



Calibration/validation of the AS_Q_FAST channels

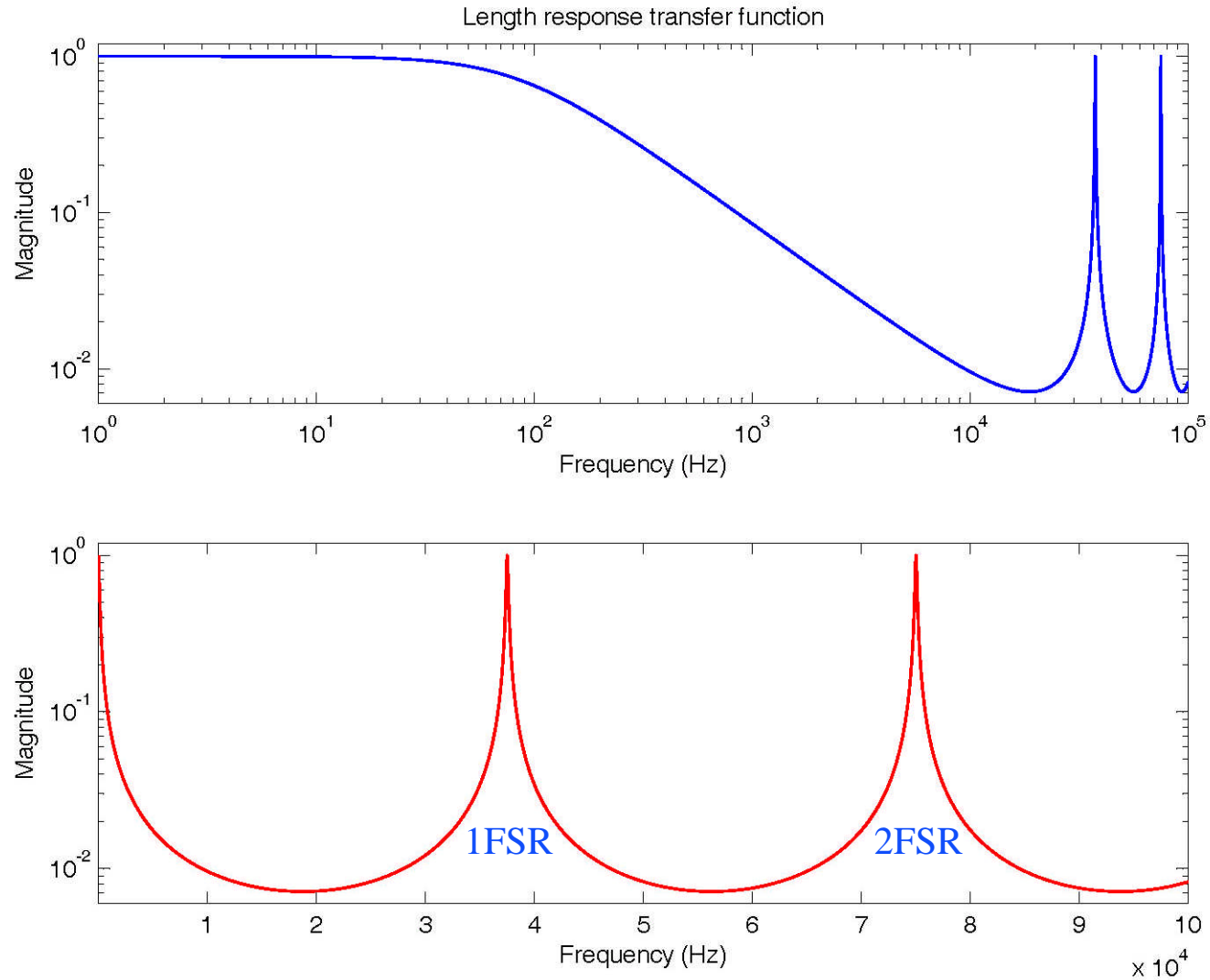
Rick Savage - LHO

Stefanos Giampanis – Univ. Rochester

(Daniel Sigg – LHO)

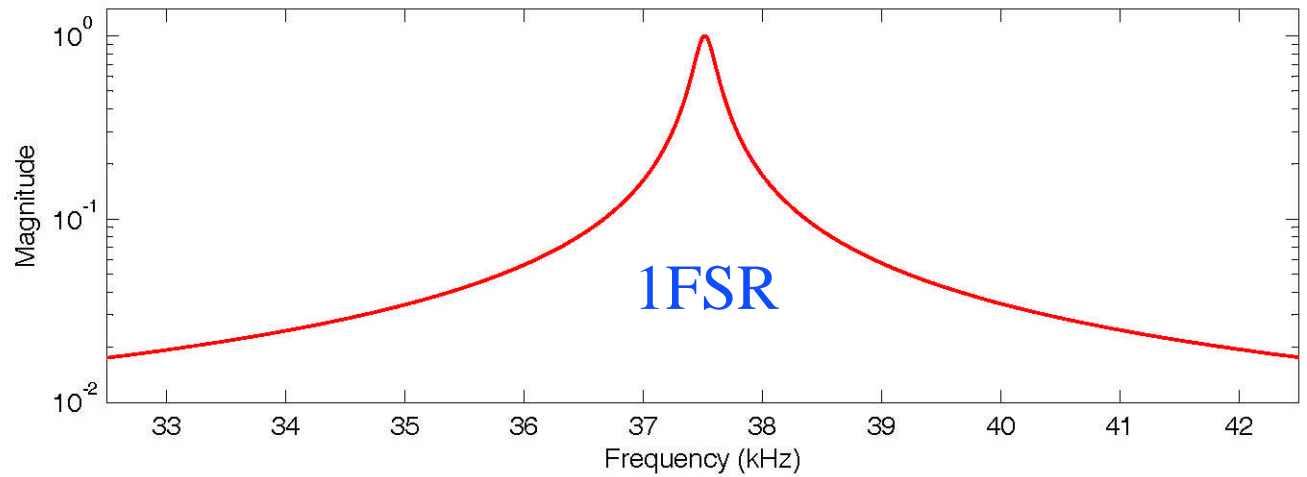
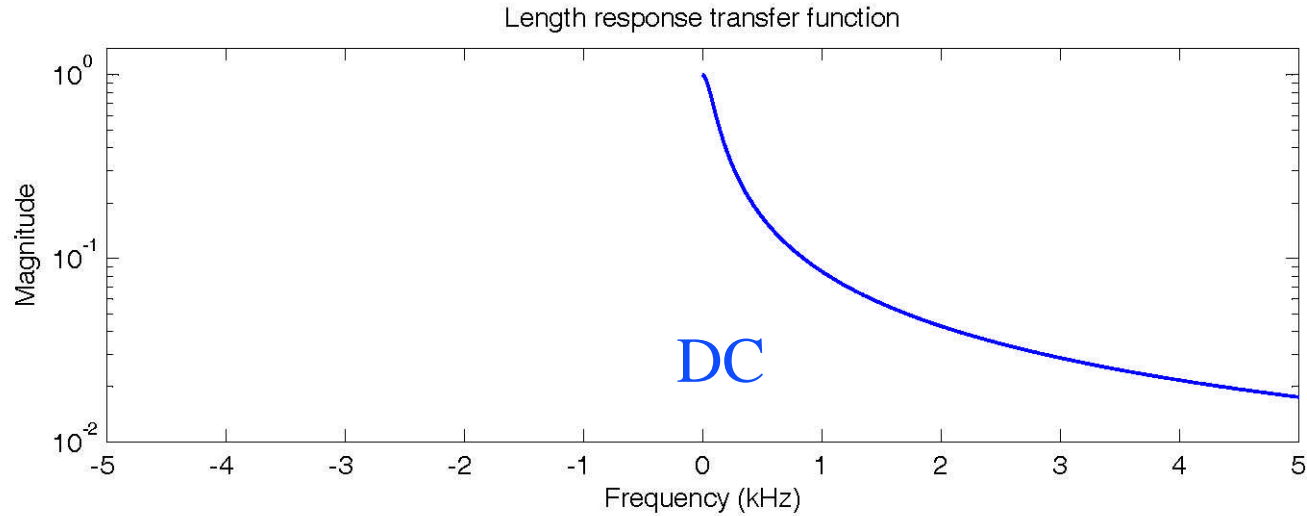


