

Targeted Searches using Q Pipeline

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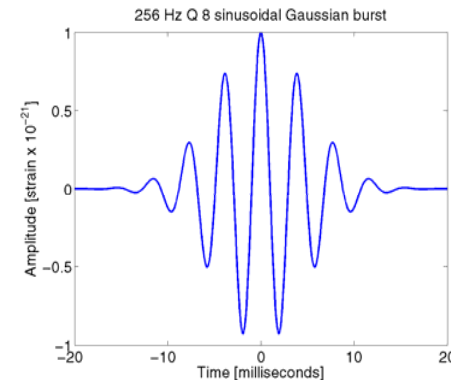
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Outline

- Overview: Burst GW Sources
- Q Transform
- Motivation
- Coherent Targeted Searches
- Consistency Checks
- Examples
- Conclusions

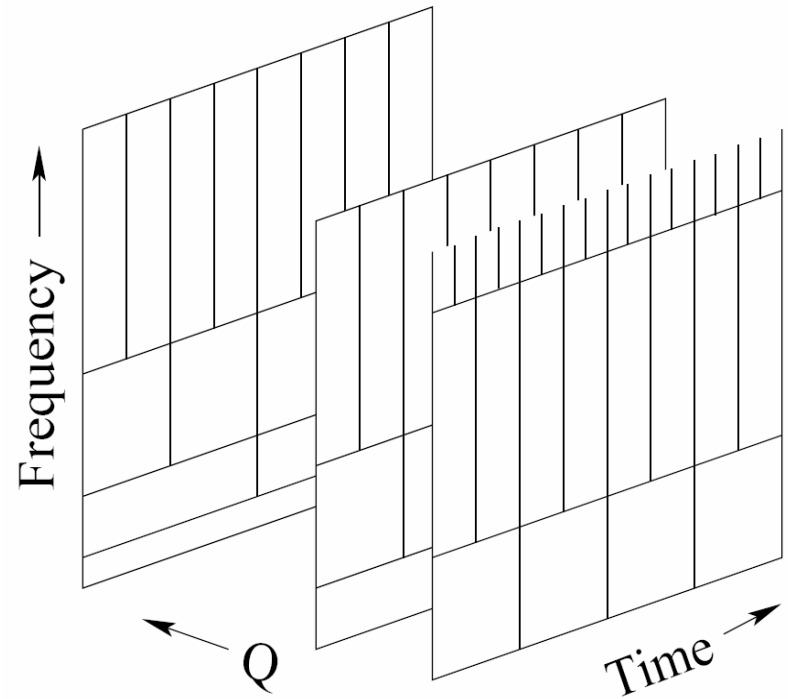
Bursts

- Unknown or crudely modeled
- “Expectations”:
 - Duration: $\sim 1\text{-}100$ ms
 - Frequency: $>\sim 500$ Hz
- Searching:
 - Matched Filter: Project data onto known waveform
 - Cross-correlation: Project data stream from one detector onto another
 - Time-Frequency: Project data stream onto a basis of waveforms designed to cover the targeted signal space



