

Current status of the Mesa beam experiment: Tilt sensitivity

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LIGO seminar

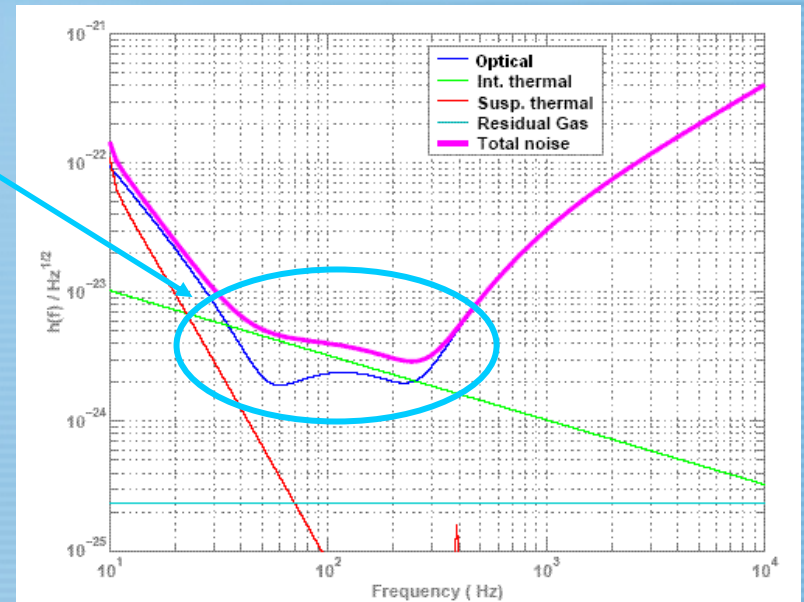
September 12 2006

- **Introduction**
- **Current Status of Mesa Beam Experiment**
- **Tilt Sensitivity**
- **Results of work done**
- **Future work**

Introduction

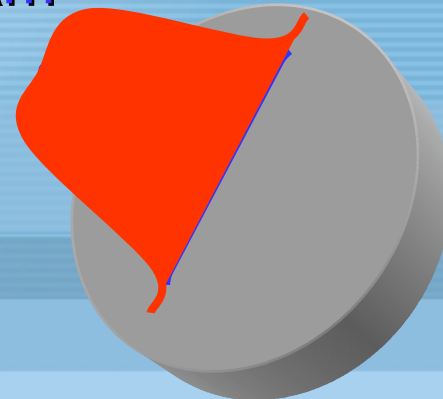
Thermal noise limits the sensitivity of gravitational wave interferometric detectors

Gaussian Beam → Mesa Beam

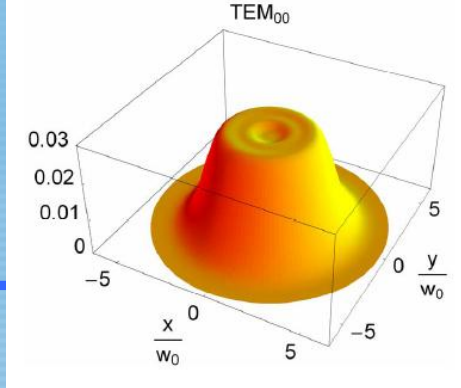


Gaussian Beam

Mesa Beam



Introduction

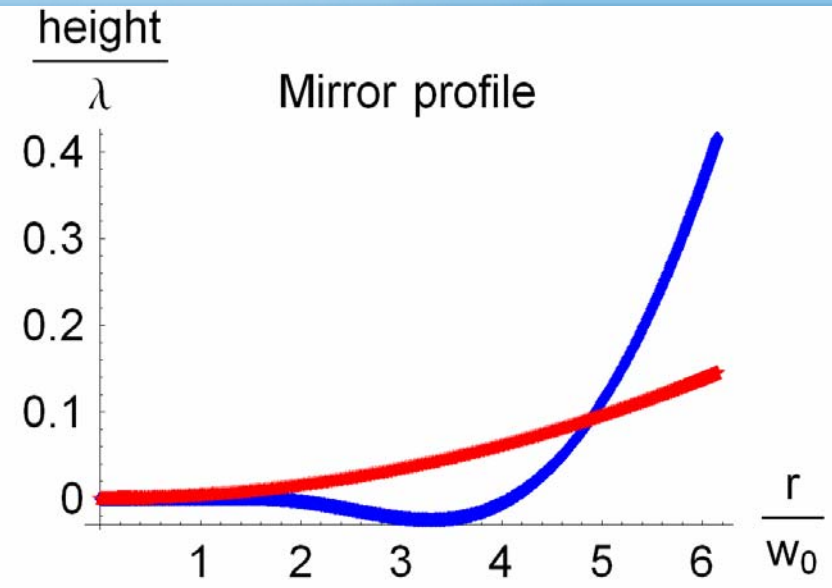
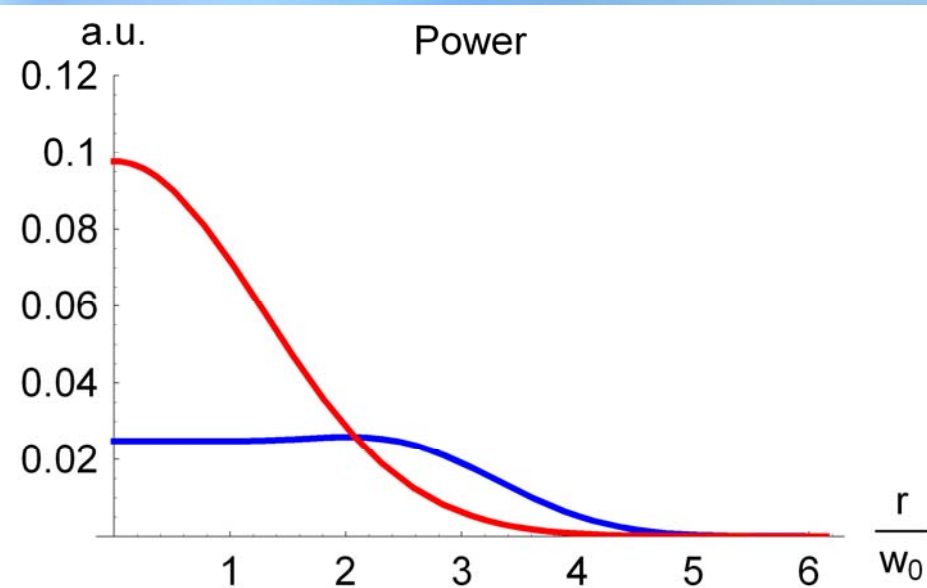


The Mesa Beam is a superposition of minimal Gaussians with

$$w_0 = \sqrt{L/k}$$

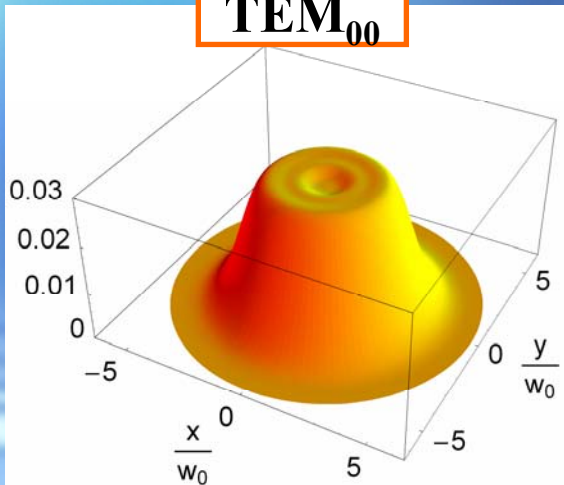
$$U(D, r) = \int_{C_D} \exp \left[\frac{- \left[(x - x_0)^2 + (y - y_0)^2 \right] [1 + i]}{2w_0} \right] dx_0 dy_0$$

