

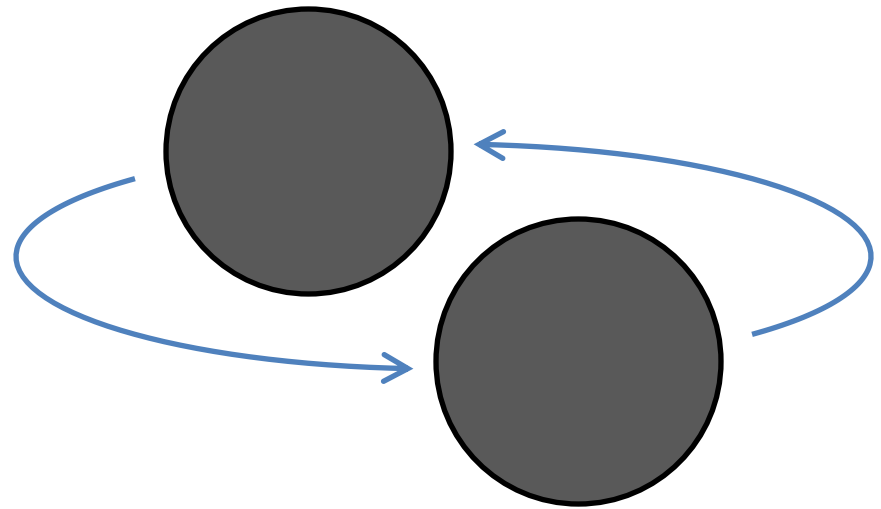
Detecting poorly understood sources with LIGO

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Detecting

- Inspiralling compact bodies are LIGO's most plausible burst source
- The search for them is highly optimised and looks only for those known waveforms
 - Known through computer simulation and analytic approximation
- It won't find the unexpected

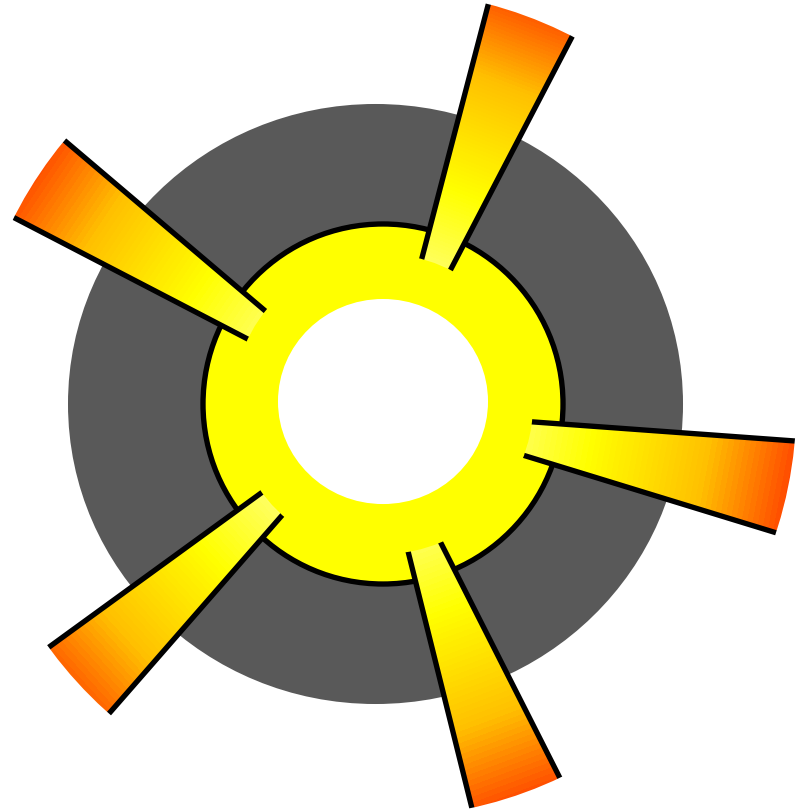


$$h(t) = \text{[red waveform]}$$

The equation shows a red waveform representing the gravitational wave signal $h(t)$. The waveform starts with a low-frequency, low-amplitude oscillation that gradually increases in both frequency and amplitude over time, characteristic of a chirp signal from an inspiralling binary system.