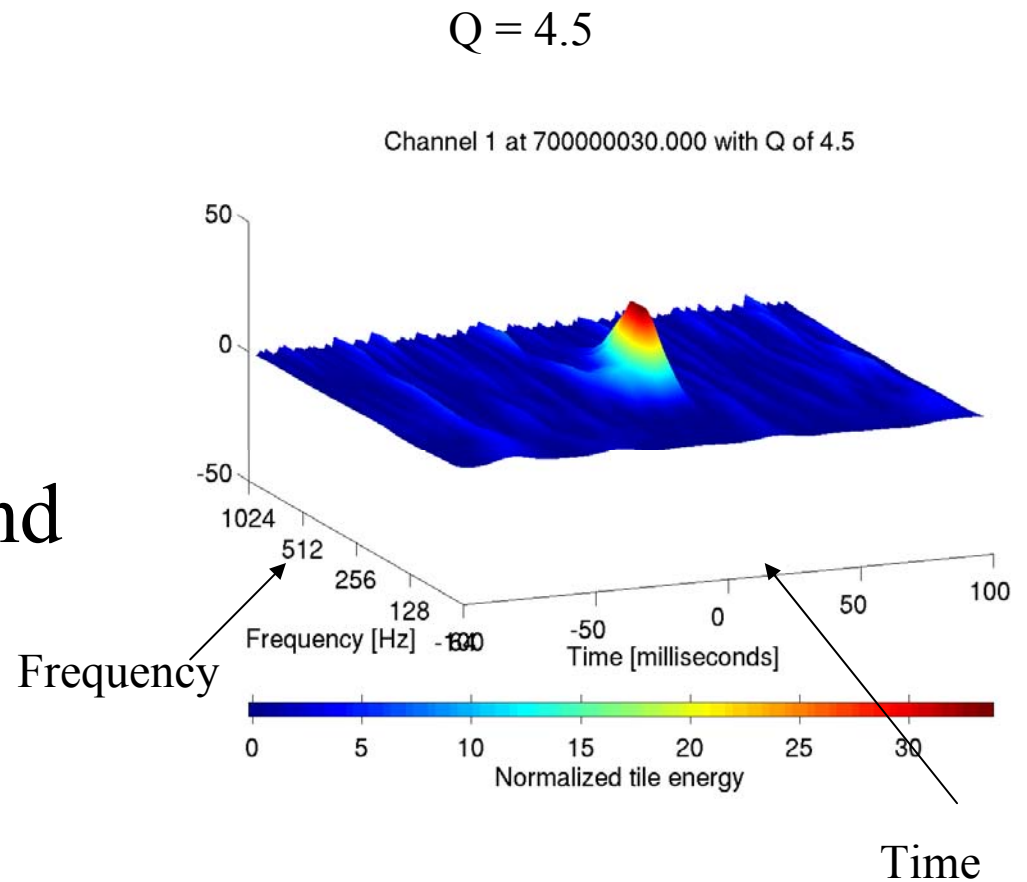
Q Transform

- Multi-resolution basis
- Time, frequency, Q
- $Q = f/\Delta f$
- Over-complete basis
- Purpose is detection not reconstruction
 - Logarithmic spacing Q
 - Logarithmic spacing f
 - Linear spacing in t



Q Transform

- Simulated Data
- SG235 + simulated noise
- Signal $Q = 16$
- Problem: Burst GW and noise can look alike
- Vetoing techniques crucial



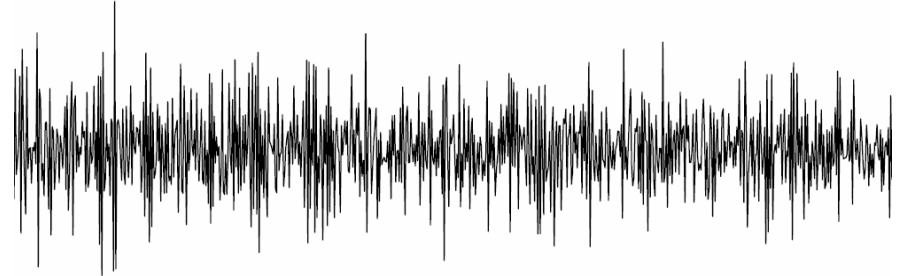
Multiple Detectors

- Coherent Sum
- Three or more needed (Gursel, Tinto, 1989)
- Achieve higher SNR using HL network
- Distinguish GW from Glitches using H1 and H2



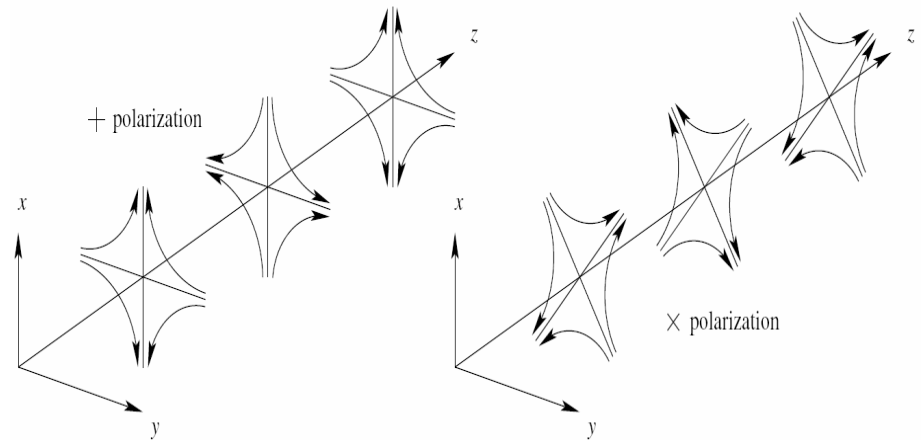
Coherent Sum

- Single detector response of form



$$h(t) = F_+(\theta, \varphi)h_+(t) + F_{\times}(\theta, \varphi)h_{\times}(t) + \eta$$

