



LSC-Virgo Project 2b Joint Burst Search

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LIGO / Virgo Scientific Collaborations Meeting March 20, 2007

LIGO-G070108-00-Z





Project 2b Definition

Followed from project 2a

(3 hours of real, non-coincident data exchanged: S4 H1 and G1 ; C7 Virgo)

Designed to involve *coincident* data exchange

Refined and launched in Fall 2006

Goals

Exchange a few days of real coincident data from all 5 detectors

Have people from both collaborations work together on analysis

Discuss this work in regular Burst Group telecons

Use data quality and veto information

Deal with a mixture of network configurations

Try to draw some conclusions about the quality of the data and the benefit of using the whole network





Coincident data from Sept. 8–11, 2006

Virgo Weekend Science Run 1 (WSR1)

LIGO + GEO S5 data

Includes 35 hours of 5-way coincident running

Data exchanged

(Marie Anne Bizouard, Patrick Sutton, Shourov Chatterji, Siong Heng, Laura Cadonati, Keith Thorne)

Frame data with different secret time shift for each site

Virgo data resampled from 20000 Hz \rightarrow 16384 Hz, 4096 Hz, 4000 Hz

Segment lists

Data quality and veto time intervals

Times of hardware signal injections

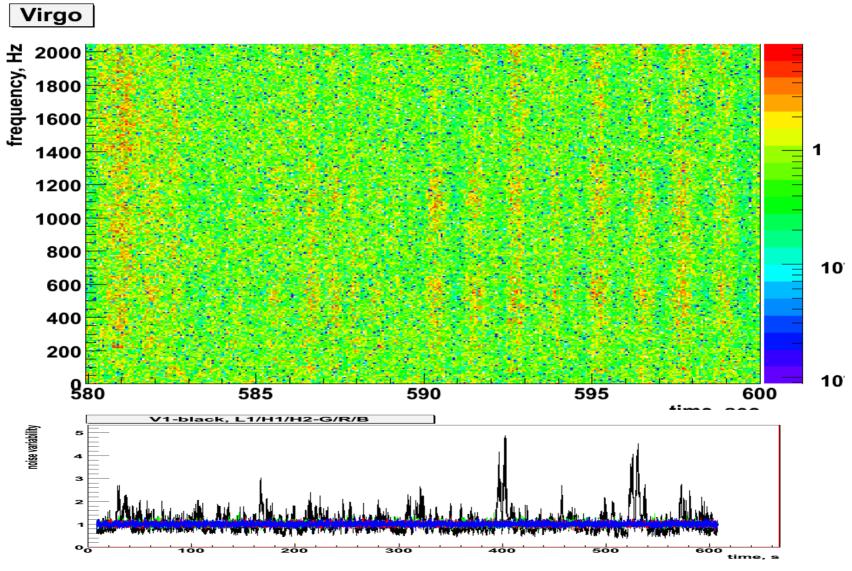
MDC frame files containing simulated signals

A laborious task to agree on file contents, produce & copy files



Non-Stationarity of Virgo WSR1 Data





LSC Meeting, 19 Mar 2007

LIGO-G070108-00-Z



Coincidence

Gianluca Guidi, Jonah Kanner

Coherent WaveBurst

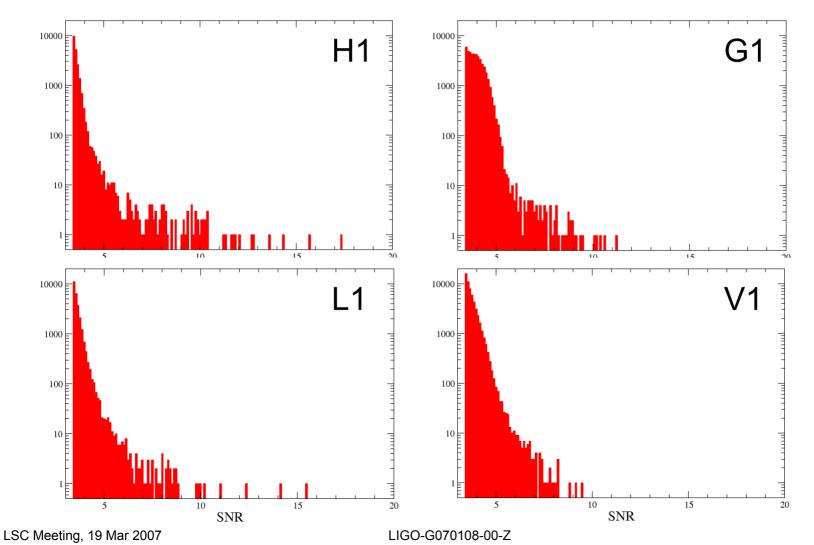
Igor Yakushin, Sergey Klimenko

X Pipeline GRB analysis

Patrick Sutton, Michal Was







6



Define coincidence requirements Look at time, frequency, ... for simulated signals Delta t between 2 IFOs - H1 V1 - hrss = 5e-21 18 ±30_Ims H1-V1 Δt 16 14 12 Counts - Total 500 10 8 2 0 -0.1 -0.08 -0.06 -0.04 -0.02 0 0.02 0.04 0.06 0.08 0.1 Delta t (s)

