

# GR14 Conference Florence, Italy August, 1995

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LASER INTERFEROMETER GRAVITATIONAL-WAVE  
OBSERVATORY

## STATUS OF LIGO

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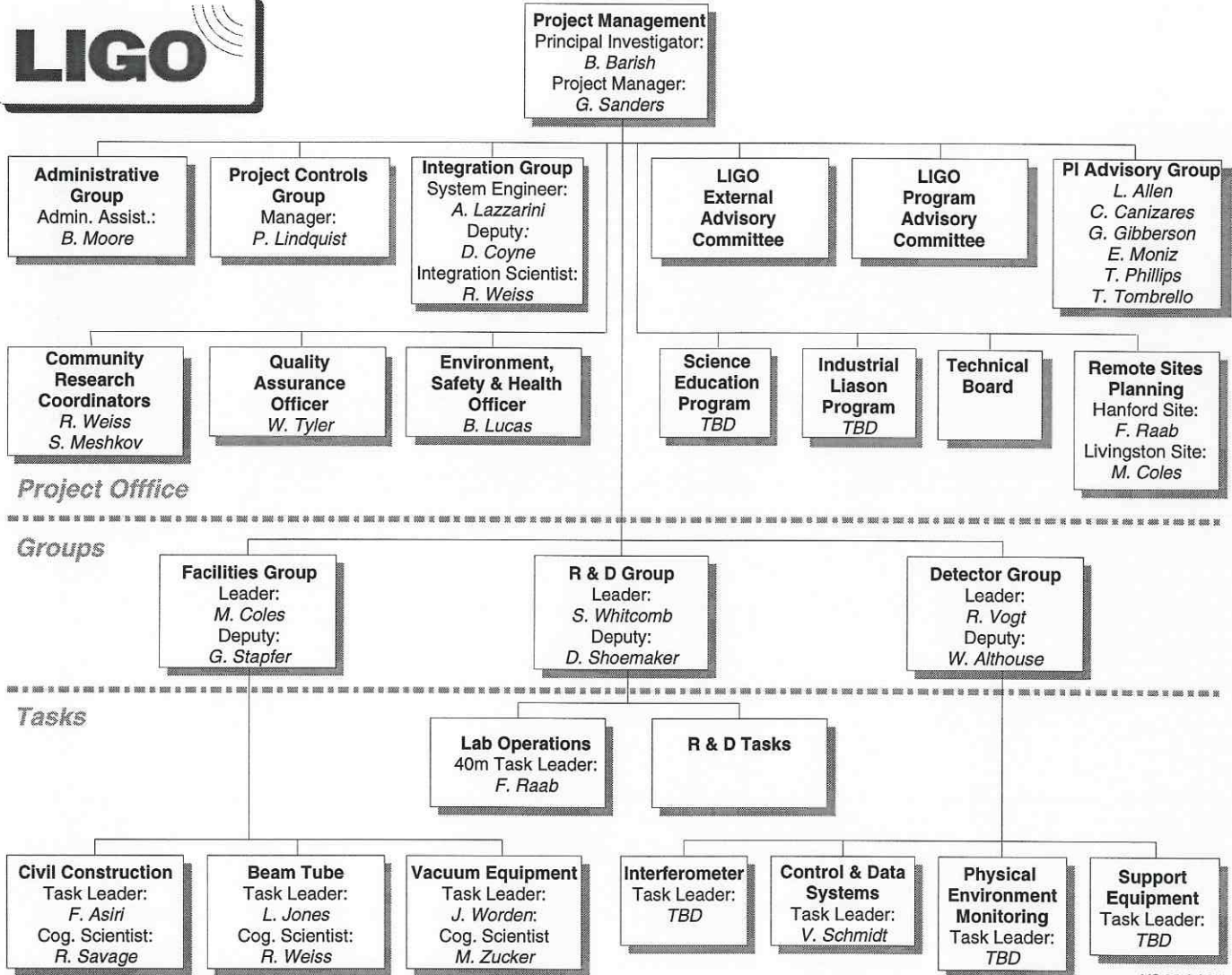
# OVERVIEW

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- LIGO DESIGN
- MAJOR LIGO SYSTEMS
  - CIVIL CONSTRUCTION
  - BEAM TUBE SYSTEMS
  - VACUUM EQUIPMENT
  - DETECTOR SYSTEMS
  - R&D
- LIGO RESEARCH COMMUNITY



# LIGO ORGANIZATIONAL STRUCTURE



VS 25 Jul 95



# LIGO DESIGN

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- OBSERVATORY CHARACTERISTICS

- TWO SITES: HANFORD, WASHINGTON & LIVINGSTON, LOUISIANA
- TWO 4KM INTERFEROMETERS & ONE 2 KM INTERFEROMETER (HANFORD)
- ARMS ORIENTED “PARALLEL” TO ONE ANOTHER
- COINCIDENT OBSERVATIONS AMONG ALL THREE INTERFEROMETERS
- INITIAL SENSITIVITY:  $h_{\text{rms}} \leq 10^{-21}$  WITHIN 100 HZ BAND CENTERED AT MAXIMUM SENSITIVITY
- OBSERVATORY EXTENSIBILITY:
  - EVENTUAL EXPANSION TO 9 INTERFEROMETERS
  - LIMITING SENSITIVITIES:
    - Naturally occurring gravity gradients (at lowest frequencies)
    - Scattered light phase noise (in the mid-frequency range)
    - Residual gas phase noise (at the highest high-frequencies)





# LIGO SITES

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## HANFORD, WASHINGTON

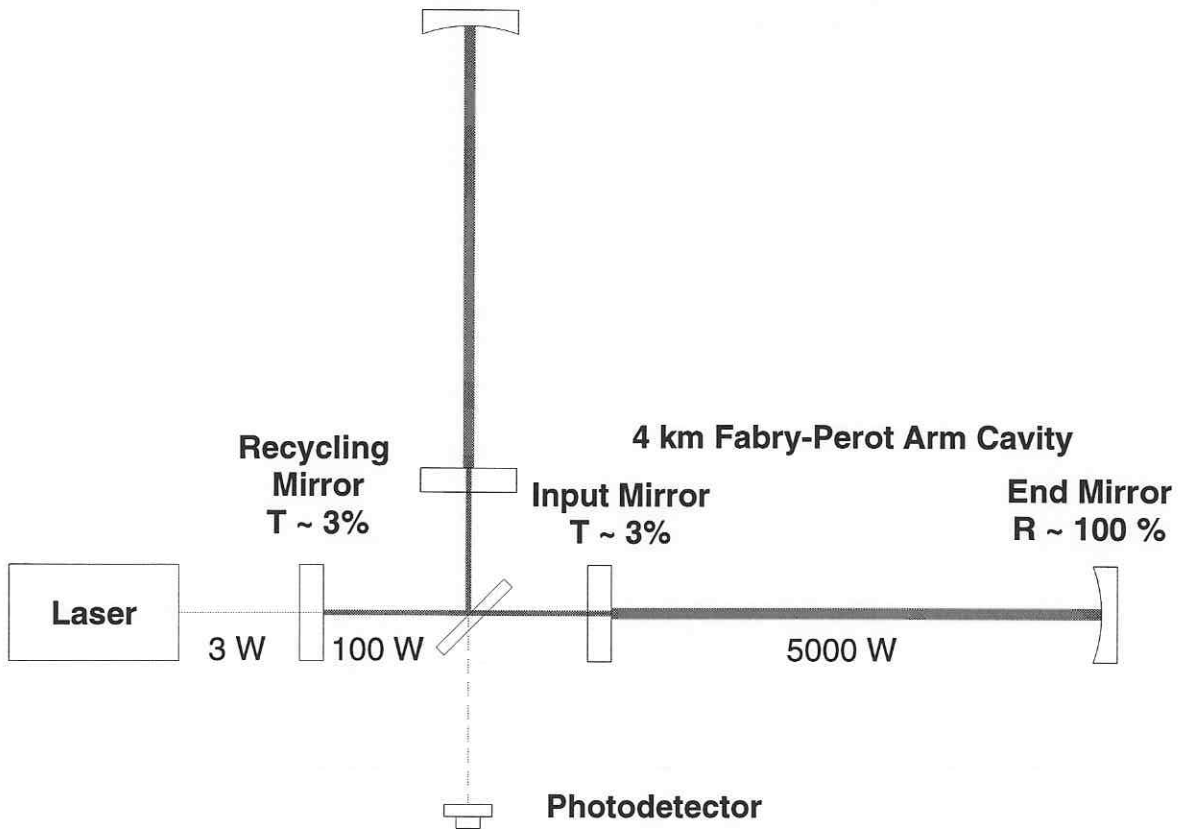
- LOCATED ON U.S. DOE RESERVATION
- TREELESS, SEMI-ARID HIGH DESERT
- APPROX. 25 KM FROM RICHLAND, WA (POPULATION :140,000)

## LIVINGSTON, LOUISIANA

- LOCATED IN FORESTED RURAL AREA
- MIXED FOREST; LOW-LYING; POOR DRAINAGE
- APPROX. 50 KM FROM BATON ROUGE, LA (POPULATION :450,000)



