

Detector Research & Development

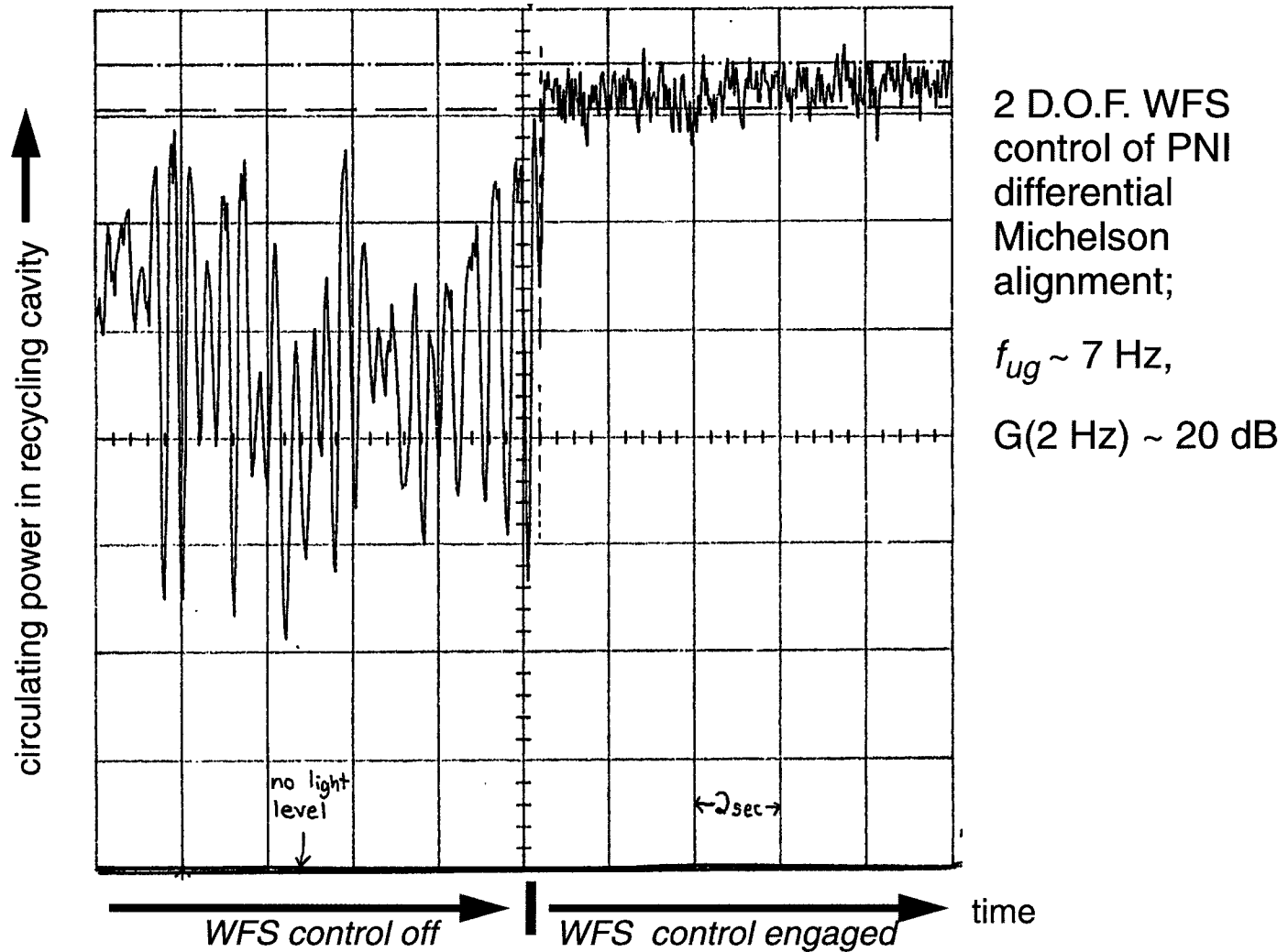
M. E. Zucker

- Phase Noise Interferometer (PNI) program
- 40-meter Interferometer program
- Fixed Mirror Interferometer (FMI) program

