Macroscopic Quantum Mechanics in Gravitational-Wave Detectors

Yanbei Chen

summarizing research done by Stefan Danilishin, Chao Li, Haixing Miao, Helge Müller-Ebhardt, Henning Rehbein, Roman Schnabel and Kentaro Somiya

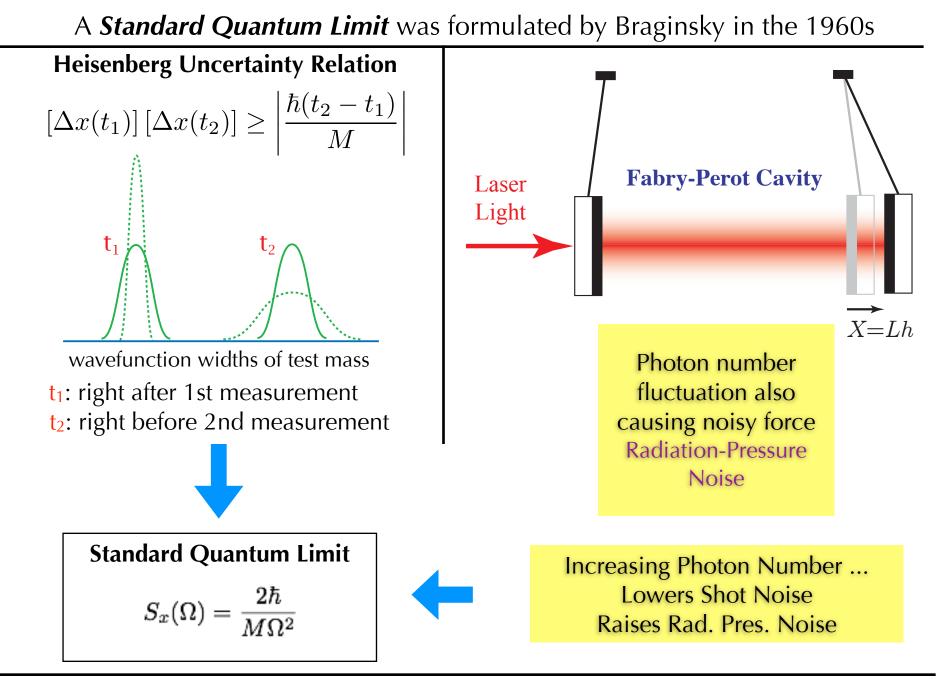
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LIGO-G070761-00-Z

Theme

- Major R&D of 3rd generation detectors directed toward surpassing the Standard Quantum Limit
- Having classical noise below the SQL makes it feasible to study **quantum mechanics** of **macroscopic test masses** (1g to 10 kg)
 - **Prepare** nearly Heisenberg Limited, **test-mass** quantum states
 - Allow them to **evolve** (survive) quantum mechanically
 - Allow verification of quantum state with sub-Heisenberg accuracy
- People (so far) involved in this collaboration
 - AEI: Yanbei Chen, Karsten Danzmann, Stefan Danilishin, Helge Müller-Ebhardt, Henning Rehbein, Roman Schnabel, Kentaro Somiya
 - Caltech: Chao Li, Yasushi Mino, Sam Waldman, Kip Thorne
 - MIT: Thomas Corbitt, Nergis Mavalvala, Chris Wipf
 - UWA: Haixing Miao
 - Moscow ...

