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# Stochastic Searches: Status and Plans

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# Status of S4 Papers

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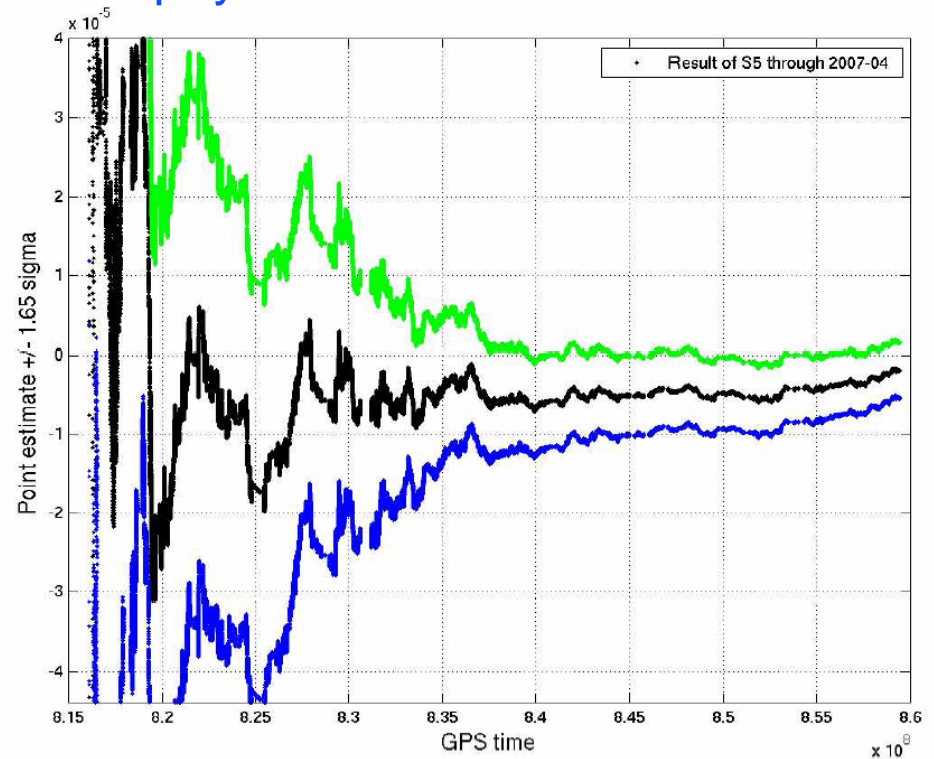
- Old news: S4 isotropic search using LHO-LLO pairs published in *The Astrophysical Journal* 659, 918 (2007).
- S4 ALLEGRO-LLO search published in *Physical Review D* 76, 022001 (2007).
- S4 radiometer paper under review in *Physical Review D*.
- S4 FSR paper (H1H2) under internal review (both calibration and stochastic reviews).



# S5 Status: H1L1 Isotropic Search

- S5 H1L1 isotropic search:  $\sigma_{\Omega} = 4.1 \times 10^{-6}$ .
  - » With unphysical time-shift.
  - » Using data up to April 1, 2007.
  - » Assuming frequency independent spectrum, and  $h=0.72$ .
- Before opening the box:
  - » Data quality cuts to be finalized
  - » Frequency mask to be finalized
- Goal: have the result by September, such that the internal review could be completed by the end of the year.
  - » *Preliminary result to be presented at the AAS meeting in January, 2008.*
- Should beat the BBN bound ( $1.4 \times 10^{-5}$  in our frequency band).