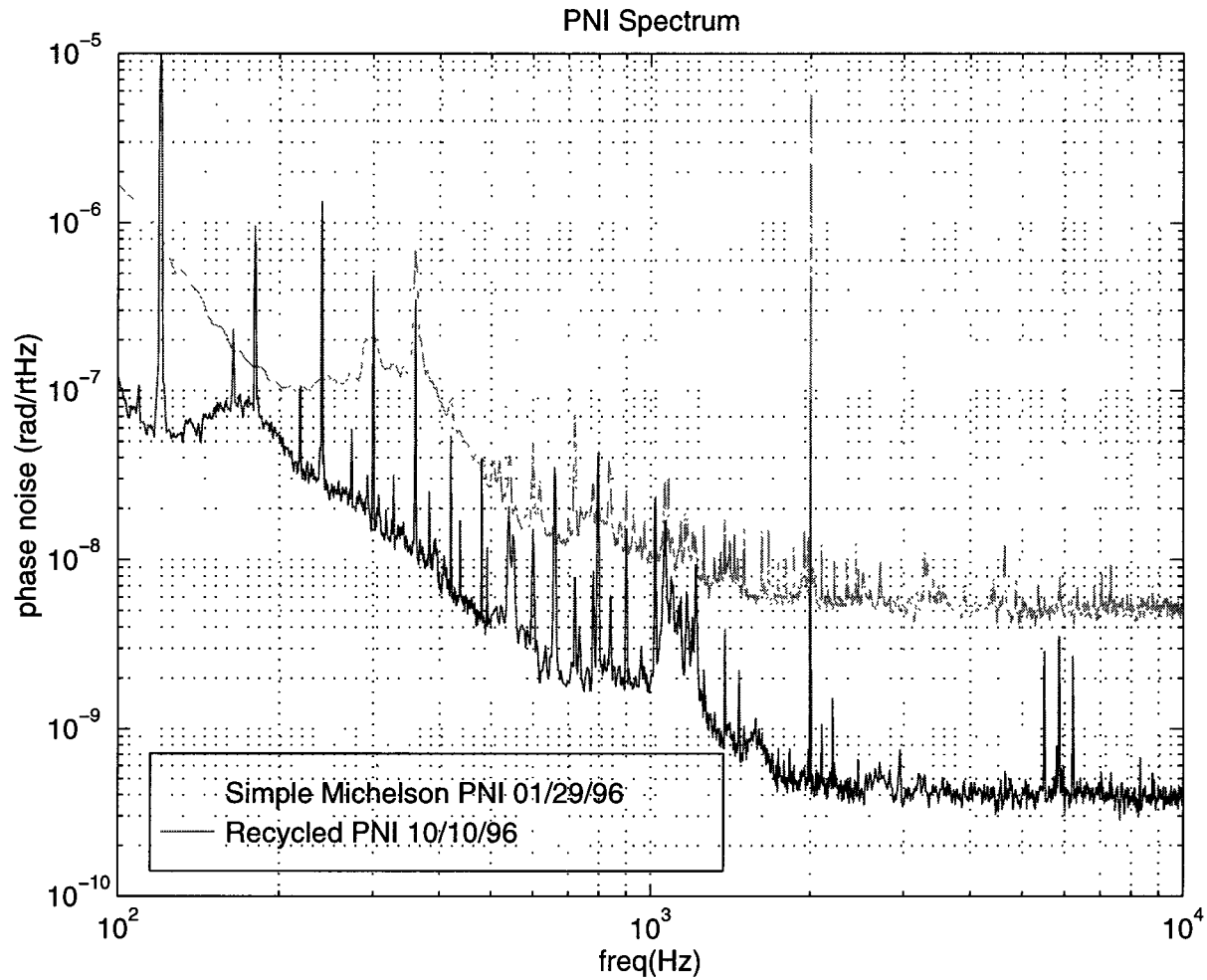
Phase Noise Interferometer (PNI)

- Goals: to demonstrate optical phase measurement sensitivity, understand technical sensing noise sources, and test LIGO length sensing/control (LSC) components
- Recent advances:
 - ›› Barry STACIS active isolation systems added for second vacuum chamber
 - ›› Recycling mirror installed, RF modulation & control systems upgraded
 - ›› Wavefront sensing control system installed for differential MI alignment
- Results:
 - ›› Recycling gain $G = 450$
 - ›› Power incident on beamsplitter 60 W (carrier only; $P_{in} = 200$ mW total)
 - ›› High-frequency phase sensitivity $\sim 3.5 \times 10^{-10}$ rad/Hz^{1/2}