High frequency length response



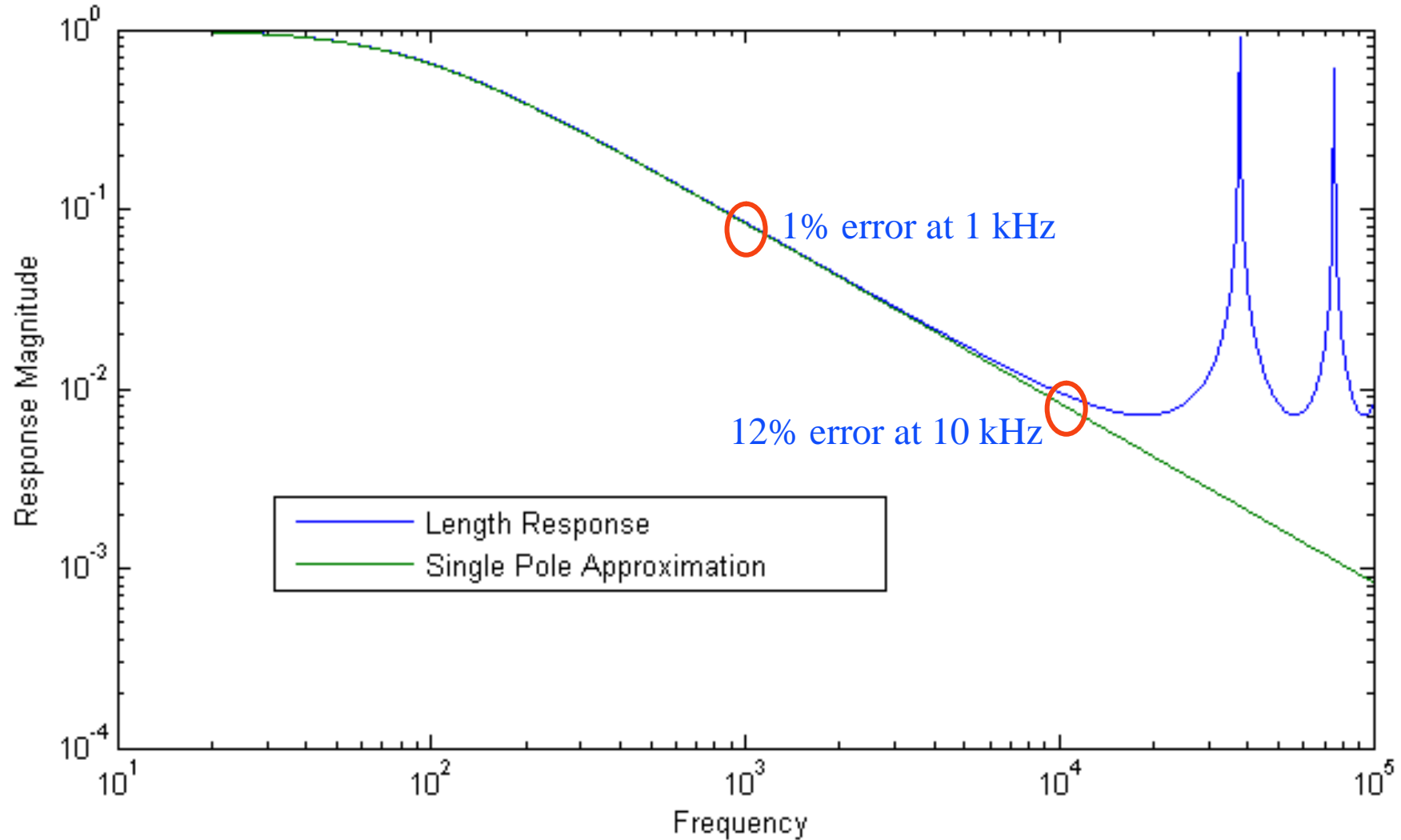


Length response TF detail





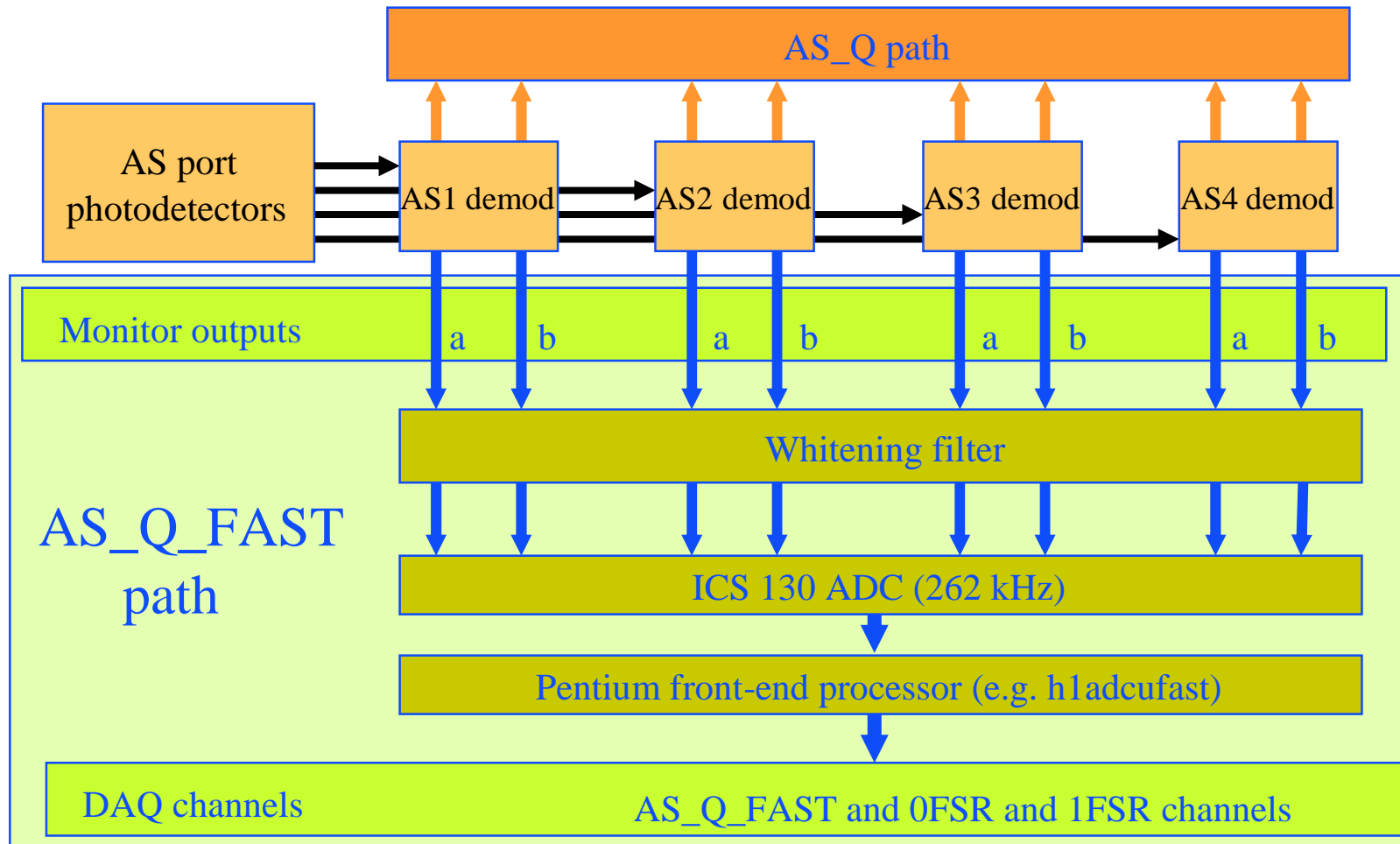
Comparison with single pole approximation



“FAST” channels

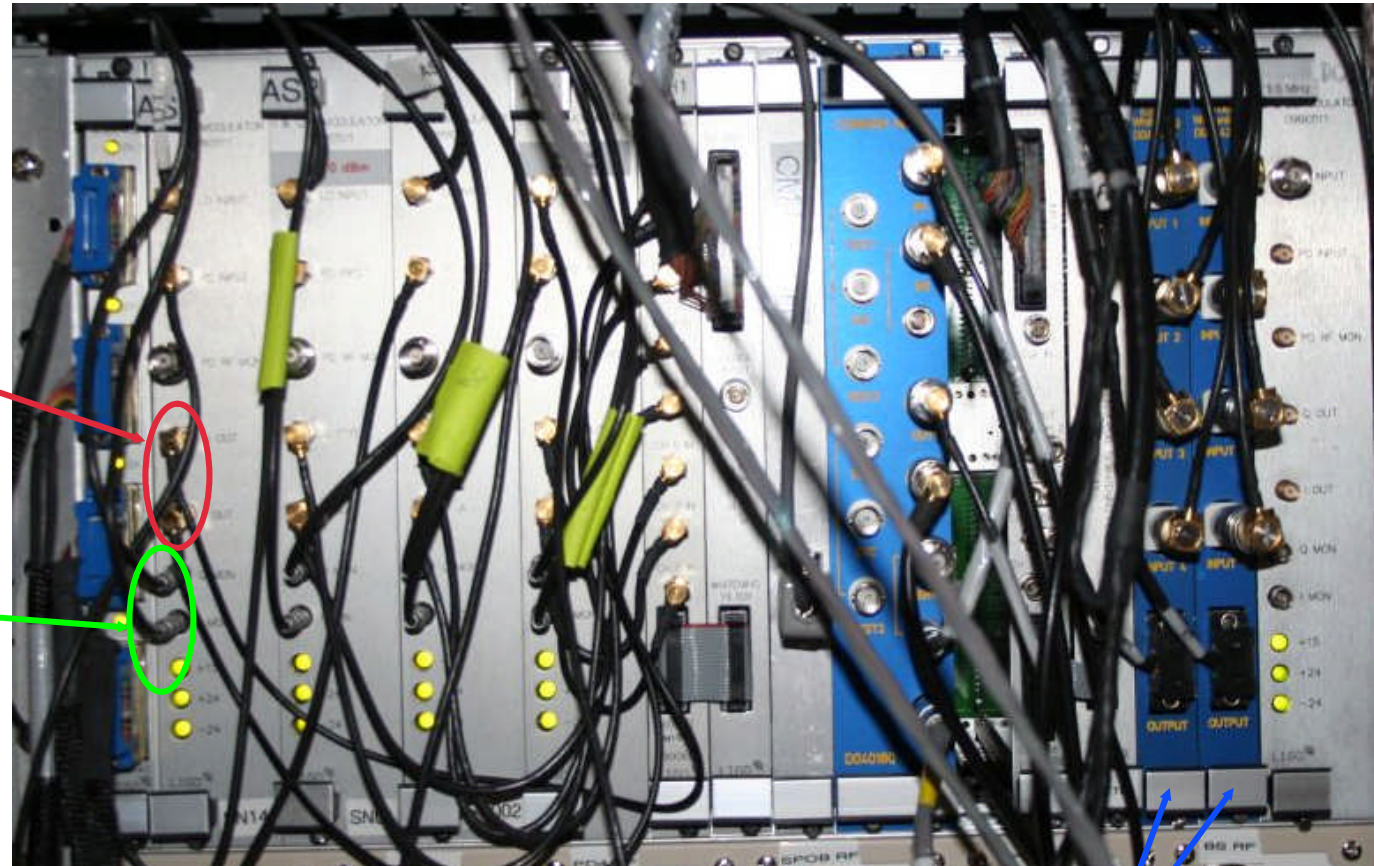
- All three interferometers have FAST (262 kHz) data acquisition systems generating AS_Q_FAST signals with 100 kHz bandwidths
- The signals are also digitally heterodyned at 1024 Hz and 37.504 kHz and downsampled to 2048 Hz to generate the 0FSR and 1FSR channels
 - » AS_Q_0FSR.mag AS_Q_1FSR.mag
 - » AS_Q_0FSR.phs AS_Q_1FSR.phs
 - » AS_Q_0FSR.real AS_Q_1FSR.real
 - » AS_Q_0FSR.img AS_Q_1FSR.img
- At LHO the full 262 kHz AS_Q_FAST channels are written to the frames. Not yet implemented at LLO.

