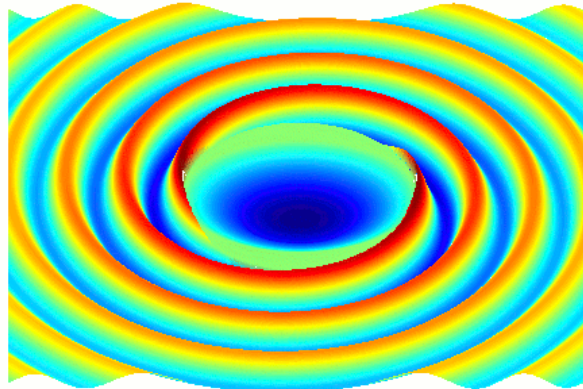


# Stackslide search for continuous gravitational waves using LIGO S4 data.



*Supported by the National Science Foundation*

*<http://www.ligo.caltech.edu>*

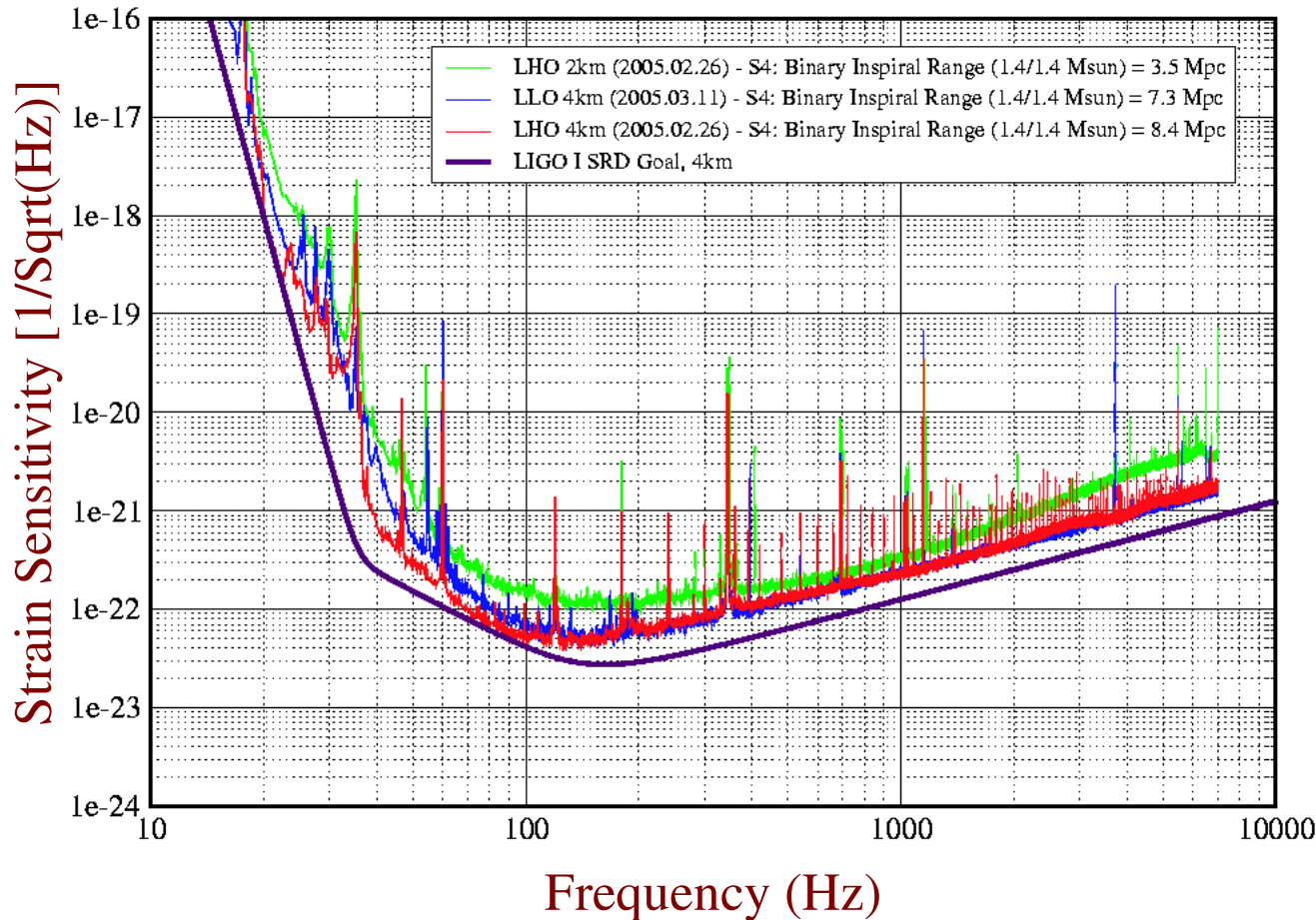
Gregory Mendell, LIGO Hanford Observatory  
on behalf of the LIGO Scientific Collaboration

