### Development for Observation and Reduction of Radiation Pressure Noise

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## 1. Objectives and Scope

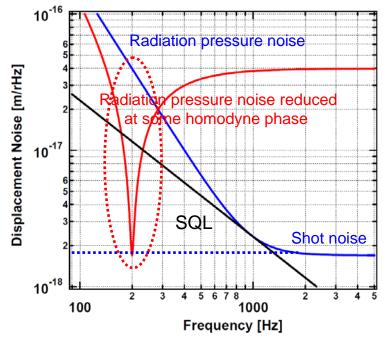
#### Objectives

- » Observation of radiation pressure noise
- » Reduction of radiation pressure noise
  - Measurement of ponderomotively squeezed vacuum fluctuations

at the best homodyne phase

#### Scope

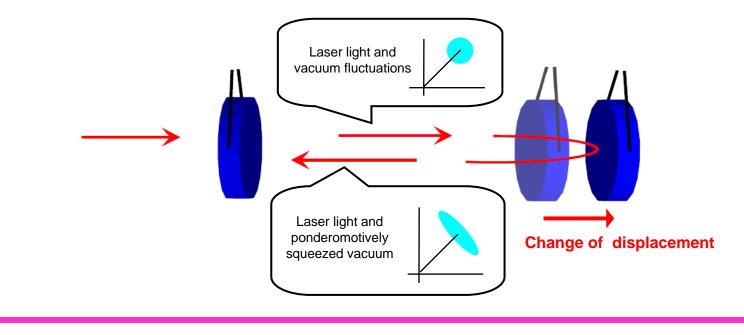
 Fabry-Perot cavities with tiny mirrors and high finesse



## Ponderomotive squeezing

#### Ponderomotive squeezing

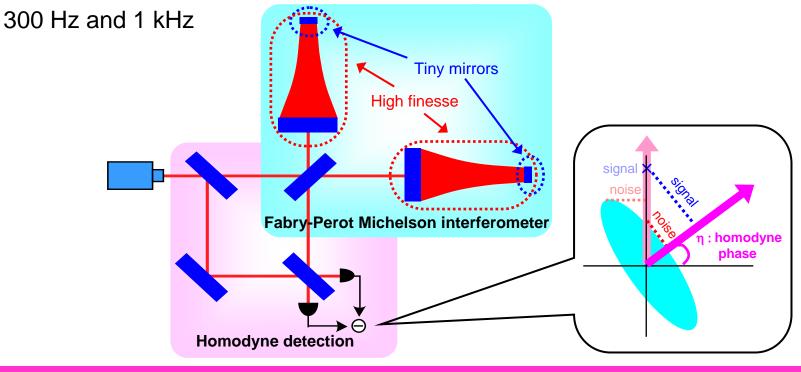
- » Amplitude fluctuation is correlated with phase fluctuation
  - Displacement changed by back-action of laser light and vacuum fluctuations
- » Squeezing with frequency dependence



