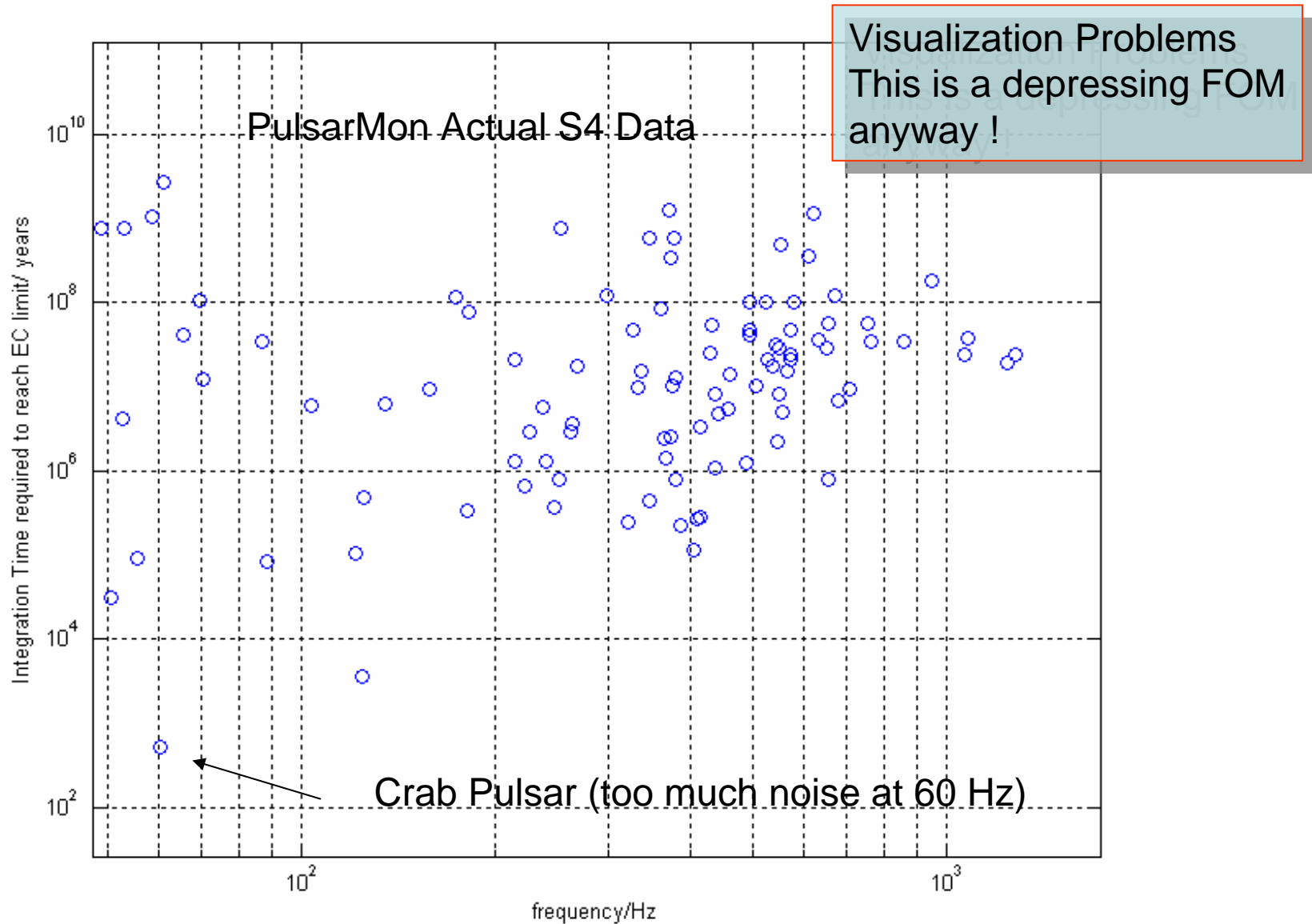
To Do:

- Fix Visualizations Problems with Scattered Plots (FOM2: IT for known Pulsars, FOM3: Ellipticity)
- Use the EasyCalibrate calibration code
(this is already done but PulsarMon EasyCalibrate version is trapped in Alvar, that is not responding right now)
- Interpolate value of h_{noise} for Crab (use two close frequencies near f_{crab})
- Independent Tests and Checks
- New FOMs ???