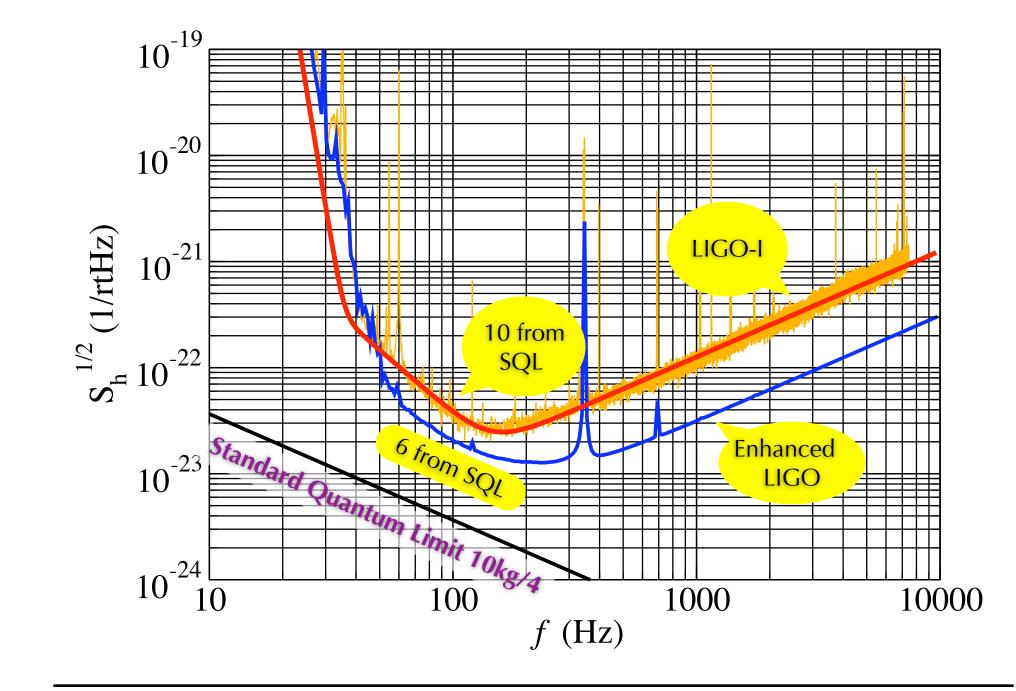
The Standard Quantum Limit



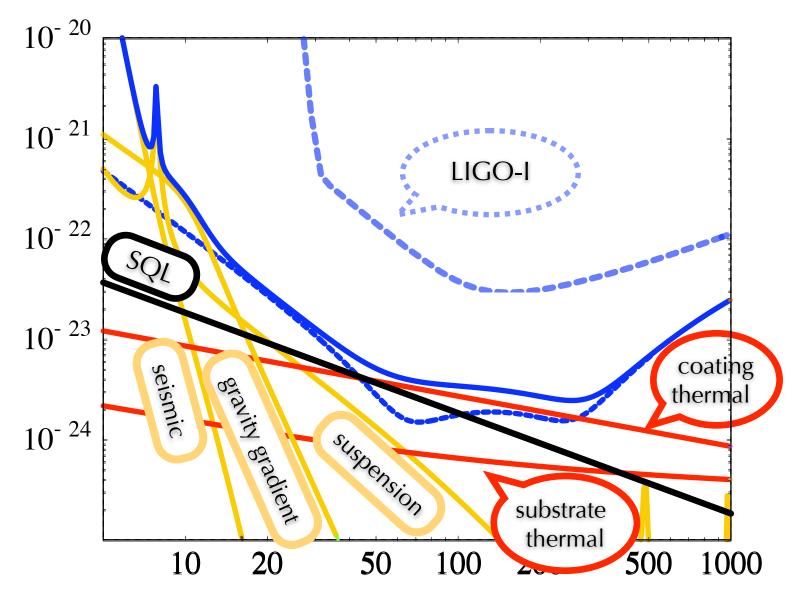
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Current Situation



Noise Budget of Advanced LIGO (2014)



Sensitivity improvement by 10, yet mass increase by 4 to 40kg

unless keep increasing mass (SQL~1/M^{1/2}), need to surpass SQL in 3rd generation detectors

Mechanical Harmonic Oscillator

• Steady-State Schrödinger Equation

$$\left[-\frac{\hbar^2}{2M}\frac{\partial^2}{\partial x^2}+\frac{M\Omega^2x^2}{2}\right]\psi(x)=E\psi(x)$$

• Ground state

$$\begin{split} E &= \frac{\hbar\Omega}{2} \\ \delta x &= \sqrt{V_{xx}} \quad = \quad \sqrt{\frac{\hbar}{2M\Omega}} \equiv \delta x_q(\Omega) \\ \delta p &= \sqrt{V_{pp}} \quad = \quad \sqrt{\frac{\hbar M\Omega}{2}} = \delta p_q(\Omega) \\ V_{xp} &= \quad 0 \end{split}$$