PNI: WFS-based Alignment Control



PNI: Progress on noise spectrum



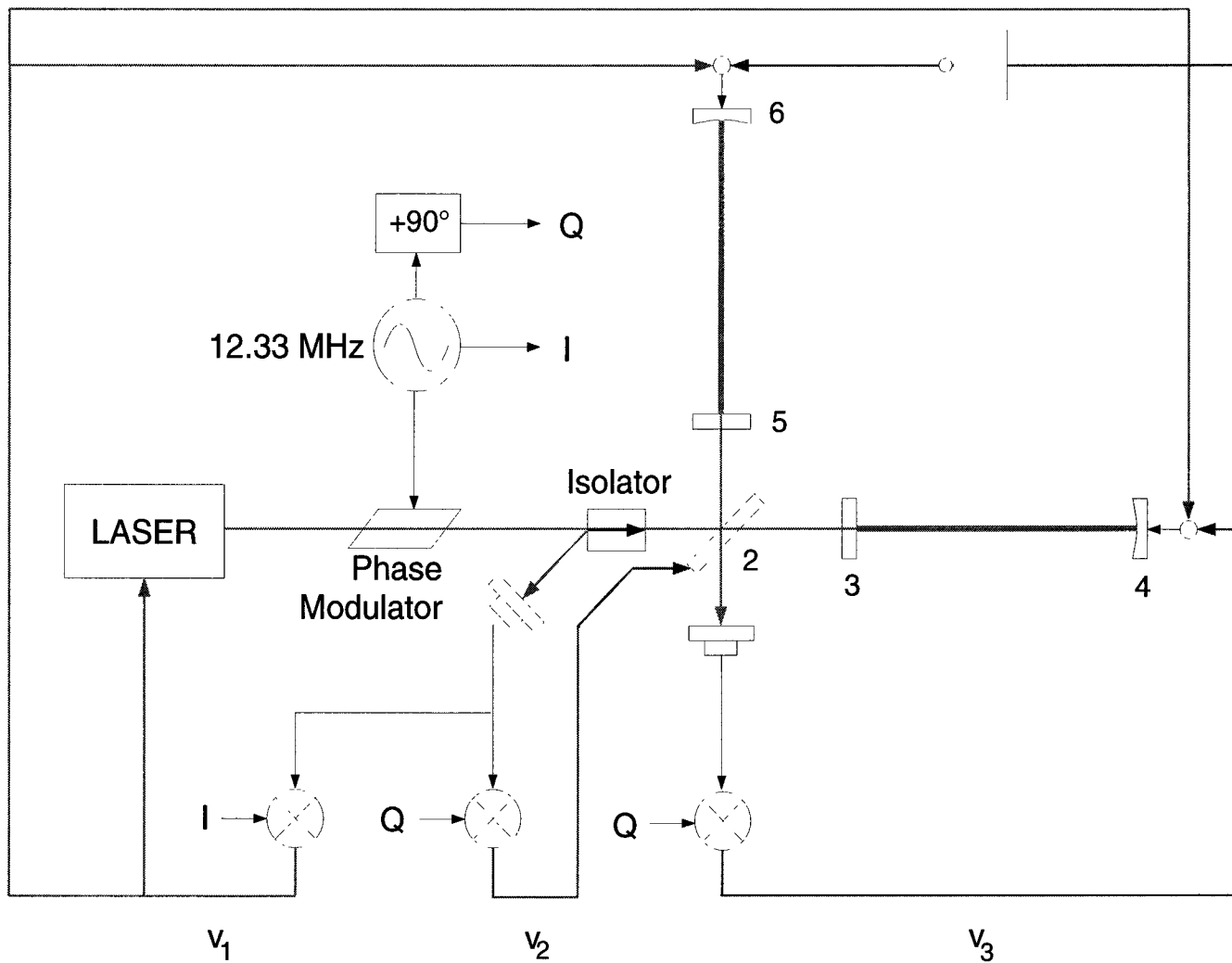
PNI: plan & schedule

- Wrapup of Ar⁺ laser experiments next month
- Conversion to Nd:YAG (NPRO, .7 W) starting in December
 - ››optics & laser ready
 - ››laser prestabilization system being assembled/tested at Caltech
- First phase: linear cavity
 - ››test PSL frequency noise
 - ››debug new frequency control servo
- Second phase: Recycled Michelson configuration (as now)
 - ›› prototype test of LSC high-power photodetector
 - ››test of LSC digital controls (still tentative)
- Tests to wrap up last quarter of '97