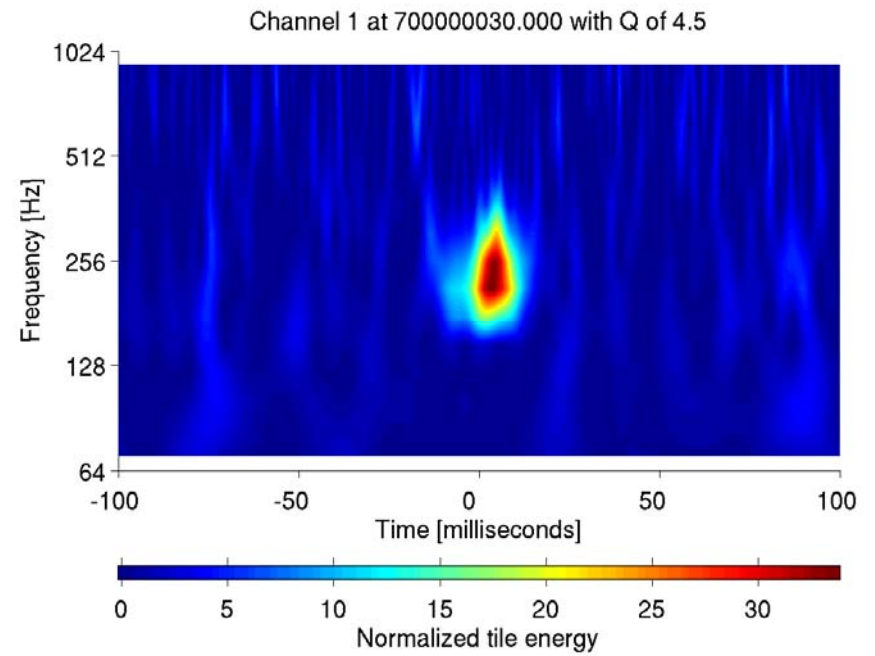
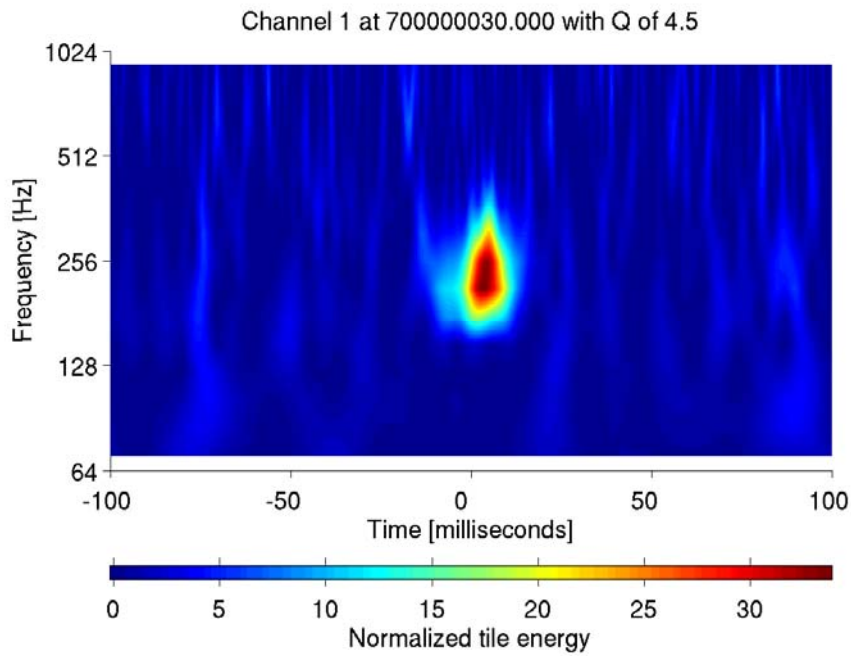
- Find set of scalar coefficients that would maximize SNR
- How much a priori information is required?