Poorly understood sources

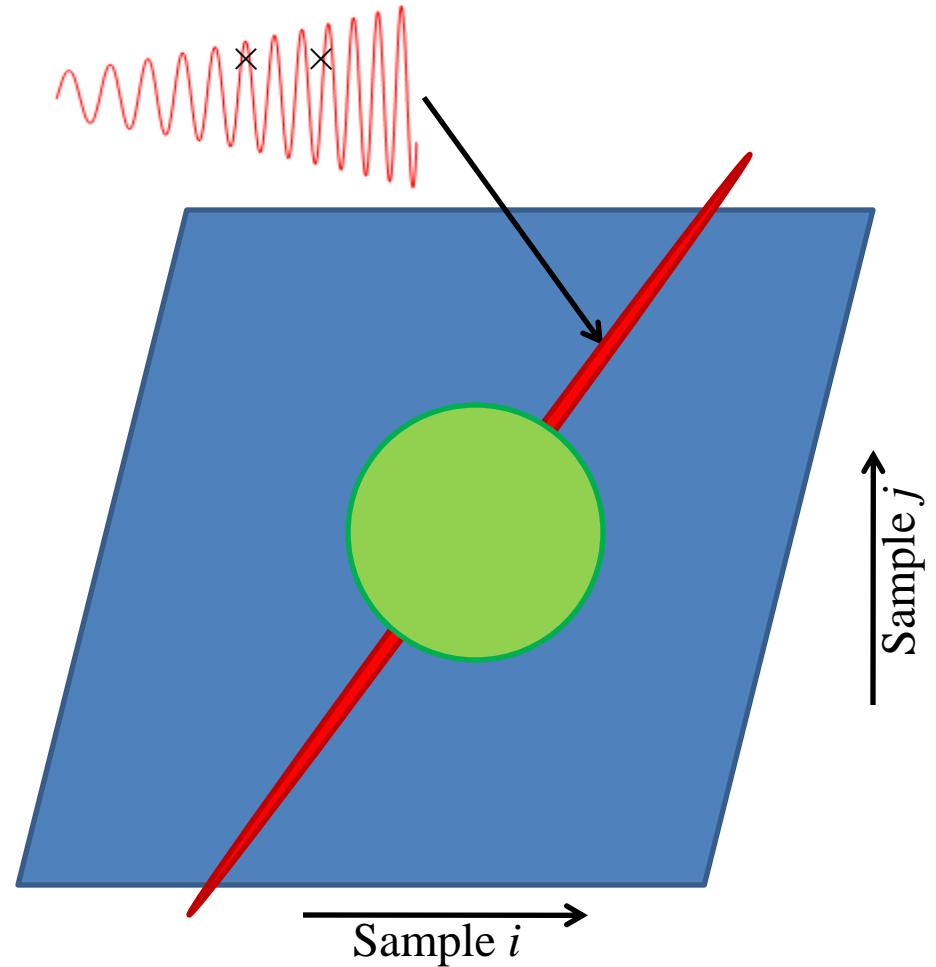
- Physical problems
 - Equation of state of very dense matter
 - Exotic sources or new physics
- Computational problems
 - Supernovae simulations
 - Inspiral simulations for some regions of parameter space
 - Enormous progress in recent years