## Running Point Estimate with Unphysical Time Shift



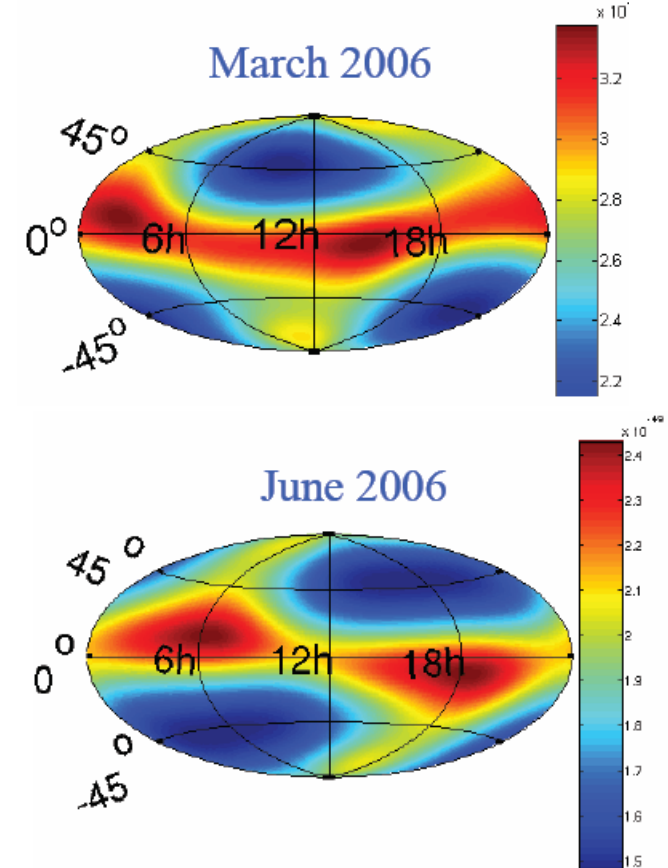


# S5 Status: Radiometer Search

- S5 radiometer search (R. Ward)
  - » Repeat of the S4 radiometer analysis.
  - » Intend to deconvolve the antenna pattern from the maps and estimate covariance matrix on the “cleaned” map.
  - » Also intend to do “narrow-band” searches in different directions on the sky.

Since IFOs are most sensitive at night, the region on highest sensitivity moves across the sky as the Earth moves around the Sun.

Map of sensitivity  $\sigma$  as a function of direction



$$\sigma(\Omega) = (1.3 - 2.27) \times 10^{-49} \text{ Hz}^{-1}$$



# S5 Status: Spherical Harmonics Decomposition

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- Attempt to produce a “best” estimate of the background distribution on the sky.
  - » Rather than assuming the background is isotropic or point-like.
- Slightly different formalism (see S. Ballmer’s talk).
- Implement as part of existing pipeline.
- Current efforts:
  - » Produce new simulation code that can inject an arbitrary map.
  - » Perform end-to-end tests of the injection code.
  - » 2 SURF students at Caltech: Madeleine Udell, Grant Meadors.
- To do:
  - » Adjust post-processing procedures.
  - » Finish extensive testing:
    - Performance of different searches for different sky distributions.

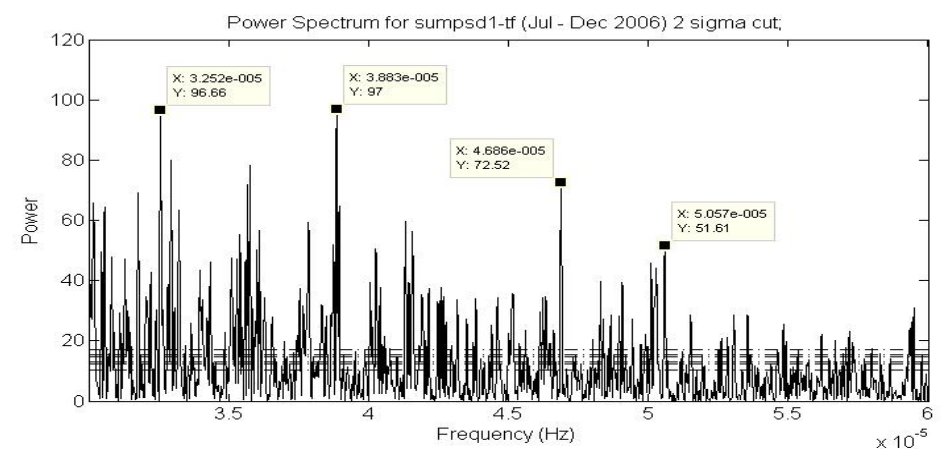
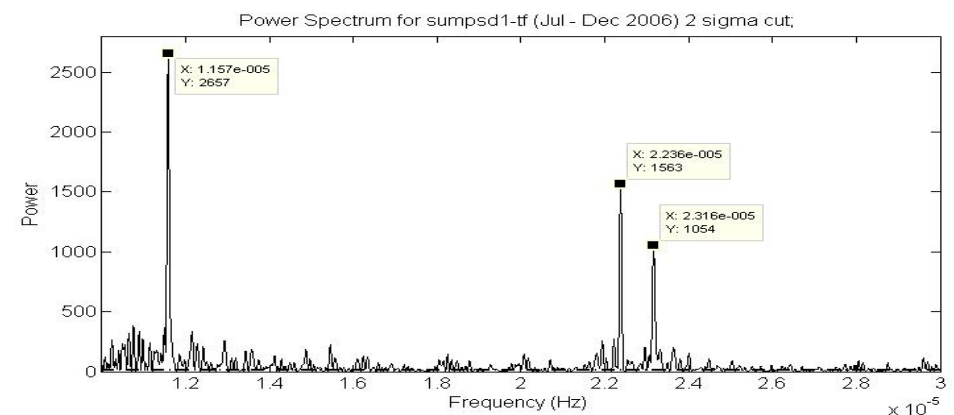