Coherent Sum

- 1) Time Shifts: GR propagates at speed c .
- 2) Amplitude Corrections: Accounts for different detector responses
- Both time shifts and detector response are a function of propagation direction
- Then for a given sky location carry out the calculations
- Sylvestre Algorithm

Time Shifts

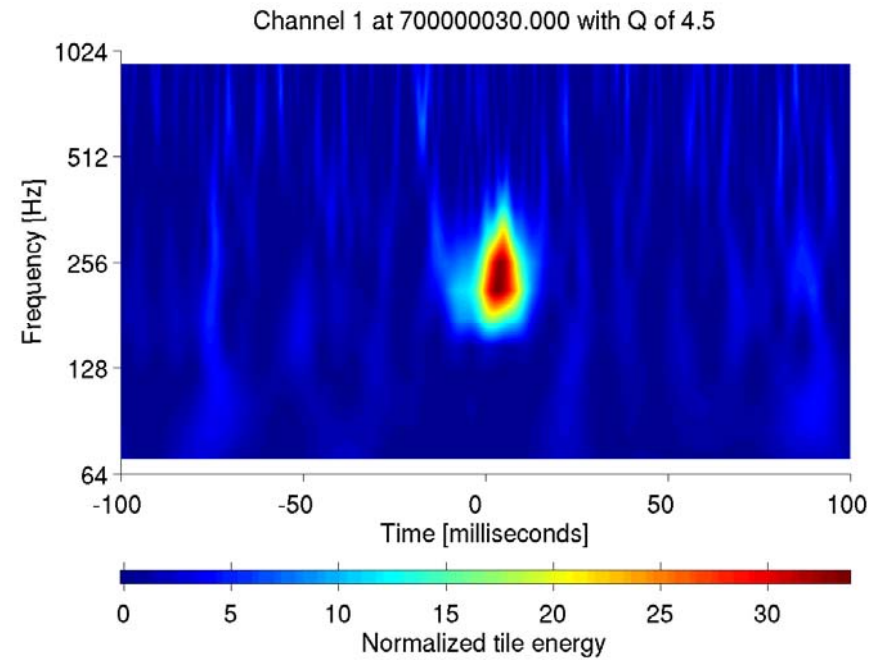
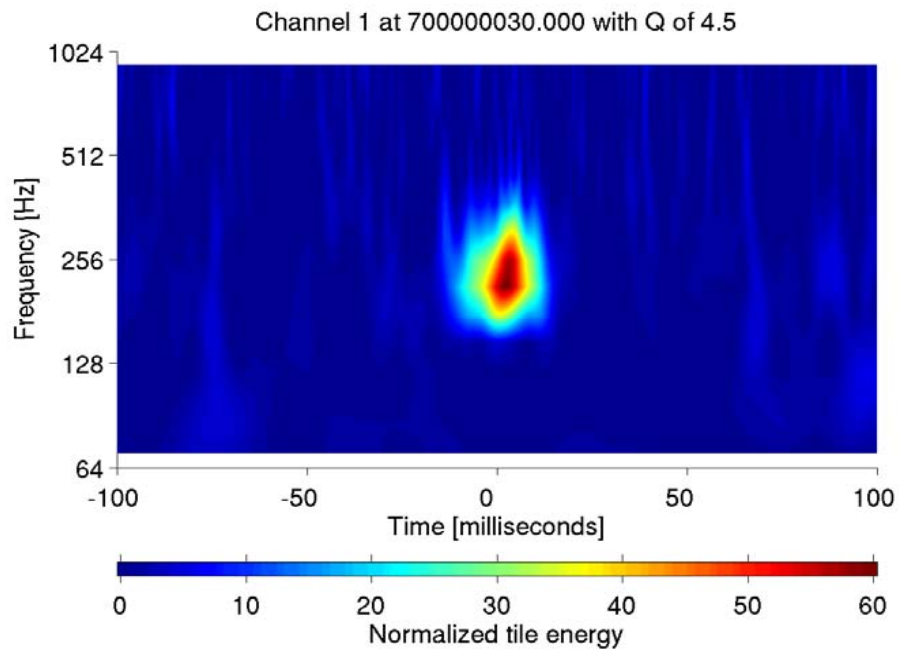