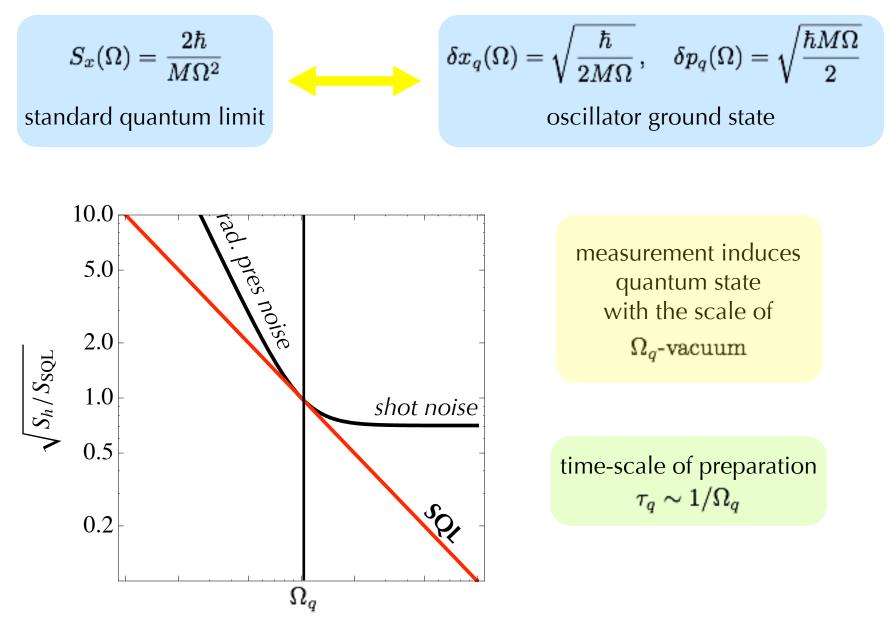
• Other Gaussian pure states

$$V_{xx}V_{pp}-V_{xp}^2=\hbar^2/4$$

 10Ω -vacuum 3 a pure state 2 a mixed state l $p/\delta p_q(\Omega)$ $\Omega/10$ -vacuum 0 Ω -vacuum -2-3 -3 -2-1 0 2 3 $x/\delta x_q(\Omega)$

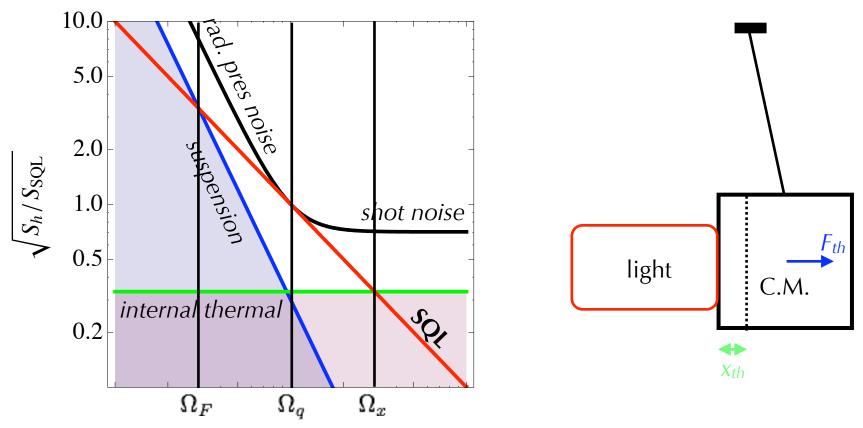
• The notion of "squeezing" and "anti-squeezing" depends on the **eigenfrequency**, which depends on the **potential well**

Connection between SQL and Ground State



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Simplified Classical and Quantum Noise Budget



- Two types of classical noise, both white
 - Force Noise: force acting on mirror center of mass (e.g., suspension thermal noise)
 - Sensing Noise: difference between what is measured and the center of mass (e.g., laser noise, internal thermal noise)
- Important time scales:

$$\tau_F \equiv 1/\Omega_F \,, \quad \tau_q \equiv 1/\Omega_q \,, \quad \tau_x \equiv 1/\Omega_x$$

Experimental Scenarios

- State preparation:
 - Takes time of $\tau_q \sim 1/\Omega_q$ to reach nearly pure state, if $\Omega_F < \Omega_q < \Omega_x$
 - Quantum entanglement between test masses [Talk by Helge Müller-Ebhardt]
- State survival under decoherence:
 - Lifetime given by $\tau_F \sim 1/\Omega_F$
 - allows
 - quantum evolution [Talk by Yanbei Chen]
 - state verification [Talks by Stefan Danilishin and Kentaro Somiya]
 - demonstrates that quantum entanglement can exist for a finite amount of time *[Talk by Helge Müller-Ebhardt]*
- Alternatively, pure **quantum state via feedback control**
 - steady-state entanglement [Talk by Helge Müller-Ebhardt]
- Before reaching sub-SQL classical noise
 - quantum optics/optical entanglement through ponderomotive squeezing [Talk by Henning Rehbein]