$$h(t) = ???$$

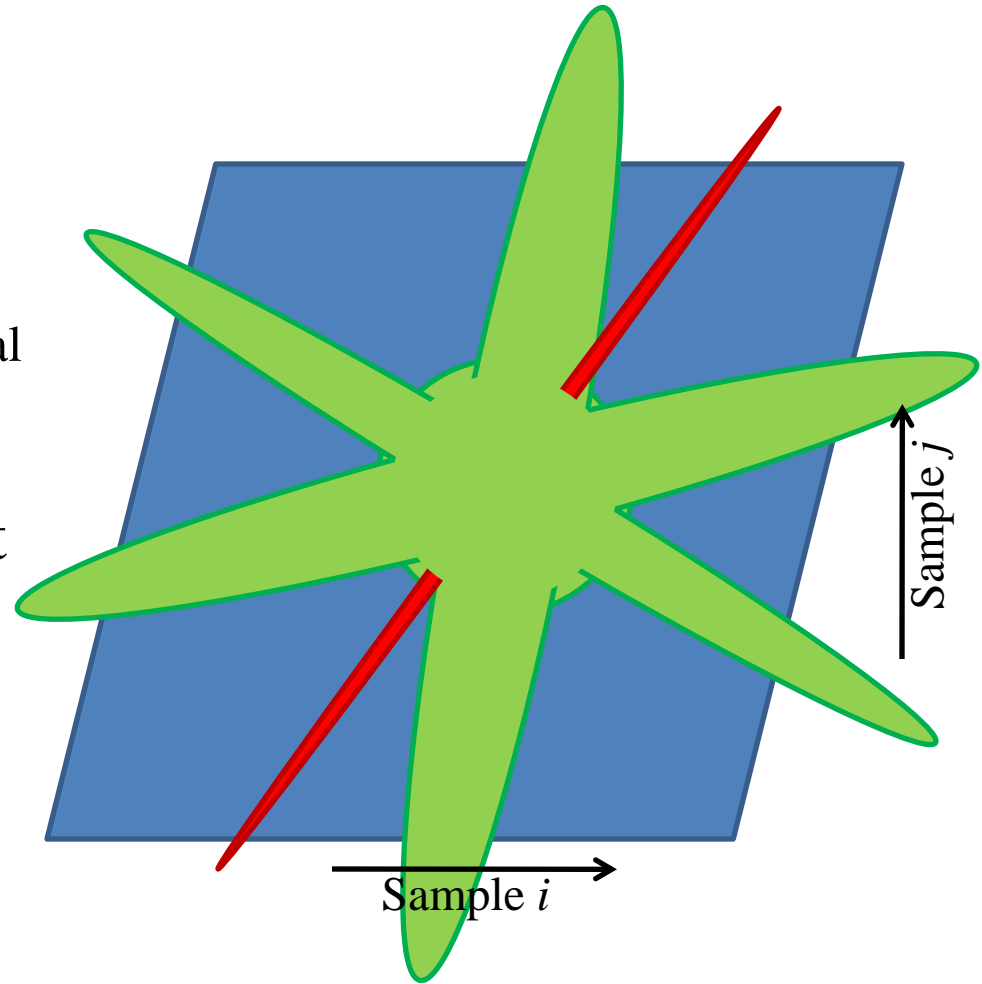
How to search for the unknown?

- **Inspirals** are a small fraction of **all** possible signals
- When we look for **all** possible signals, we are effectively limited to rejecting the instrumental **noise** hypothesis