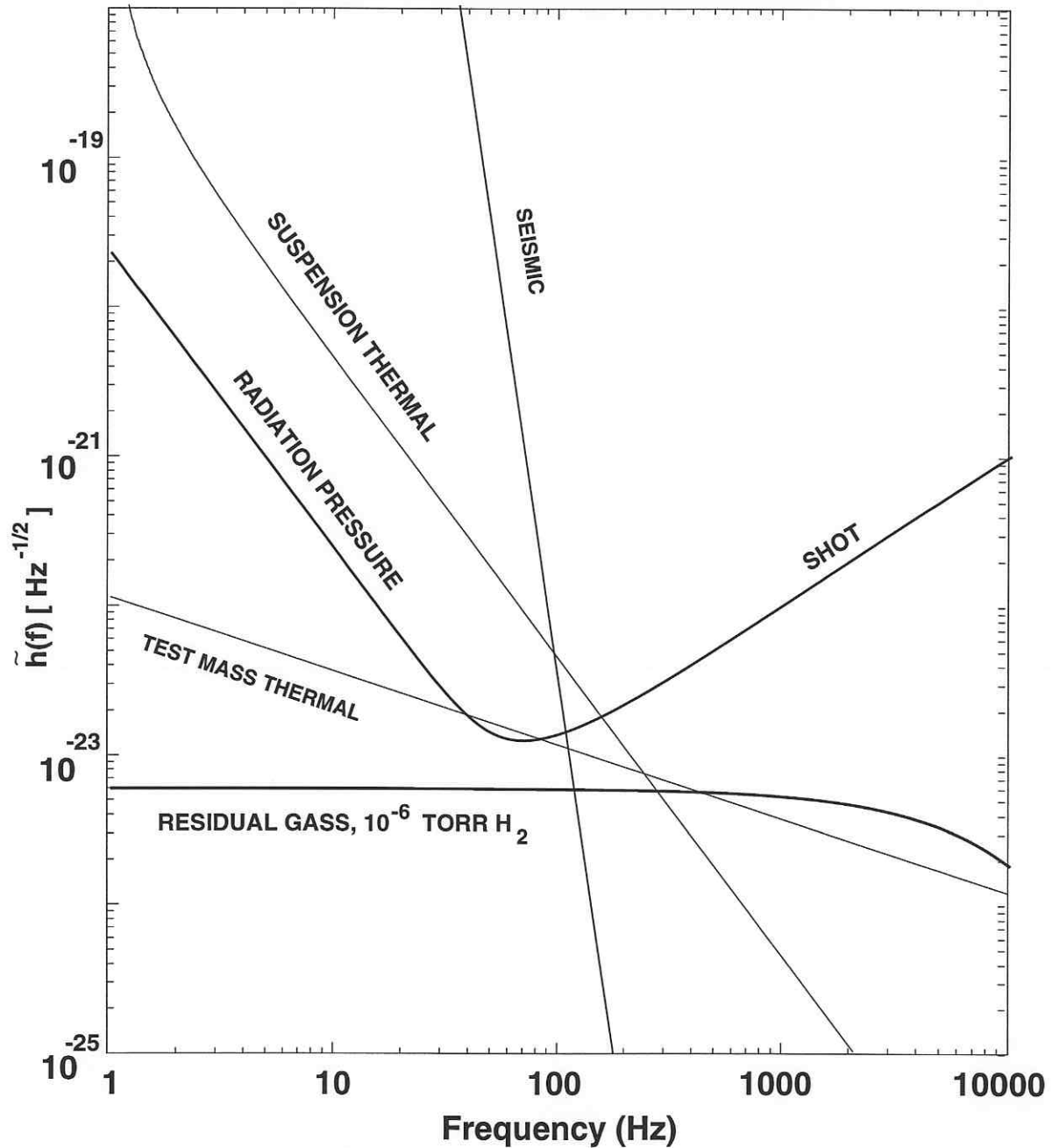
# INITIAL INTERFEROMETER CONFIGURATION



- FABRY-PEROT ARM CAVITIES
- MODEST INPUT POWER (2 - 3 w)
- INITIAL LASER:  $\text{Ar}^+$   $\lambda = 0.5145 \mu\text{m}$  POWER RECYCLING
- MODEST RECYCLING FACTOR (  $\gamma \sim 30\text{X}$  )
- MODEST CAVITY FINESSE (  $\eta \sim 50$  )