Supported by *nearly-flat* mirrors with **“Mexican-hat”** profile

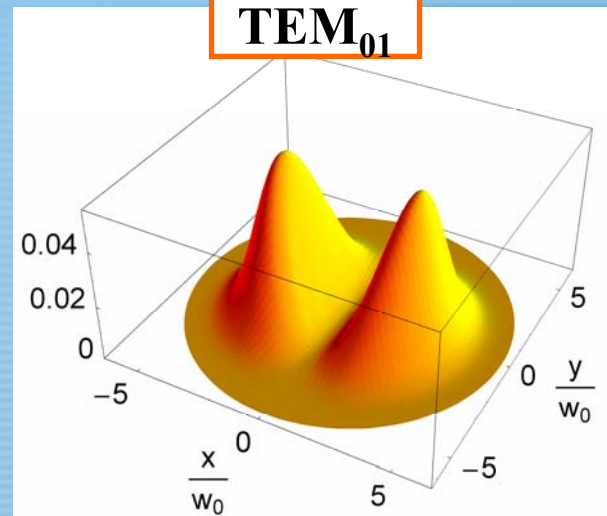


Introduction

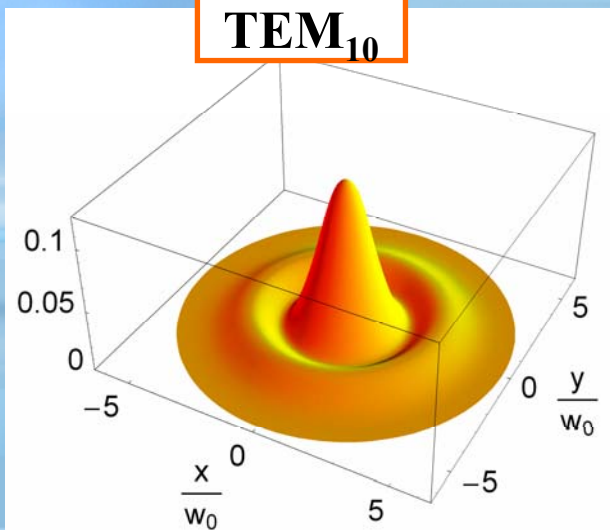
TEM₀₀



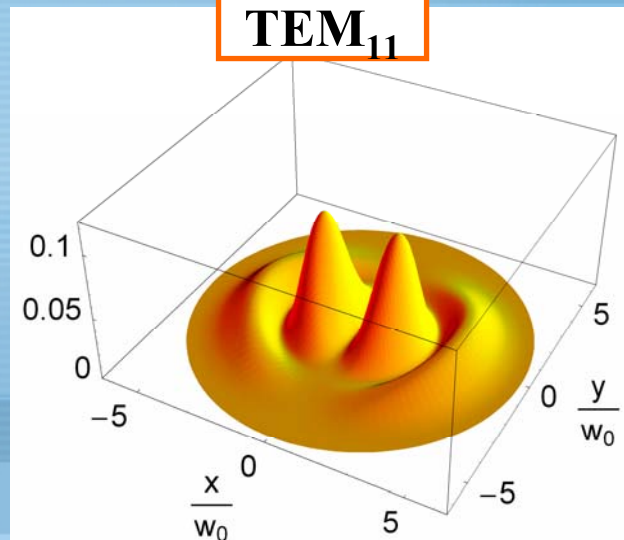