# Fourth Science Run (S4) Sensitivity

(February 22, 2005 – March 23, 2005)

Strain Sensivities for the LIGO Interferometers

Best Performance for S4 LIGO-G050230-02-E



S5 is currently running at design sensitivity!



# Sources of Continuous Gravitational Waves

(This talk shows results for isolated sources only.)

Search can detect any periodic source.

Upper limits are set on gravitational-wave amplitude,  $h_0$ , of rotating triaxial ellipsoid. →

Credits:

A. image by Jolien Creighton; LIGO Lab Document G030163-03-Z.

B. image by M. Kramer; Press Release PR0003, University of Manchester - Jodrell Bank Observatory, 2 August 2000.

C. image by Dana Berry/NASA; NASA News Release posted July 2, 2003 on Spaceflight Now.

D. image from a simulation by Chad Hanna and Benjamin Owen; B. J. Owen's research page, Penn State University.

**A** Mountain on neutron star

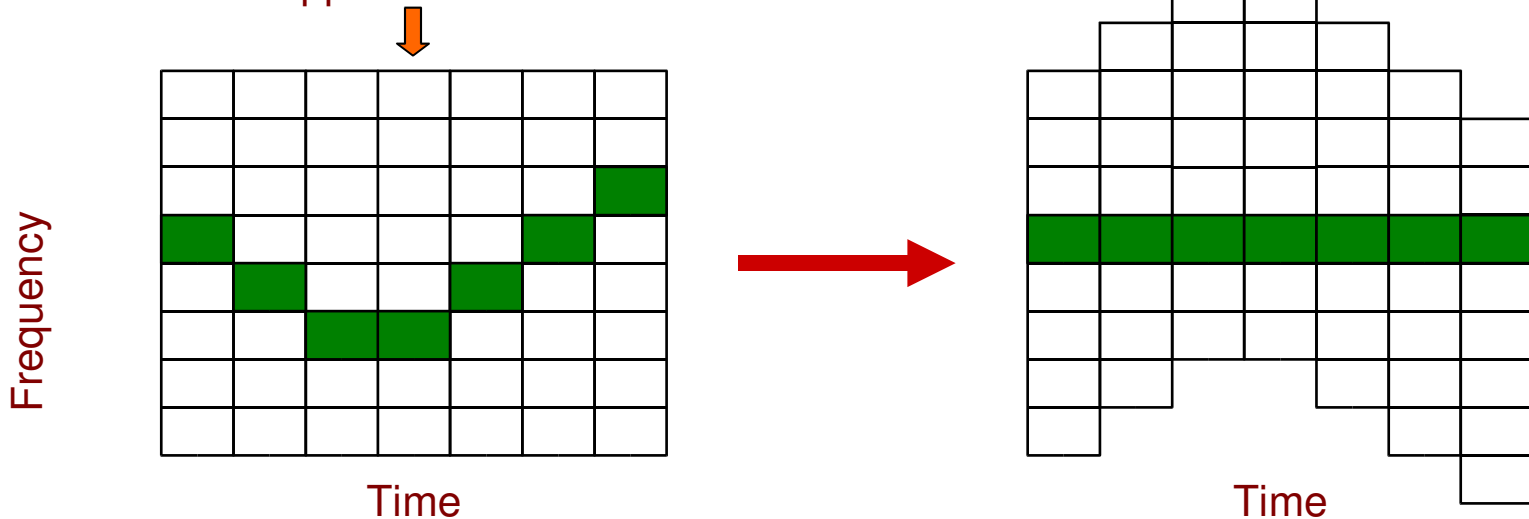
**B** Precessing neutron star

**C** Accreting neutron star

**D** Oscillating neutron star

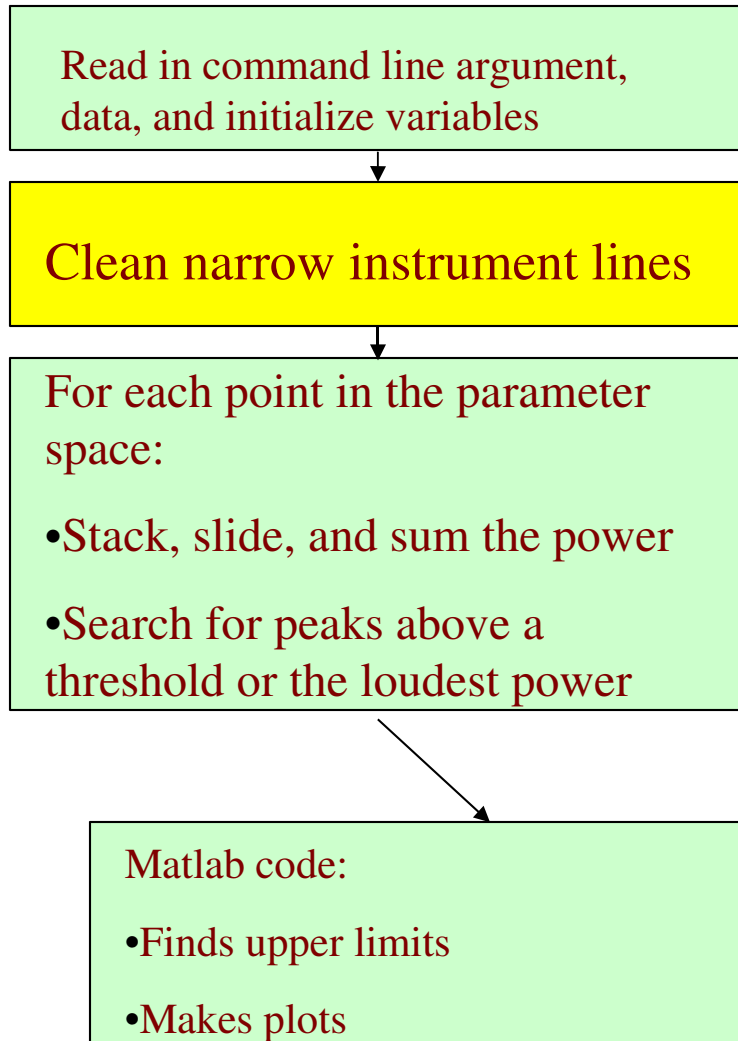
- Break up data into segments; FFT each, producing Short (30 min) Fourier Transforms (SFTs) = coherent step.
- StackSlide: stack SFTs, track frequency, slide to line up & add the power weighted by noise inverse = incoherent step.
- Other semi-coherent methods:
  - Hough Transform: Phys. Rev. D72 (2005) 102004; gr-qc/0508065.
  - PowerFlux: see next talk, W11.00005
- Fully coherent methods:
  - Frequency domain match filtering/maximum likelihood estimation (C7.00001; W11.00006)
  - Time domain Bayesian parameter estimation (C7.00002)
- **Improvements and hierarchical pipeline under development.**

