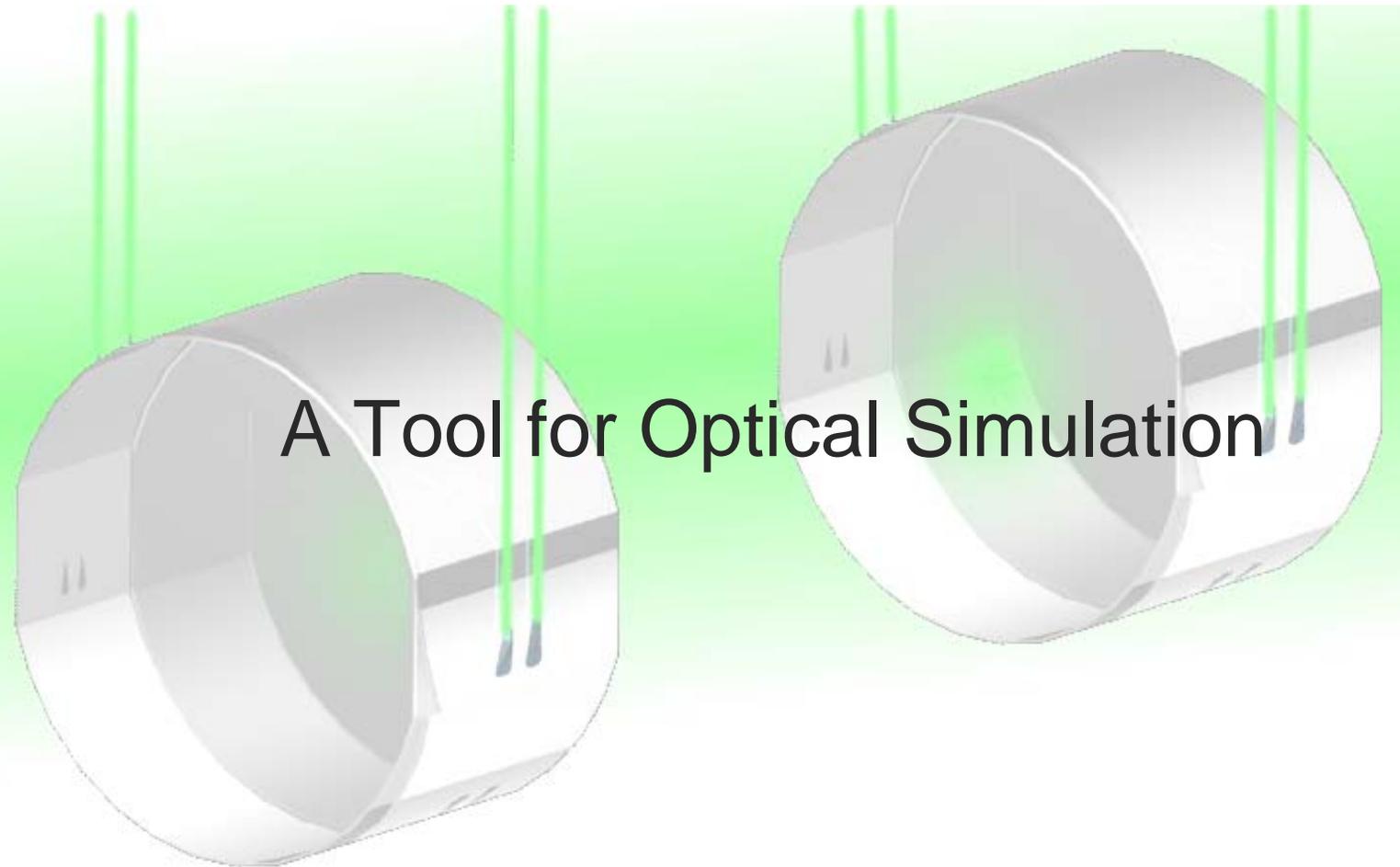
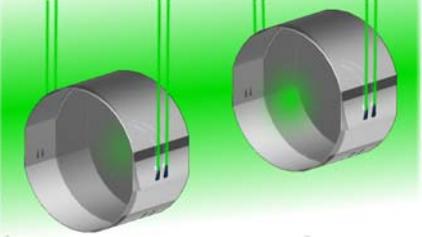