TEM₀₁



TEM₁₀

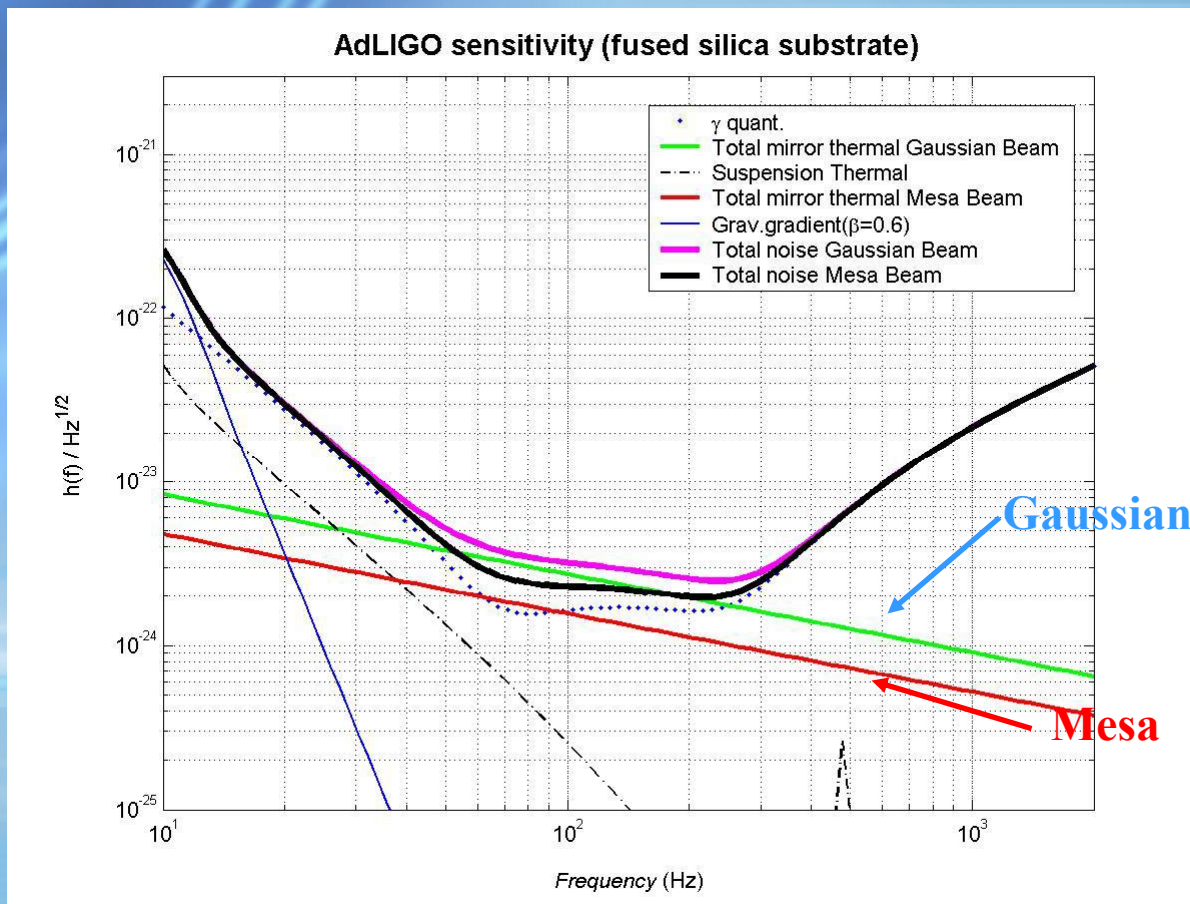


TEM₁₁



Introduction

Excepted Thermal noise reduction

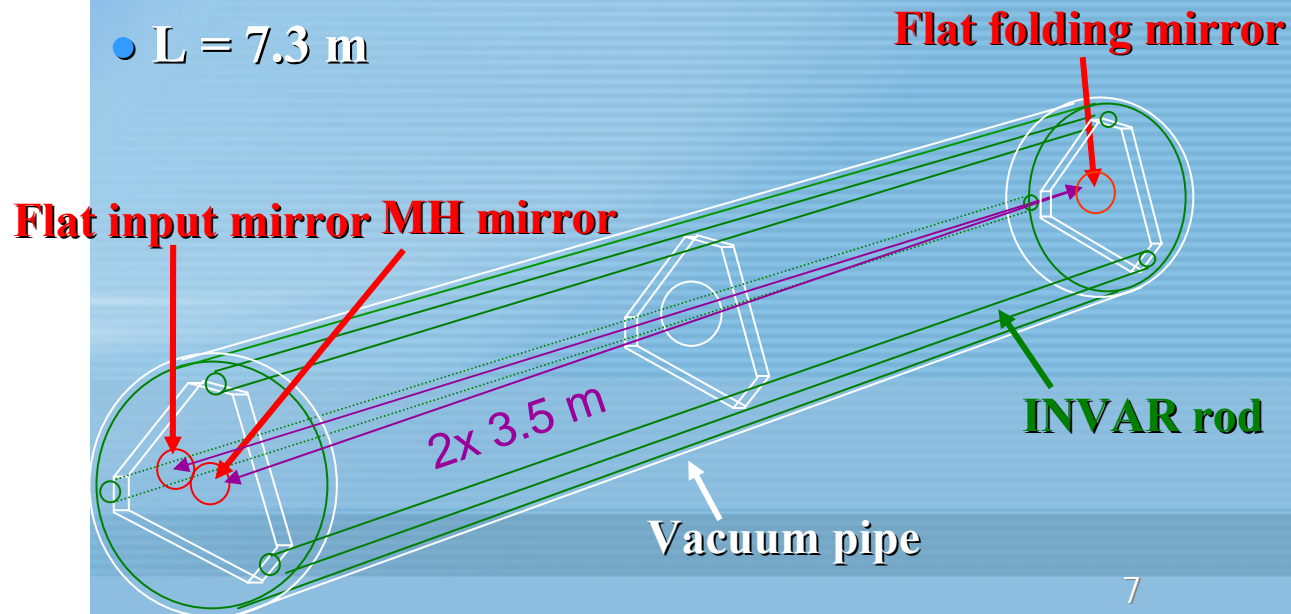
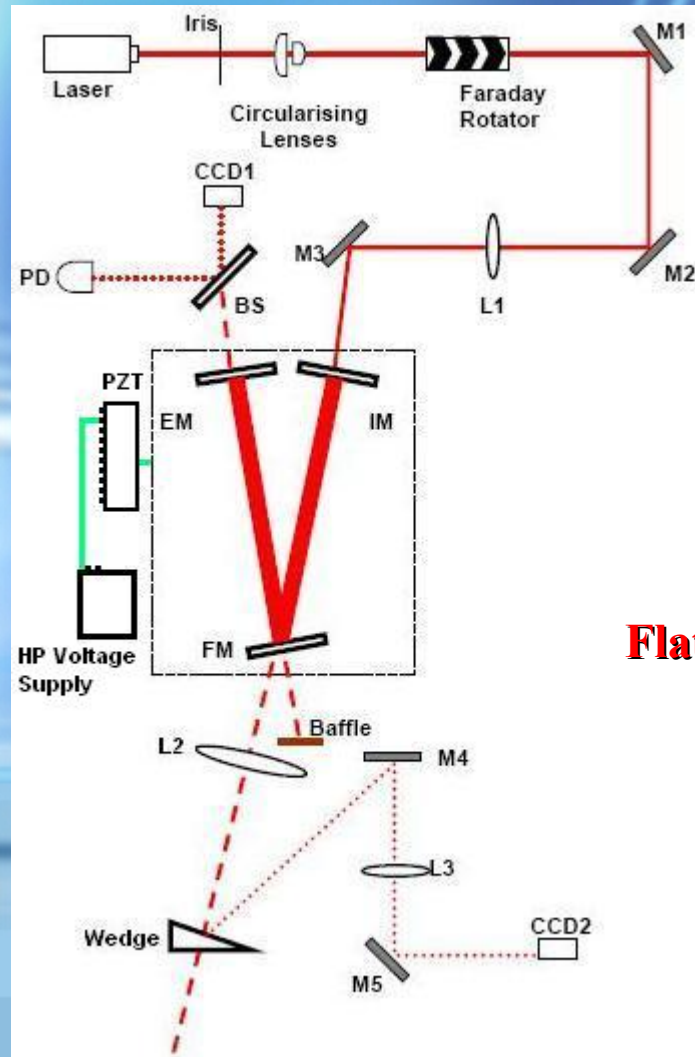