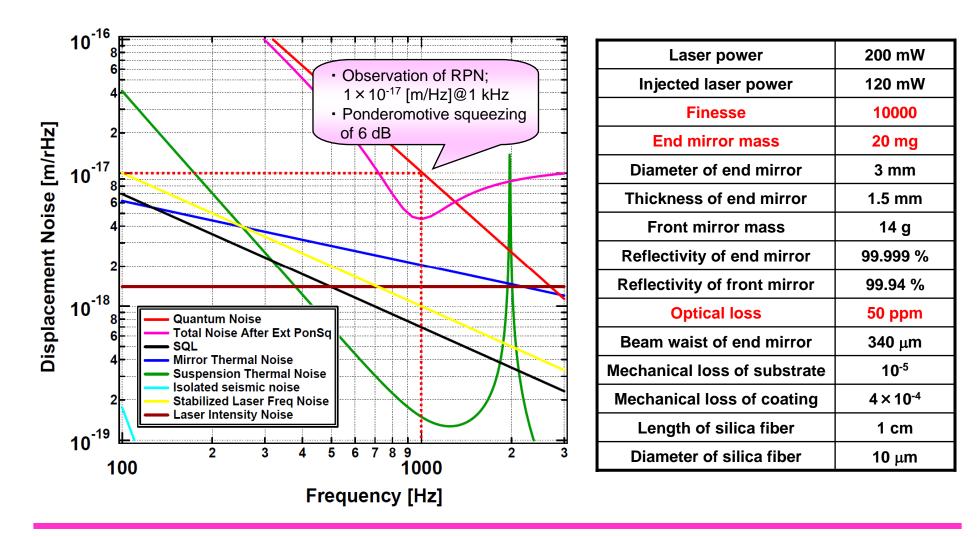
## **Conceptual design**

#### Conceptual design

- » Fabry-Perot Michelson interferometer with homodyne detection
  - Fabry-Perot cavities with tiny mirrors and high finesse
- » Observation and reduction of radiation pressure noise between



## Noise Budget



## 2. Experiment and Current status

#### Step1

- » Assembly of experimental setup
  - 20 mg mirror suspended by a 10  $\mu m$  silica fiber, Fixed front mirror

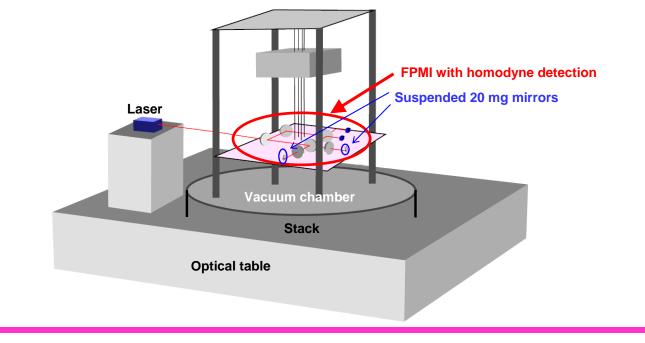
### Step2

- » Achievement of sensitivity of  $1 \times 10^{-18}$  [m/Hz] at 1 kHz
  - Replacement to suspended front mirrors
- Step3
  - » Observation of radiation pressure noise
    - Replacement to high reflectivity mirrors
- Step4
  - » Reduction of the radiation pressure noise

## **Experimental setup**

#### Experimental setup

- » Fabry-Perot Michelson interferometer with homodyne detection is suspended by a double pendulum
  - ✓ Assembly of a Fabry-Perot cavity
    - 20 mg mirror suspended by a  $10\mu m$  silica fiber



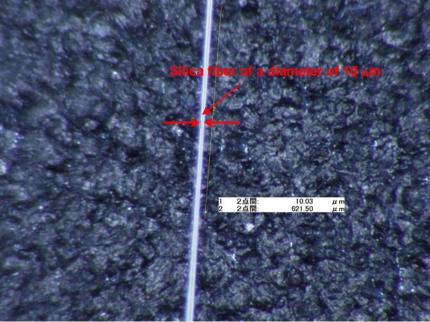
## 20 mg mirror and 10 $\mu m$ silica fiber

#### 20 mg mirror with a diameter of 3mm



#### Silica fiber of 10μm in diameter

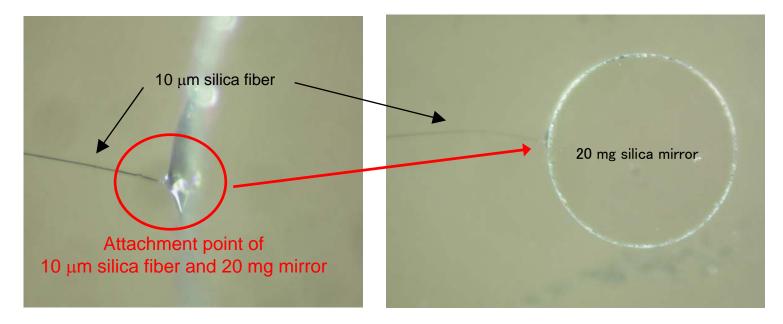
» For lower suspension thermal noise



## Silica fiber attached on 20 mg mirror

#### Silica fiber attached on 20 mg mirror

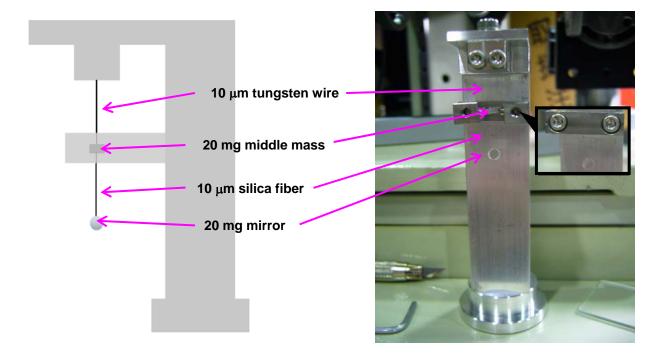
- » Glued with UV cured resin
  - Q-factor of suspension with the suspended mirror should be measured
- » We may investigate welding using CO2 laser