Optickle

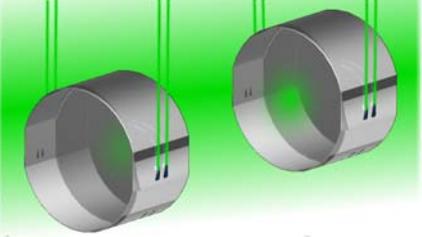
A Tool for Optical Simulation

The image features two 3D cutaway diagrams of optical cavities, likely representing the mirrors of a gravitational wave detector. Each cavity is a cylindrical component with a central lens and a central hole. Two parallel green laser beams are shown entering from the top and reflecting off the inner surfaces of the cavity. The background is a light green gradient.



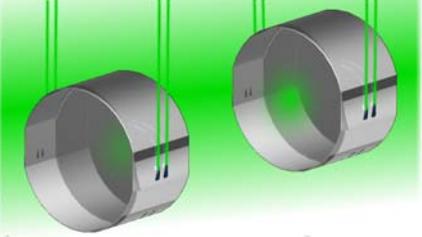
What is already out there?

- Finesse
 - Many modes
- Static FFT models
 - Non ideal optics
- E2E, Siesta
 - Time domain
- Optickle
 - TEM00 and 01 only
- Optickle
 - Ideal optics
- Optickle
 - Frequency domain



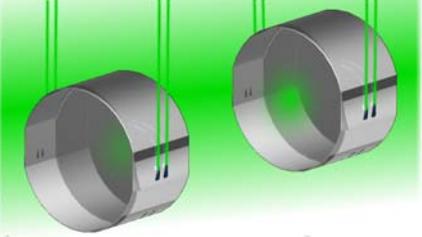
Why Optickle?

- Includes radiation pressure directly
- Computes quantum noises starting from vacuum fluctuations at all loss points
- Matlab based
 - Commonly used environment, especially for control system development and noise analysis (Simulink interface recently added)
 - Convenient plotting and post-processing
 - Avoids many portability problems



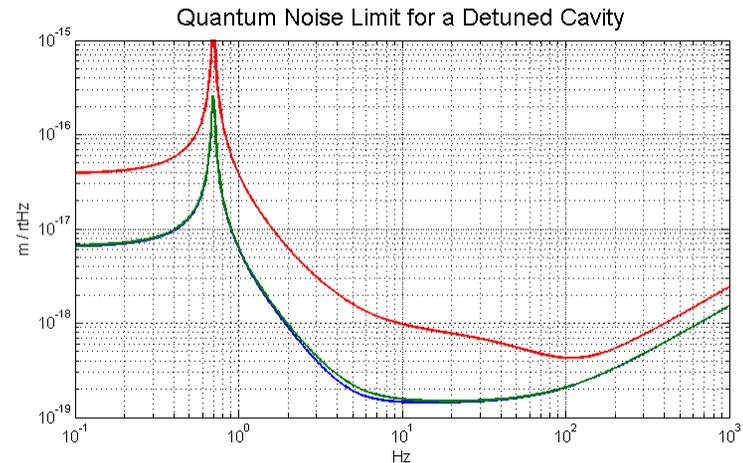
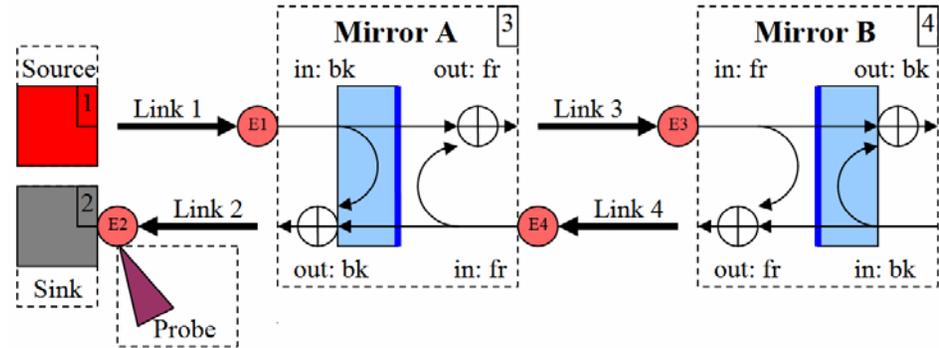
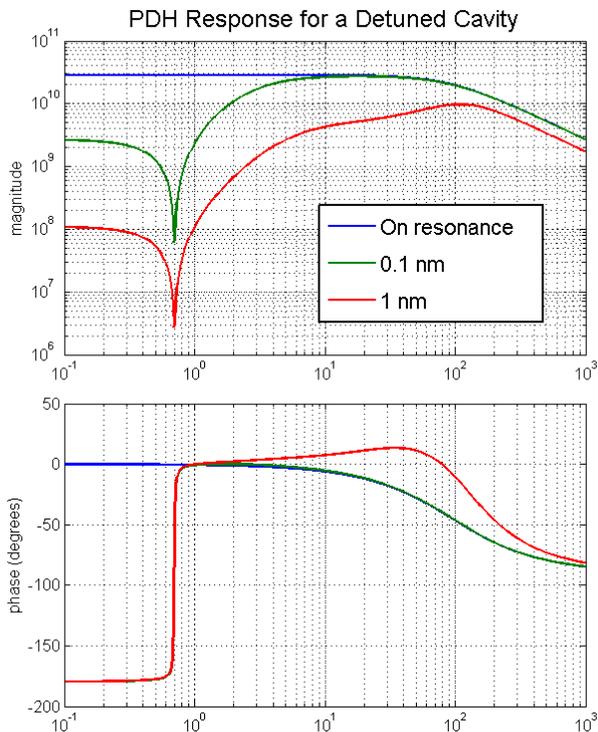
What can it do?