Fabry-Perot Cavity

Reshaping the FP arm cavity mirrors it is possible to obtain a flat top beam

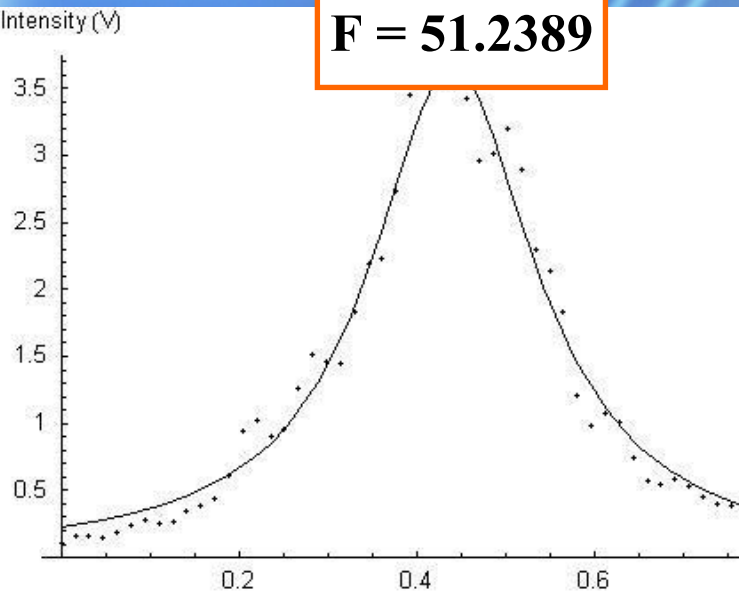
Caltech prototype

- Rigid, folded and suspended
- $L = 7.3 \text{ m}$

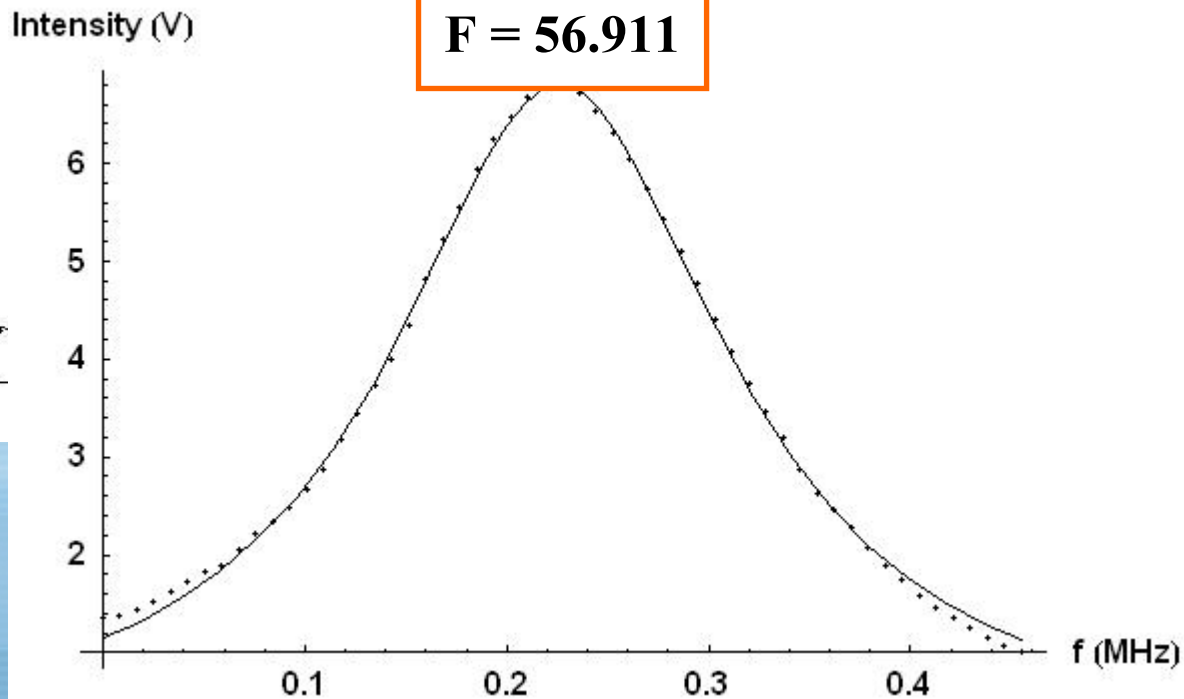


Cavity Finesse

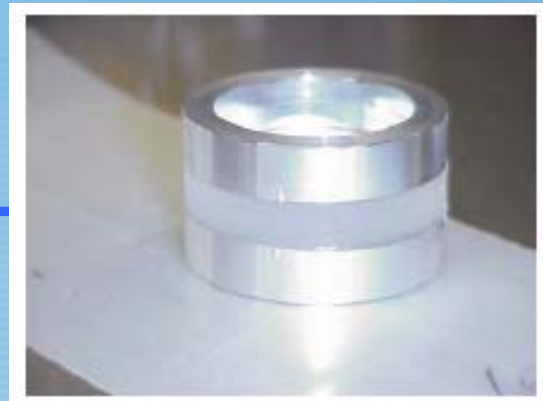
F = 51.2389



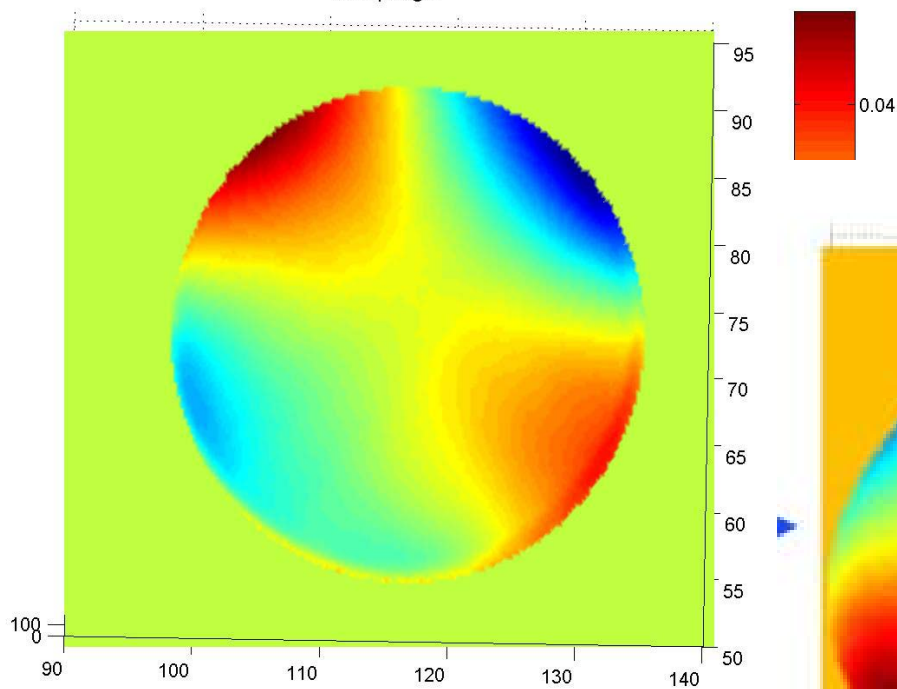