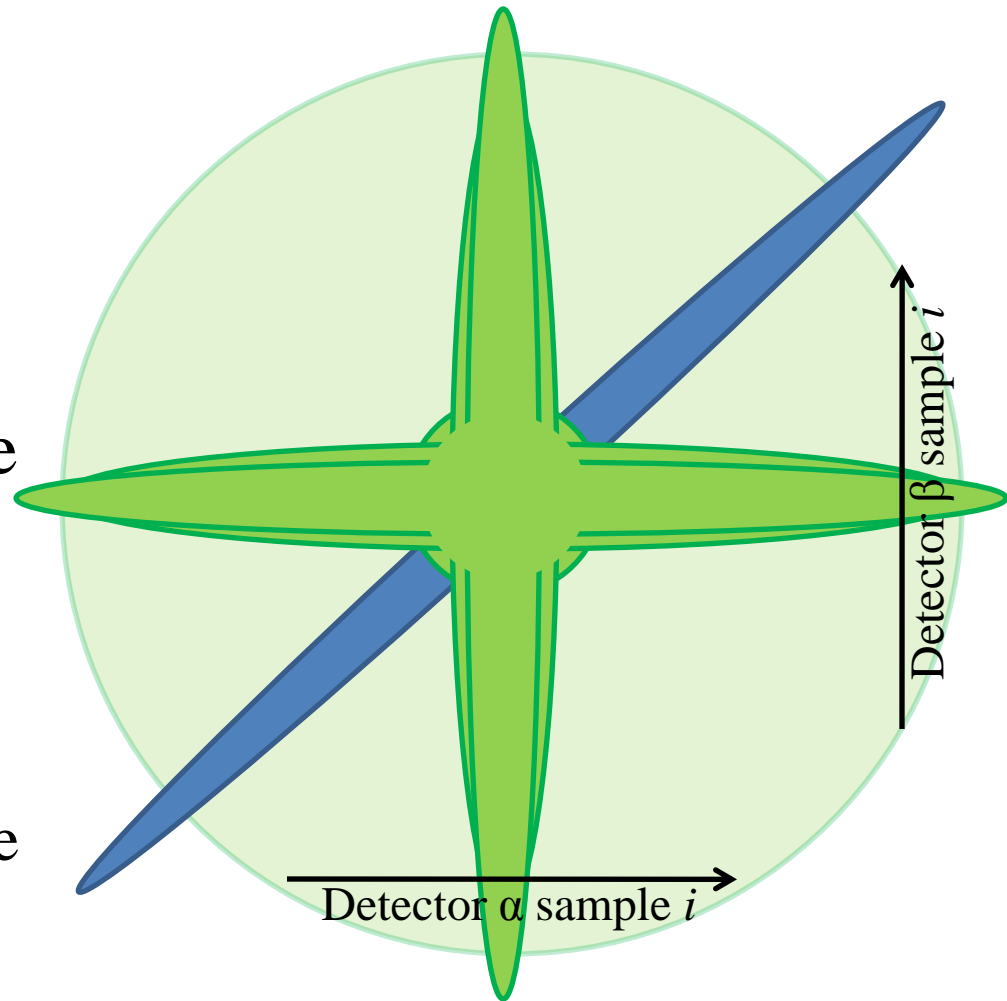
Incompletely understood detector

- LIGO experiences bursts of noise called “glitches”
 - Physical environment
 - Can veto things that occur at the same time as physical glitches
 - Unknown???
- The “glitchy” noise is not predictable enough that we can confidently pronounce **something** not a glitch
 - Rare for a glitch to look like an **inspiral**



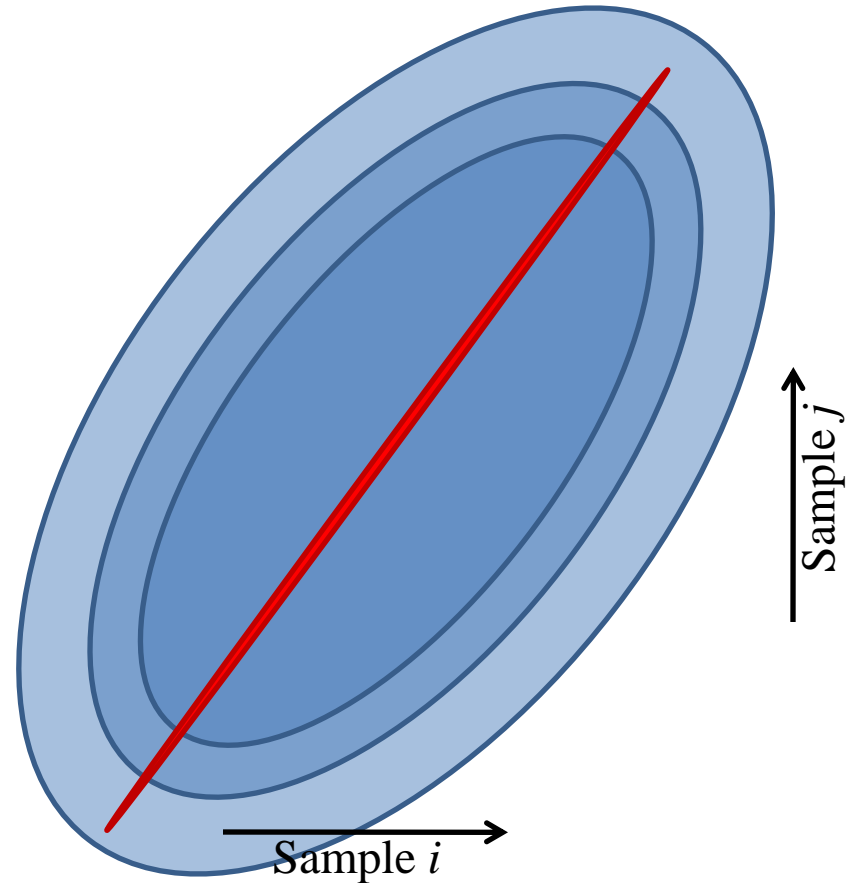
Intersite correlations

- If we consider three or more observatories, **signals** are constrained to lie on a subspace*
- Independent **glitches** are constrained to lie on other subspaces
 - Glitches in different detectors at the same time with the same shape are very, very rare