Track Doppler shift and  $df/dt$

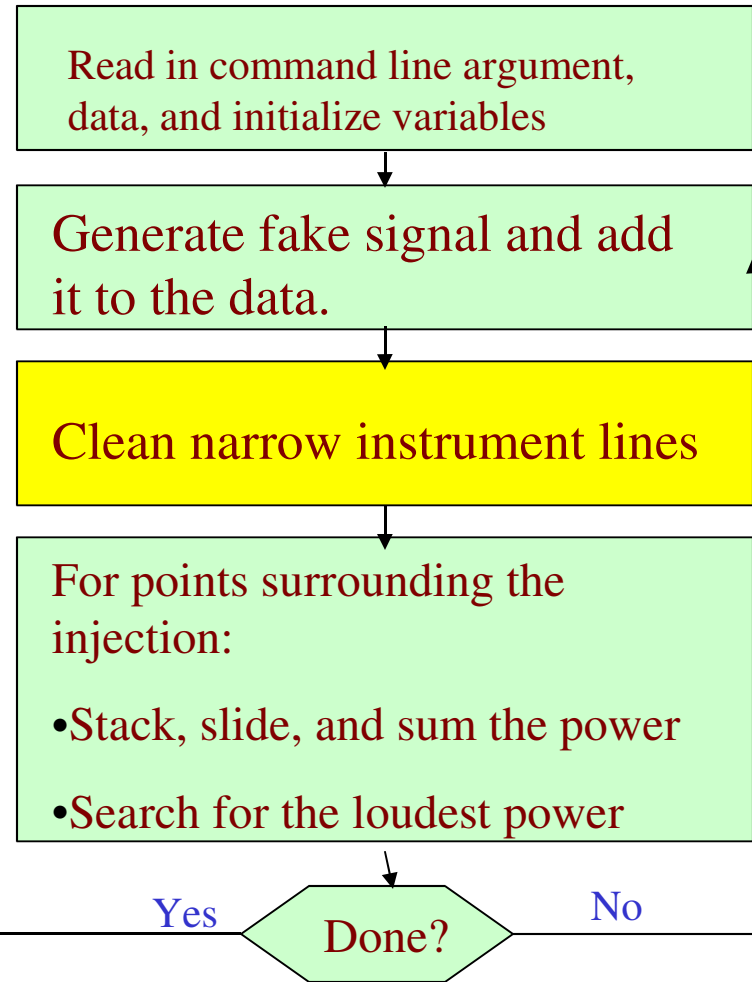


# The StackSlide S4 Pipeline

## Search



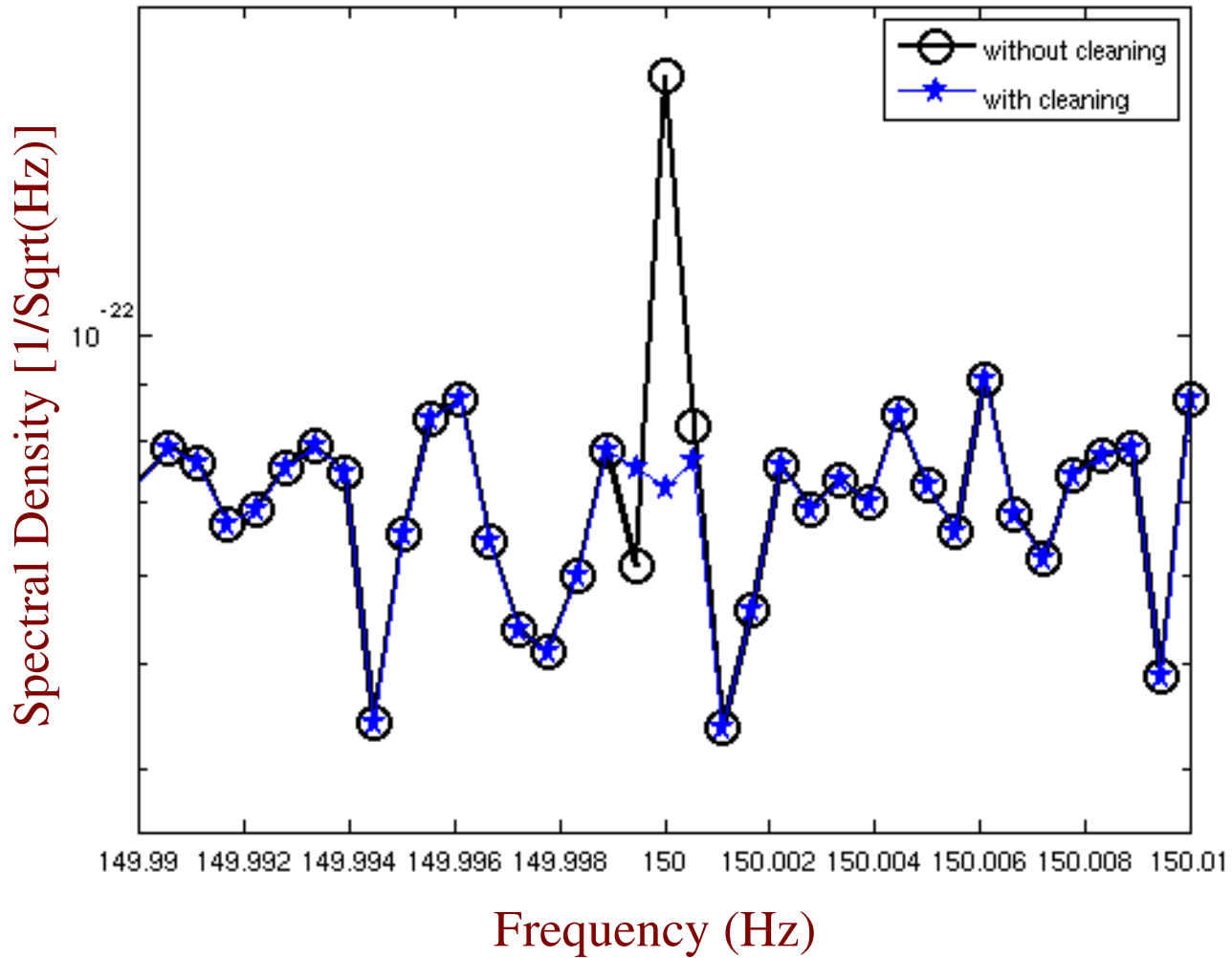
## Monte Carlo Simulation



# Line Cleaning