AS signal flow

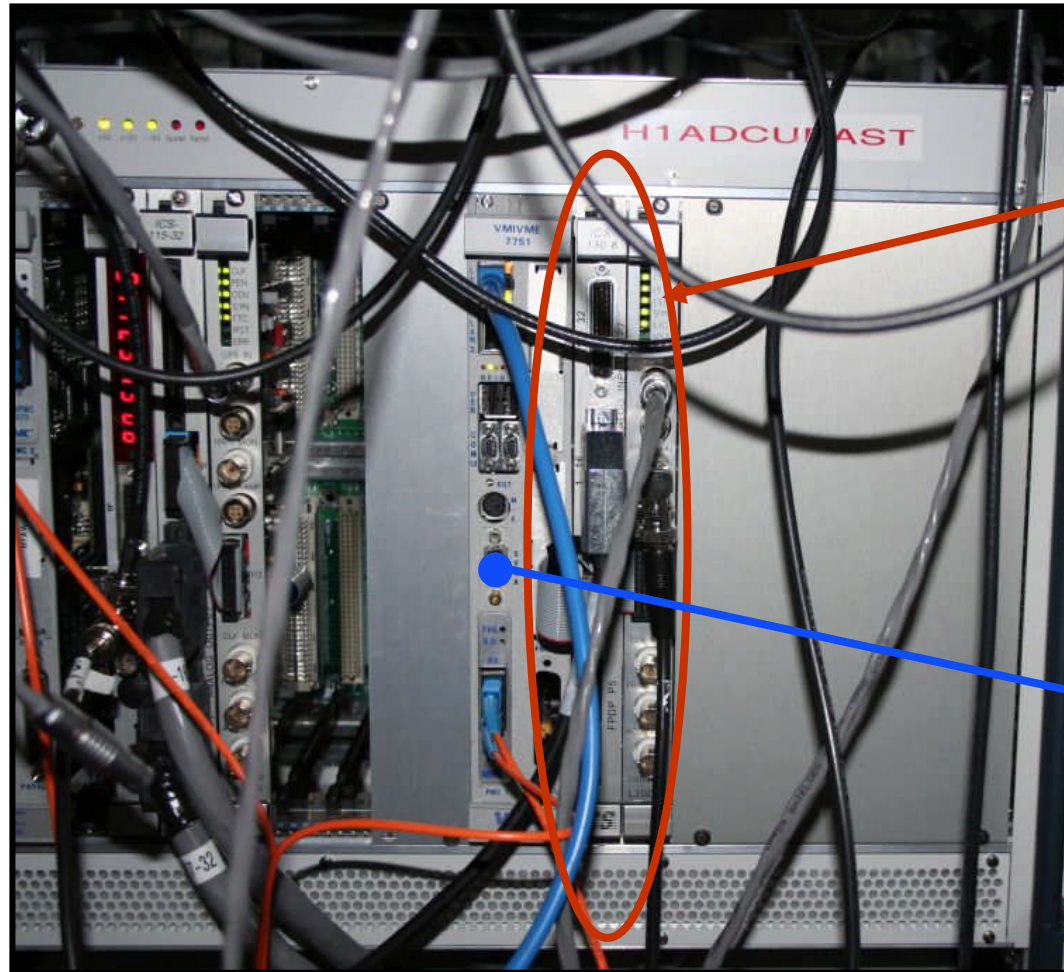


ASn
outputs

ASn-mon
outputs
(x 2 gain)