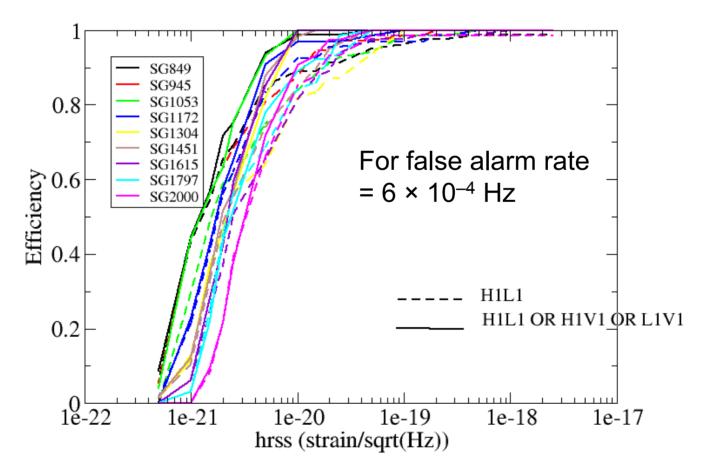
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(0)



Efficiency for OR of 2-fold coincident triggers

H1L1 vs (H1L1 OR H1V1 OR L1V1)



((O))VIRG



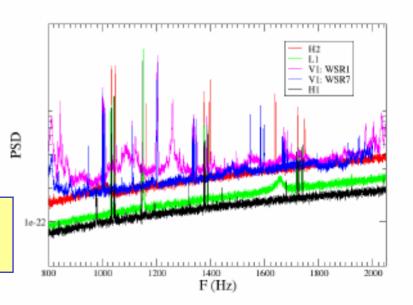


 performance at FA rate of 1μHz h_{rss} errors ~15% 					
network	hrss@50% sg361q9	hrss@50% sg849q9	hrss@50% sg1615q9	live time sec	
H1xH2	11x10 ⁻²²	16x10 ⁻²²	31x10 ⁻²²	182772	
L1xH1xH2	8x10 ⁻²²	14x10 ⁻²²	37x10 ⁻²²	157599	
L1xH1xH2xV1	9x10 ⁻²²	17x10 ⁻²²	40x10 ⁻²²	1040 <mark>62</mark>	
L1xH1xH2xG1	9x10 ⁻²²	16x10 ⁻²²	41x10 ⁻²²	140351	
L1xH1xH2xV1xG1	9x10 ⁻²²	16x10 ⁻²²	42x10 ⁻²²	102907	

both sensitivity and stationarity of the noise are critical for a detector to be useful in the network

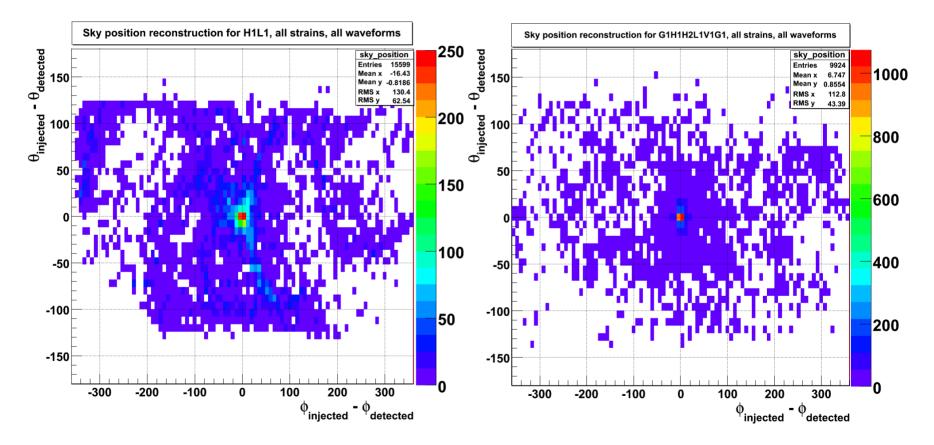
Adding Virgo to network doesn't improve hrss @50% upper limits for this data set

S.Klimenko, G070090 -00-Z , March 19, 2007, LSC





Including Virgo and GEO improves sky position reconstruction





A fictitious GRB during WSR1: Sky position favorable for GEO & Virgo

GPS trigger time = 841896355	Site	F_+^2+F_x^2
right ascension = 217.5255	Н	0.2598
declination = -28.7510	${ m L}$	0.3364
	V	0.8356
	G	0.7691

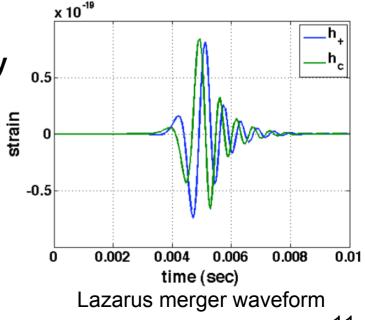
Use X Pipeline with "hard constraint" likelihood

Use off-source and simulation results to tune null energy vs. incoherent energy consistency test

hrss 90% C.L. upper limits:

LIGO-only: 4.6 × 10⁻²²

LIGO-Virgo: 3.7 × 10⁻²²







The Project 2b exercise has been (and still is) useful

Coherent WaveBurst analysis is pretty much complete Would like to implement complete coincidence analysis pipeline

We have had to deal with a number of practical issues

WSR1 data is glitchy

Value added to these analyses is marginal (?)

More recent Virgo data is better!