

Max-Planck-Institut Für Gravitationsphysik (Albert-Einstein-Institut)

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Outline

1.1-1

2

Summary of: All-reflective Interferomtry (with multilayer coatings)

Outlook to: High reflective diffractive structures

(without multilayer coatings)





### Motivation #1



3



Interferometers beyond LIGOII (MW of power)

Transmission → Absorption → Thermal problems



All-reflective interferometers:

- Allow opaque test mass materials
- New interferometer topologies

## MI with all-reflective beamsplitter



# Ę

## What kind of gratings?



#### Grating equation:

$$d(\sin\theta_m + \sin\theta_{\rm in}) = m\lambda$$

#### Period defines # of orders & direction



### Coating below

 $d >> \lambda$ : scalar approach  $d \sim \lambda$ : rigorous theories (RCWA, Modal method)  $d << \lambda$ : effective medium theories (ETM)

![](_page_4_Picture_9.jpeg)

#### Coating on top

![](_page_5_Figure_0.jpeg)

A. Bunkowski et al, Applied Optics (in press) http://ao.osa.org/upcoming\_pdf.cfm?id=66481

![](_page_6_Picture_0.jpeg)

![](_page_6_Picture_1.jpeg)

![](_page_6_Figure_2.jpeg)

![](_page_6_Figure_3.jpeg)

Need high efficiency

(which is hard to achieve)

99.62% achieved (Finesse of 1580)

#### 2nd order Littrow

![](_page_6_Figure_8.jpeg)

Only low efficiency

(3 ports are more Complicated)

A. Bunkowski et al, Applied Optics (in press) http://ao.osa.org/upcoming\_pdf.cfm?id=66481

### 3-port couplers

![](_page_7_Picture_1.jpeg)

![](_page_7_Figure_2.jpeg)

### Angle resolved scattering

![](_page_8_Figure_1.jpeg)

### Grating beneath coating

![](_page_9_Figure_1.jpeg)

![](_page_9_Figure_2.jpeg)

T. Clausnitzer et. al., Optics Express, 13, 4370 (2005)

### Grating beneath coating II

![](_page_10_Figure_1.jpeg)

T. Clausnitzer et al, Optics Express, 13, 4370 (2005)

![](_page_11_Picture_0.jpeg)

![](_page_11_Picture_1.jpeg)

#### Summary:

- High quality all-reflective components
- Demonstration of new cavity coupling concept
- Reduction of scattering loss

#### Outlook:

- Further reduction of loss
- Scale to large test mass
- Suspended 10 m all-reflected cavity to be built in Glasgow later this year

### Motivation #2

I.H

Several fundamental noise sources as currently estimated for advanced LIGO:

![](_page_12_Figure_3.jpeg)

S. Penn et.al, LIGO-P050049-00-R (2005)

## **Beating Coating Noise**

![](_page_13_Picture_1.jpeg)

The high refractive material *tantala*— $(Ta_2O_5)$  has been identified to be the main source of *coating thermal noise*.

#### Some ideas so far:

- Less tantala in stack
  - → Innocenzo Pinto's talk
- Doping of tantala with TiO<sub>2</sub>
  → Harry *et al*, Appl. Opt. **45**, 1569 (2006)
- Total internal reflection
  - → Adalberto Giazotto's talk
  - → V. Braginsky and S. Vyatchanin, Phys. Lett. A **324**, 345 (2004)
- Double mirrors:
  → F. Ya. Khalili, Phys. Lett. A 334, 67 (2005)

![](_page_13_Figure_11.jpeg)

![](_page_13_Figure_12.jpeg)

![](_page_14_Figure_0.jpeg)

![](_page_15_Figure_0.jpeg)

see for example: Sharon et al, J. Opt. Soc. Am. A, 14 (1997)

![](_page_16_Picture_0.jpeg)

![](_page_16_Figure_1.jpeg)

![](_page_16_Figure_2.jpeg)

Liu et al, Opt. Lett. 23, 1556 (1998)

## Design width of reflection peak

![](_page_17_Figure_1.jpeg)

d: period g: groove depth r: ridge width s: film thickness r/d: fill factor

![](_page_17_Figure_3.jpeg)

.

### Design width of reflection peak

![](_page_18_Figure_1.jpeg)

A. Bunkowski Nano-structured Optics for GW Detectors 19

Broadband mirror for 1550nm

![](_page_19_Figure_1.jpeg)

A. Bunkowski Nano-structured Optics for GW Detectors

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![](_page_20_Picture_0.jpeg)

### Summary:

- High reflectivity is possible with single layer
  Outlook:
- Check fabrication tolerances, finite size effects...
- Design, build and test gratings
- Think of better suited wave guide structures

### great, greater, grating

T. LAR