whitening filters for ICS 130

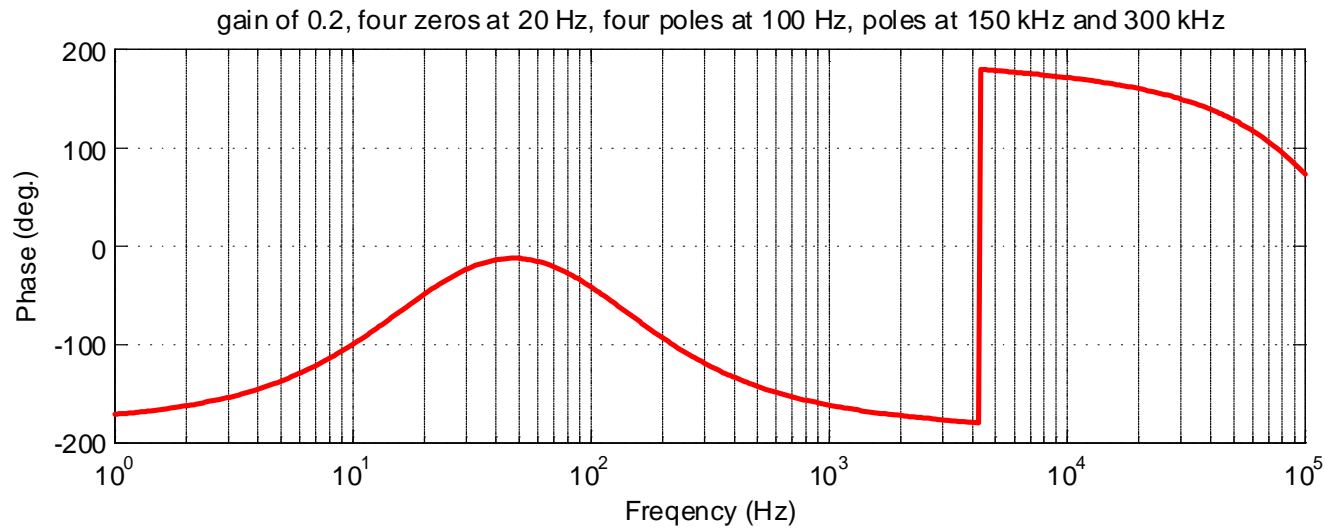
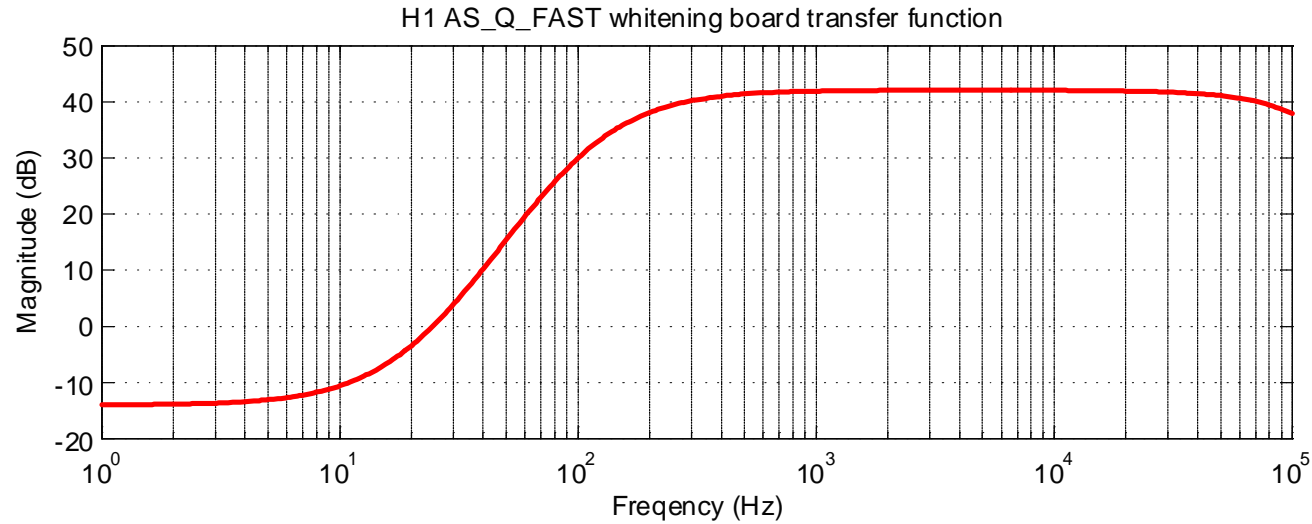


ICS 130 – 8 channels
running at 262,144 Hz

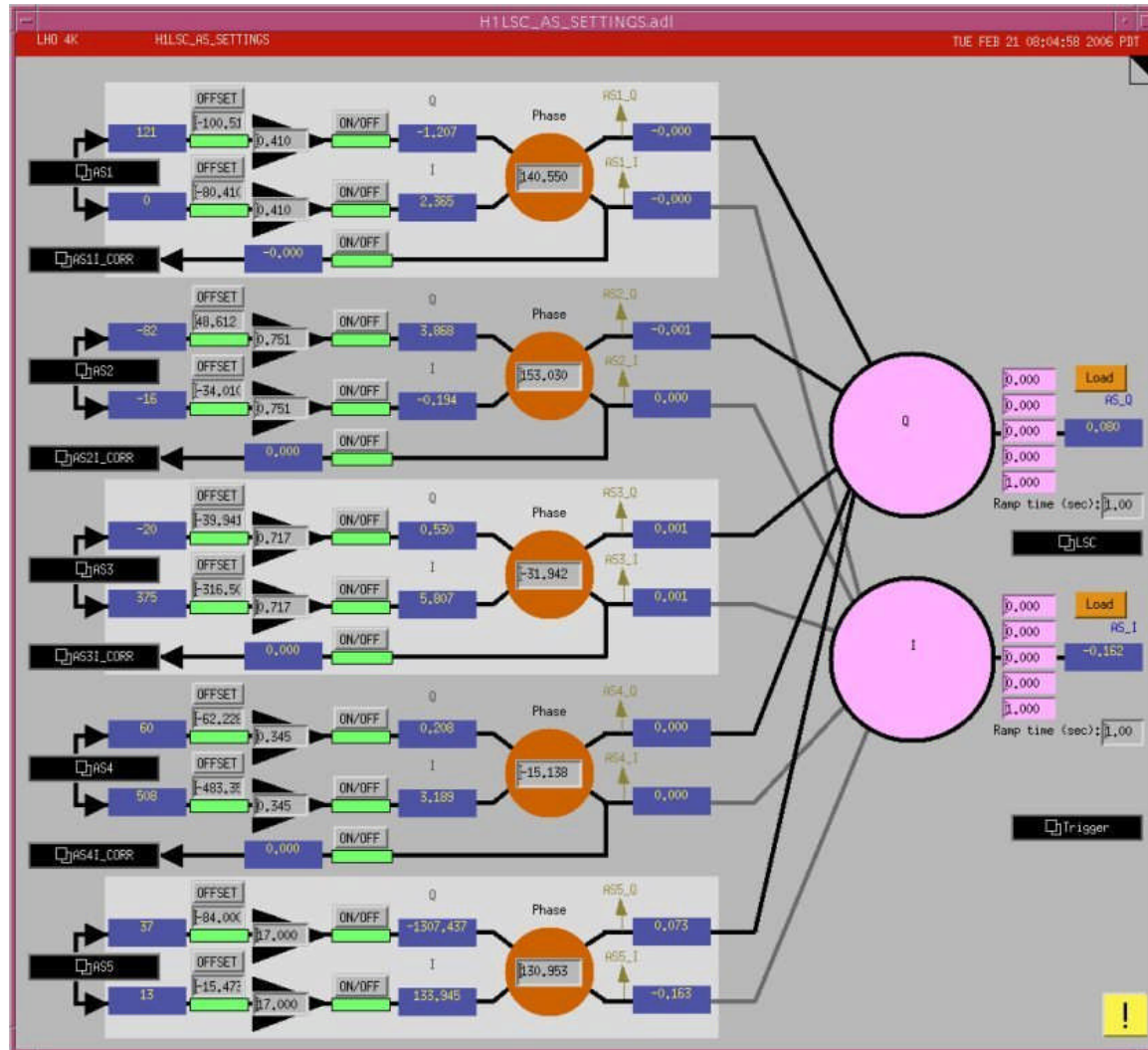
Pentek code generates
the AS_Q_FAST and
hetrodyned channels