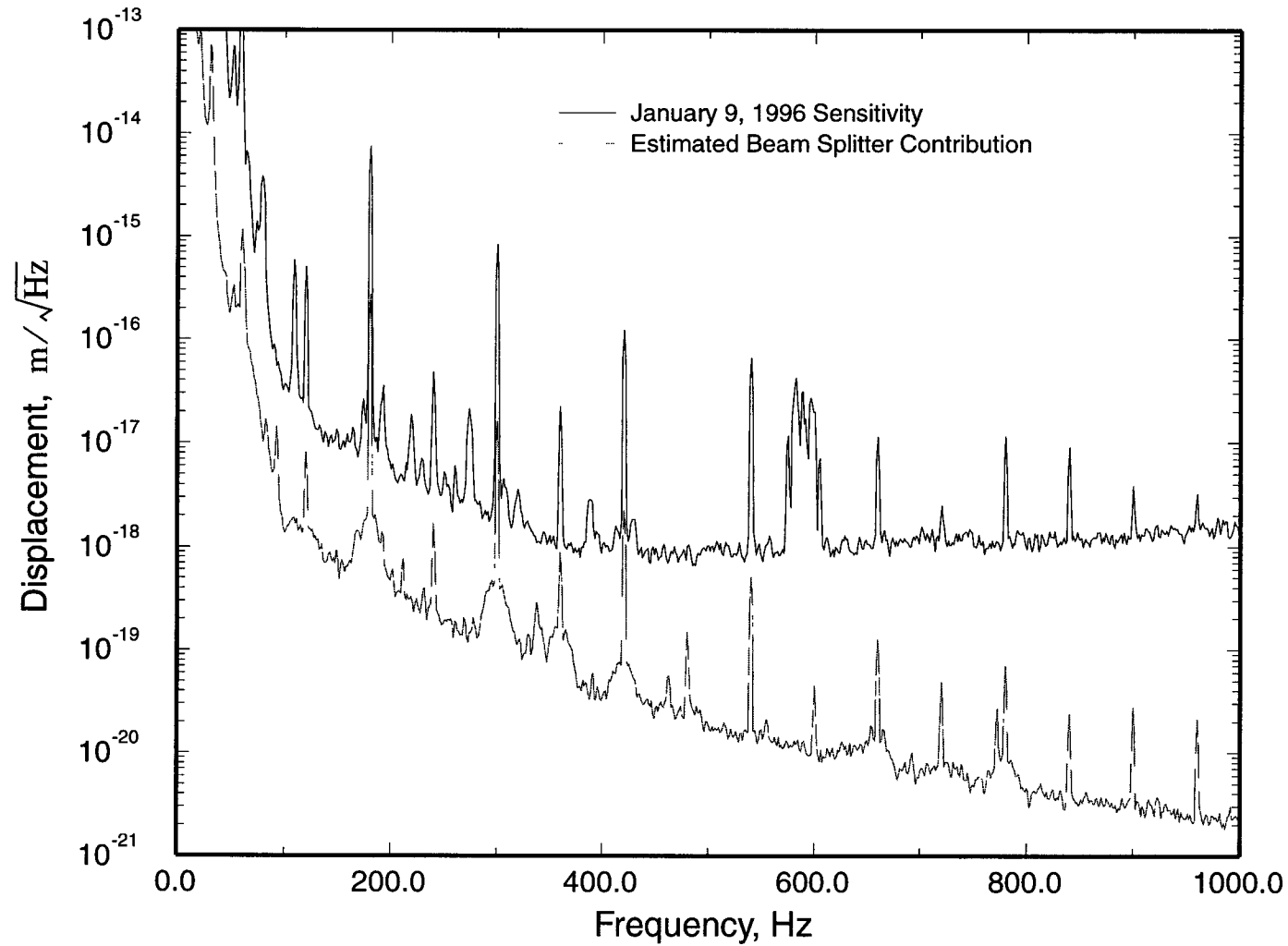
40-m Optical Beam Recombination

- Focus on role as a LIGO configuration testbed; reduced emphasis on displacement noise
- Key first step toward LIGO power-recycled configuration
- Explored new features & noise couplings:
 - ›› Coupled control discriminants (nondiagonal readout)
 - ›› Sign reversals in Michelson differential readout during lock acquisition
 - ›› Greater dependence on uniform mirror figure (new alignment constraint)
 - ›› First-order sensitivity to beamsplitter motion
- First test of a recombined Fabry-Perot Michelson interferometer at high sensitivity