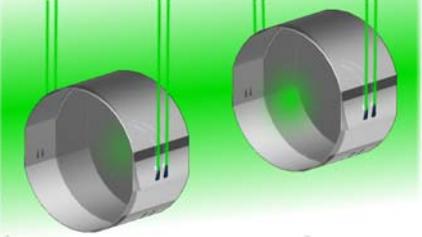
- Simulates a well aligned, well matched optical system
- Modular structure allows for many interferometer configurations
- Computes longitudinal and angular transfer functions, including radiation pressure
- Computes DC signals, and quantum noises



Longitudinal Example: Detuned Cavity

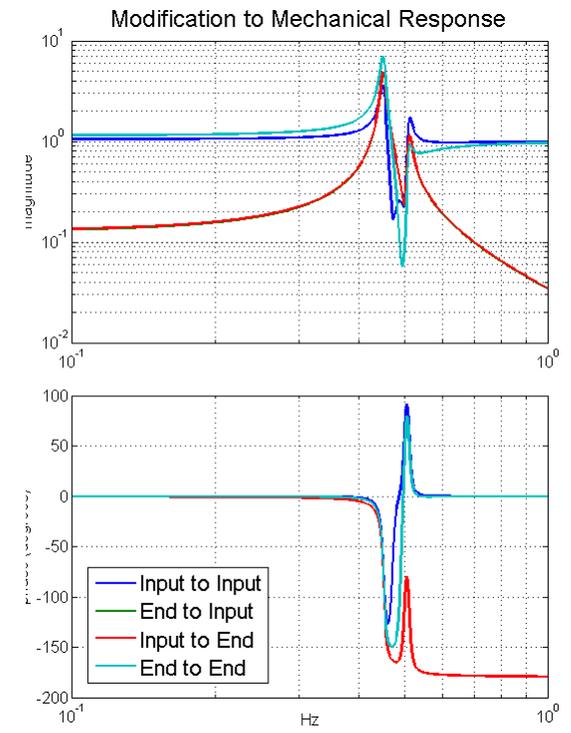
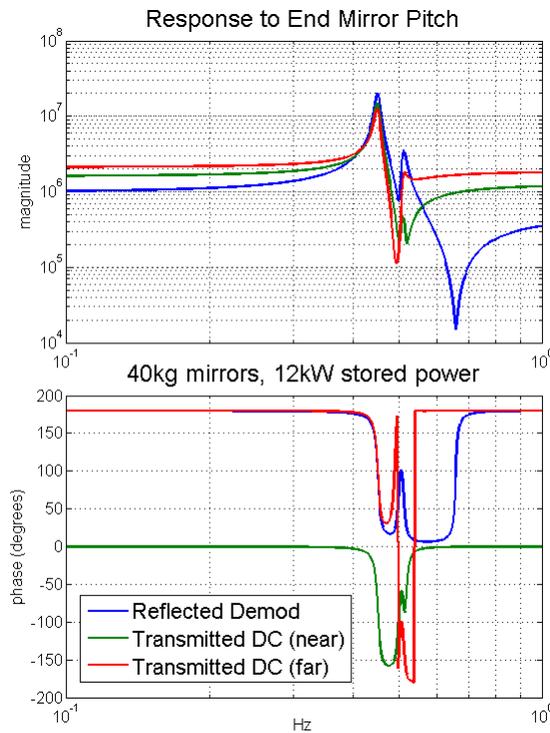
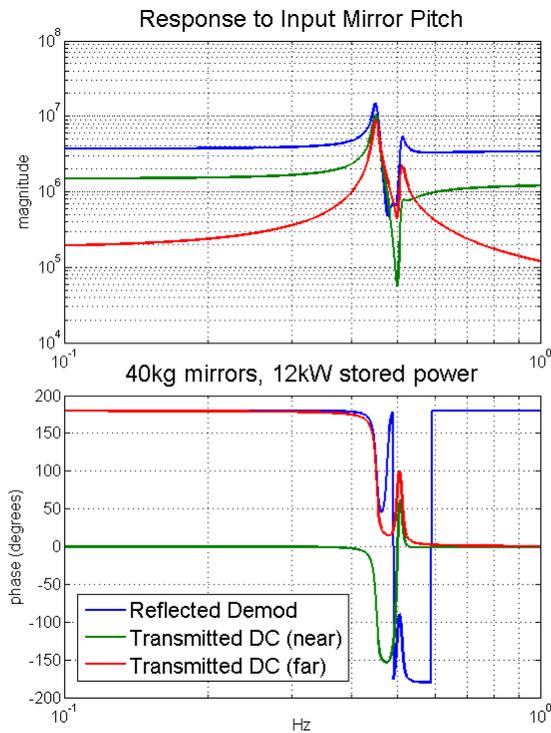
- Radiation pressure effects
- Quantum noise