# S5 Status: H1H2 FSR Search

By Chad Forrest

- In the process of repeating the S4 analysis procedure on S5 data.
- Also investigating diurnal variations at FSR:
  - » Integrate H1 PSD in the band  $\pm 200$  Hz around FSR (for every 64-sec frame).
  - » Calculate frequency power spectrum of this time series using Lomb-Scargle algorithm.
- Observe 1<sup>st</sup> and 2<sup>nd</sup> harmonics of the solar daily frequency (not of sidereal day!)
- Other lines not understood, but could be related to the tidal servo.
- Plans: Repeat on the whole dataset, and try to determine upper limits for localized sources (such as the galactic center etc).

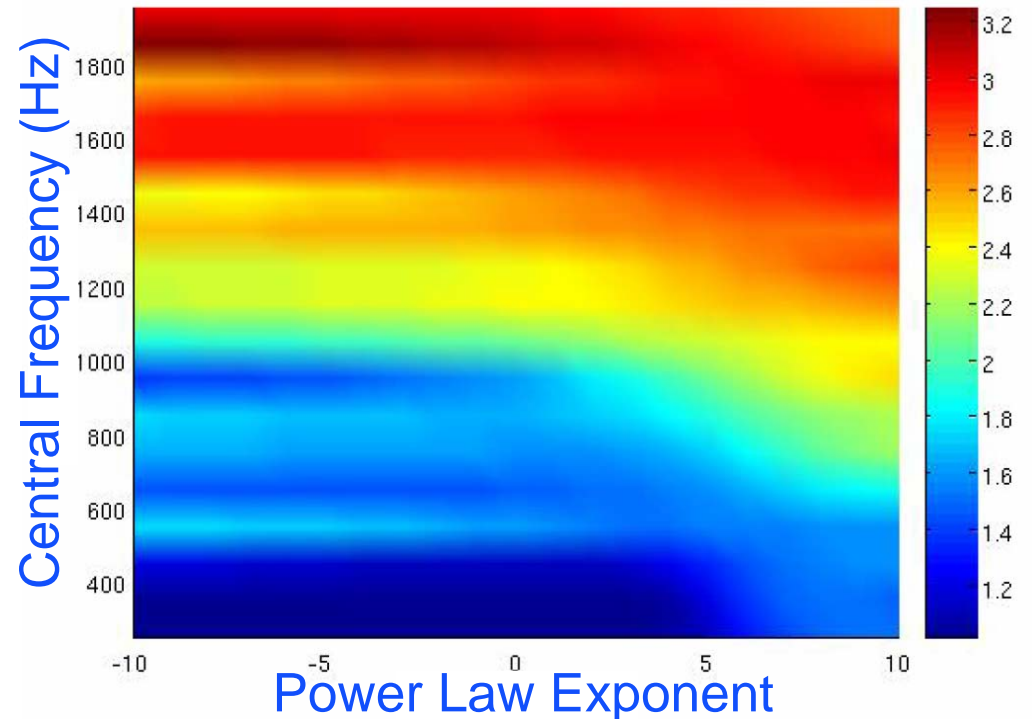
LIGO-G070509-00-D





# S5 Status: LSC-VIRGO

- Cross-correlating LSC-VIRGO pairs could improve the sensitivity and the robustness of the network, especially at high frequencies.
- Considering both isotropic and directional (radiometer) searches with S5 data.
- Analysis of real data to start soon.