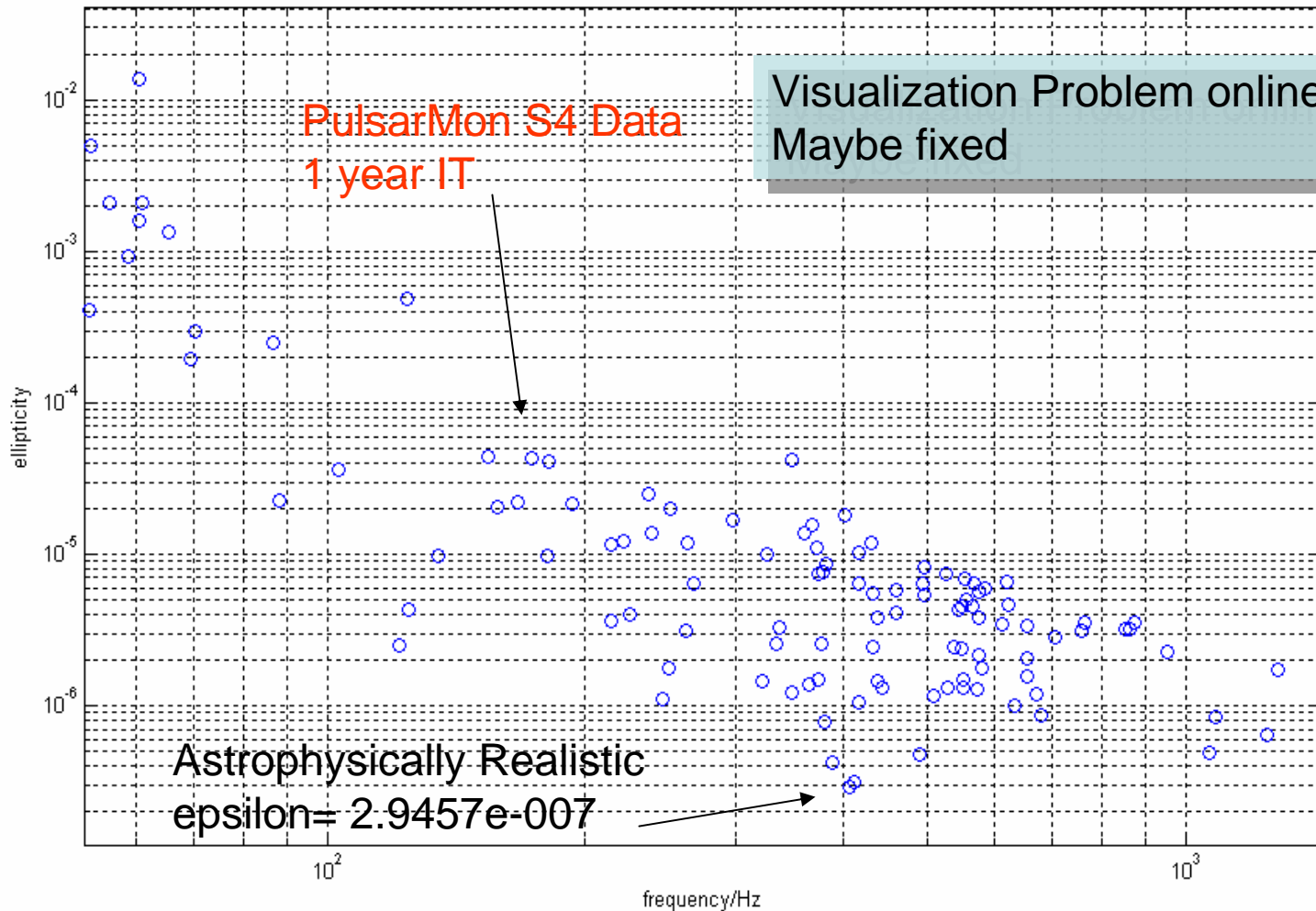
FOM 1: Strain Sensitivity & know Pulsars h_{EC}