## 20 mg mirror suspension

#### Double suspension with 20 mg mirror

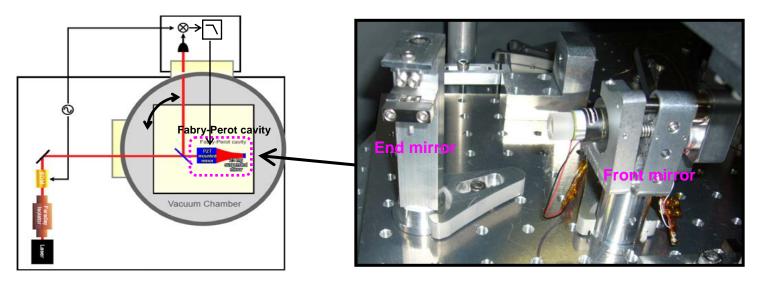
- » Middle mass of 20 mg is damped using eddy current damping
  - Very high Q-factor of suspension



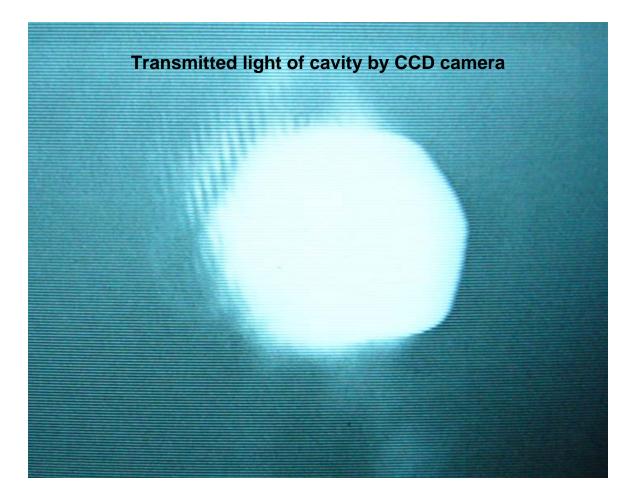
## Fabry-Perot cavity

#### Fabry-Perot cavity

- » Cavity length; 6cm, Finesse; 1400
  - Front mirror; fixed mirror mounted on PZT
  - End mirror; suspension with 20 mg mirror
    - Effective diameter of the mirror; 2 mm
- » Optical table is suspended by a double pendulum



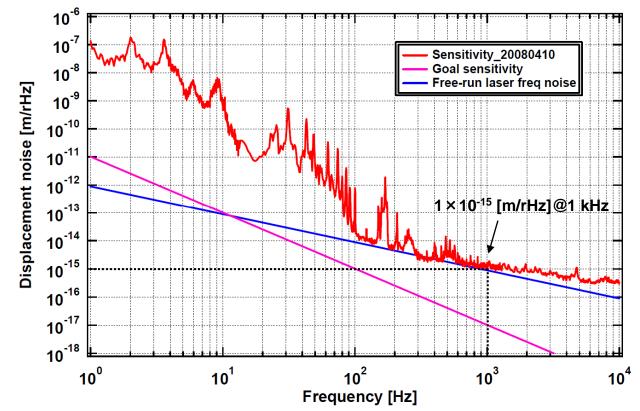
## Cavity is locked



## Current sensitivity

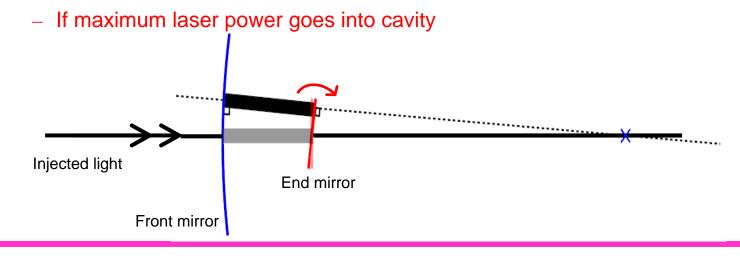
#### ■ Current sensitivity; 1 × 10<sup>-15</sup> [m/rHz]@1 kHz