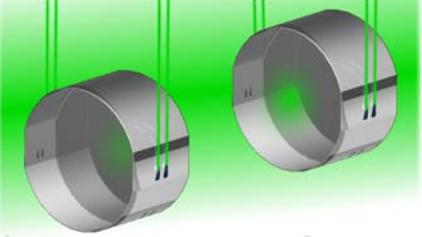




Angular Example: Flat-Curved Cavity

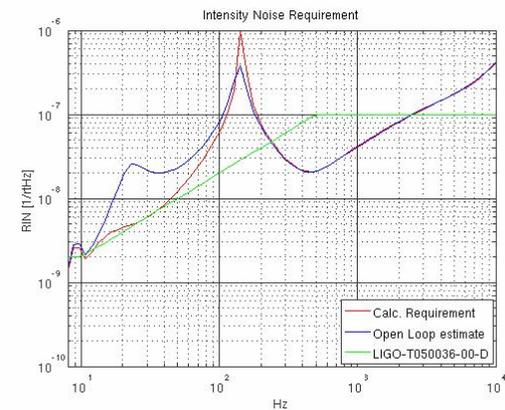
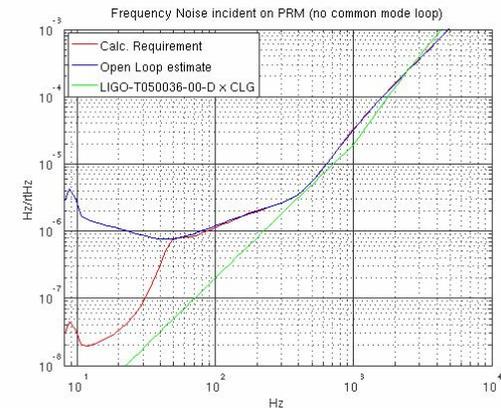
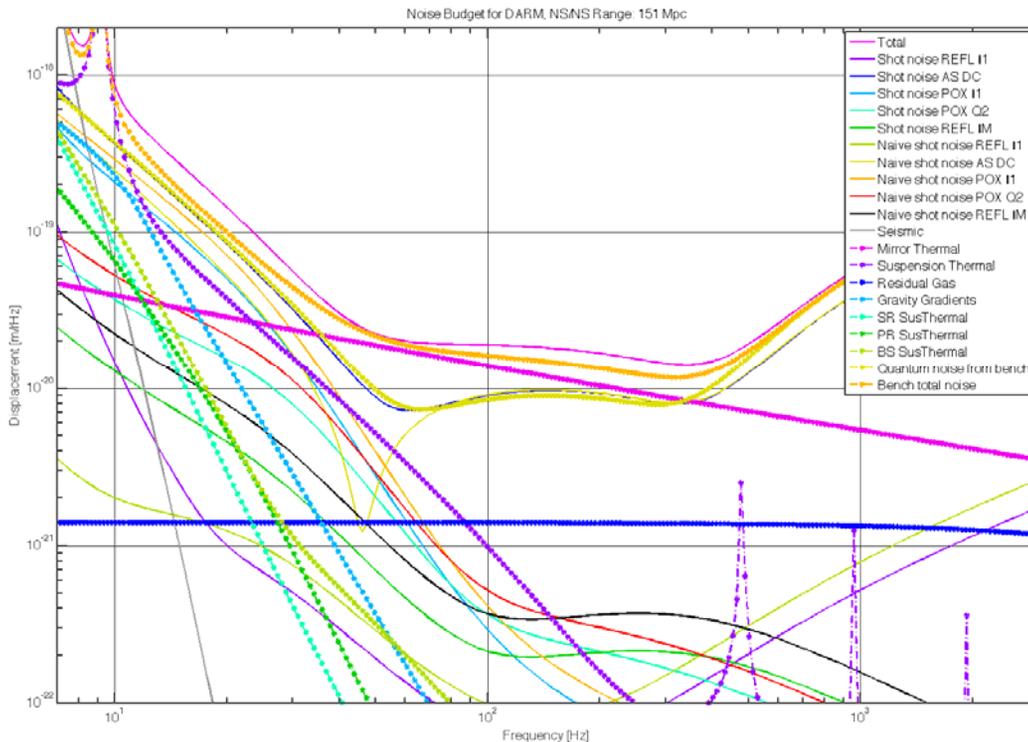
- Radiation pressure effects present in angular response

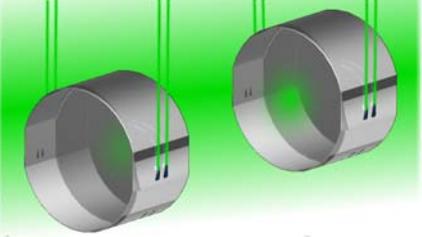




Some Applications (Stefan Ballmer)