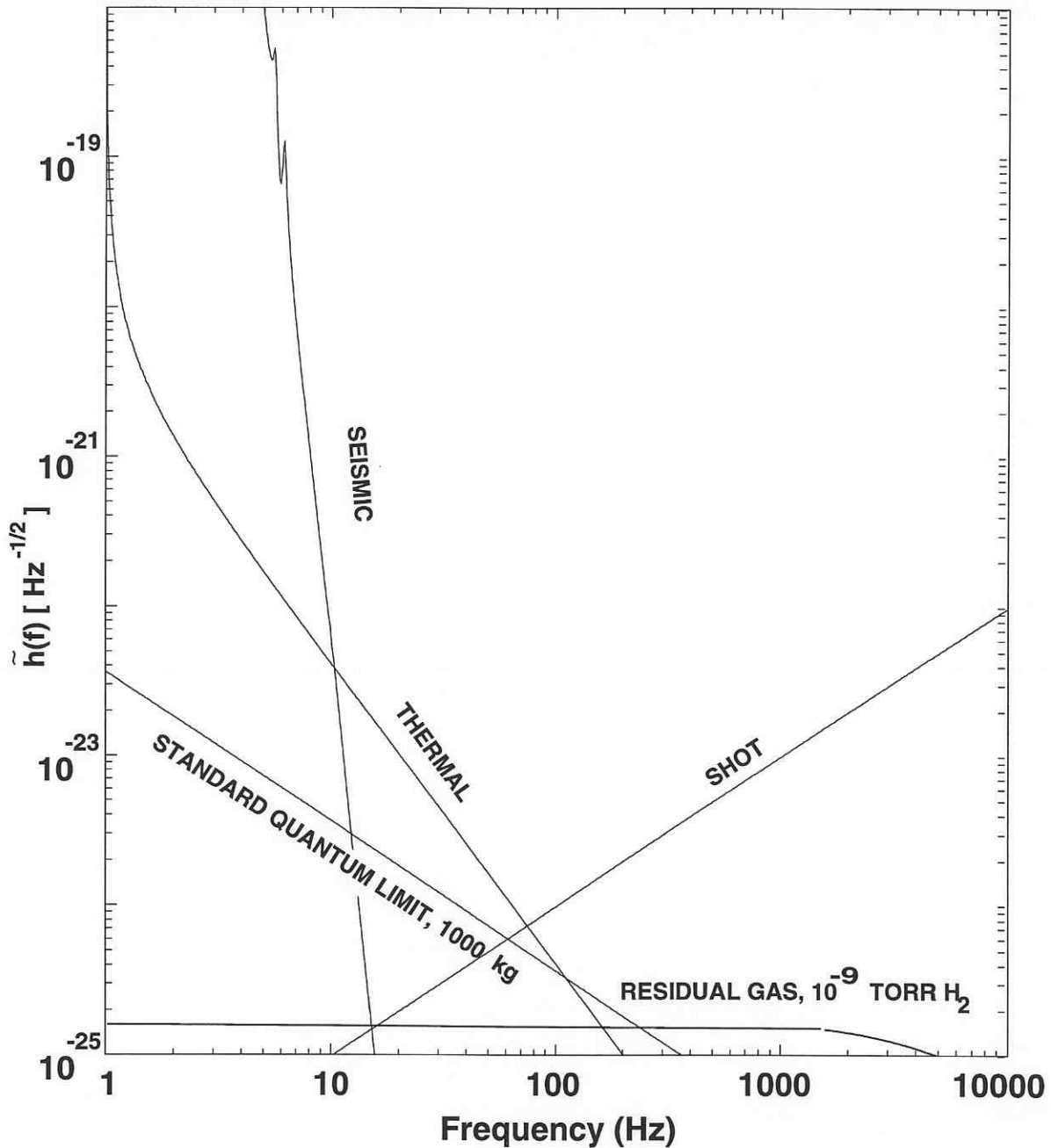
# INITIAL DESIGN PERFORMANCE GOAL

INITIAL INTERFEROMETER SENSITIVITY

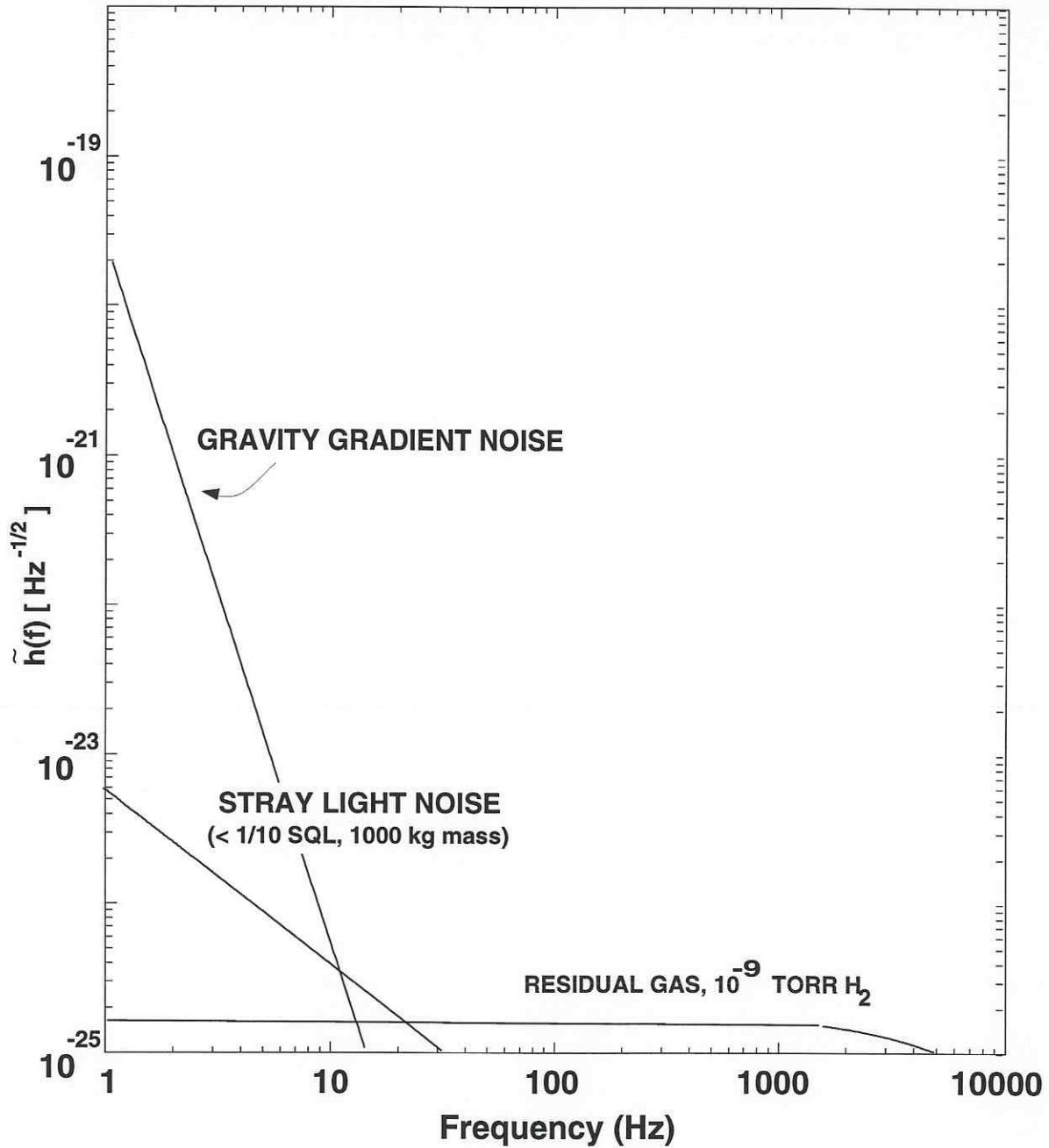


# ADVANCED DESIGN PERFORMANCE GOAL

## ADVANCED INTERFEROMETER, BROADBAND RECYCLING



# LIMITING PERFORMANCE DUE TO FACILITIES





# LIGO DESIGN (cont.)

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- OBSERVATORY EXTENSIBILITY (CONT.):
  - EXTREME CARE IN DESIGNING QUIET FACILITIES
    - site selection
    - design considerations
  - APERTURE & BAFFLING DESIGNED TO MINIMIZE SCATTERED STRAY LIGHT
  - VACUUM CHARACTERISTICS OF BEAM TUBES AND VACUUM EQUIPMENT DESIGNED TO PROVIDE ULTRA-HIGH VACUUM OPERATING CONDITIONS.
- OPERATIONAL CONSIDERATIONS
  - ASTROPHYSICAL RESEARCH
    - HIGH ON-LINE AVAILABILITY
    - MULTIPLE MODES OF OPERATION:
      - 3X ;  $T_{\text{online}} > 75\%$
      - 2X: (WA-LA);  $T_{\text{online}} > 85\%$
      - 1X:  $T_{\text{online}} > 90\%$
  - ALLOCATION OF OBSERVING TIME BY PI/PAC/NSF
  - TIME FOR DEVELOPMENT OF IMPROVED DETECTORS



# LIGO DESIGN (cont.)

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- DATA FORMATS
  - COMPATIBILITY WITH OTHER GRAVITATIONAL WAVE DETECTORS & PARTICLE DETECTORS
  - SAMPLE RATE AND PRECISION COMMENSURATE WITH SCIENTIFIC MISSIONS
- ENABLING RESEARCH & FACILITIES
  - LONG-TERM DEVELOPMENT OF ADVANCED DETECTORS
  - UNIVERSITY-BASED INTERFEROMETER FACILITIES
  - LIGO RESEARCH COMMUNITY



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# MAJOR LIGO SYSTEMS

## CIVIL CONSTRUCTION



# CIVIL CONSTRUCTION

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- CHARACTERISTICS

- INFRASTRUCTURE

- LARGE AND CLEAN LABORATORY (CLEAN CONSTRUCTION PRACTICES; HIGH BAY)
    - BEAM TUBE ENCLOSURES TO PROTECT THIN-WALLED TUBE
    - OFFICE SPACE/LABORATORY SPACE TO ACCOMMODATE STAFF AND VISITORS
    - ROADS, FOUNDATIONS, STRUCTURES, ETC.

- REQUIREMENTS

- REMOTE SITES WITH SEISMIC STABILITY -- THICK SLAB
    - FEW INTERNAL NOISE SOURCES, EMI/EMF BACKGROUNDS, ETC.
    - CLEANLINESS



# CIVIL CONSTRUCTION

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- STATUS (cont.)
  - BOTH SITES UNDER DEVELOPMENT & IMPROVEMENT;
  - WASHINGTON SITE
    - GRADED TO FINAL TOPOGRAPHY
    - NOW AWAITING SETTLEMENT BEFORE PROCEEDING WITH BUILDING CONSTRUCTION
  - LOUISIANA SITE
    - LOGGED AND CLEARED
    - GROUND BREAKING TOOK PLACE 7 JULY
  - INITIAL CONCEPTUAL DESIGN OF FACILITIES DEVELOPED BY LIGO





# CIVIL CONSTRUCTION

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- **STATUS (cont.)**

- **ARCHITECTURAL & ENGINEERING FIRM COMPLETING THE CONCEPTUAL DESIGN AND DEFINING REQUIREMENTS**
  - RALPH M. PARSONS CO.
  - BEAM TUBE ENCLOSURE CONCEPT COMPLETE AND BEING DESIGNED
  - FACILITY CONCEPTUAL DESIGN AT 100% COMPLETENESS (6/30/95)
  - FULL DESIGN REVIEW IN 11/95
- **A&E WILL OVERSEE CONSTRUCTION OF FACILITIES AT BOTH SITES**



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# MAJOR LIGO SYSTEMS

## BEAM TUBES



# BEAM TUBES

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- CHARACTERISTICS

- MODULE LENGTHS OF 2 KM
- TUBE DIAMETER - 1.22 M (1 M CLEAR APERTURE AFTER BAFFLING)
- BAFFLING TO REDUCE SCATTERED LIGHT PHASE NOISE
- VERY LOW ALLOWED AIR LEAKAGE:  
 $F < 10^{-9}$  ATM CC/S, HE
- VERY LOW OUTGASSING:
  - $P_{\text{ADVANCED}} < 10^{-9}$  TORR (ALL RESIDUALS);
  - $J[\text{H}_2]: < 10^{-13}$  TORR LITERS/CM<sup>2</sup>/S  
 $J[\text{H}_2\text{O}]: < 10^{-15}$  TORR LITERS/CM<sup>2</sup>/S
  - PARTIAL PRESSURES FOR CO + CO<sub>2</sub> + H<sub>2</sub>N C<sub>N</sub> + ... MUST BE EVEN LOWER
- QUALITY CONTROL AND CLEANLINESS MUST BE PURSUED DILIGENTLY THROUGHOUT FABRICATION AND INTEGRATION PROCESS

# BEAM TUBES

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- STATUS

- DESIGN & QUALIFICATION PHASE AWARDED TO CHICAGO BRIDGE & IRON IN 8/93
- FINAL DESIGN REVIEW COMPLETED 4/94
- PROTOTYPE FABRICATION - COMPLETE 5/95
  - QUALIFICATION TEST 1/95 - 3/95
  - DEMONSTRATED SUCCESSFUL CLEANING PROCEDURE
  - DEMONSTRATED SUCCESSFUL LEAK TEST PROCEDURE
    - SPIRAL WELDS
    - JOINT WELDS
    - BELLOWS EXPANSION JOINT
    - $F < 10^{-11}$  ATM CC/S
  - DEMONSTRATED SUCCESSFUL BAKEOUT PROCEDURE
    - RESIDUAL OUTGASSING RATES FOR H<sub>2</sub>, H<sub>2</sub>O BELOW REQUIRED LEVELS:
      - H<sub>2</sub>:  $< 10^{-13}$  TORR LITERS/CM<sup>2</sup>/S
      - H<sub>2</sub>O:  $< 10^{-15}$  TORR LITERS/CM<sup>2</sup>/S
    - RESIDUAL OUTGASSING RATES FOR CO<sub>2</sub>, H<sub>2</sub>N<sub>2</sub>C<sub>N</sub> BELOW MINIMUM ACCEPTABLE RATES;

# BEAM TUBES

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- PRODUCTION OF 16 KM TUBES FOR LIGO
  - QUALIFICATION TEST REVIEW HELD 4/95
  - REVIEW BOARD ENDORSED ALL ASPECTS OF TESTS DEMONSTRATING READINESS FOR PRODUCTION
  - RFP FOR PRODUCTION PHASE ISSUED 1 JUNE 1995
  - REVIEWING PROPOSAL RECEIVED IN JULY
  - CONCURRENT RESEARCH INTO BAFFLING MATERIAL OPTICAL PROPERTIES AND BAFFLE MECHANICAL DESIGN





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# MAJOR LIGO SYSTEMS

## VACUUM EQUIPMENT



# VACUUM EQUIPMENT

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- CHARACTERISTICS

- THE LIGO VACUUM *SYSTEM* (VACUUM EQUIPMENT + BEAM TUBES) PROVIDES:
  - A CLEAR APERTURE FOR THE INTERFEROMETERS --VERY LARGE APERTURE GATE VALVES TO ISOLATE 1.22 M BEAM TUBES
  - A CLEAN ENVIRONMENT FOR THE PRECISION OPTICS.
  - A LOW PRESSURE IN ORDER TO MINIMIZE DIFFRACTION AND ACOUSTIC COUPLING.
  - EXTENSIVE CONTROLS AND MONITORING EQUIPMENT TO ENSURE SAFE OPERATION
  - LARGE PUMPING SPEEDS AND VOLUMES -- BEAM TUBE PUMPING SOLELY FROM 4KM ENDS
- THE LIGO VACUUM *SYSTEM* WILL BE THE WORLD'S LARGEST HIGH PERFORMANCE VACUUM SYSTEM WITH A PUMPED VOLUME OF ROUGHLY 20,000 M<sup>3</sup>
- MOSTLY CONVENTIONAL VACUUM HARDWARE