Whitening filter transfer function



ASPD summation for AS_Q and AS_Q_FAST

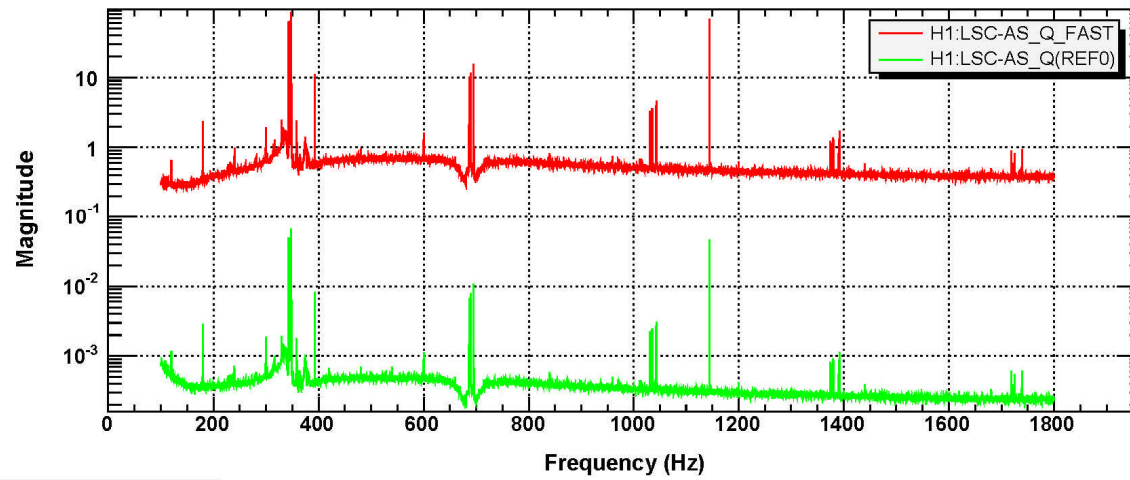


- » Similar calculation done in adcufast
- » Presently, startup.cmd must be updated manually every time medm settings are changed
- » LLO adjusts ASPD matrix coefficients for each lock stretch
- » Working toward reading epics settings on the fly.