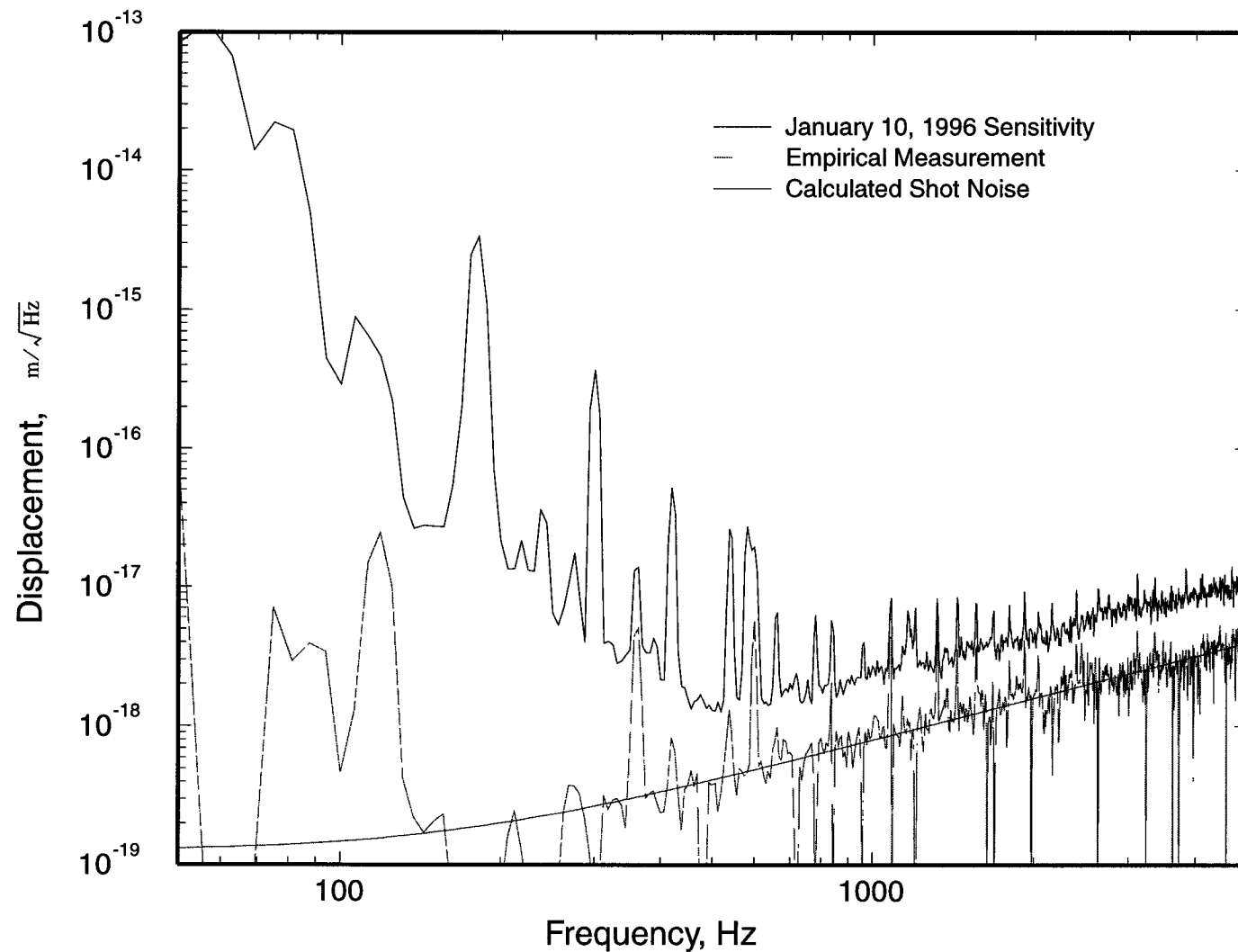
40-m: Recombined Control Topology



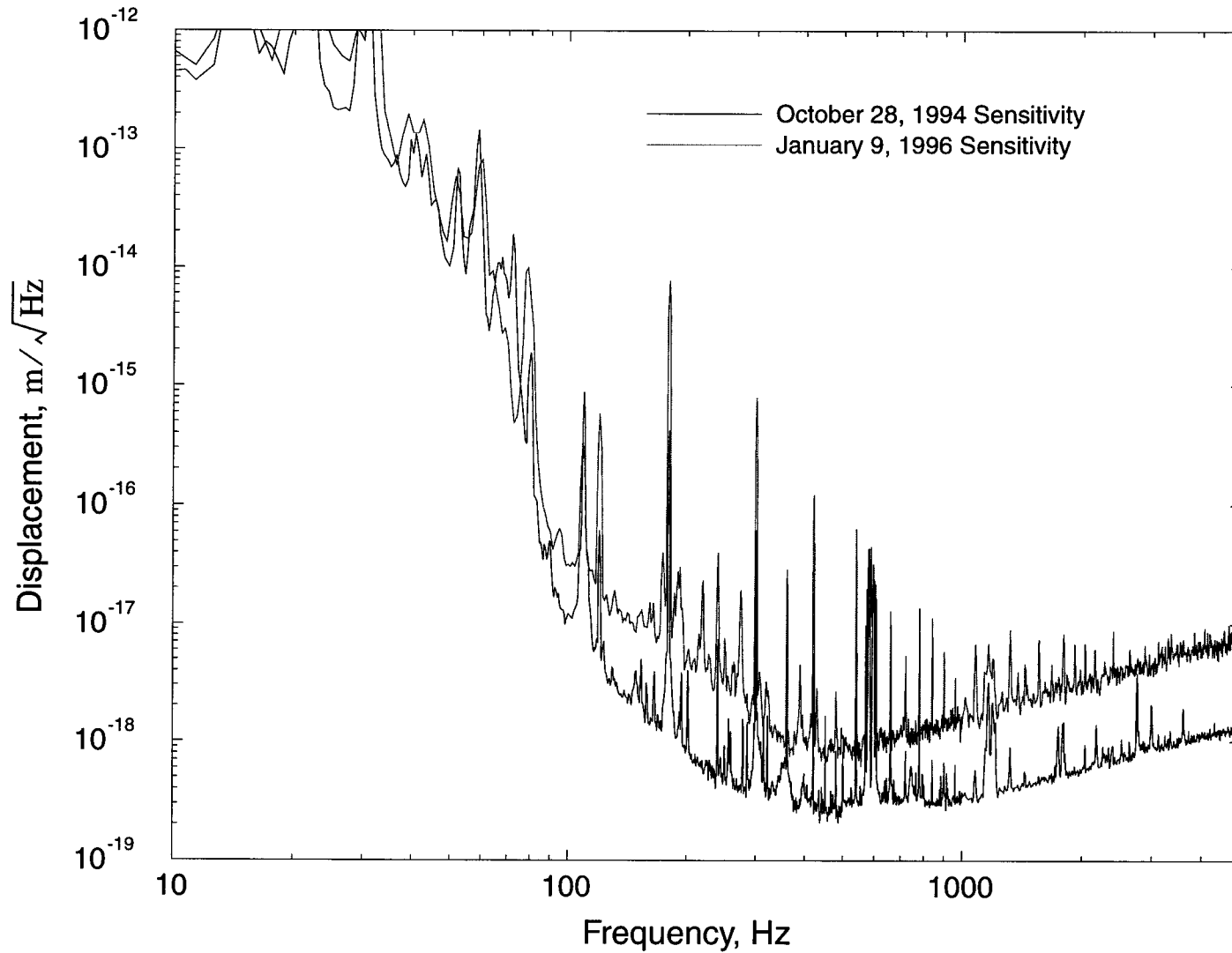
40-m: Beamsplitter Motion Sensitivity



Calculated and measured shot noise



40-m: Sensitivity Comparison



40-m: Suspension prototype test

- Trial of LIGO-type suspension at high displacement-sensitivity
 - ›› Single-loop suspension
 - ›› Integrated sensor/actuators for pitch, yaw, position
 - ›› High-Q attachments
- Integrated test of prototype suspension control electronics
 - ›› Dynamic range, noise
 - ›› Diagnostics & tuning/setup functions
- Existing 40-m suspensions limit sensitivity, repeatability & ability to generalize other tests (=>other 3 pending)
- Significant impact on SUS Preliminary Design

40-m: Power Recycling

- Program concurrent with LIGO LSC design phase; results support LSC final design
- Focus on validating
 - ››cavity lock acquisition sequence
 - ››alignment technique
 - ››modeling codes and design tools
 - ››control electronics prototypes
- Integrated system tests
- Diagnostics and commissioning exercises
- Training

