





Laboratoire Kastler Brossel

Thermal and quantum noises in interferometric measurements



Tristan Briant

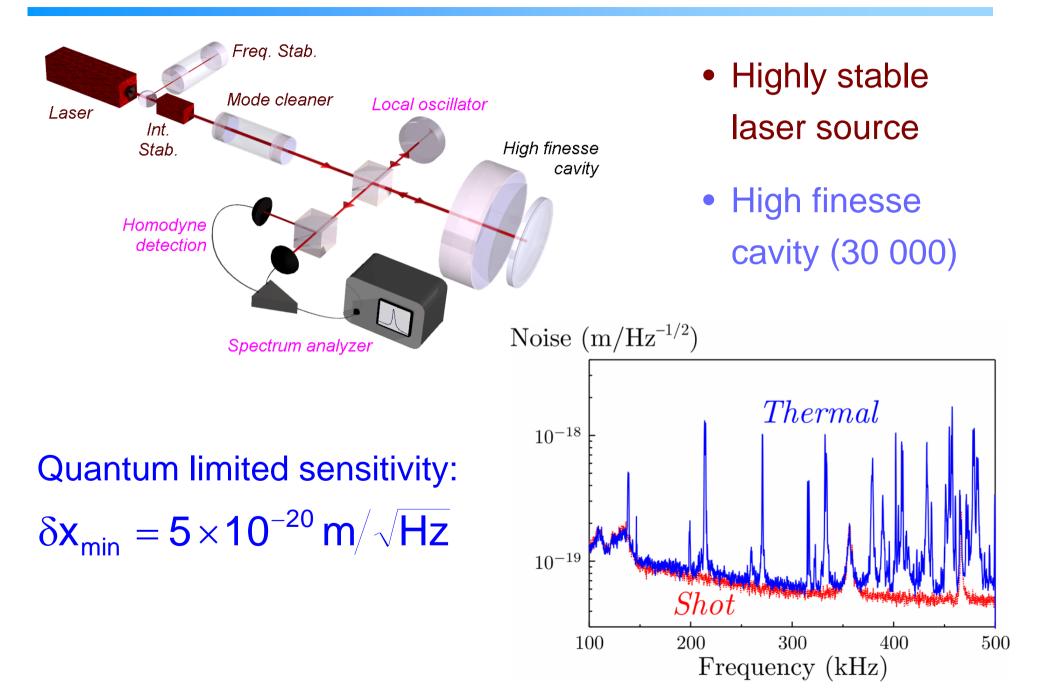
- A. Heidmann M. Pinard
- J.-M. Courty P.-F. Cohadon

O. Arcizet T. Caniard J. Le Bars LIGO-G060300-00-Z

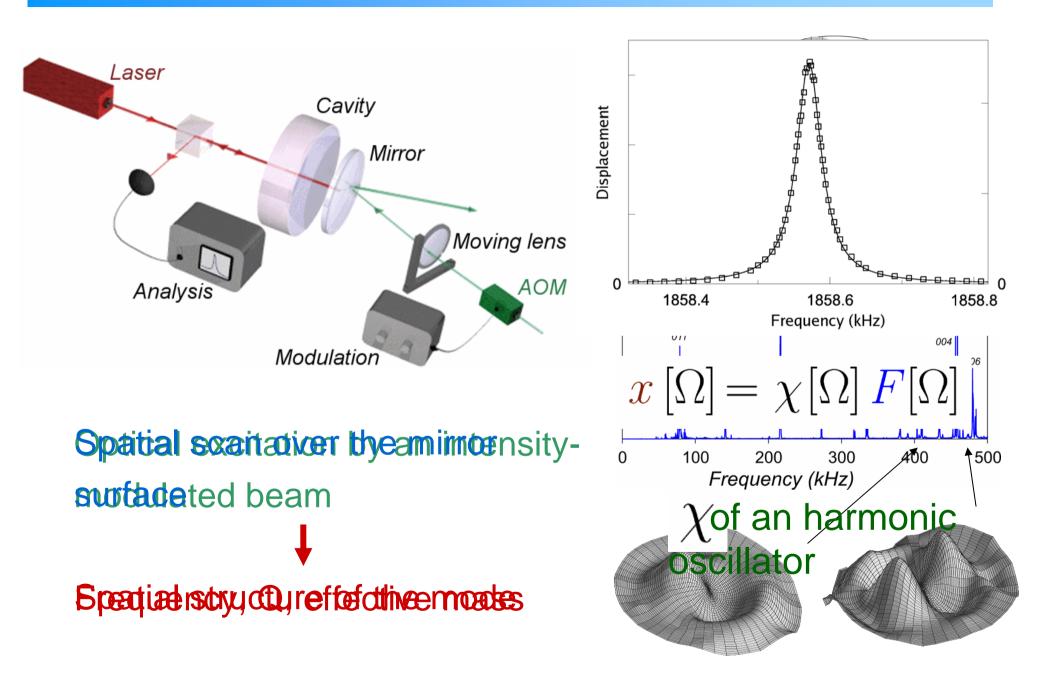
Noises reduction in interferometric measurements