Method

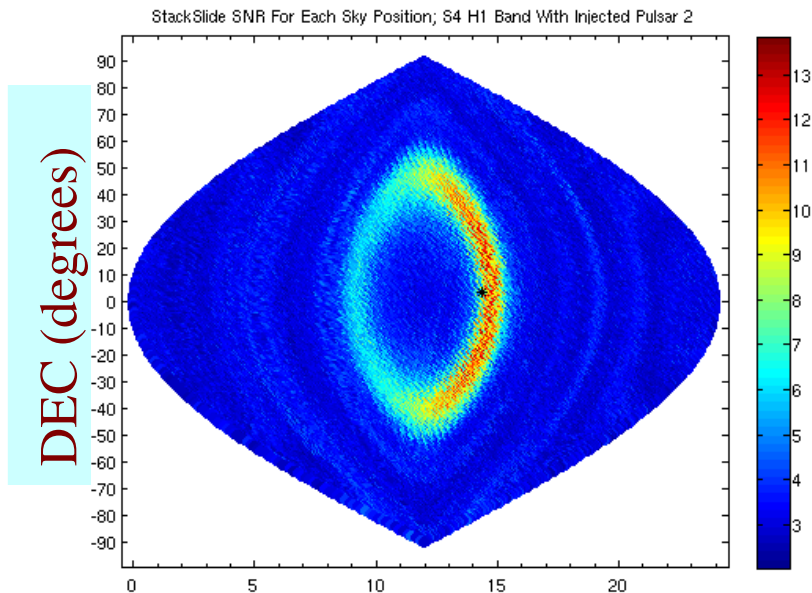
(Only clean known lines; those wider than 0.021 Hz are not cleaned.)

L1 spectral density estimated by averaging power from 10 SFTs without and with line cleaning