# VACUUM EQUIPMENT (cont.)

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- STATUS

- SPECIFICATIONS DEFINED -- SCIENCE REQUIREMENTS REVIEW 8/95
- RFP FOR DESIGN PHASE AND FABRICATION PHASE ISSUED 12/95
  - 3-MONTH EFFORT TO COMPLETE SYSTEM DESIGN
- TWO CONTRACTORS SELECTED FOR DESIGN PHASE
  - CHICAGO BRIDGE & IRON
  - PROCESS SYSTEMS INTERNATIONAL, INC.
- DESIGN STUDY FROM 4/1/95 - 6/28/95
- PDRs HELD AT CALTECH 26, 27 JUNE 1995
- PROCESS SYSTEMS INTERNATIONAL, INC. SELECTED LAST WEEK



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# MAJOR LIGO SYSTEMS

## DETECTOR



# DETECTOR SYSTEM

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- **CHARACTERISTICS**

- **BASELINE CONFIGURATION EMPLOYS PROVEN TECHNIQUES**
  - 40M AND 5M SCALE SYSTEMS (AT CIT & MIT)
  - OPTICS EXPERIMENTS ON RECOMBINATION AND RECYCLING BEING PERFORMED IN R&D
  - PRECISION ENGINEERING
- **COLLABORATIVE STUDIES WITH INDUSTRY TO DEVELOP MIRROR COATING AND POLISHING TECHNOLOGIES**
- **BUILT-IN DESIGN FLEXIBILITY FOR LATER IMPROVEMENTS AS R&D RESULTS BECOME AVAILABLE**
  - DUAL WAVELENGTH COATINGS TO PERMIT CHANGE FROM AR<sup>+</sup> TO Nd:YAG LASER TECHNOLOGY
- **PASSIVE & ACTIVE SEISMIC ISOLATION**
- **ADVANCED CONTROL & DATA SYSTEM USING STATE-OF-THE-ART LAN/WAN ARCHITECTURES**





# DETECTOR SYSTEM (cont.)

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- STATUS

- DETECTOR DESIGN UNDER WAY
- KEY INITIAL DESIGN REVIEWS HELD:
  - PRESTABILIZED LASER SUBSYSTEM
    - DESIGN READINESS REVIEW IN 2/95
    - PRELIMINARY DESIGN READINESS REVIEW 6/95
    - PROTOTYPE SYSTEM BEING READIED FOR 40 m INTERFEROMETER INTEGRATION
  - ALIGNMENT CONTROL SYSTEM DESIGN READINESS REVIEW HELD 5/95
  - SUSPENSION SUBSYSTEM DESIGN READINESS REVIEW HELD 6/95
  - CORE OPTICS COMPONENTS DESIGN READINESS REVIEW TO BE HELD IN 7/95
  - MODE-CLEANER PROTOTYPE BEING READIED FOR INSERTION INTO 40 m SYSTEM
  - ACTIVE SEISMIC ISOLATION (STACIS SYSTEM FROM BARRY CONTROLS, INC.) SUCCESSFULLY DEMONSTRATED AT MIT



# DETECTOR SYSTEM (cont.)

## PRESTABILIZED LASER (PSL)

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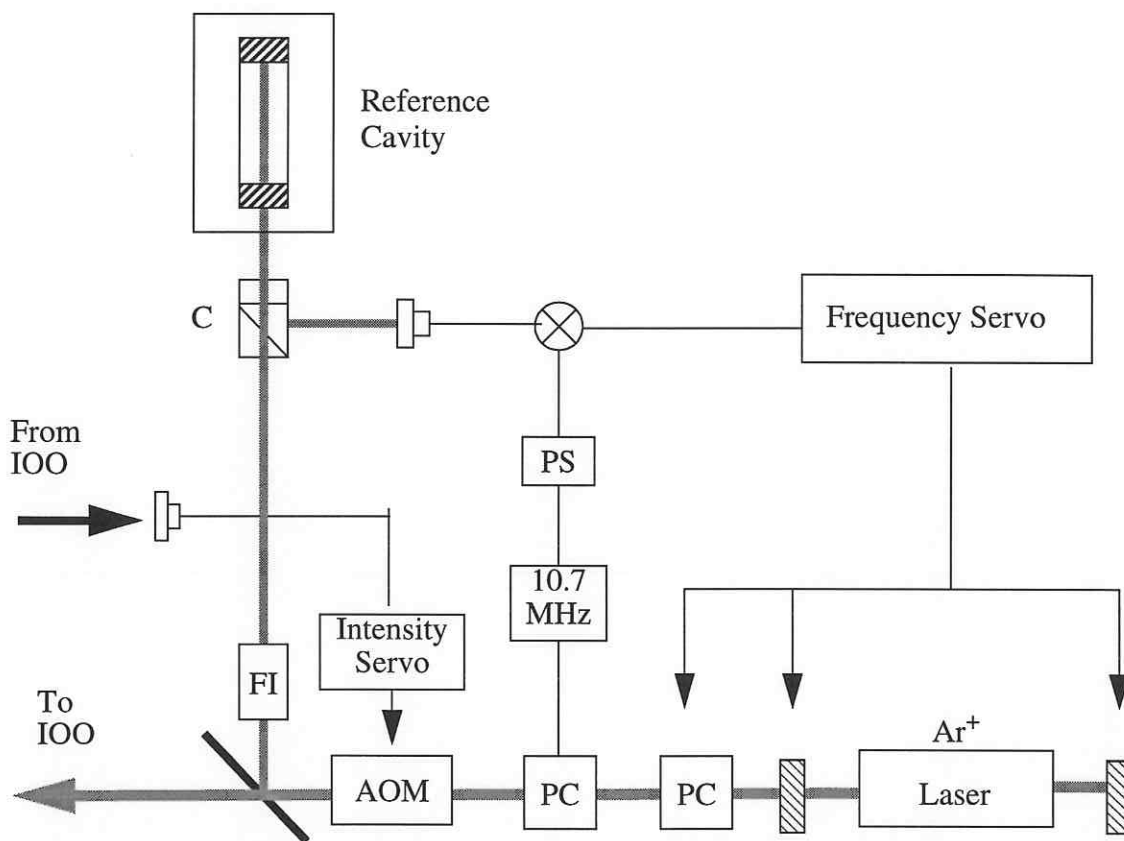
- CHARACTERISTICS

- PROVIDES SINGLE-MODE, SINGLE-FREQUENCY BEAM > 4 W TO INPUT OPTICS SUBSYSTEM
- SERVO LOOPS TOP STABILIZE LIGHT:
  - FREQUENCY STABILIZATION TO  $< 1 \text{ Hz}/\sqrt{\text{Hz}}$  @ 100 Hz (GAIN OF  $10^5$ )
  - INTENSITY (AMPLITUDE) FLUCTUATIONS  $< 10^{-7} \text{ W}/\sqrt{\text{Hz}}$  @ 100 Hz (GAIN OF  $10^3$ )
  - BEAM ANGULAR JITTER TO  $< 200 \mu\text{rad}/\sqrt{\text{Hz}}$  @ 100 Hz
  - POLARIZATION FLUCTUATIONS IN THE BEAM TO  $< 10^{-6} 1/\sqrt{\text{Hz}}$  @ 100 Hz
- PSL IS BUILT AROUND A CUSTOMIZED COMMERCIALY AVAILABLE AR<sup>+</sup> LASER
  - LIGO IS PRESENTLY CONDUCTING A TRADE STUDY EVALUATING DESIRABILITY/FEASIBILITY OF ADOPTING SOLID STATE LASER TECHNOLOGY TO REPLACE PLASMA TUBE LASER [Nd:YAG 10 w @ 1.064  $\mu\text{m}$  or 5w @ 0.532  $\mu\text{m}$ ]

# DETECTOR SYSTEM (cont.)

## PRESTABILIZED LASER (PSL)

PC: Pockels Cell  
 AOM: Acousto Optical Modulator  
 FI: Faraday Isolator  
 C: Optical Circulator  
 PS: Phase Shifter



⊗ RF Mixer  
 ▨ PZT actuated mirror  
 □ Photodiode

— light beam  
 - - - electrical signal

# DETECTOR SYSTEM (cont.)

## PRESTABILIZED LASER (PSL)

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- STATUS

- FREQUENCY & INTENSITY SERVOS OPERATIONAL UNDER LIGO-TYPE COMPUTER CONTROL (CDS SYSTEM)
- PSL INTERFACE USES EPICS-BASED CONTROL SYSTEM:
  - LASER CONTROL
  - SERVO CONTROL
  - DATA MONITORING & ARCHIVAL
  - ALARM HANDLING
- PSL PROTOTYPE TESTING UNDERWAY (6/95 - 9/95)
  - NOISE PERFORMANCE
  - AVAILABILITY
  - ENVIRONMENTAL SENSITIVITY
    - SEISMIC NOISE
    - ACOUSTIC NOISE
    - THERMAL VARIATION





# DETECTOR SYSTEM (cont.) PRESTABILIZED LASER (PSL)

**MK2 PSL1 Control Panel**

**PS\_SUBSYSTEM1**

Mode	State
MANUAL	MANUAL
AUTO	AUTO

**FSS\_LCU**

LCU Mode	Search	Scan
MANUAL	OFF	High
AUTO	FULL	Auto-Scan
	ON	Bypass

Visibility Monitor: 0.0, 50.7, 100.0

**Laser\_LCU**

LCU Mode	Laser PS	Laser Power (Watts)	2.00	1.95
MANUAL	OFF	Laser Current (Amps)	24.6	41
AUTO	ON	Laser Voltage (Volts)		470

**PSE\_LCU**

LCU Mode	PSA Output	Reference PM	1.00
MANUAL	OFF	PSA Output	2.00
FULL	ON	Intensity Loss	

**Laser LCU**

Laser Current Monitor: 0 to 100 (Amps), Value: 40

Laser Voltage Monitor: 0 to 1000 (Volts), Value: 484

Laser Power Monitor: 0.00 to 10.00 (Watts), Value: 2.01

Laser Current Adjust: 0.0, 24.6, 70.0

Laser Power Adjust: 0.00, 2.00, 10.00

LCU Mode: MANUAL, AUTO

Purge Gas: OFF, ON

Cooling Water: OFF, ON

Control Mode: Current, Power

Laser PS: OFF, ON

Head Cover Intik.