- » A factor of 10 at 100 Hz, 100 at 1 kHz to reach the goal sensitivity
  - Laser power is a factor of 100 smaller than the maximum laser power because of angular anti-spring effect by radiation pressure



# 3. Angular anti-spring effect by radiation pressure

- Principle of angular anti-spring effect by radiation pressure
  - Resonant axis is changed from center of mirror because of mirror fluctuations
    - Tiny mirror is suspended by single silica fiber
    - Rotation is increased by radiation pressure  $\rightarrow$  angular anti-spring effect
  - » Torque by radiation pressure is a factor of 100 larger than restoring force

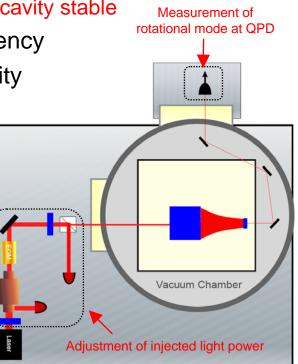


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## Measurement of angular anti-spring effect

#### Measurement of angular anti-spring effect

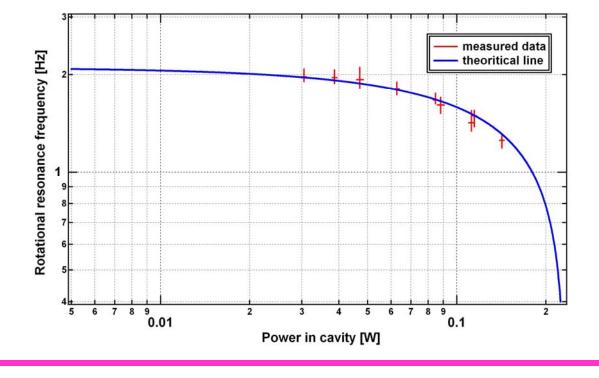
- » Radiation pressure force is smaller than restoring force
  - Injected laser power is adjusted to make the cavity stable
- » Measurement of rotational resonance frequency of suspended mirror at some power in a cavity
  - QPD is put on transmitted port
- Resonance frequency became smaller
  with higher laser power



## Preliminary result

#### Plot of rotational resonance frequency over power in cavity

- » Power calibration is fit
- » The angular instability is observed with the increased power

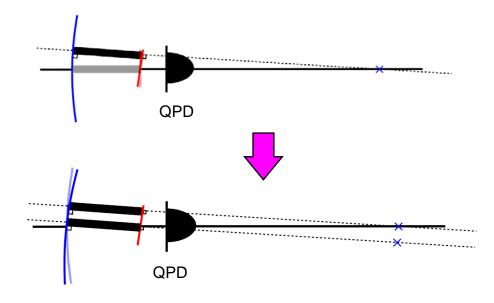


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# Solution for angular instability by radiation pressure

#### Solution for angular motions by radiation pressure

- 1. QPD is put at transmitted light port to read how much resonant axis is moved from the center
- 2. Center of curvature is moved to get back the axis to the center
- 3. Feedback the QPD signal to front mirror



## Summary and Future plans

#### Summary

- » Observation and reduction of radiation pressure noise between 100 Hz and 1 kHz
  - Radiation pressure noise level of  $1 \times 10^{-17}$  [m/rHz] at 1kHz
  - Ponderomotive squeezing of 6 dB is expected
- » Current status
  - 20 mg mirror is suspended by single 10  $\mu m$  silica fiber
  - Current sensitivity is  $1 \times 10^{-15}$  [m/rHz] at 1kHz
  - Measurement of anti-spring effect by radiation pressure
- Near future plans
  - » Fabry-Perot Michelson interferometer will be built
    - Reach a sensitivity of  $1 \times 10^{-18}$  [m/rHz] at 1kHz
      - Front mirror will be suspended

## The End