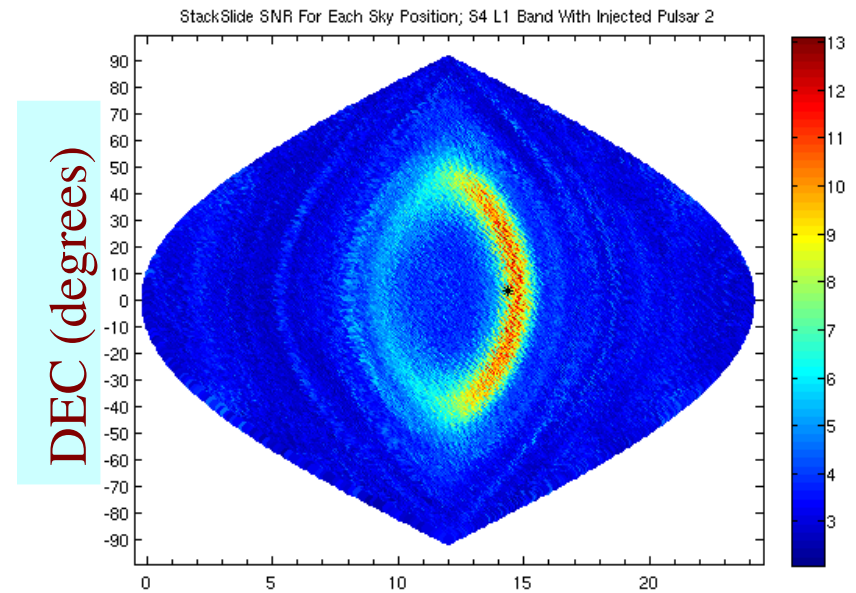


# Analysis of Hardware Injections

Fake gravitational-wave signals corresponding to rotating neutron stars with varying degrees of asymmetry were injected for parts of the S4 run by actuating on one end mirror. Sky maps for the search for an injected signal with  $h_0 \sim 7.5e-24$  are below. Black stars show the fake signal's sky position.



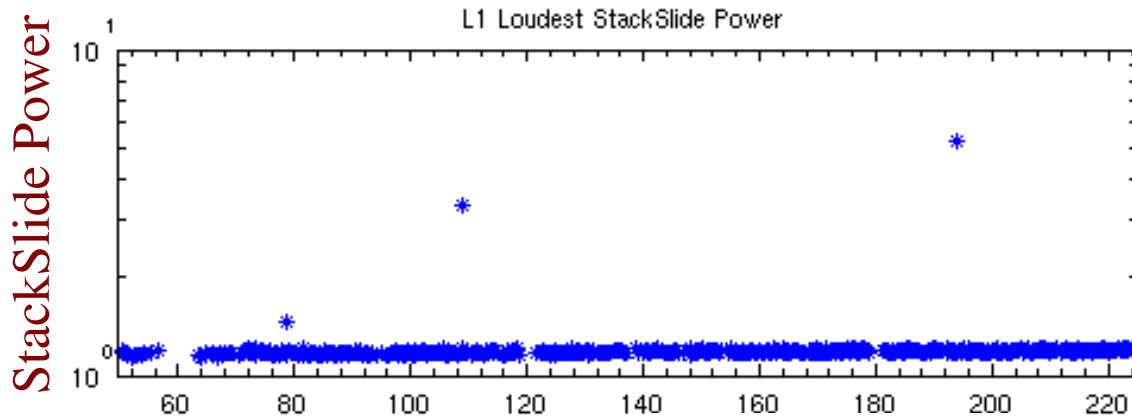
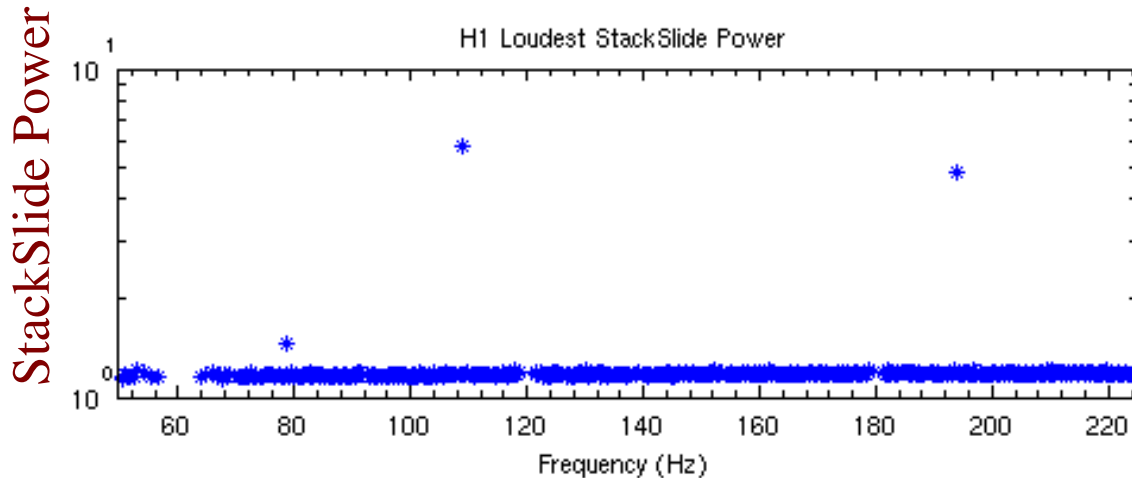
RA (hours)