Laser Key Switch

Tube Fill Status

Laser Low Water Flow

Laser High H2O Temp.

Regulator Fault

PS On/Off Request

Power Supply Status

Laser Emission

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# MAJOR LIGO SYSTEMS

## RESEARCH AND DEVELOPMENT



# RESEARCH & DEVELOPMENT

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- **OBJECTIVES**

- **PHASED TIME LAG BETWEEN R&D TASKS WHICH DEMONSTRATE ENABLING TECHNOLOGIES AND DETAILED TASKS FOR LIGO DETECTOR DESIGNS:**
  - THERMAL NOISE INVESTIGATIONS
  - SUSPENSIONS
  - SEISMIC STACKS
  - TEST MASSES
  - OPTICAL CONFIGURATION DEFINITION STUDIES
  - RECOMBINATION;
  - POWER RECYCLING;
  - PHASE NOISE STUDIES AT HIGH RECYCLING POWER
  - ALIGNMENT SENSING & CONTROL;
  - LENGTH SENSING & CONTROL;
  - ACTIVE/PASSIVE SEISMIC ISOLATION
- **VALIDATING DESIGN PROTOTYPES:**
  - TRIANGULAR MODE CLEANER
  - PRESTABILIZED LASER SUBSYSTEM
- **LONGER-TERM EFFORT TO ESTABLISH FOUNDATIONS FOR FUTURE INTERFEROMETER ADVANCES**



# RESEARCH & DEVELOPMENT

## THERMAL NOISE STUDIES

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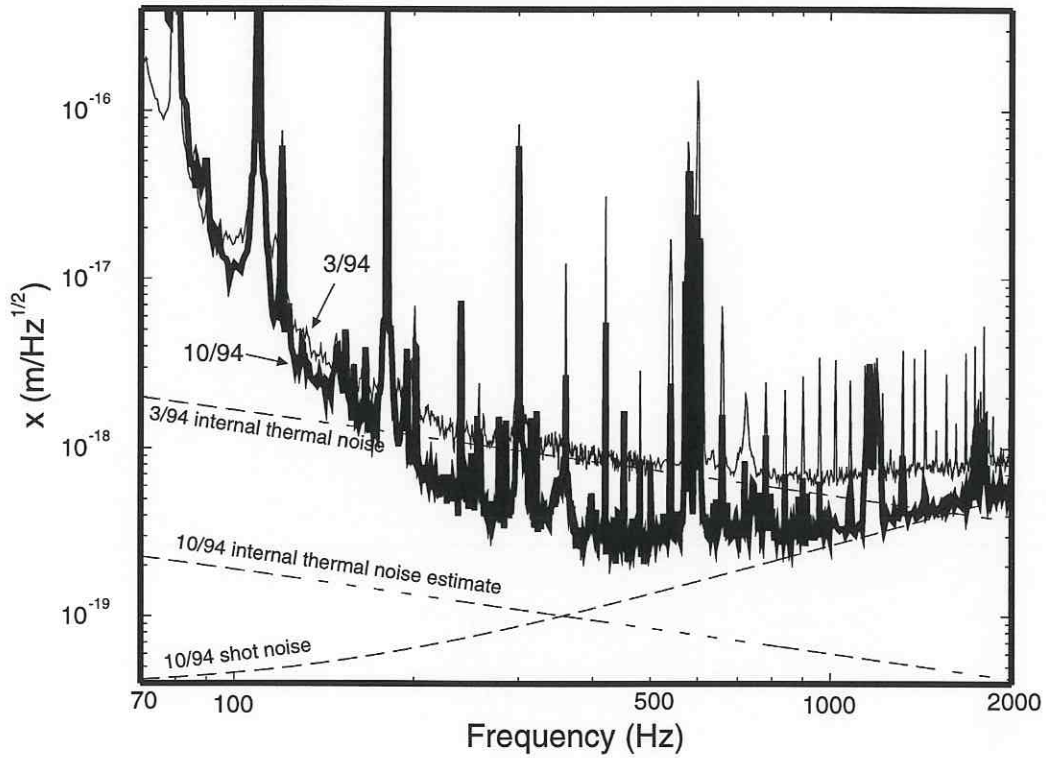
- THERMAL NOISE OF SUBSTRATE, PENDULUM SIGNIFICANT FOR LIGO
  - NO DIRECT EXPERIMENTAL OBSERVATIONS TO VERIFY MODELS
  - OPPORTUNITY TO OBSERVE IN 40M BY CHANGE OF MASSES
- THERMAL NOISE IN COMPOUND MASSES CAREFULLY CALCULATED
  - GOOD AGREEMENT WITH 40 M DATA
- NEW MONOLITHIC MASSES STUDIED SEPARATELY
  - ATTENTION TO COUPLING OF CONTROL/SENSING MAGNETS/FINS, SUSPENSION WIRES,...
  - INSTALLED INTO 40M FALL OF 1994
- LATEST (AND BEST) SPECTRUM TO DATE CONSISTENT WITH THERMAL NOISE MODEL