Loop shapes and noise couplings for Advanced LIGO design

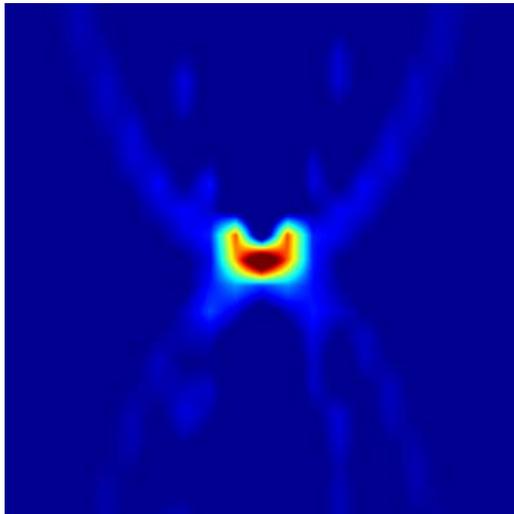




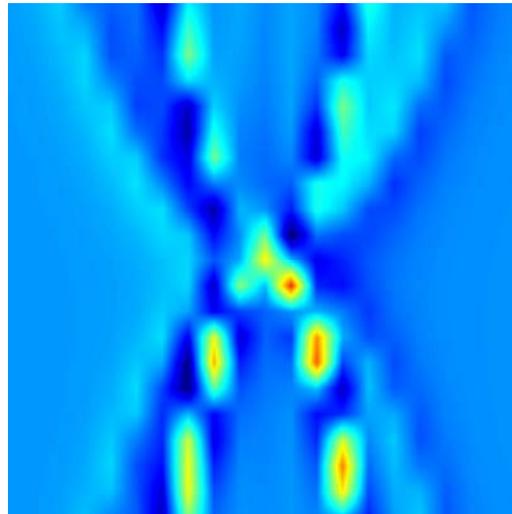
Some Applications (Lisa Barsotti)

- Phase space exploration for Advanced LIGO locking algorithm

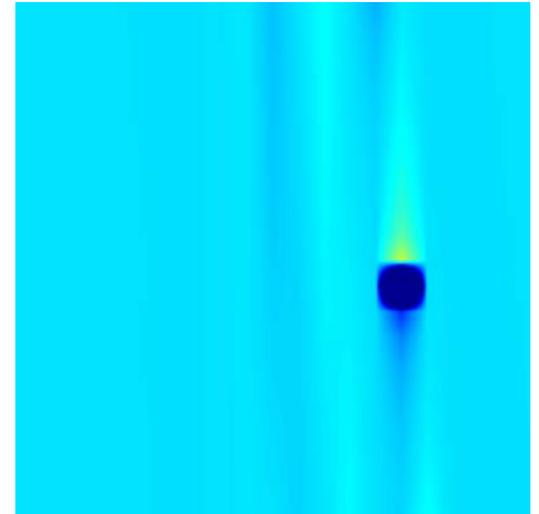
REFL31I vs PRCL

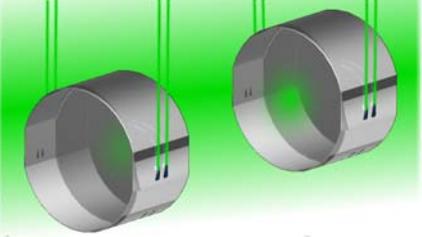


REFL31Q vs MICH



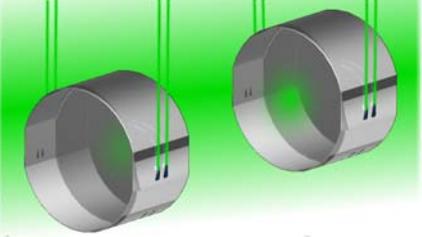
REFL32I vs SRCL





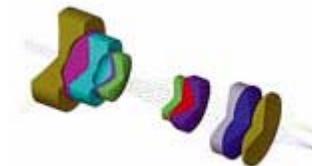
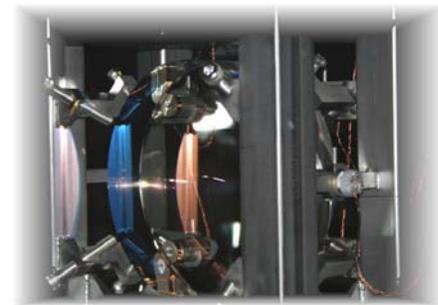
Conclusion

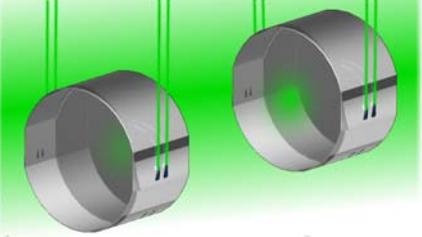
- Optickle was written to facilitate advanced interferometer design
 - Radiation pressure effects are large in both longitudinal and angular responses
 - Quantum noise is not easily divided into shot noise and radiation pressure noise
- Optickle is being used in Advanced LIGO design (Stefan, Lisa, Rana, ...)
- Development is complete!