RA (hours)

## PRELIMINARY

## S4 StackSlide “Loudest Events” 50-225 Hz



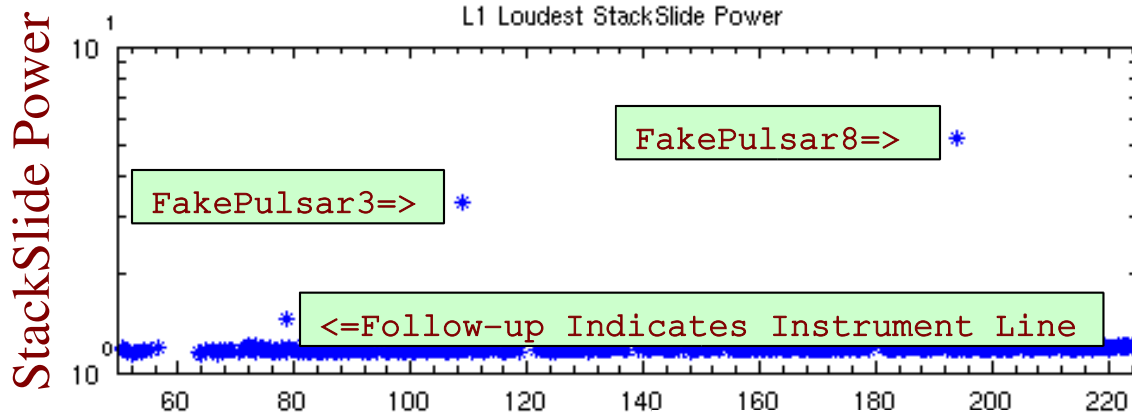
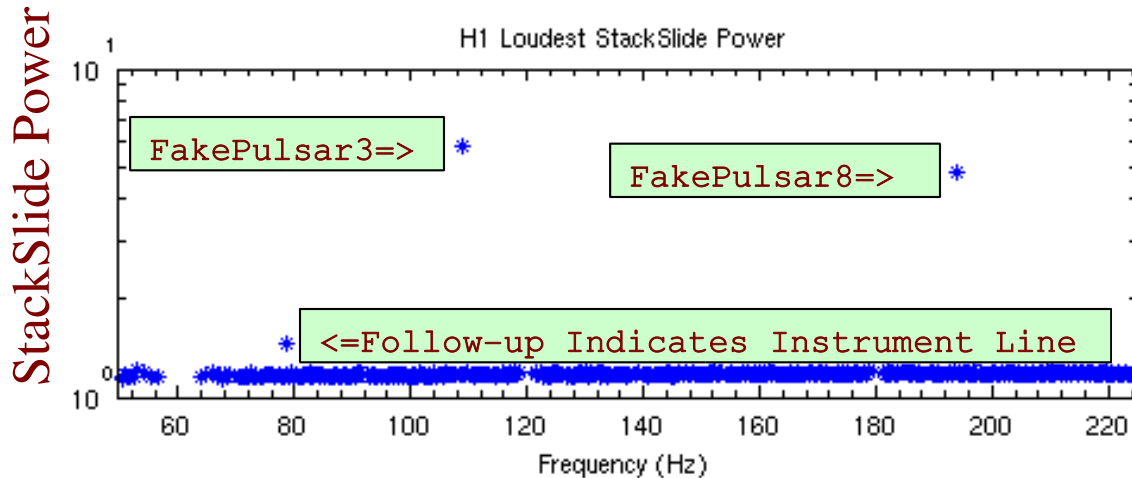
Frequency (Hz)

- Searched 450 freq. per .25 Hz band, 51 values of  $df/dt$ , between 0 &  $-1e-8$  Hz/s, up to 82,120 sky positions (up to  $2e9$  templates). The expected loudest StackSlide Power was  $\sim 1.22$  (SNR  $\sim 7$ )
- Veto bands affected by harmonics of 60 Hz.
- Simple cut: if SNR  $> 7$  in only one IFO veto; if in both IFOs, veto if  $\text{abs}(f_{\text{H1}} - f_{\text{L1}}) > 1.1e-4 * f_0$