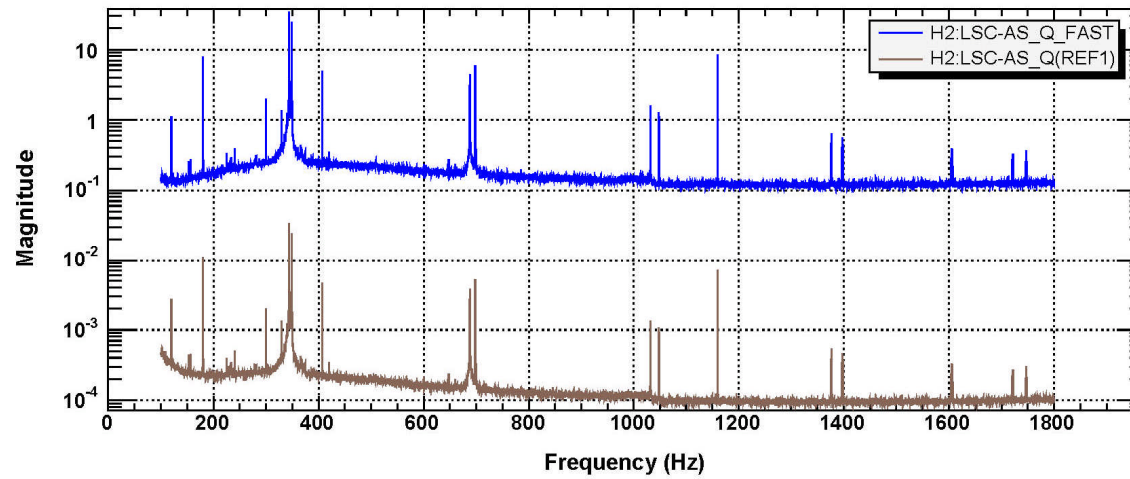


AS_Q : AS_Q_FAST comparison

Power spectrum



Power spectrum



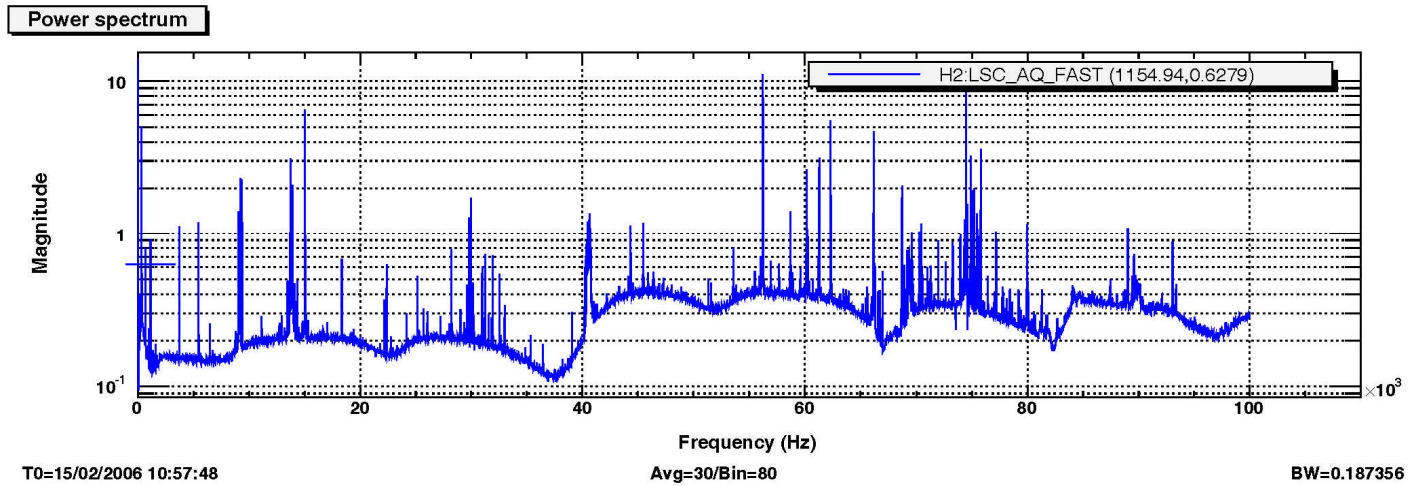
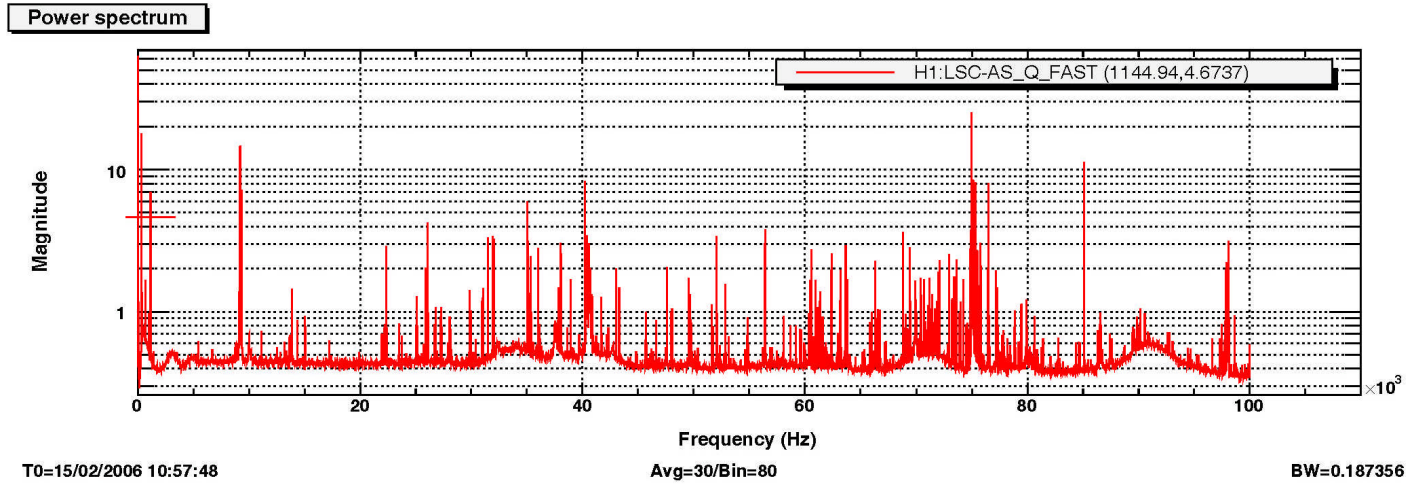
T0=19/03/2006 20:25:40