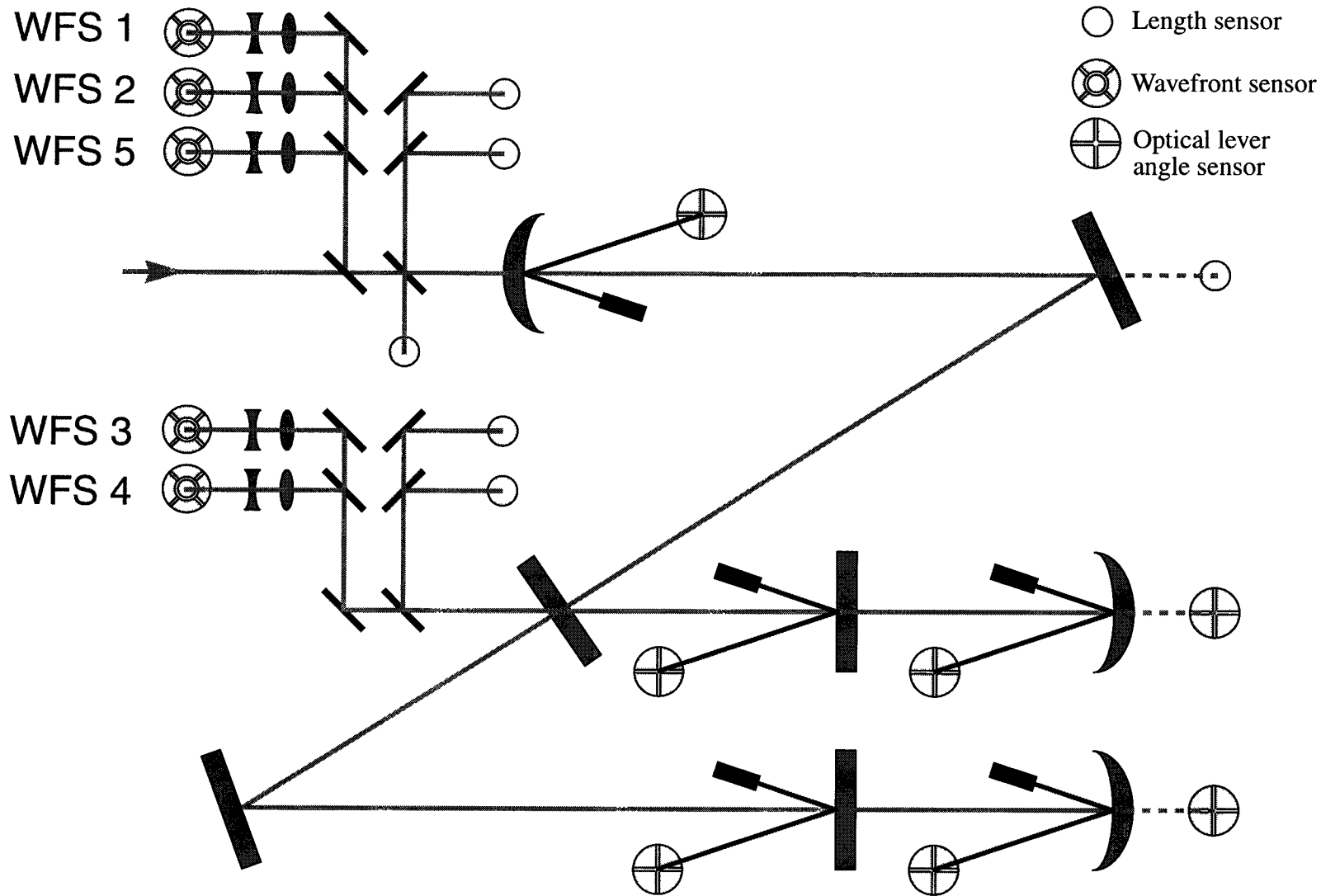
40-m: Recycling Status/Plans

- Installed higher-transmission input couplers
 - ›› target recycling factor of 5 ($T_{in} = 5600$ ppm)
 - ›› installation complete, currently shaking down
- Next: reconfigure vacuum envelope & input optics layout
 - ›› scheduled to start in November; offline preparations underway
 - ›› new side chamber & seismic isolation for expanded input optics
 - ›› new beamsplitter to be installed (in LIGO SOS suspension prototype)
 - ›› new RF modulation frequency to satisfy resonant condition in final stage
- Final stage: install recycling mirror
 - ›› Currently on track for March

FMI: Wavefront Sensing Research

- Goals:
 - ›› Validate Modal Model and its predictions for sensitivity of Wavefront Sensing (WFS) angle readouts, a critical technology for LIGO
 - ›› Develop WFS sensor and signal processing hardware and software
 - ›› Test concepts on a “full-configuration” power-recycled Fabry-Perot michelson
- Apparatus now complete
 - ›› Prestabilized Ar⁺ laser, LIGO-like RF length control
 - ›› Multifrequency phase modulation + frequency-shifted subcarrier generator
 - ›› Tabletop interferometer with PZT tip/tilt and fast/slow piston mirror actuators; aux. laser diode optical lever angle calibrators
 - ›› 5 WFS prototype heads & demodulator modules, VME digital signal processing system

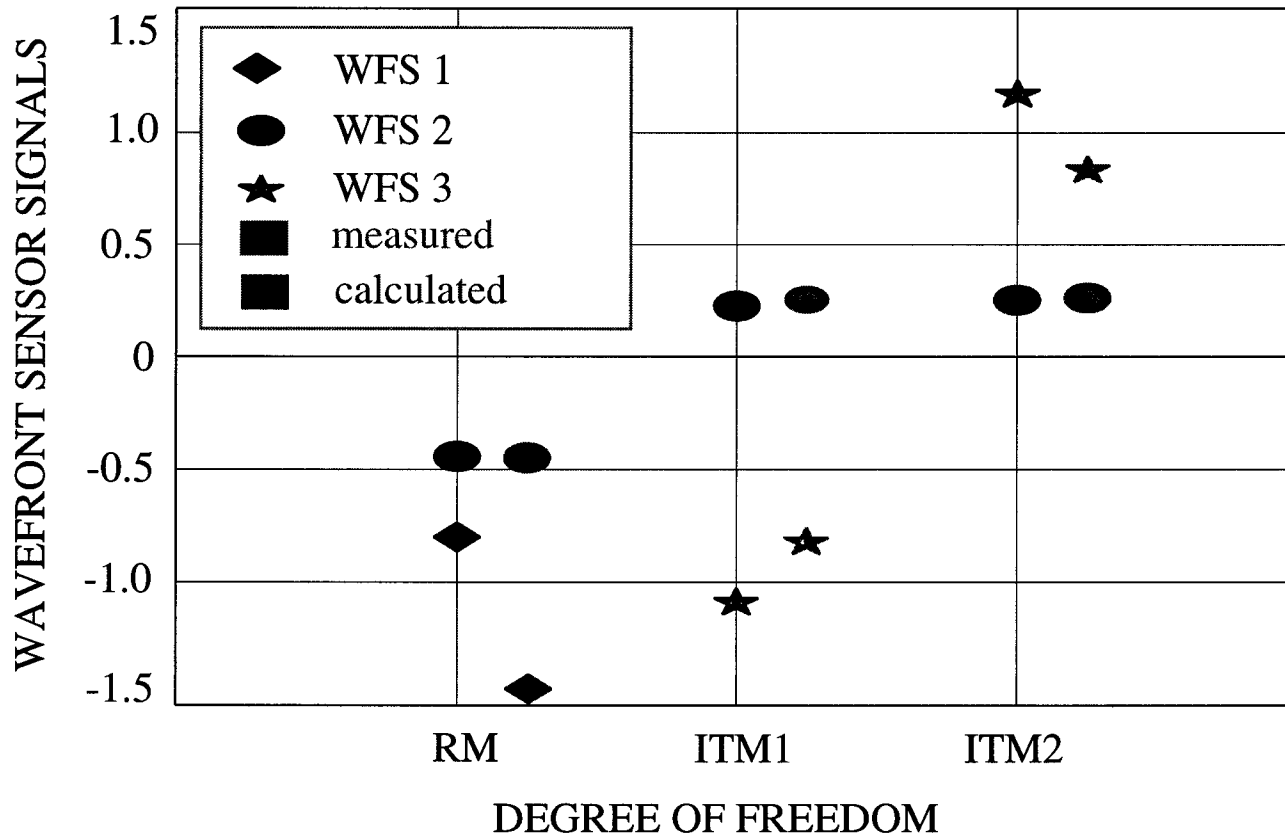
FMI Configuration (schematic)