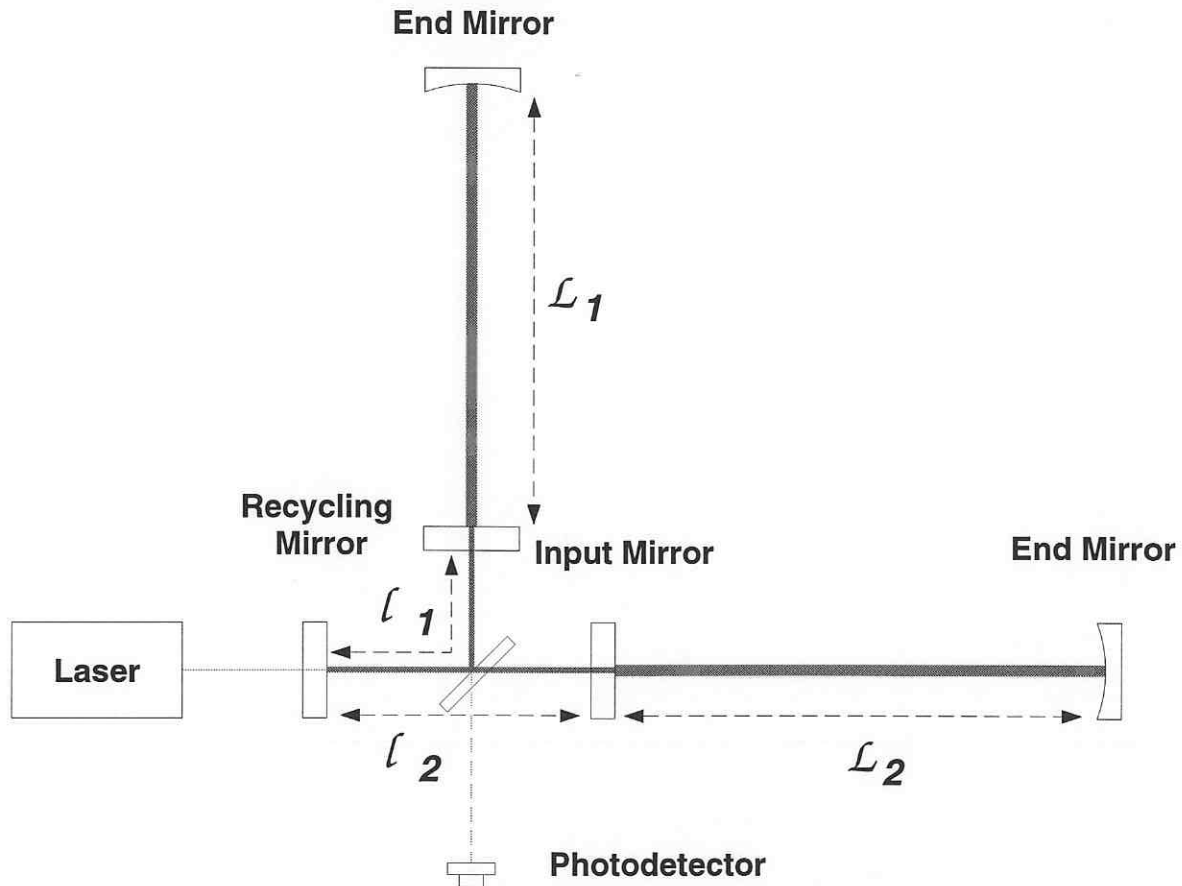
# RESEARCH & DEVELOPMENT

## THERMAL NOISE STUDIES

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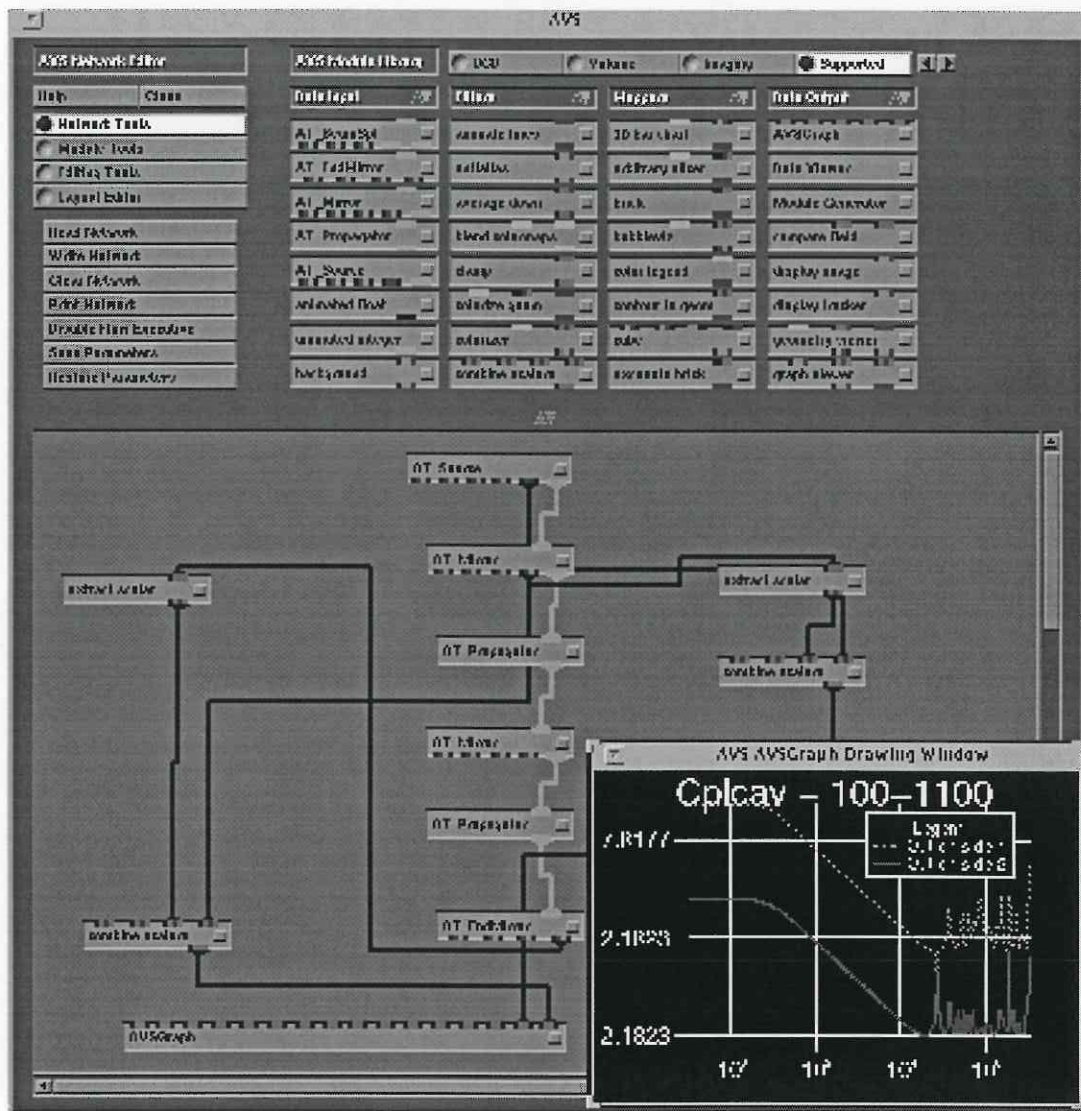
# RESEARCH & DEVELOPMENT OPTICAL CONFIGURATION STUDIES



- MUST CONTROL AT LEAST FOUR DEGREES OF FREEDOM:
  - PLACEMENT OF SIGNAL-SENSING ELEMENTS
  - PHOTODETECTORS AND MODULATORS
- CONCURRENT TRADE STUDY BEING CONDUCTED FOR OPTICAL CONFIGURATION
- MODIFYING THE 40M TO SUPPORT R&D STUDIES

# INTERFEROMETER CONFIGURATION DEFINITION

## MODELING





# RESEARCH & DEVELOPMENT

## OPTICAL CONFIGURATION

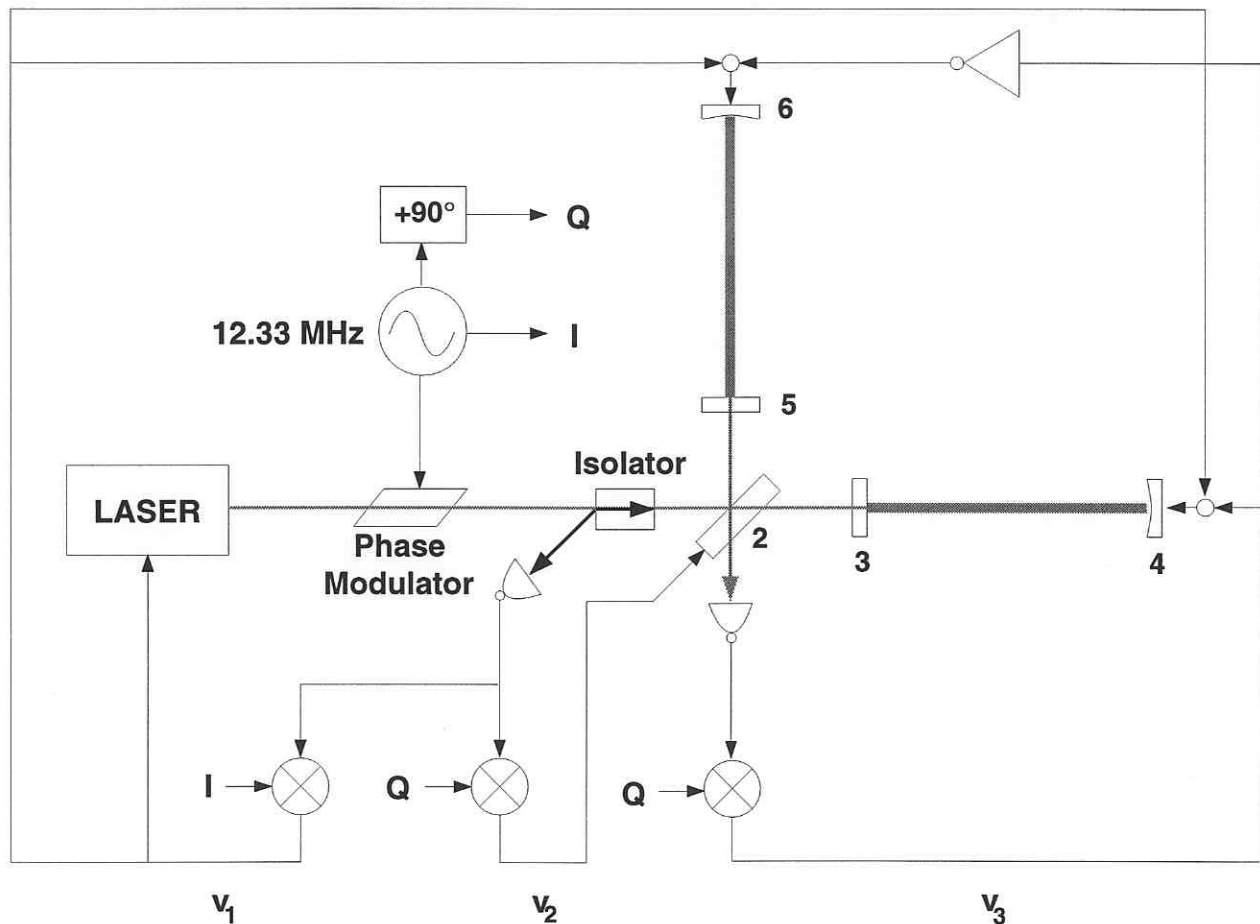
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- RECOMBINATION & POWER RECYCLING
  - MODELING OF LIGO WELL ADVANCED
  - EAGER TO HAVE EXPERIMENTAL VERIFICATION
    - LARGE & SMALL SIGNAL INTERFEROMETER DYNAMIC RESPONSE FUNCTIONS
    - IMPROVED COMMON-MODE NOISE REJECTION FOR DISPLACEMENT STUDIES
  - STATUS:
    - SERIES OF SMALL-SCALE EXPERIMENTS TO VALIDATE DIFFERENT ASPECTS OF MODELS COMPLETE (DOCTORAL THESES)
    - ARM ASYMMETRY INTRODUCED FOR MODULATION TECHNIQUE
    - SYSTEM HAS ACQUIRED LOCK AND ARMS HAVE BEEN RESONANT
    - PRESENTLY IN SHAKEDOWN TO REPEAT PREVIOUSLY ACHIEVED LEVEL OF NOISE PERFORMANCE
    - RECYCLING WILL FOLLOW ONCE RECOMBINED SYSTEM IS FULLY CHARACTERIZED



# RESEARCH & DEVELOPMENT

## OPTICAL CONFIGURATION



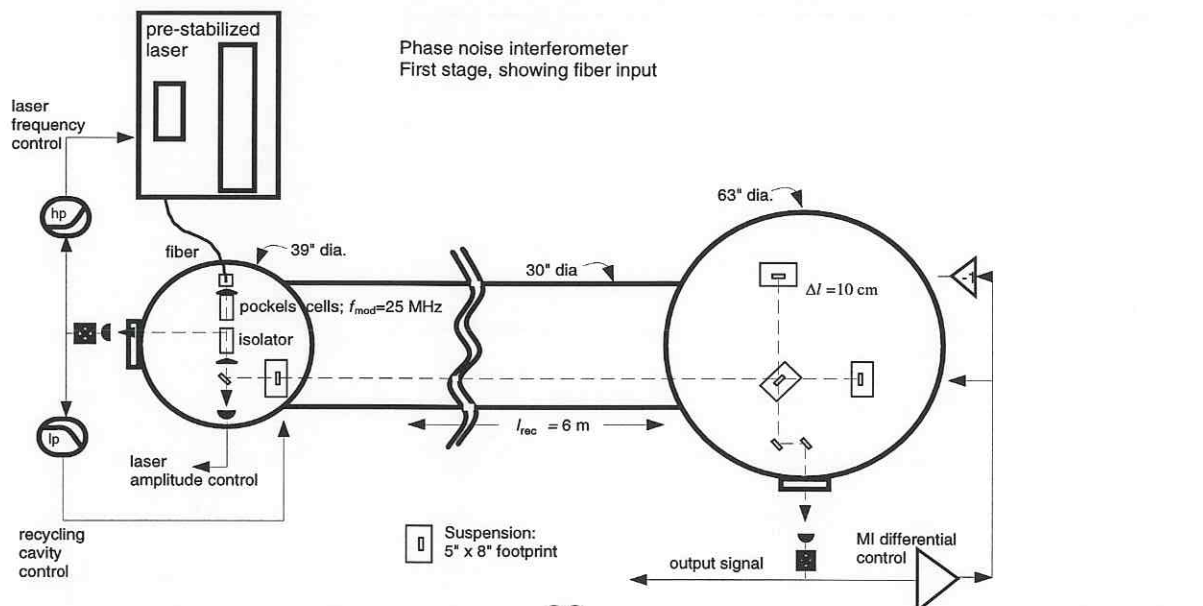
- RECOMBINED CONFIGURATION COMPLETE: FIRST LOCK ACQUISITION ACHIEVED MAY/JUNE
- PRESENTLY UNDERGOING SHAKEDOWN TO RECOVER SENSITIVITY ACHIEVED EARLIER