*Avg=10/Bin=21

BW=0.0117166



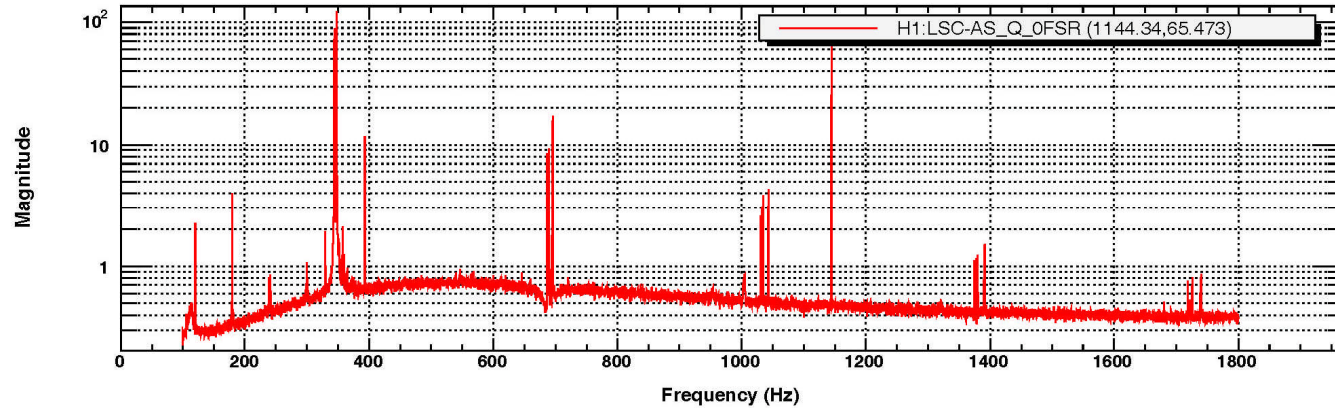
AS_Q_FAST signals during full ifo. lock





0FSR signals

Power spectrum

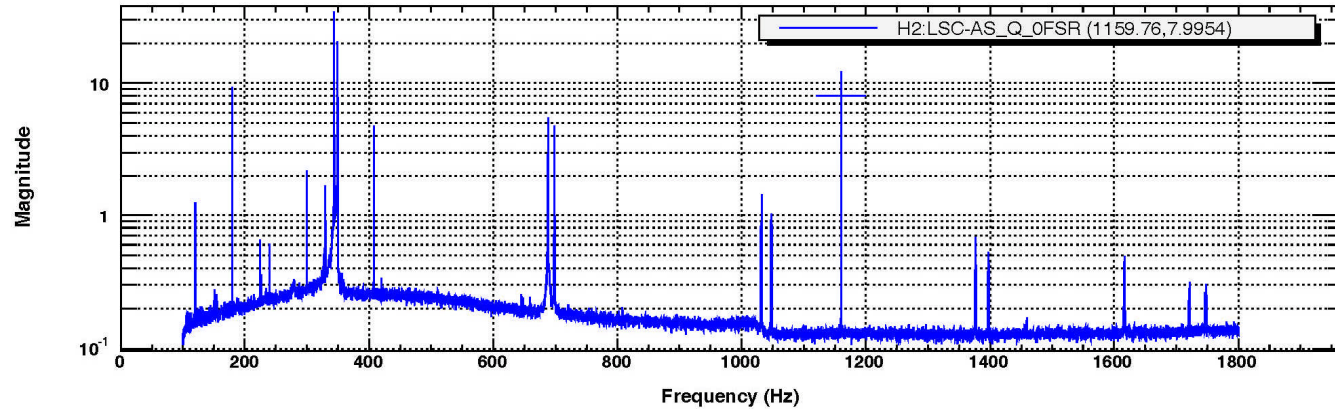


T0=15/02/2006 10:57:48

Avg=10/Bin=21

BW=0.0117178

Power spectrum



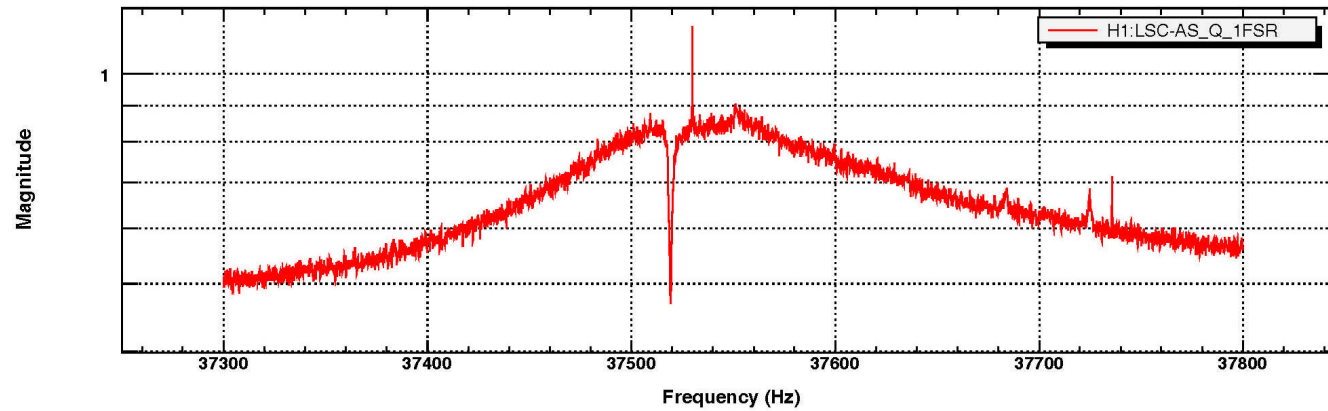
T0=15/02/2006 10:57:48

Avg=10/Bin=21

BW=0.0117178

1FSR signals

Power spectrum

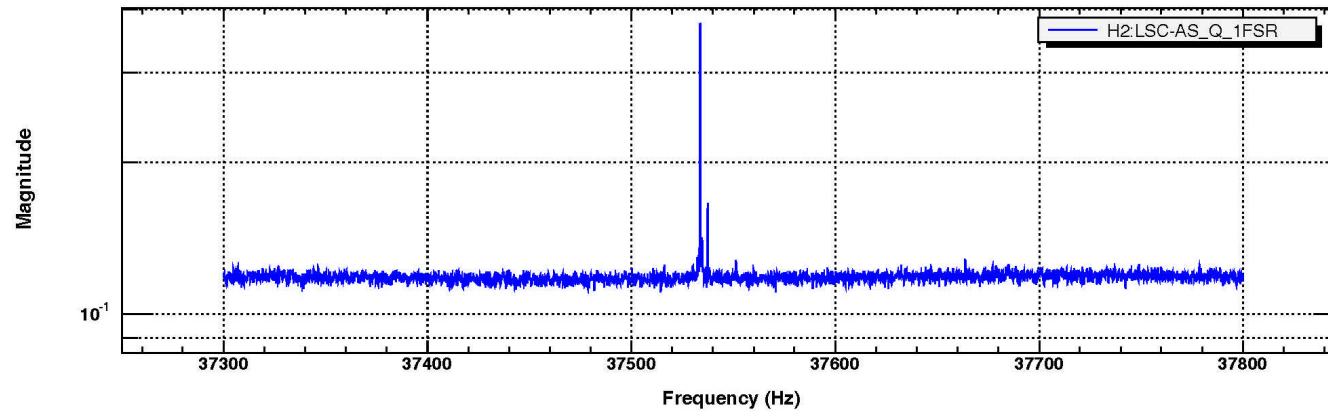


T0=15/02/2006 10:57:48

Avg=1000

BW=0.1875

Power spectrum



T0=15/02/2006 10:57:48

Avg=1000

BW=0.1875

- Drive test mass at 37.5 kHz
 - » Attempted during Oct. 2005
 - Factor of five larger value than expected
 - Electronics not designed to be driven at this frequency
 - Uncertainty in force/displacement too high
- Photon calibrator
 - » Now driving at ~ 1600 Hz
 - To resolve peak in ASD factor at the level of a factor of a few above the background requires hundreds of seconds of integration
 - At 37.5 kHz one data point would require ~ one year integration time
- Thermally-excited TM internal mode vibrations
 - » Modeling (and more thought) required



LIGO

“Blind” extrapolation of low-freq. calibration

- Use cal line at ~ 1150 Hz
- Replace cavity pole approximation with length response given by $H_L(f)$
- Extrapolate to 100 kHz without HF measurement
 - » Analytic model used to date for single FP cavity - probably gets differential mode right , but not leakage of common mode effects into AS port – doesn’t model noise floor.
 - » Use full interferometer model such as Finesse to study noise couplings at high frequencies.
- Effort underway to implement “blind” calibration