- Exhaustive study of the thermal noise with a high finesse cavity
- Observation of radiation pressure effect back-action noise cancellation
- Sensitivity improvement by cavity detuning (self cooling effect, observation of the cavity instability)

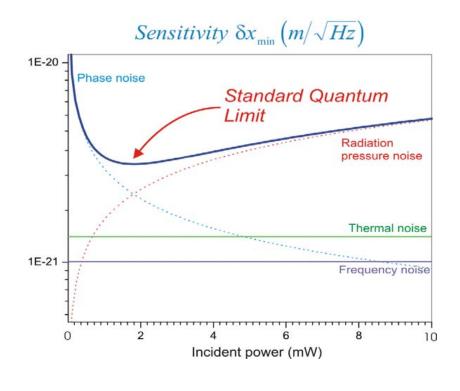
Very sensitive optomechanical sensor



Characterization of internal modes



Observation of the Standard Quantum Limit Test of quantum noise reduction schemes Squeezed light, quantum locking...

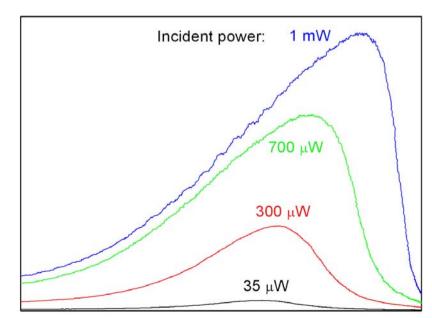


Main experimental improvements:

- Better mechanical response
 M≈10mg Q>10⁶
- Better sensitivity
- Low temperature operation

Current status of the experiment





New cavity with a finesse 230 000

Intracavity intensity increased by a factor of 10

Improvement of the laser frequency locking

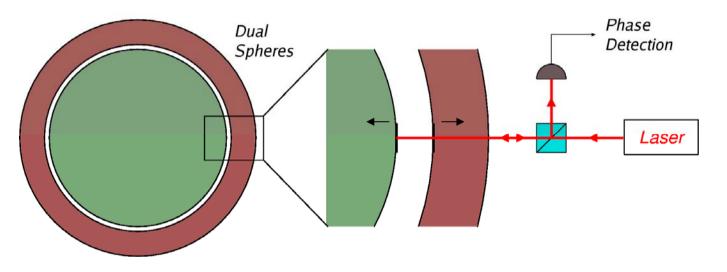
> Cryogenic setup (4K)