# RESEARCH & DEVELOPMENT

## PHASE NOISE RESEARCH

### • OBJECTIVES:

- DISPLACEMENT SENSITIVITY STUDY UNDER WAY WITH 40M FACILITY
- LIGO REQUIRES HIGH PHASE SENSITIVITY ( $< 10^{-10}$  RAD/ $\sqrt{\text{Hz}}$ )
  - BEST TO DATE ACHIEVED WITH 1W OF OPTICAL POWER ( $< 10^{-9}$  RAD/ $\sqrt{\text{Hz}}$  @ MPI/GARCHING)
  - LIGO REQUIRES ROUGHLY 75W
- PHASE NOISE STUDIES WILL DEMONSTRATE REQUIRED SENSITIVITY
  - CONFIRMS UNDERSTANDING OF SHOT NOISE
  - NEED TO SOLVE TECHNICAL PROBLEMS (PHOTODETECTOR IRRADIANCES)
  - NON-RECYCLED MICHELSON CONFIGURATION INITIALLY TO BE USED (1 W) ; RECYCLING TO BE ADDED (15 W) ; MODE CLEANER WILL EVENTUALLY BOOST POWER TO 70 W.





# RESEARCH & DEVELOPMENT

## PHASE NOISE RESEARCH

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- STATUS

- SYSTEM HAS BEEN INTEGRATED AT MIT
- INCLUDES ACTIVE ISOLATION SYSTEM FROM BARRY CONTROLS, INC. (STACIS) BEING EVALUATED FOR EVENTUAL INCLUSION INTO LIGO
- IN-AIR INTERFERENCE HAS BEEN OBSERVED
- BEING READIED FOR EVACUATION AND SHAKE-DOWN TESTING



# RESEARCH & DEVELOPMENT

## ALIGNMENT SENSING & CONTROL

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- ALIGNMENT REQUIREMENTS FOR LIGO

- INTERFEROMETER USES  $TEM_{00}$  LIGHT
- REQUIRE GOOD OVERLAP BETWEEN INPUT LIGHT, OPTIC AXIS
- HIGH CONTRAST IN MICHELSON INTERFERENCE
- LEADS TO REQUIREMENT OF  $<10^{-7}$  RADIANS ALLOWED MIS-ALIGNMENT
- OR...  $<1\%$  (POWER) EXCITATION OF  $TEM_{01}$  MODE

- ALIGNMENT SENSING

- PHASE MODULATION OF THE LIGHT
- SIMILAR TECHNIQUES TO THE GW READOUT SYSTEM
- SPATIALLY SENSITIVE DETECTORS, PLACED AT CRITICAL POINTS
- SENSES PHASE SHEAR IN THE BEAMS IN THE INTERFEROMETER
- SUBTLE RELATIONSHIP BETWEEN SENSORS AND MIRROR ANGLES



# RESEARCH & DEVELOPMENT

## PHASE NOISE RESEARCH

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- STATUS

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# RESEARCH & DEVELOPMENT

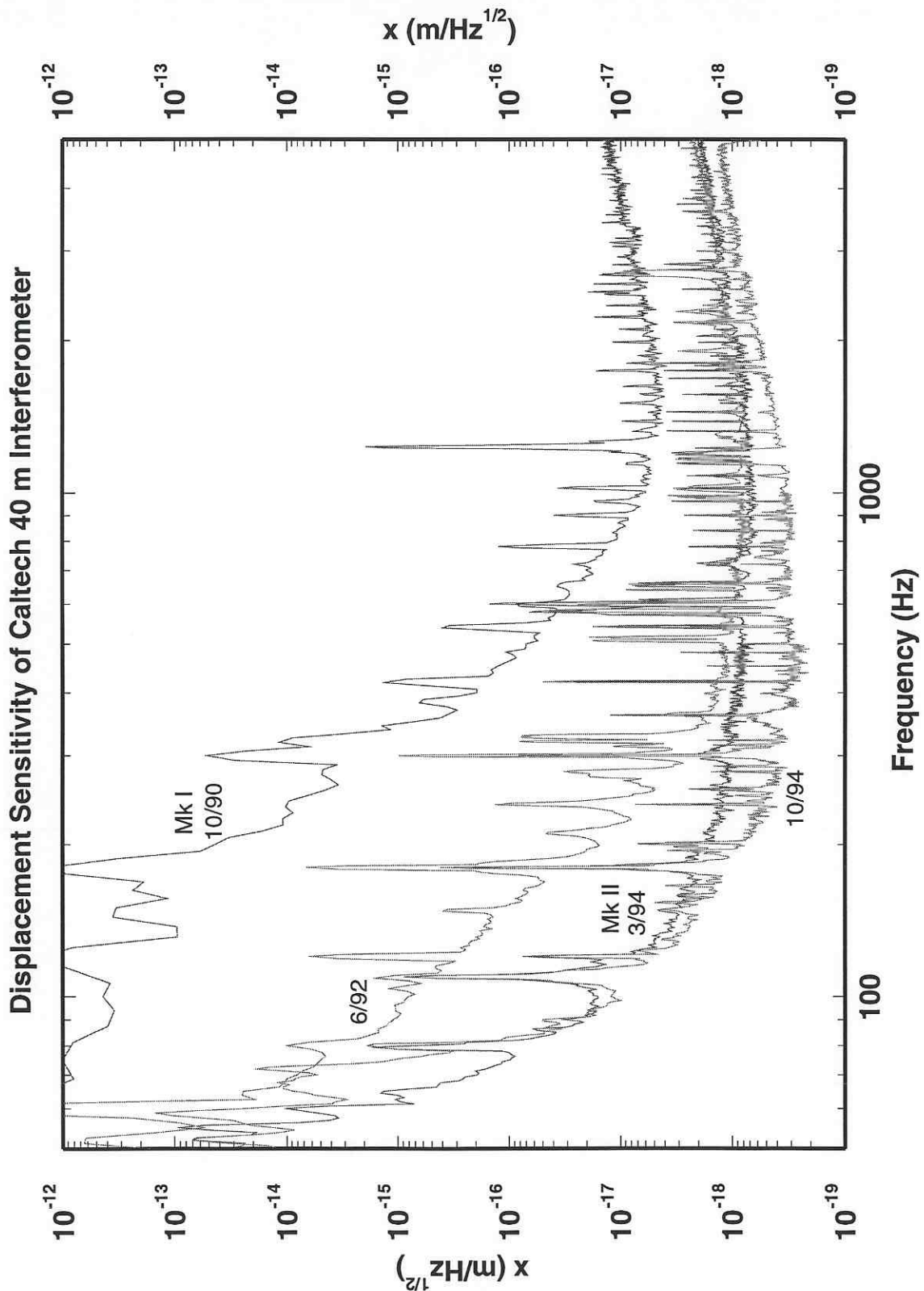
## ALIGNMENT SENSING & CONTROL (cont.)

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- MODELING THE ALIGNMENT SENSITIVITY AND SENSING SYSTEM
  - USE MATRIX TECHNIQUES AND SCATTER TO ESTIMATE OFF-DIAGONAL RELATIONSHIPS
    - BREAK LIGHT INTO  $TEM_{00}$ ,  $TEM_{01}$  MODES, PARAXIAL APPROXIMATION
    - PROVIDES BOTH REQUIREMENTS AND DETAILED DESIGN INFORMATION
- EXPERIMENTAL VERIFICATION OF MODEL
  - SUB-SCALE EXPERIMENT WITH COMPLETE LIGO LENGTH, ANGLE READOUT CONFIGURATION
  - QUANTITATIVE RESULTS EXPECTED IN EARLY '96



# IMPROVEMENTS IN 40m INTERFEROMETER SENSITIVITY



# RESEARCH & DEVELOPMENT

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- STATUS

- SEISMIC NOISE ISOLATION

- PASSIVE SEISMIC STACK UNDERSTOOD
- SOFTER ELASTOMER SPRINGS WILL BE USED (RTV VS VITON)
- SITE CHARACTERISTICS VERY QUIET AND CHARACTERIZED
- BASELINE DESIGN FOR DETECTOR WILL MEET LIGO GOAL

- THERMAL NOISE

- PHYSICAL PROCESSES HAVE BEEN IDENTIFIED AND CHARACTERIZED
- SUSPENSION REQUIREMENTS RELAXED IN FAVOR OF TIGHTER INTERNAL MODE REQUIREMENTS
- BASELINE DESIGN FOR DETECTOR WILL MEET LIGO GOAL

- SHOT NOISE

- LIGO GOAL IS 10X BETTER THAN DEMONSTRATED TO DATE
- REQUIRES CHARACTERIZATION AND INTEGRATION OF DIFFERENT ELECTRO-OPTICAL COMPONENTS AND SUB-SYSTEMS -- CIT/40M ; MIT/5M FACILITIES:
  - RECYCLING
  - RECOMBINATION
  - HIGHER POWER