FMI: Status

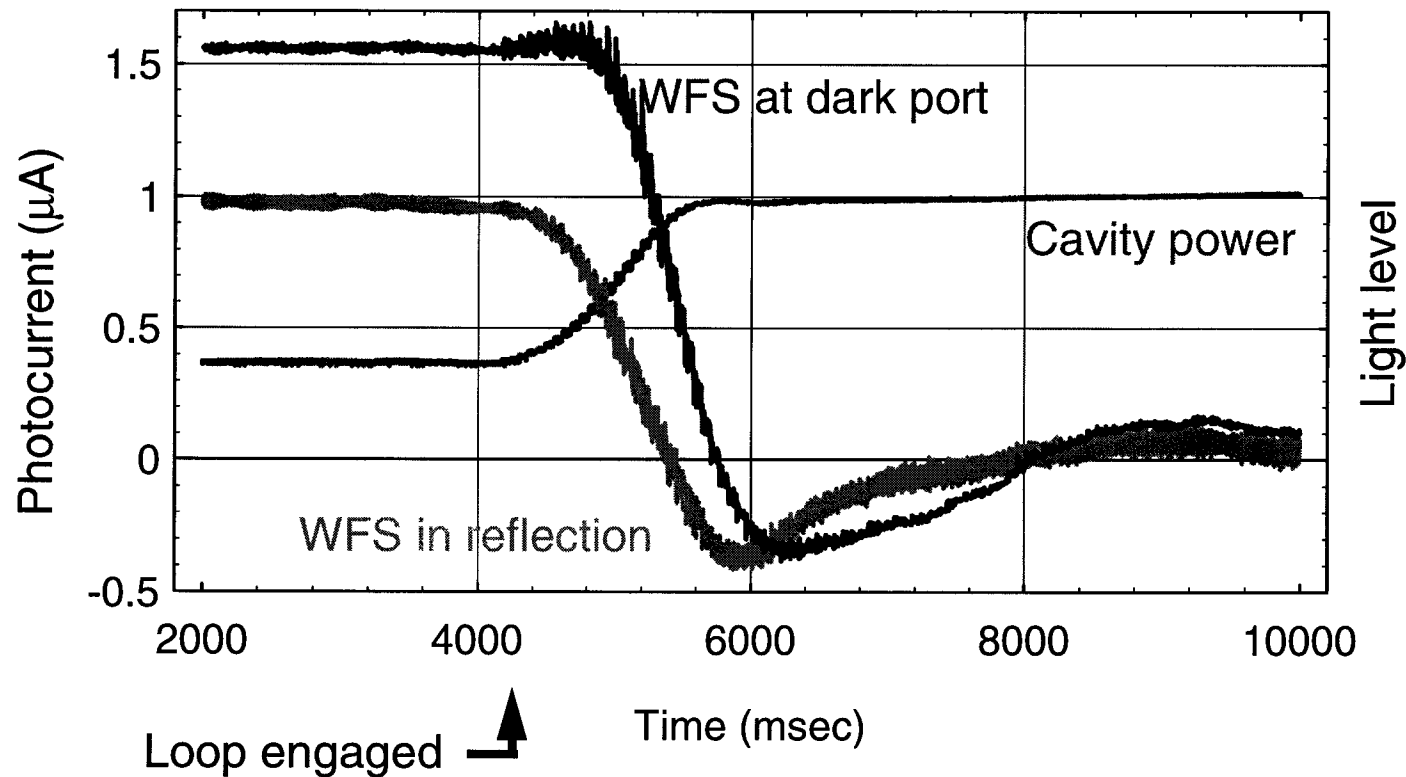
- Good preliminary results with arm cavities disabled (power-recycled simple Michelson only)
- WFS prototype hardware performance consistent with LIGO ASC requirements
- Successful trial of digital MIMO control system; correctly optimized all 6 degrees of freedom (d.o.f.)
- Now bringing arm cavities online for complete data run with all 10 d.o.f.

FMI: Preliminary Results vs. Model



FMI: Digital WFS Control Test

- Closed Loop Control of a Recycled Michelson Interferometer



1 of 1

LIGO-G950000-00-M