F = 56.911



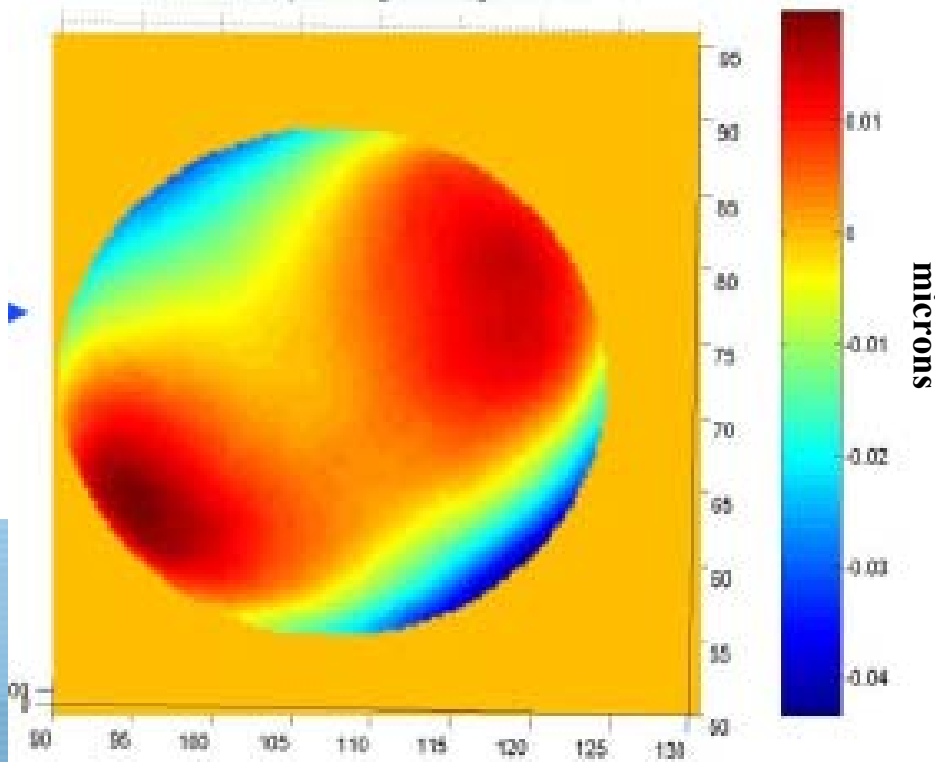
Mirror maps



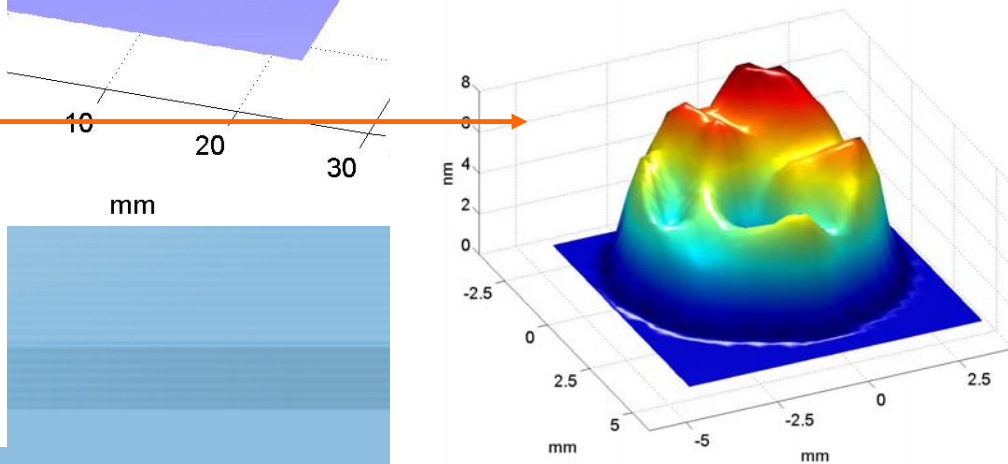
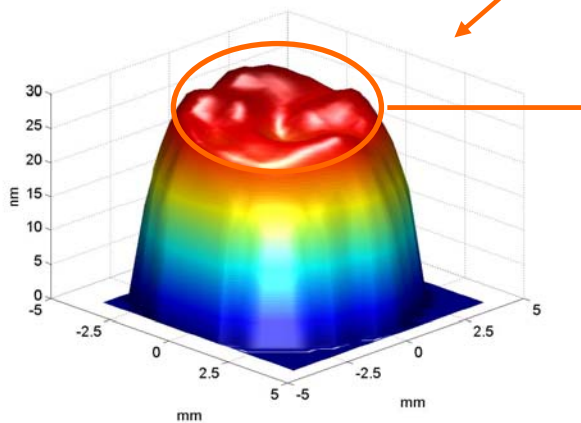
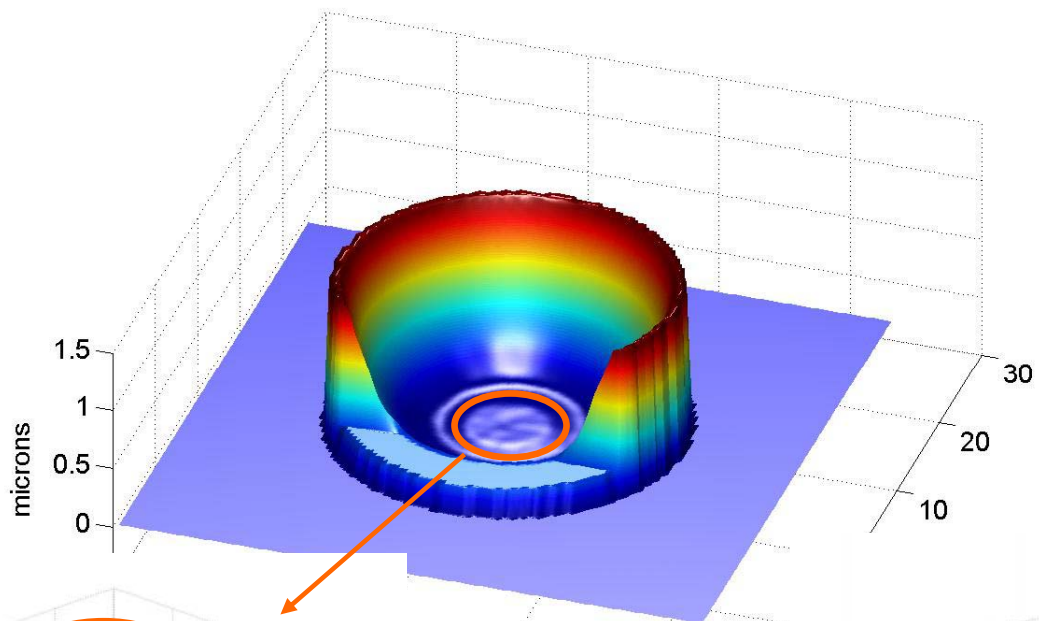
Clamps tight



MHM3 Central portion - Tight screws, glued mount

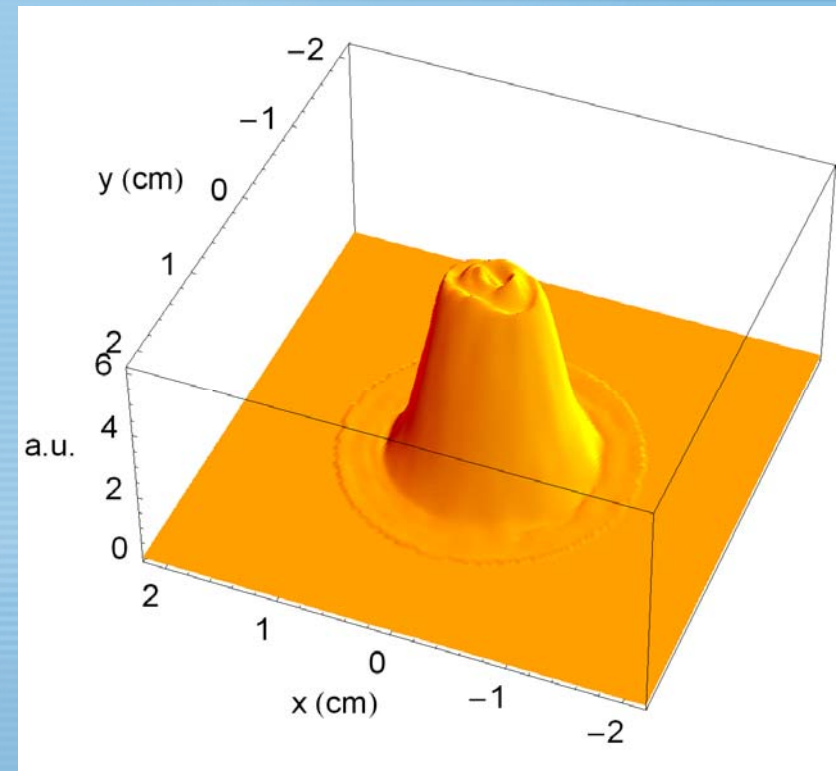
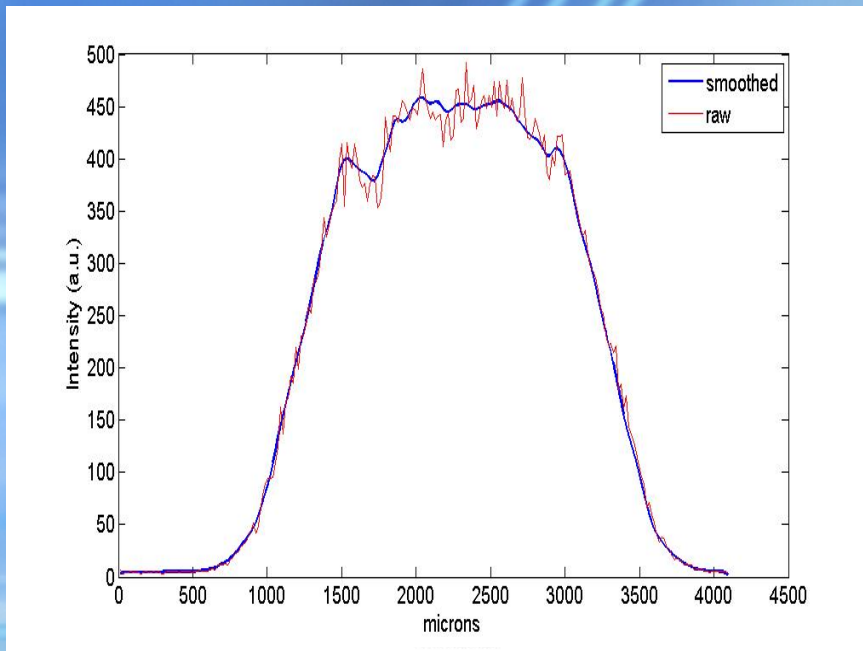


