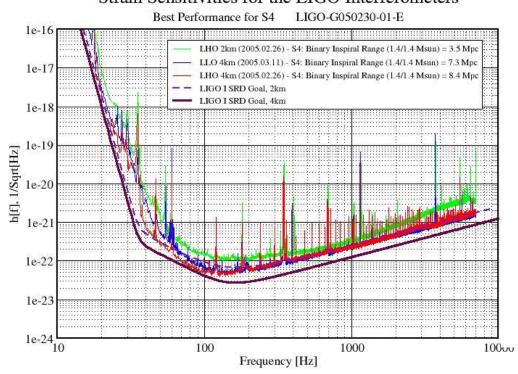
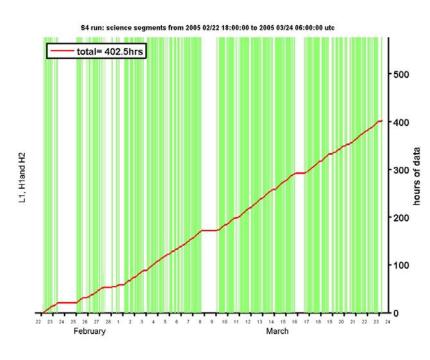


S5 Run Performance Goals

Strain Sensitivities for the LIGO Interferometers





Peter Fritschel & Michael Zucker

PAC18 meeting LIGO Livingston Observatory 18 May 2005



Overview

- S4 a success
 - » Sensitivity good (and astrophysically relevant)
 - » Machines reliable and well characterized
 - » Commissioning, operations, analysis teams in top form
- What are our expectations for S5?
- What are the principal challenges?



Interferometer Duty Factor

| Run | S2 | S3 | S4 | S5 Target (proposed) | SRD goal |
|-------|-----|-----|-----|----------------------|-------------|
| L1 | 37% | 22% | 75% | 85% | 90% |
| H1 | 74% | 69% | 81% | 85% | 90% |
| H2 | 58% | 63% | 81% | 85% | 90% |
| 3-way | 22% | 16% | 57% | 70% | 75% |

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Interferometer Duty Factor

- HEPI works; L1 no longer vulnerable to logging, μseism
- What's the remaining background of downtime?
 - » H1, H2: high winds uncorrelated with L1, ~ 10% uptime loss
 - » L1: extreme weather *uncorrelated*, ~ 1-3% loss
 - » All 3: global earthquakes (magnitude > 5) correlated, ~ 1-3%
 - » All 3: equipment maintenance *correlated*, ~ 7-10%
 - e.g., liquid nitrogen deliveries, HVAC maintenance
 - current experience bottoms at about 1/2 day per week
 - » All 3: hardware & software failure uncorrelated, high @ run start
 - Latter 1/4 of S4 uptimes averaged 87% (L1), 86% (H1), 90% (H2)
- What are realistic goals?
 - » 3 x 85% @ 70% triple looks achievable without heroic investments
 - » Approaching 3 x 90% @ 75% triple goal may require a wind noise solution, near 100% equipment reliability, and much more experience

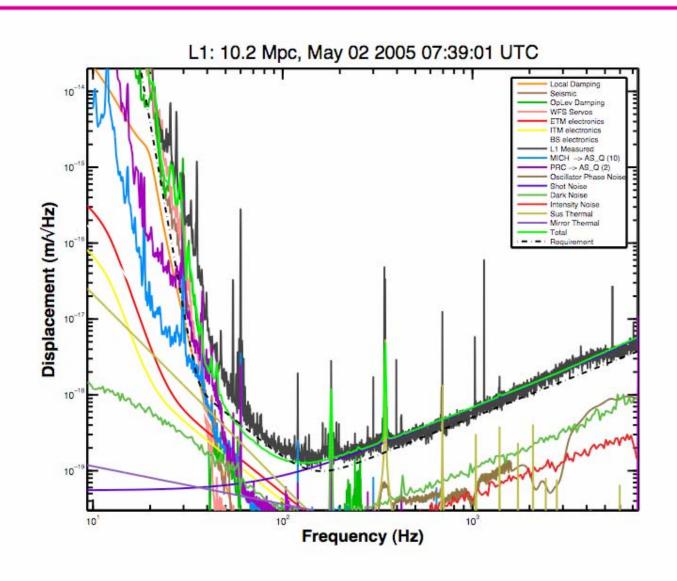


S4 Sensitivity

QuickTime™ and a TIFF (Uncompressed) decompressor are needed to see this picture.



L1 Noise Budget 5/2/05





Noise goals

- At this writing we are largely through our post-S4 upgrade lists
- Main remaining tasks:
 - » Optic replacement on H1
 - » (and/or) TCS boost to handle anomalous TM heating
 - » May need other upgrades (e.g., more PD's) depending on results
- Power up req'd on all three machines
 - » Still running at 2-4 W vs. design goal of 6 W
- After resolution of H1 optics:
 - » principal effort is to get comfortable with high power and work ancillary noise issues

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Challenges

Technical:

- » H1 lossy optic R&R and/or TCS upgrade critical for high P operation
- » High power operation not yet fully demonstrated
- » Low-frequency noise not understood
- » 60 Hz spikes still an issue

Schedule:

- » LLO Science Education Center construction starting September!
 - Wait for earthmoving & concrete to finish, or start run at diminished duty factor?
 - Post-steelwork construction activity not expected to interfere as much

Operations:

- » Site and interferometer reliability, maintenance in "steady state"
- » Insufficient ops staff to maintain watch (except in "burst mode")

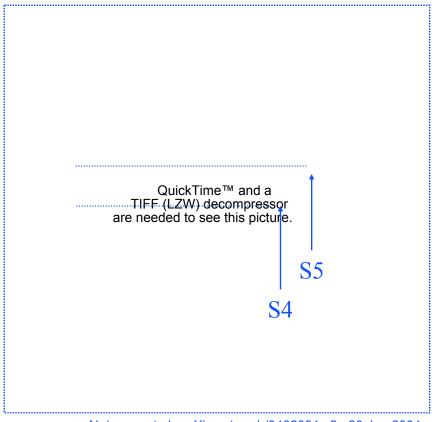
LSC Community:

- » Science monitoring model not viable for long run
- » Are we ready to drink from the data firehose?



Perspective

BNS inspiral: well into VIRGO cluster



Nutzman et al., arXiv:astro-ph/0402091 v2, 28 Jun 2004



Conclusion

- Tasks in progress should net about a factor of 1.4 in strain sensitivity with respect to 'S4 best' (~ 10-12 Mpc binary inspiral range for H1 and L1, ~ 5-6 Mpc for H2)
- Assuming no new surprises, we think we know how to get ~70% triple coincidence uptime
- => Current (preliminary) strain and duty factor performance already look sufficient to achieve primary S5 science objectives
- New "Run Planning" protocol affords reasonable structure for managing controlled improvements during an extended run

but....

Some challenges remain (not all technical)