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# LIGO RESEARCH COMMUNITY





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## LIGO COLLABORATIVE PROGRAM

- PURPOSE AND GOALS
  - TO ENHANCE THE PROBABILITY OF THE (EARLY) SUCCESSFUL DETECTION OF GRAVITATIONAL WAVES (GW) AND TO OPEN THE FIELD OF GW ASTROPHYSICS
  - TO INCREASE PARTICIPATION OF RESEARCHERS IN THE SCIENTIFIC & TECHNICAL CHALLENGES OF GW DETECTION
  - TO DEVELOP A RESEARCH COMMUNITY FOR THE LIGO FACILITIES (BOTH SITES & UNIVERSITY RESOURCES)
  - TO ENHANCE THE EFFECTIVENESS OF THE SCIENTIFIC CONSTITUENCY FOR THE FIELD OF GW ASTROPHYSICS



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## COMMUNICATION WITH SCIENTIFIC COMMUNITY

- ORGANIZATION OF SESSIONS/PRESENTATIONS ON GW DETECTION AT MEETINGS OF SCIENTIFIC & TECHNICAL SOCIETIES
  - APS (4/94)
  - SNOWMASS APS PLANNING MEETING FOR PARTICLE & NUCLEAR ASTROPHYSICS (7/94)
  - MARCEL GROSSMAN MEETING (7/94)
  - OPTICAL SOCIETY OF AMERICA ANNUAL MEETING (10/94)
  - ASPEN WINTER PHYSICS CONFERENCE SERIES (1/95)
  - INTERNATIONAL MEETING ON GENERAL RELATIVITY & GRAVITATION (7/95)
  - ASPEN WINTER PHYSICS CONFERENCE SERIES (1/96)
  - CORAL GABLES CONFERENCE (1/96)



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## COMMUNICATION WITH SCIENTIFIC COMMUNITY (cont.)

- LIGO-HOSTED TOPICAL WORKSHOPS ON TECHNOLOGIES RELEVANT TO GW DETECTION & TOPICAL WORKSHOPS ORGANIZED BY LIGO, VIRGO, GEO600, AIGO, AND LIGO AFFILIATES:
  - THERMAL NOISE (1/94)
  - COALESCING BINARY WAVEFORMS & DATA ANALYSIS (1/94)
  - BEAM TUBE BAFFLE DESIGNS AND ANALYSES (1/95)
  - OPTICS, MODELING, FABRICATION, TESTING
  - INTERFEROMETER CONFIGURATIONS
  - DATA ANALYSIS (WINTER '95)
  - SEISMIC ISOLATION
  - LASER SOURCES
  - APPLICATIONS OF SQUEEZED LIGHT



# LIGO RESEARCH COMMUNITY

## COMMUNICATION WITH SCIENTIFIC COMMUNITY

(cont.)

- WWW ON-LINE LIGO RESEARCH RESULT & PROJECT STATUS
- WWW ON-LINE PUBLICATIONS LIST <http://www.ligo.caltech.edu>



### LIGO Project Home Page

California Institute of Technology

Massachusetts Institute of Technology

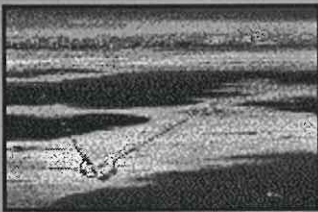
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Welcome to the experimental WWW home page of the LIGO project.

*Please note that this is an experimental set-up; send any comments or observations which you may have to [webmaster@ligo.caltech.edu](mailto:webmaster@ligo.caltech.edu)*

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The Laser Interferometer Gravitational-Wave Observatory (LIGO) project is a pioneering effort to design and construct a novel scientific facility – a gravitational-wave observatory – that will open a new observational window on the universe.



- LIGO FACILITIES INTERFACE DOCUMENT





# LIGO RESEARCH COMMUNITY

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## ORGANIZATIONAL STRUCTURE

- LIGO PRE-PROGRAM ADVISORY COMMITTEE (LPPAC)
  - SELF-DISSOLVING GROUP
    - FORMED BY LIGO MANAGEMENT WITH NSF CONCURRENCE
    - COMPOSED OF INDIVIDUALS EXPERIENCED IN GW RESEARCH, OPERATION OF SCIENTIFIC FACILITIES
    - MEMBERS:
      - P. SAULSON (SYRACUSE UNIVERSITY) - -CHAIR
      - L.S. FINN (NORTHWESTERN UNIVERSITY)
      - A. GIAZOTTO (UNIVERSITY OF PISA/VIRGO)
      - J. HALL (JILA)
      - W. HAMILTON (LSU)
      - C. PRESCOTT (SLAC)
      - A. RUEDIGER (MPI-GARCHING/GEO)
  - ADVISES LIGO:
    - COMPOSITION, CHARTER, & ROLE OF THE LIGO PROGRAM ADVISORY COMMITTEE (LPAC) WHICH WILL REPLACE IT
    - ASSISTS IN SETTING UP LIGO EXECUTIVE COMMITTEE (OVERSEES LIGO RESEARCH COMMUNITY ACTIVITIES)
    - PROPOSES EXTERNAL ADVISORY COMMITTEE (EAC)
      - PROVIDES TECHNICAL ADVICE TO LIGO PI RE: LIGO UPGRADES& OTHER TECHNICAL ISSUES



# LIGO RESEARCH COMMUNITY

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## ORGANIZATIONAL STRUCTURE (cont.)

- LIGO PROGRAM ADVISORY COMMITTEE (LPAC)
  - WILL ADVISE LIGO MANAGEMENT ON:
    - EVALUATION OF PROPOSALS BY POTENTIAL COLLABORATORS
    - SCHEDULING & PHASING OF INSTALLATION OF DETECTORS, OBSERVATIONAL PROGRAMS, & FUTURE ENHANCEMENTS; FOR OTHER FACILITY USES BY INTERNAL & EXTERNAL RESEARCHERS
    - SCHEDULING & PHASING OF FUTURE ENHANCEMENTS FOR THE FACILITIES
    - LIGO DATA ANALYSIS PROGRAM
    - COORDINATION OF OBSERVATIONS WITH OTHER GRAVITATIONAL WAVE & ASTRONOMICAL FACILITIES
    - INTERNATIONAL COLLABORATIONS FOR DEVELOPING TECHNOLOGIES & JOINT OPERATIONS



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## ORGANIZATIONAL STRUCTURE (cont.)

- LIGO RESEARCH COMMUNITY:
  - SCIENTIFIC & ENGINEERING BODY WITH INTERESTS IN GW DETECTOR RESEARCH TO BE CARRIED OUT BY LIGO
  - INTERNAL & EXTERNAL MEMBERSHIP
  - PARTICIPANTS IN COLLABORATIVE PROGRAMS OF RESEARCH & DEVELOPMENT
    - (E.G., INFORMAL GROUP ON SOURCES, WAVEFORMS, & COMPUTATIONAL ALGORITHMS, FORMED AT THE COALESCING BINARY WORKSHOP IN 1/94 -- MEETING NOW FOR 2ND TIME)
- LIGO EXTERNAL RESEARCH COORDINATOR
  - INTERNAL (TO LIGO) POINT OF CONTACT FOR INDIVIDUALS WISHING TO CARRY OUT COLLABORATIVE PROGRAMS WITH LIGO
  - WILL ADVISE POTENTIAL COLLABORATORS AND DIRECT THEM TO KNOWLEDGEABLE & INTERESTED PERSONS WITHIN LIGO PROGRAM OFFICE





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## ORGANIZATIONAL STRUCTURE (cont.)

- MODES OF COLLABORATION WITH LIGO
- INDIVIDUALS:
  - VISITORS
    - BRIEF PERIODS OF RESEARCH & STUDY AT THE UNIVERSITIES (CIT/MIT)
    - LIGO TEAM MEMBERS VISIT OTHER INSTITUTIONS FOR BRIEF PERIODS
  - RESEARCH INDEPENDENT OF, BUT RELEVANT TO LIGO (E.G. ADVANCED DETECTORS)
  - RESEARCHERS INTERESTED IN ENABLING TECHNOLOGY DEVELOPMENT FOR LIGO ITSELF:
    - HARDWARE
    - SOFTWARE
    - MODELS
  - RESEARCHERS INTERESTED IN DEVELOPING NEW ADVANCED DETECTORS FOR EVENTUAL INSTALLATION ON LIGO
  - DATA USERS: MINING OF ARCHIVED DATABASES
  - RESEARCHERS INTERESTED IN MAKING JOINT/ COORDINATED ASTROPHYSICAL SEARCHES WITH LIGO



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## ORGANIZATIONAL STRUCTURE (cont.)

- **INSTITUTIONS:**

- MEMORANDA OF UNDERSTANDING (MOUs)
- CALTECH RELATIVITY THEORY GROUP (THORNE) - COMPLETE
  - SOURCE IDENTIFICATION & ALGORITHM DEVELOPMENT
- AUSTRALIAN CONSORTIUM FOR INTERFEROMETRIC GRAVITATIONAL ASTRONOMY (ACIGA) - COMPLETE
  - DISPLACEMENT SENSORS FOR INITIAL LIGO TEST MASS STABILIZATION
- CU/JILA (BENDER) - IN PROCESS
  - LOW FREQUENCY ISOLATION PROJECT
- VIRGO - IN PROCESS
  - TECHNOLOGY EXCHANGES; JOINT DEVELOPMENT & ANALYSES
- PENN STATE GRAVITATIONAL PHYSICS GROUP (CUTLER) - PENDING
  - DATA ANALYSIS & MODELING
- EXTANT BUT TO BE RE-WRITTEN:
  - SAULSON (SYRACUSE)
  - FINN (NORTHWESTERN)
  - STANFORD
  - BRAGINSKY (MSU)
- **ACTIVITIES WOULD REQUIRE PROPOSALS TO LPAC/NSF/FUNDING AGENCIES FOR REVIEW &**