Mirror maps



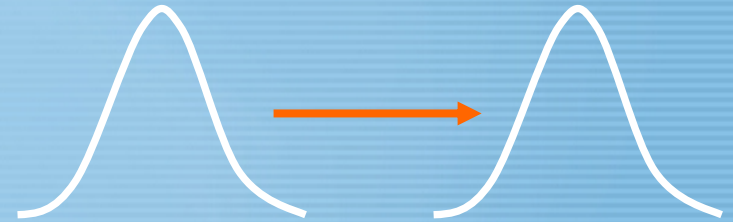
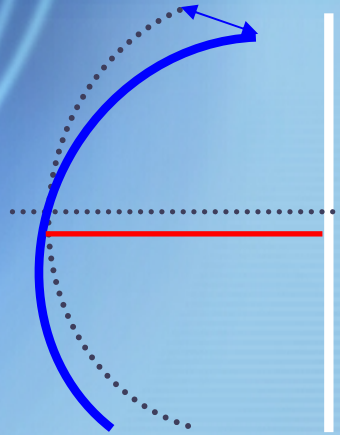
Current status

Production of acceptable flat beam w/ imperfect optics



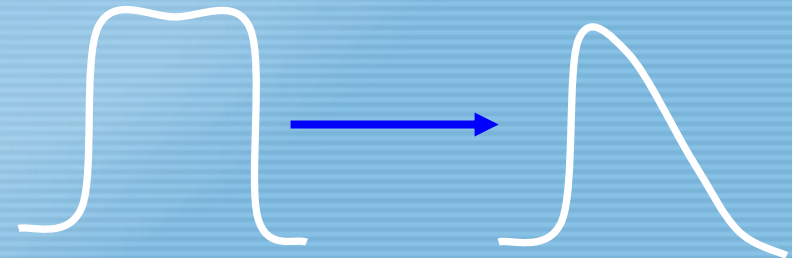
Tilt

Tilts of **spherical** mirrors
translate optical axis



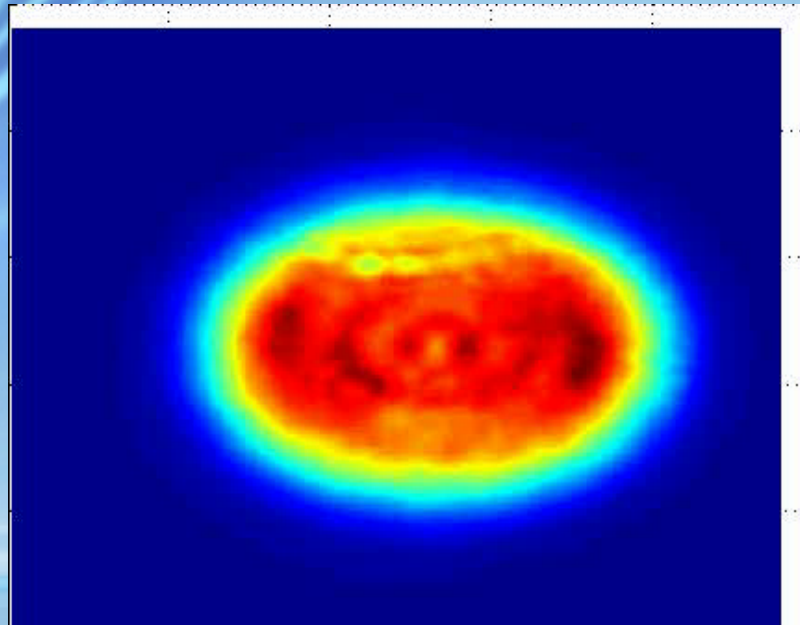
Tilts of **MH** mirrors:

- change in the **optical axis**
- resonant beam **phase front change** with the alignment



Tilt Sensitivity

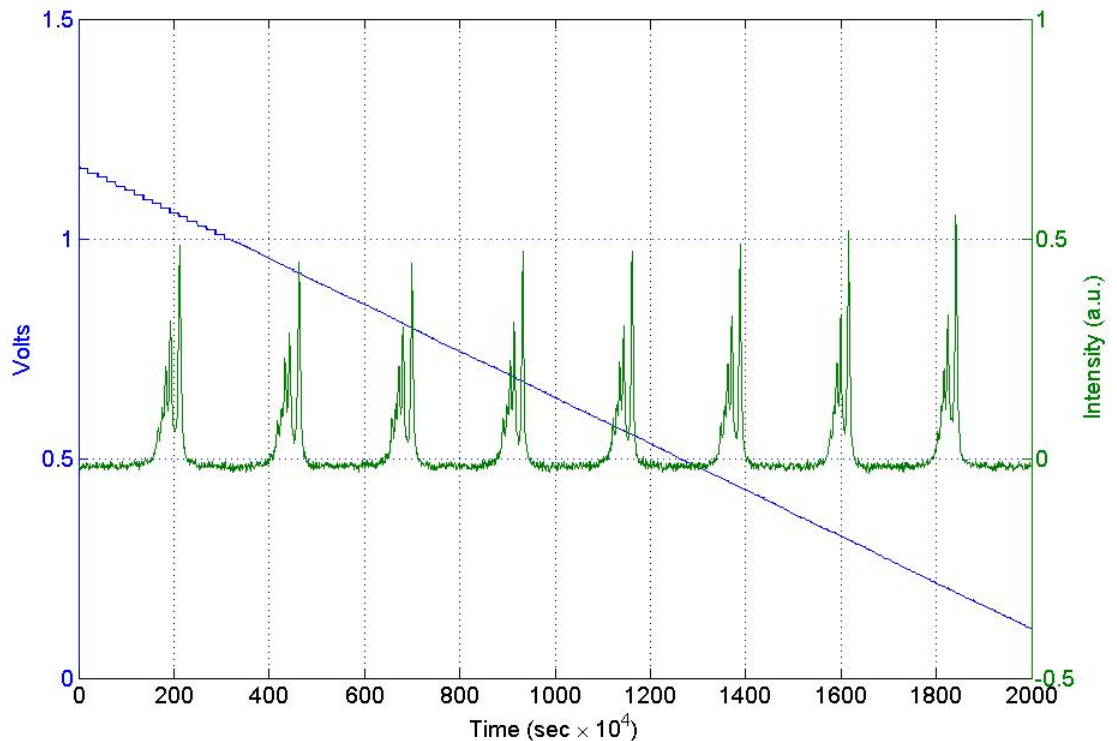
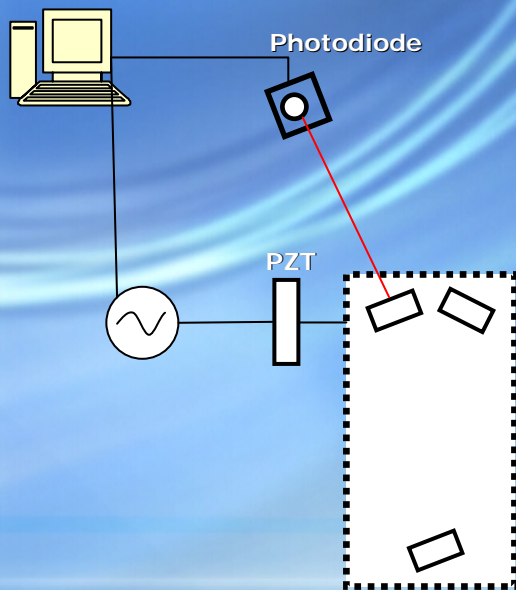
Investigate the susceptibility of mesa field to tilts of MH mirror