Optickle Modules

- Mirror
- Beam Splitter
- Source
- Sink
- Modulator
- RF Modulator
- Telescope

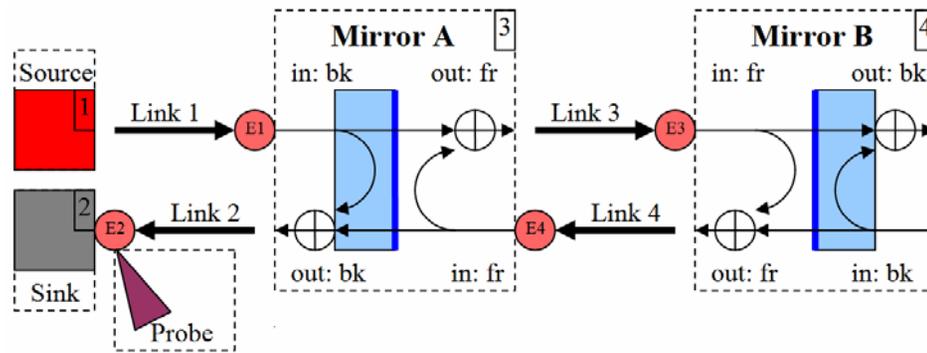


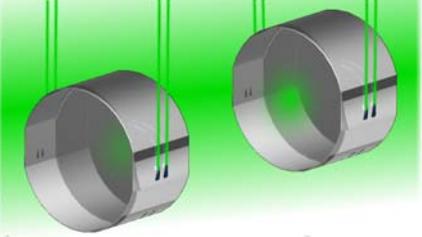


Optickle Glue

■ Links

- Used to define the relationships between the optical modules
- Connect from one module output to another module input, with some length

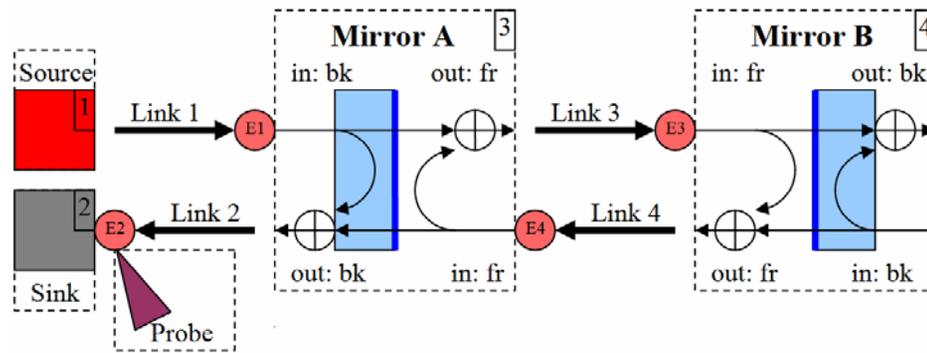


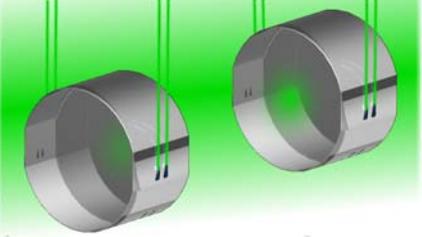


Optickle Output

■ Probes

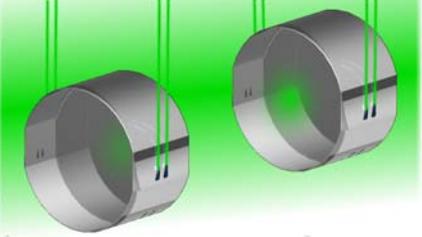
- Used to extract information about the fields in the simulation
- Often attached to the input of a Sink, thus making a photo-detector





Optickle Interface

- Direct
 - Vector of DC signals from probes
 - Matrix of transfer functions from all drives to all ports (probes and drives)
 - Quantum noise at all ports
- Indexing can be painful
 - Functions for getting indexes by name
- Control loops added externally
 - e.g., looptickle



Optickle Interface

■ Simulink

- Reduces transfer matrix to the inputs and outputs specified by the user
- Control systems are easier
- Coherent noises are easier