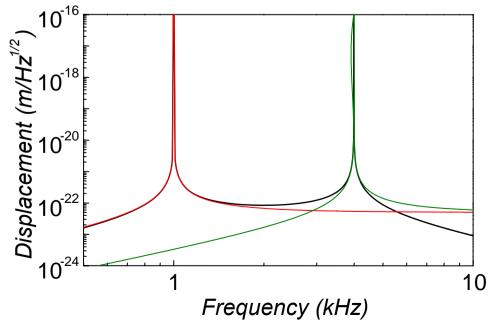
Laser frequency

Back action noise evading

First appears in the context of dual resonant detectors



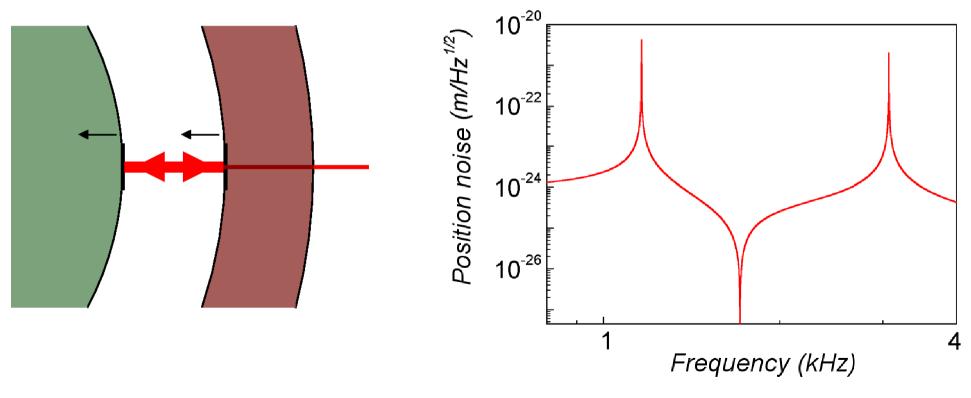
Between fundamental resonances, the responses of the spheres to the gravitationnal wave are opposite



Back action noise evading

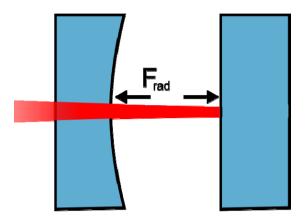
Between resonances, the responses to the radiation pressure are in phase

Back action noise cancellation



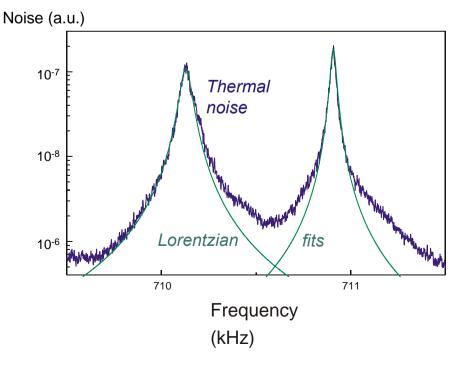
Large improvement of the sensitivity

Radiation pressure noise cancellation



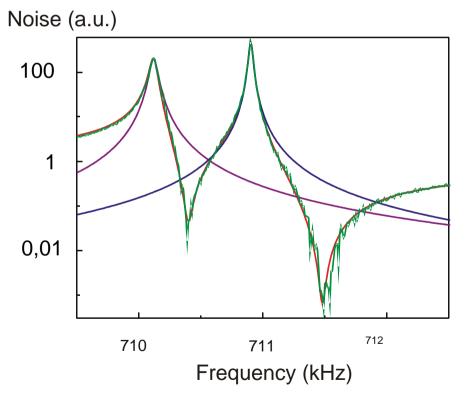
Similar effect in Fabry Perot cavity with two merely identical mirrors

Thermal noise of both mirror are uncorrelated Contributions of each mirror are added



Cancellation of radiation pressure effect

A second cross-polarized laser beam is injected into the cavity



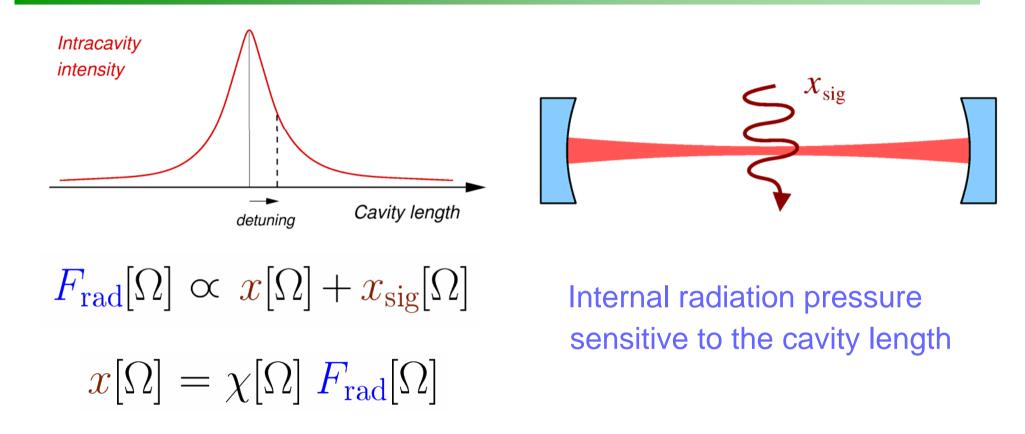
Frequency scan of a classical excitation