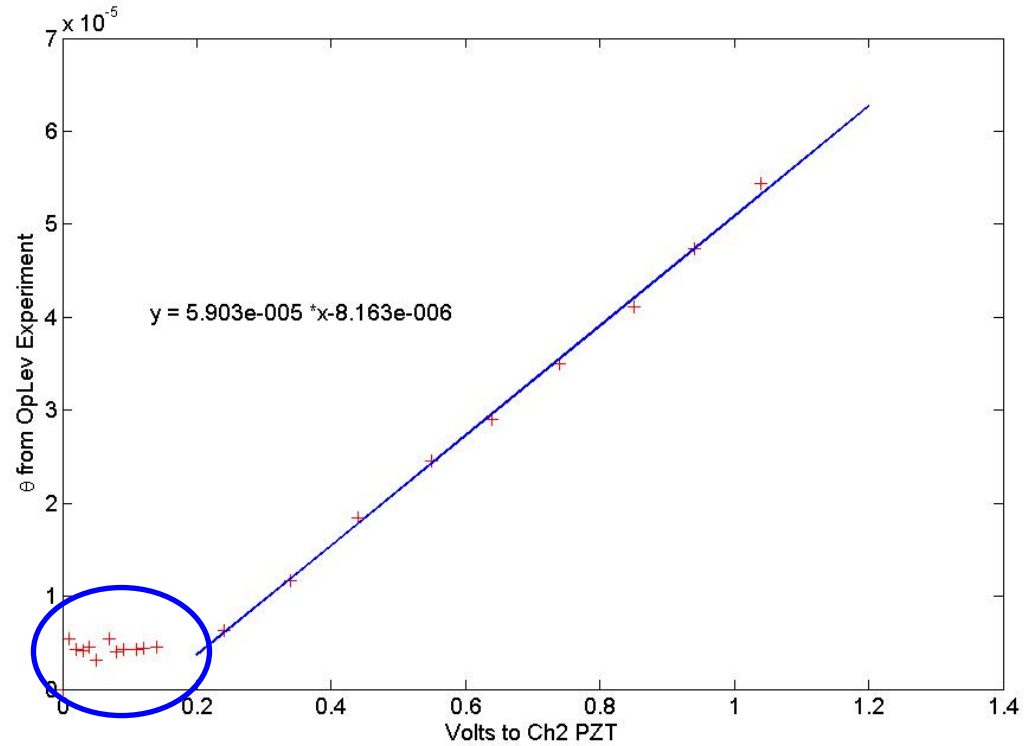
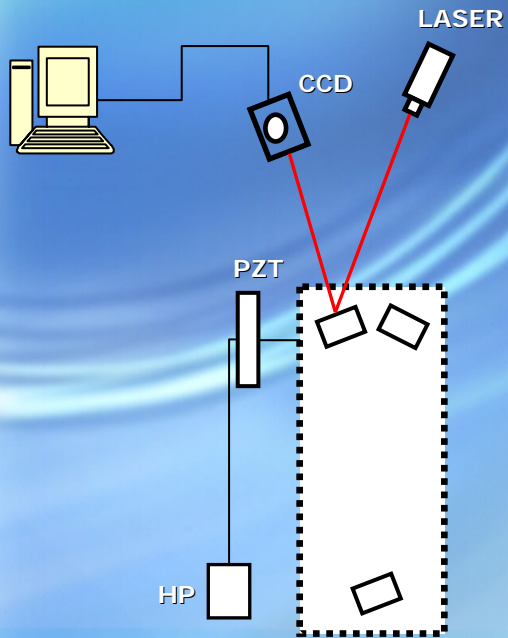
Tilt measurement, first approach: PZT

PZT procedure:

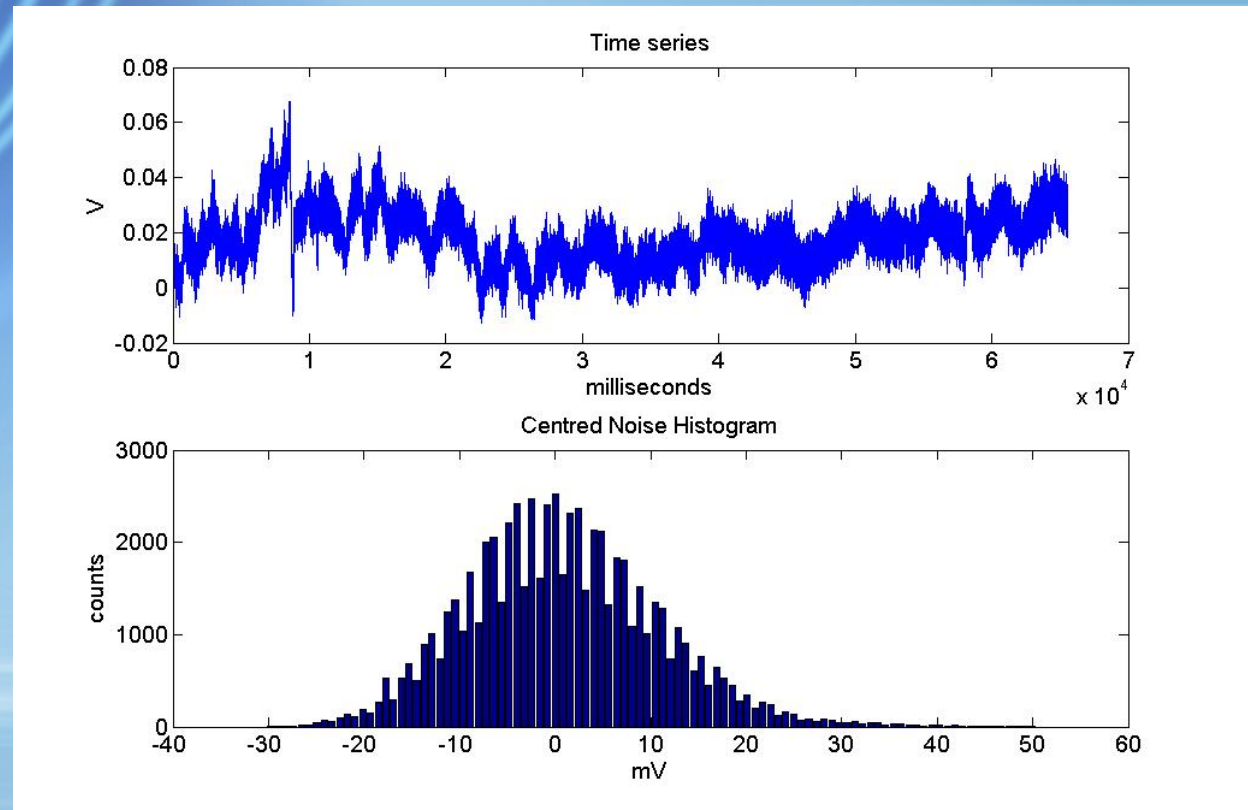
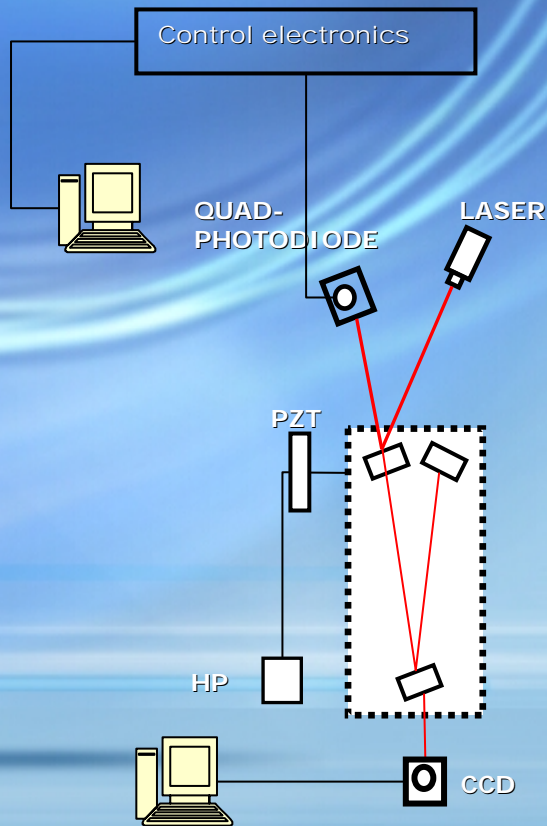
- Lock the cavity
- Sweep one mirror



LIGO Tilt measurement, Second Approach: Optical Lever & CCD



Tilt measurement, Third Approach: Optical Lever & Quad-Photodiode

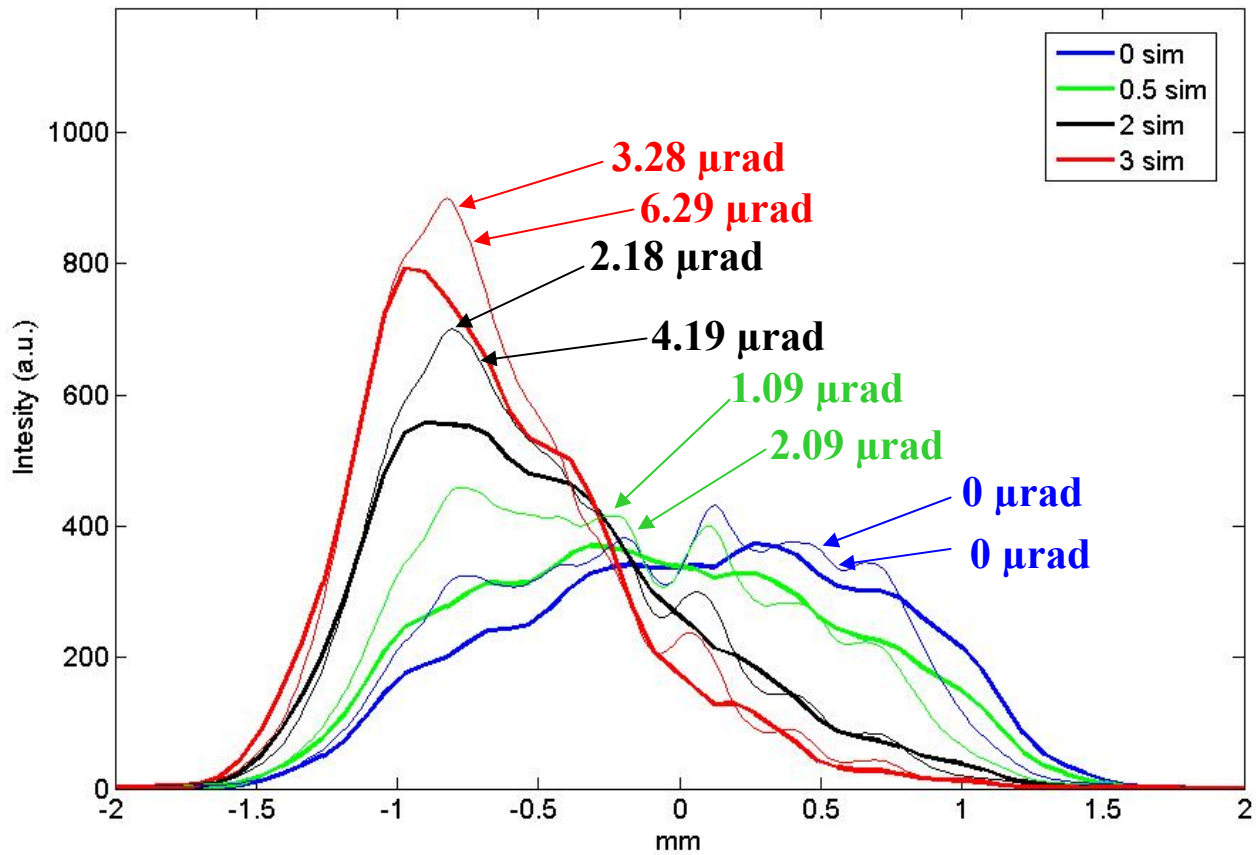
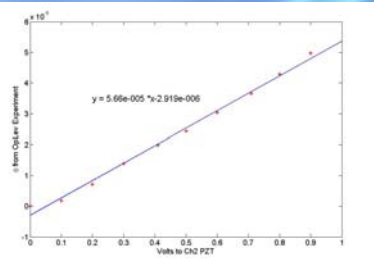
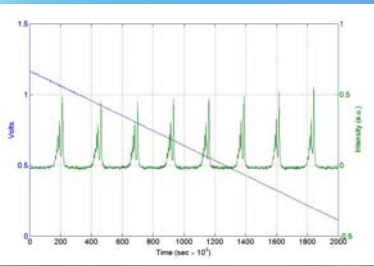


Laser off

Results

Oplever

PZT



Future Work

Caltech

- Improve methods for tilt measurement
 - Subtract beam jitter away
 - Fibre-coupled laser
 - Triggered beam capture with lock-in detection
- Power Recycling
- New Flat Mirrors (LMA)

Future Work

Possibly Elsewhere

- Build a nearly Concentric MH mirror cavity
- Thermal noise measurements (MH-TNI)

Thanks...

- Riccardo De Salvo
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