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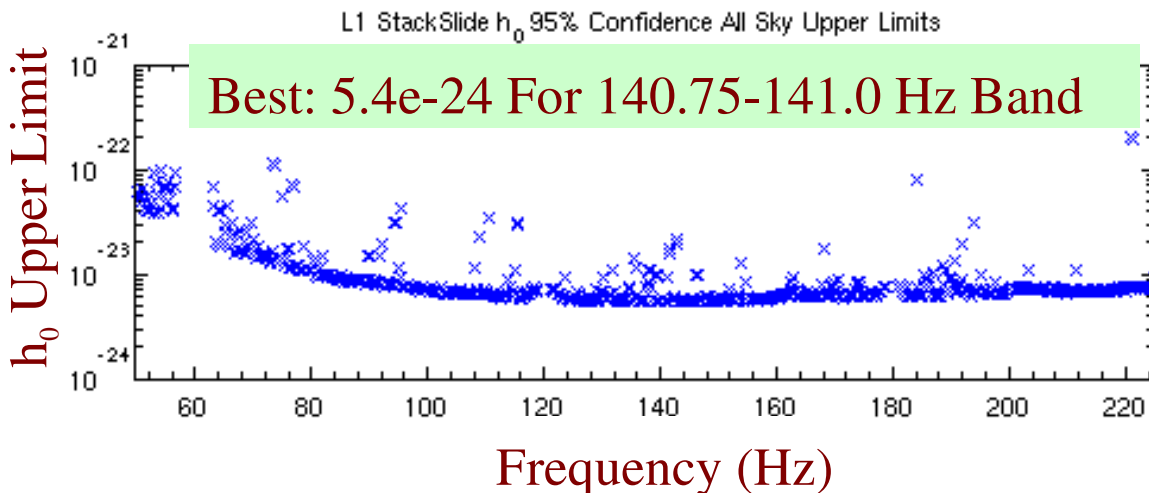
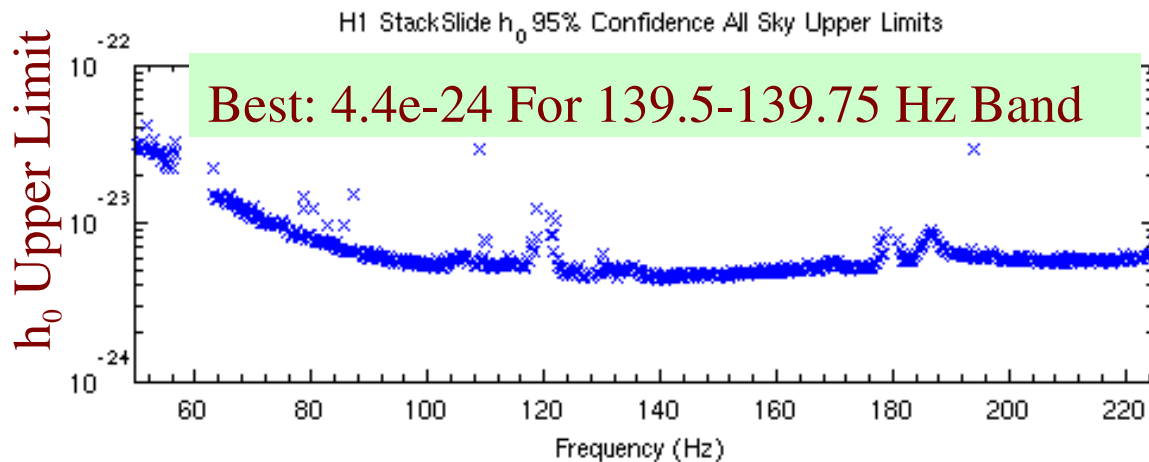


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**PRELIMINARY**

# S4 StackSlide $h_0$ 95% Confidence All Sky Upper Limits 50-225 Hz



Note that  $h_0$  is the gravitational-wave amplitude for a rotating triaxial ellipsoid. Monte Carlo injections of triaxial sources over the search parameter space & source orientations set the ULs.

# Conclusions

- **These results are interesting, but much better results are on the way!**
- **Improvements:**
  - **S5: ~ 2x better sensitivity, 12x or more data.**
  - **Increase time baseline of coherent step**
  - **Multi-ifo code**
  - **Refined coincidence steps**
- **Hierarchical pipeline under development that combines coherent and semi-coherent searches.**



END APS SLIDES