Presence of two dips

- Peak-Peak cancellation
- Peak-Baseline cancellation

Noise reduction by a factor of 200

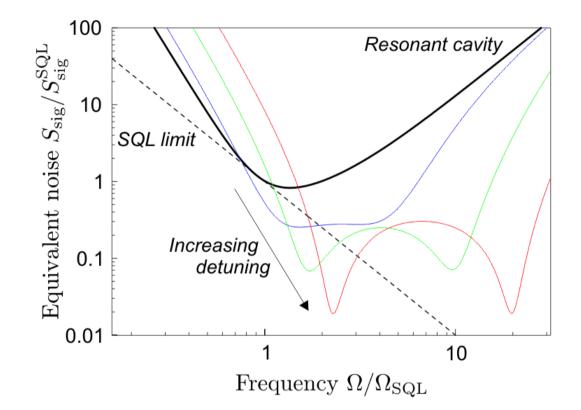
Change of mirror dynamics by cavity detuning



Two effects:

- Amplification of the signal
- modification of the mirror dynamics

Sensitivity improvement in GW detector



Similar effect in a gravitational interferometer with a detuned signal-recycling cavity

Test with our experiment

Effect on the thermal noise

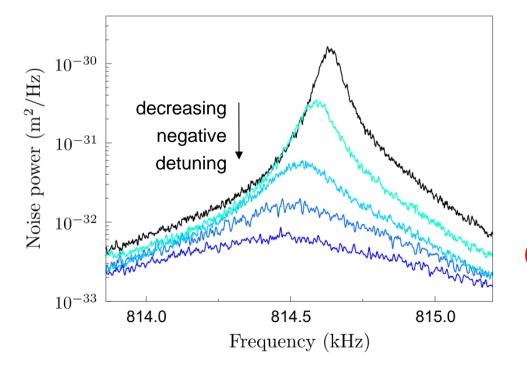
Taking into account the cavity bandwidth, the internal radiation pressure force is dephased

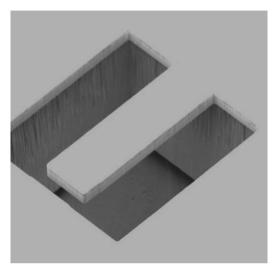
The viscous force modifies the damping rate of the mirror: $\Gamma \longrightarrow \Gamma_{\rm eff}$

Modification of temperature by a factor $\Gamma/\Gamma_{\rm eff}$

Thermal noise reduction by self cooling effect

Experiment with a mirror coated on a micro mechanical system (MEMS)



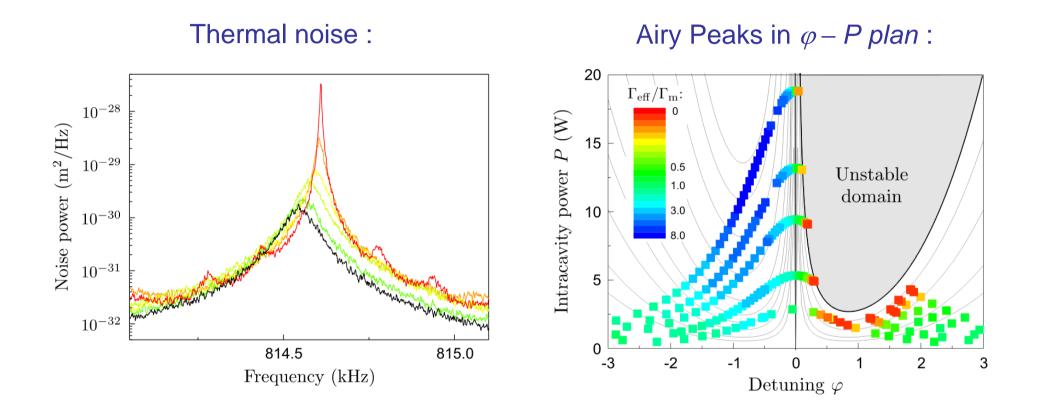


 $1 \text{ mm} \times 400 \, \mu \text{m} \times 60 \, \mu \text{m}$ $M \simeq 100 \, \mu \text{g}$

Reduction of the effective temperature

Cooling of the mode down to 10 K

Observation of the cavity instability



Heating up to 2000 K

First experimental demonstration of a radiation pressure instability in a open optical cavity

Conclusion

What's done : • Exhaustive study of thermal noise

- Observation of radiation pressure cancellation
- Demonstration of self cooling effect and observation of instability
- Next steps : Experimental observation and tests of quantum noise
 - Comprehensive study of quantum noise, search for innovative quantum optics methods