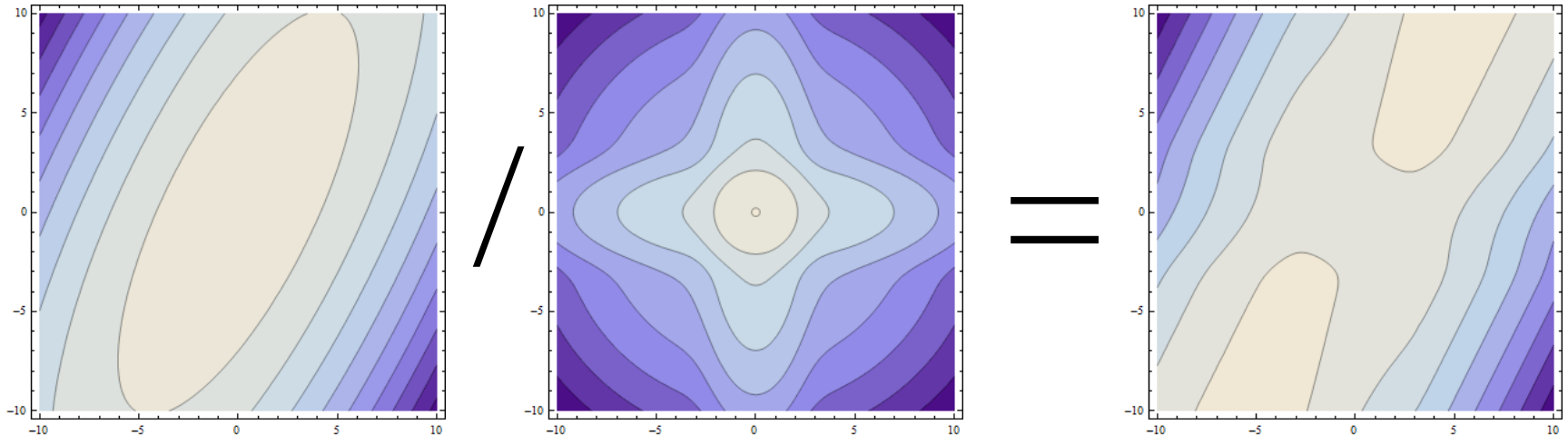
Semi-known search

- We do know something about “unknown” bursts
 - Uniform spatial distribution implies power law amplitude distribution
 - Simulations suggest morphology if not exact waveform
 - Et cetera