Amplitude Scaling

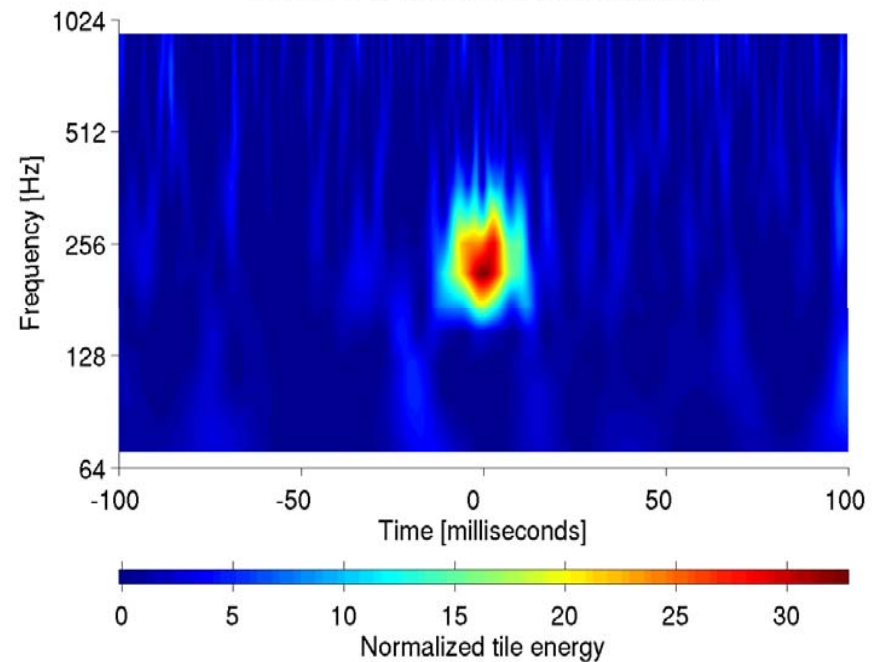
- Scale each input by an unknown coefficient
- Calculate the signal and noise power of combined response
- Using the Lagrangian method maximize for SNR
- Eigenvalue problem, easily solved
- Required information about wave:
 - 1) $h_+ \cdot h_x$ Inner product of polarizations
 - 2) $|h_+|/|h_x|$ Ratio of power

Example:

S

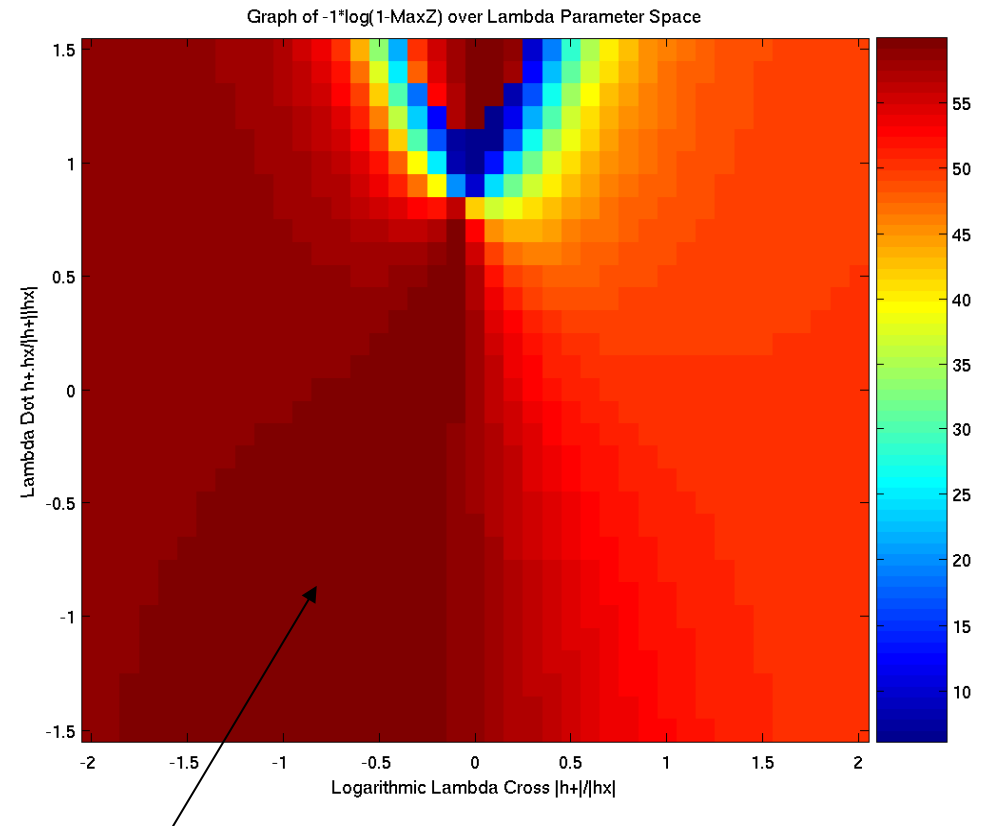


H



L

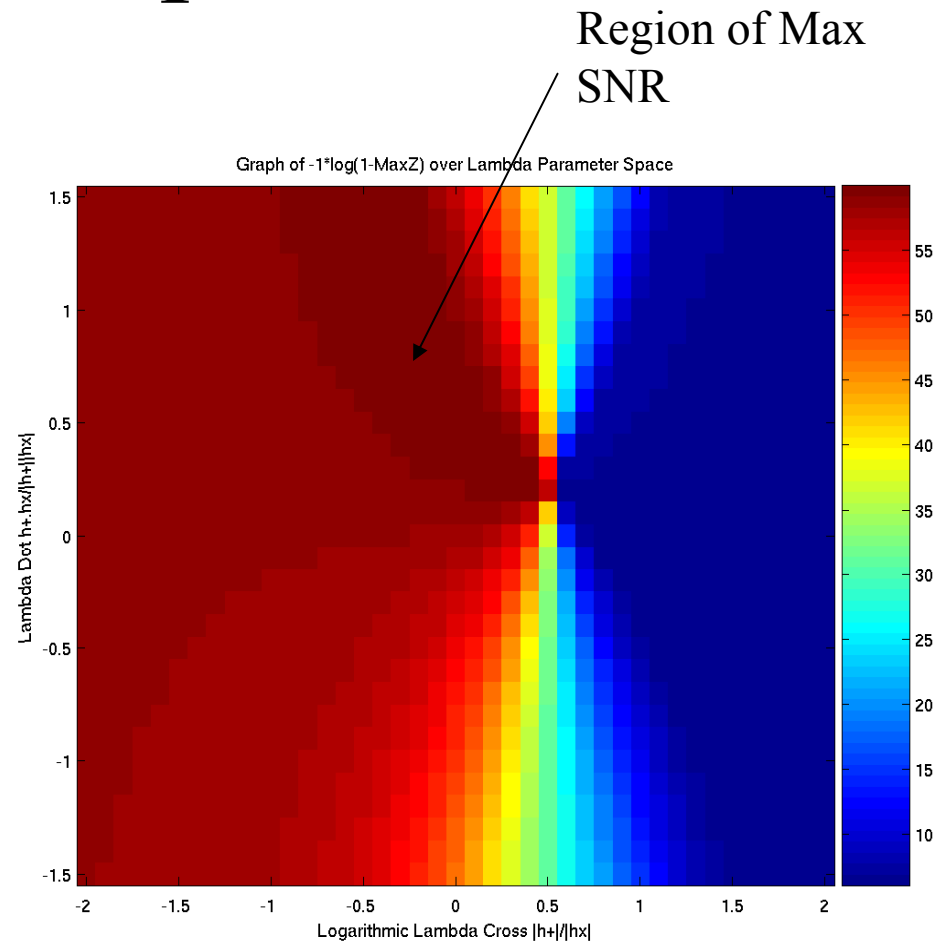
Parameter Space



Region of Max SNR

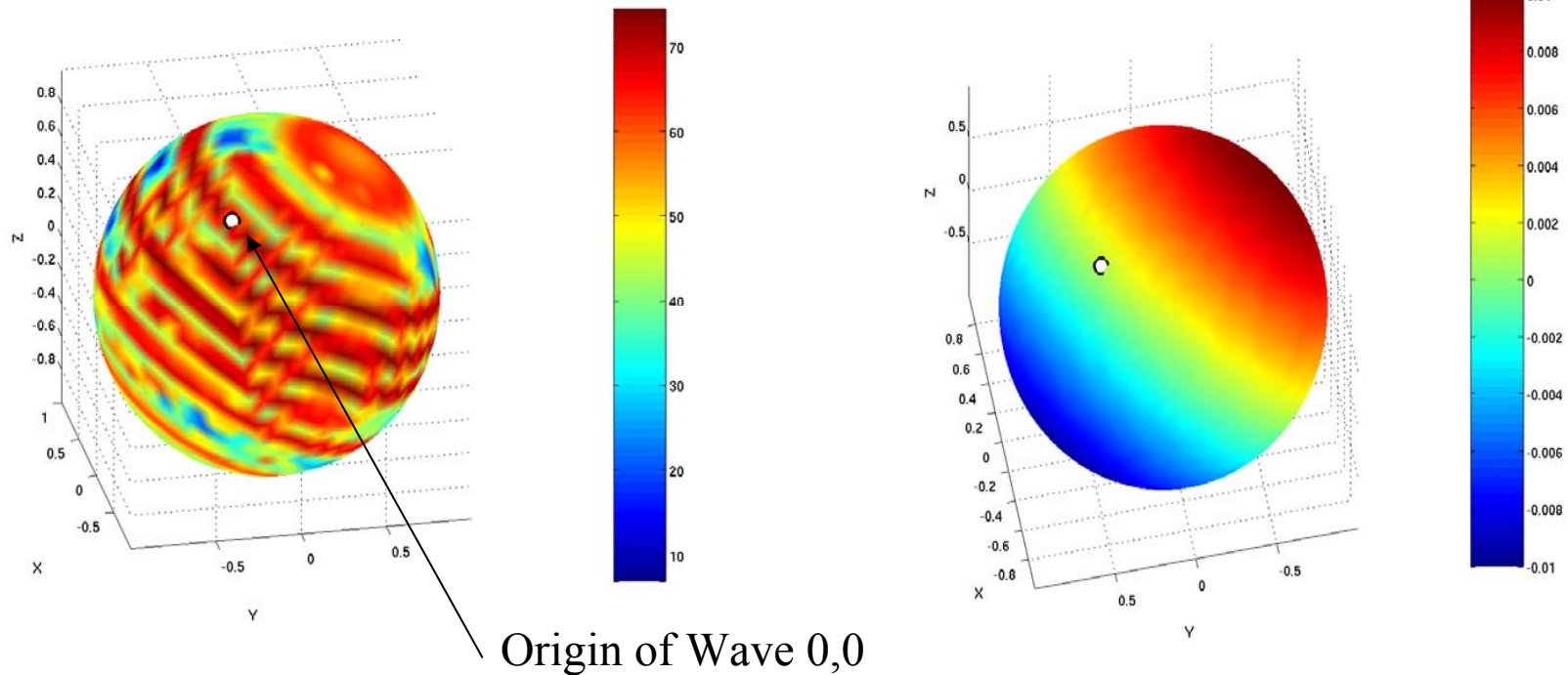
Parameter Space

- Polarization changed
- Redefinition of parameters
- However, search over the space yields same max SNR value