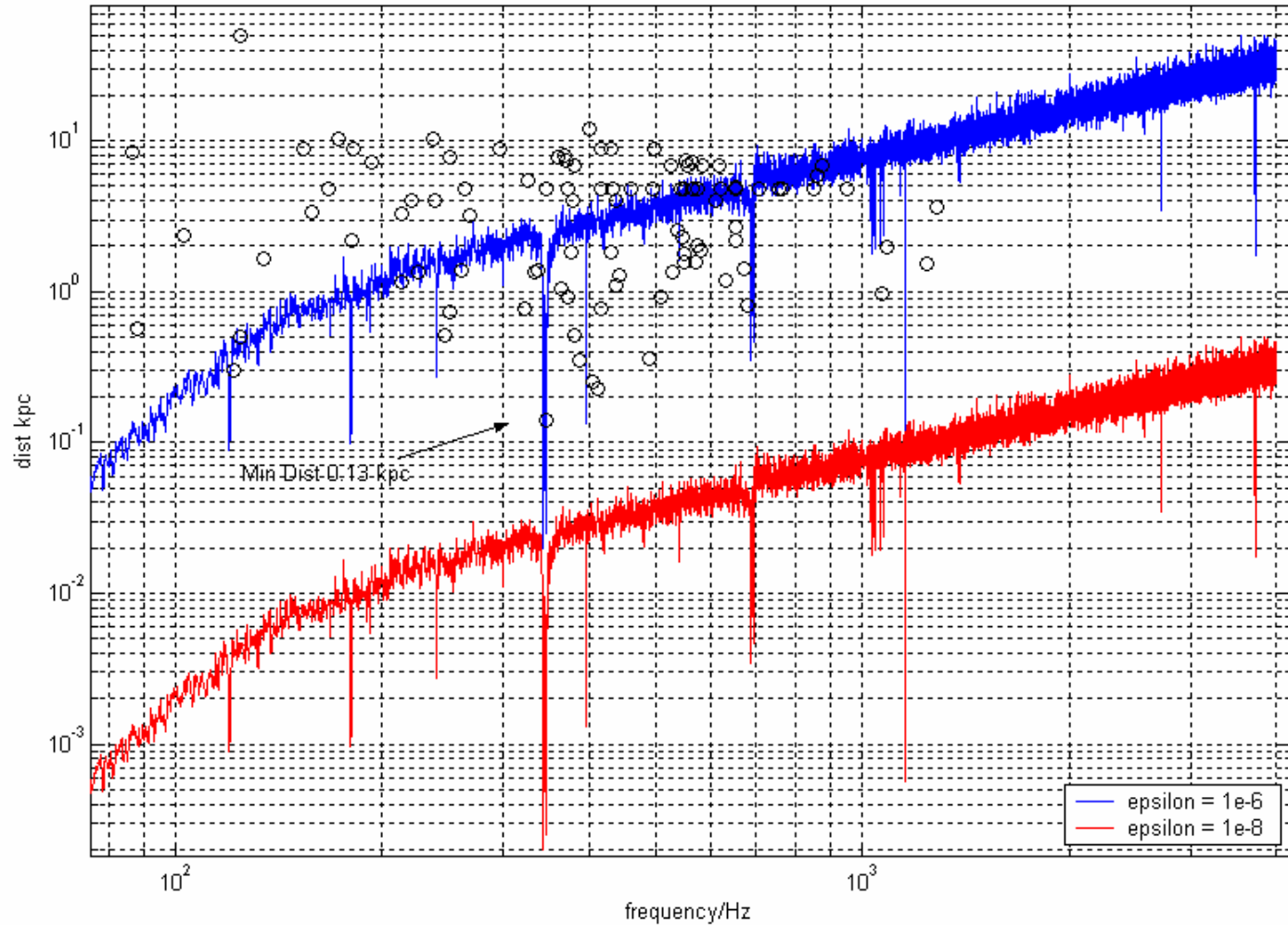
FOM 2: Integration needed to reach EC limit



FOM3: Ellipticity Sensitivity for known Pulsars



FOM 5: Range for Given Ellipticity ??



Conclusions:

- PulsarMon is producing time series and continuous Power Spectra (Crab IT, Ellipticity for 500 Hz, 1kpc pulsar, Noise Strain for 1 IT, Ellipticity vs. Frequency For unknown pulsars at 1kpc)
without any major problems (still some bothering memory leaks)
- Works to be done for Scattered Plots (FOM1: h_{EC} for known pulsars, FOM 2: IT for known pulsars, FOM3: ellipticity for known pulsars)
- EasyCalibrate Implementation
- Interpolation of h_{noise} for Crab
- New FOMs ??

Thanks to Keith, John and Patrick for advise and guidance !