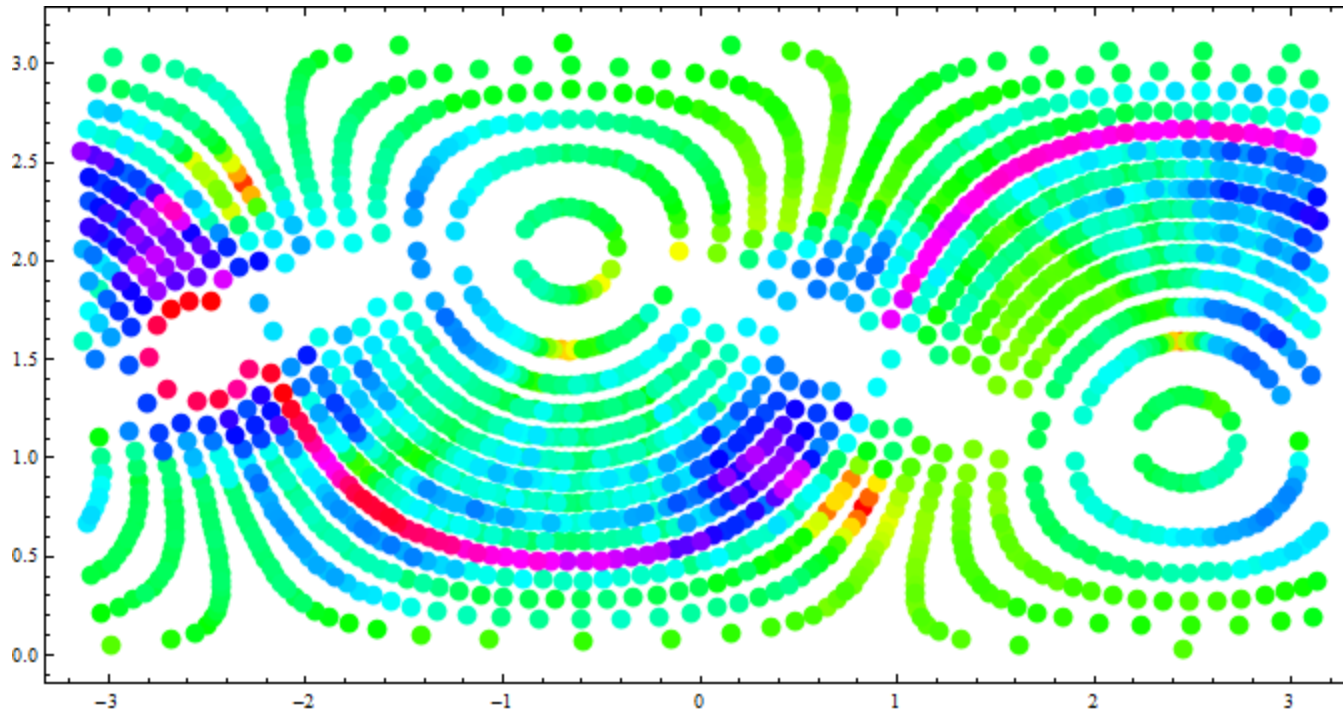
Bayesian synthesis

- Consider all data from all detectors
- Best bet signal model
 - Very general to express ignorance
- Best bet noise model
 - Long tails for glitches
- All desired features occur naturally
 - Efficient for quite flexible class of priors



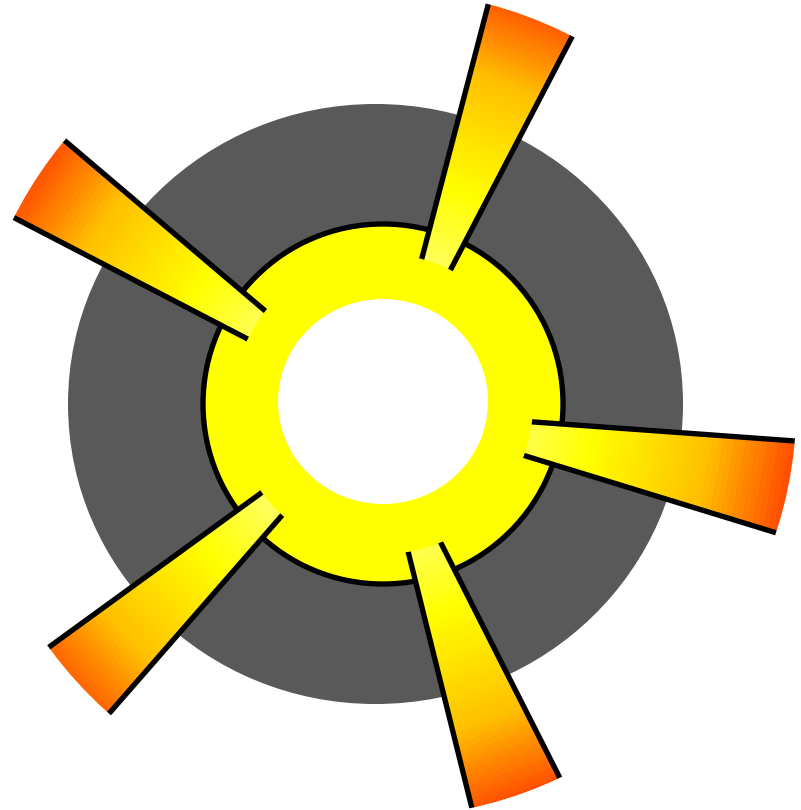
Directionality

- Have to do the search above for each direction, which brings its own fascinating problems...
 - Sky-map of the plausibility of a signal over directions



Conclusions

- Challenging!
 - Necessary if we want to be surprised by LIGO
- Information improves our sensitivity
 - Bayesian inference provides the framework to turn our knowledge into a search



$$h(t) = ???$$