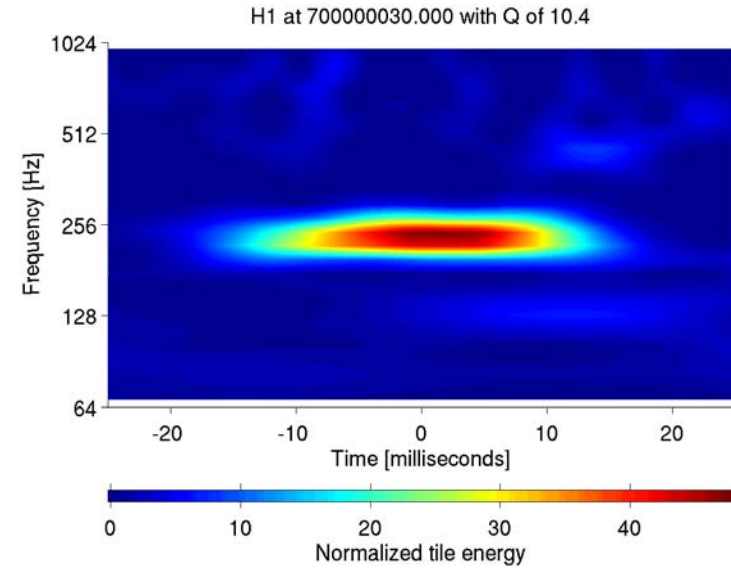
All Sky Searches

- Origin unknown
- Conduct searches over all sky

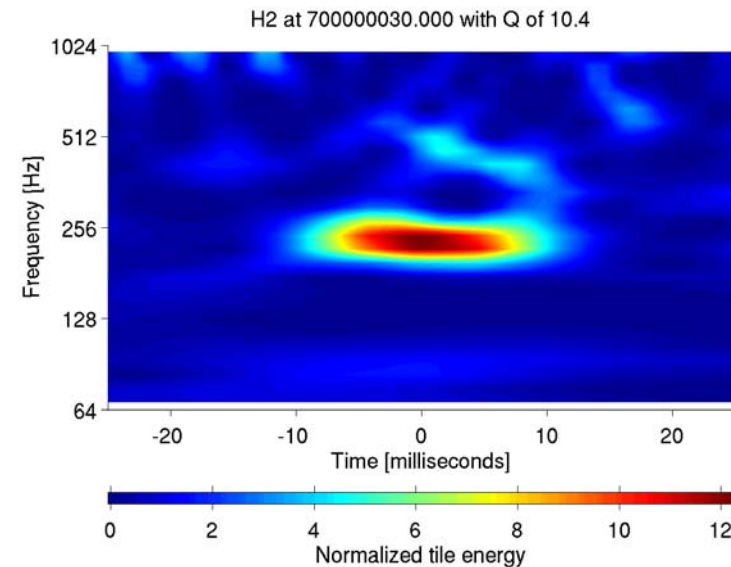


Consistency Checks

- Impossible to distinguish between GW and Glitches with two detectors (HL)
- H1 and H2 can be used for consistency checks
- Subtract out signal if true GW