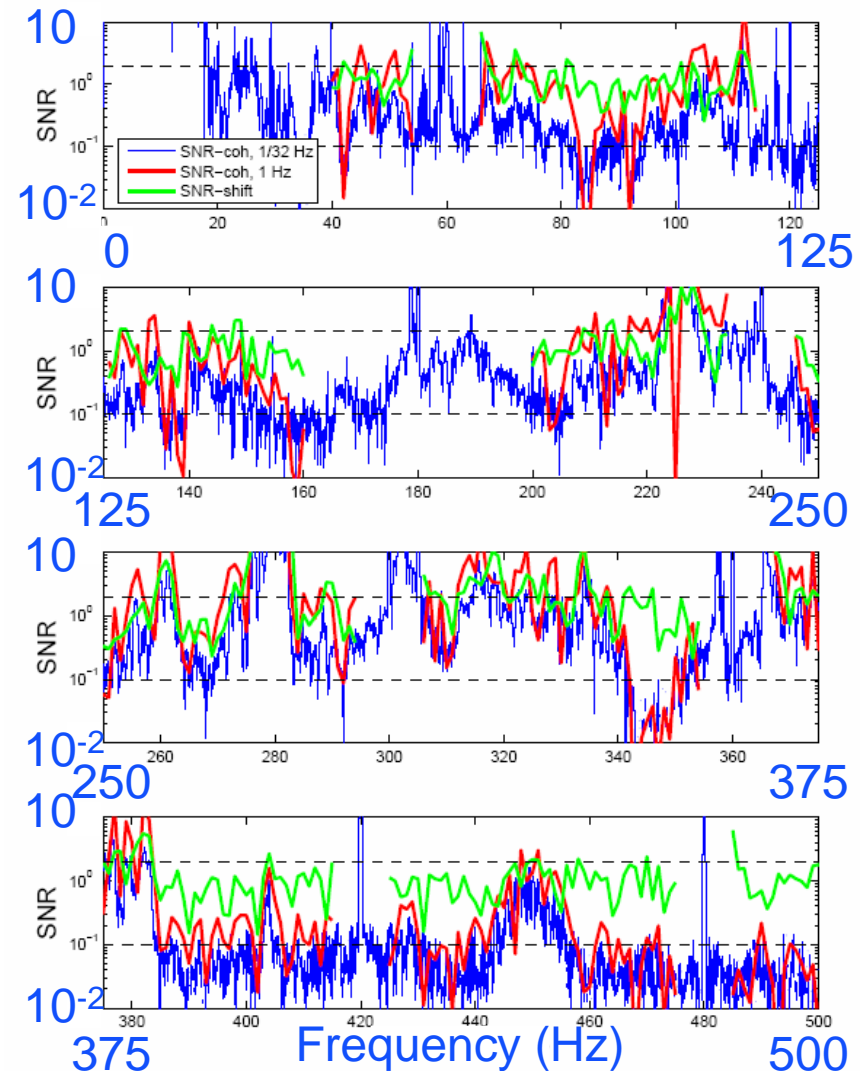
Upper limit improvement factor (WSR10) as a function of

- Power law exponent of the model
- Central frequency of a frequency band used for the analysis, width=250 Hz.



# S5 Status: H1H2 Isotropic Search

- Developed 2 techniques for identifying contaminated frequency bands:
  - » IFO-PEM coherence
  - » Time-shift
- IFO-PEM coherence method can be used to estimate the “residual” contamination in the good frequency bands.
  - » Can evaluate the method using contaminated frequency bands.
- Preliminary results (using data up to April 2006) were at the level:
  - »  $\sigma_{\Omega} = 1.4 \times 10^{-6}$
- Current efforts:
  - » Assess time-variability of both techniques.
  - » Bring the IFO-PEM coherence calculation closer to stochastic algorithm in terms of how the data is handled (frequency resolution, windowing etc).







# Pipeline Upgrade

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- Produce stochastic intermediate data (SID).
  - » Includes Power and Cross Spectral Densities of an IFO pair.
  - » Calculated for every ~minute.
- Individual searches to use SID rather than time-frames.
- Could also be used for:
  - » Detector characterization.
  - » Search for transients on the time-scale of minutes/hours.
  - » Search for narrow-band correlations.
- First-pass SID already produced on the CIT cluster.
  - » Currently checking the data quality.
- Collapse (i.e. average) SID to 1 sidereal day.
  - » Collapsed data could be stored on a laptop...



# Publication Plans

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- Preliminary S5 LHO-LLO isotropic result to be presented at the AAS Meeting in January 2008.
  - » Relies on the review of S5  $h(t)$  data!
- S5 LHO-LLO isotropic result to be published in a short paper (all of S5).
  - » Beating the BBN bound, implications for some models.
- S5 Radiometer paper to be published in a short paper, ~mid 2008.
  - » Improved sensitivity, deconvolving antenna pattern...
  - » If time-scale is similar, publish radiometer and isotropic results together.
- S5 H1H2 isotropic paper should be long, detailing the analysis procedure.
- S5 Spherical Harmonic Decomposition paper.
  - » Likely to be long, detailing the analysis procedure.
- S5 FSR paper.
  - » Likely to be long, detailing the analysis procedure.
- LSC-VIRGO paper
  - » Isotropic, using several baselines, probably focusing on higher frequencies.
  - » Radiometer search using several baselines?