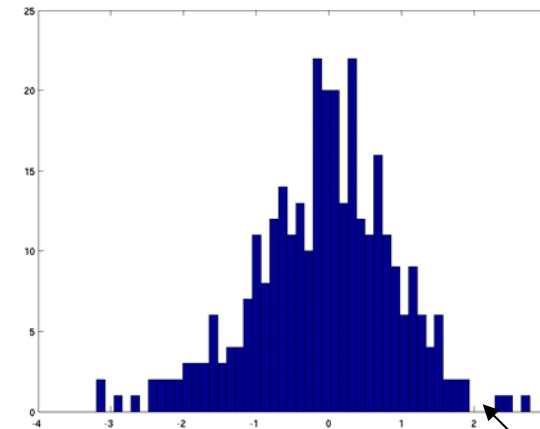
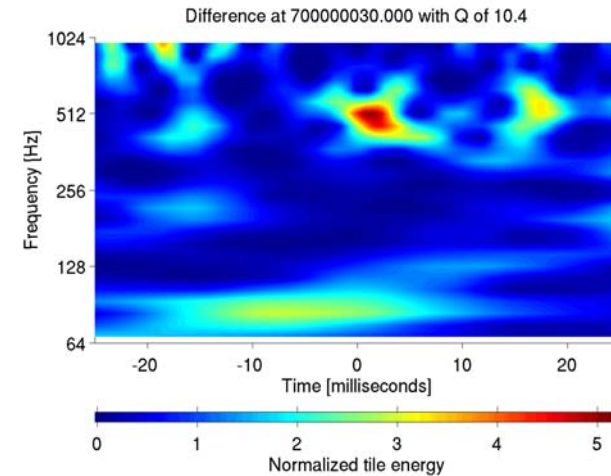
H1



H2

Q Statistics, GW

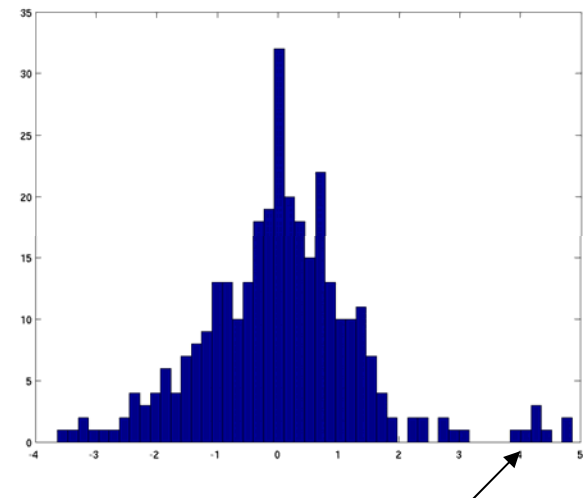
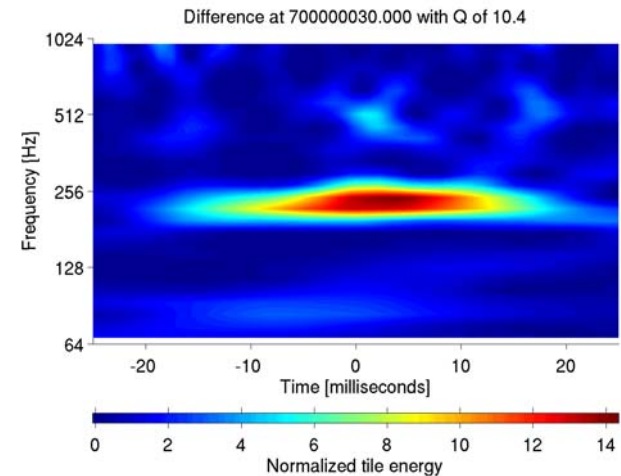
- H1 H2 Consistency Check
- After subtraction Gaussian distribution expected in each component of coefficient
- Lilliefors test used
- Hypothesis test on whether GW or Glitch



Max 3

Q Statistics, Glitch

- Simulated Glitch
- Subtraction not possible (complex coefficients)
- Reflected in the statistics of the difference transform



Conclusions

- Coherent Searches
- Coherent Sum: Improvement in SNR
- Consistency Checks using H1 and H2
- Future Work: Special class of waves Q Difference to distinguish between glitches and GW in HL network
- Run on real data