

QRPN Experiment with Suspended 20mg Mirrors

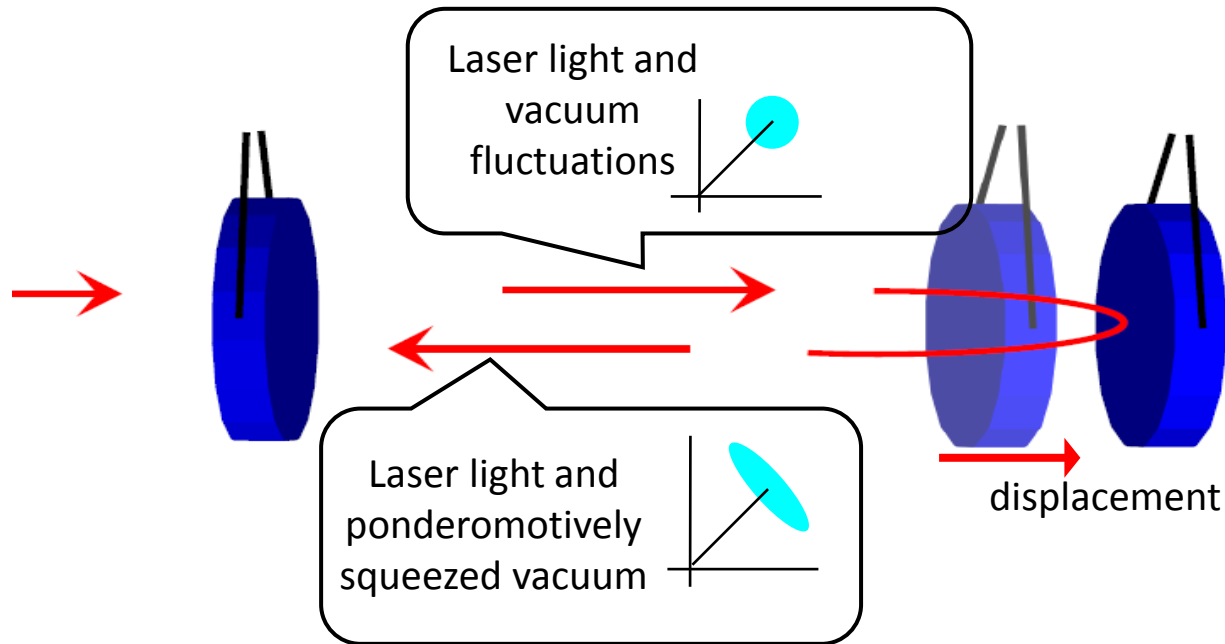
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K. Agatsuma, S. Sakata, T. Mori, S. Kawamura

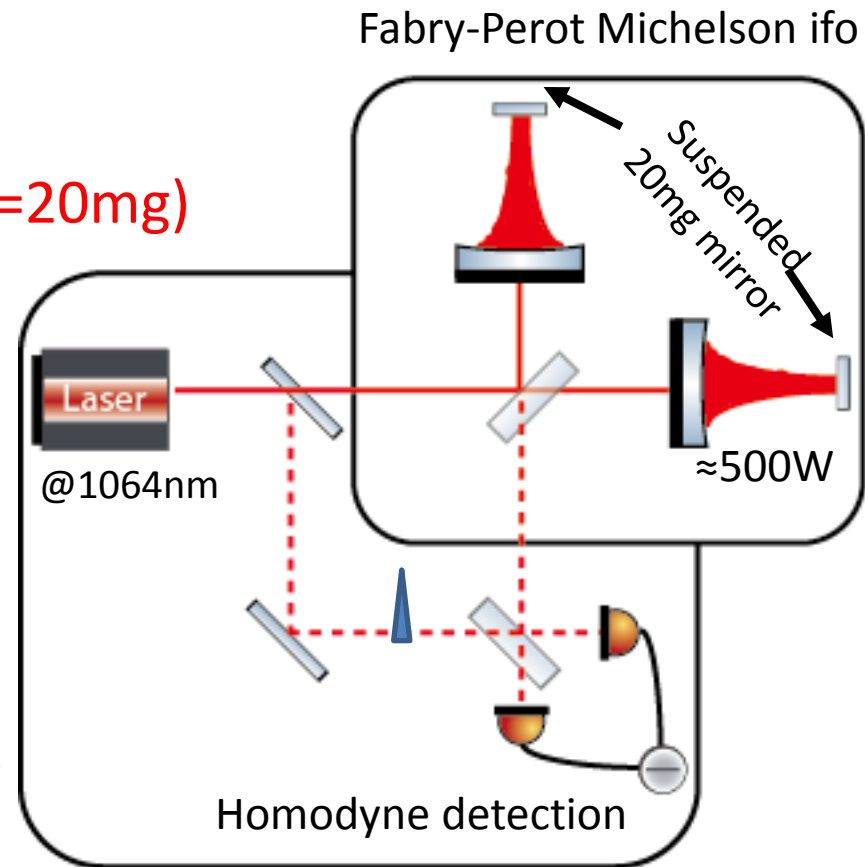
Objectives

- Observation/Reduction of Quantum Radiation Pressure Noise
 - Back-action correlates amplitude and phase fluctuations
 - Ponderomotive squeezing

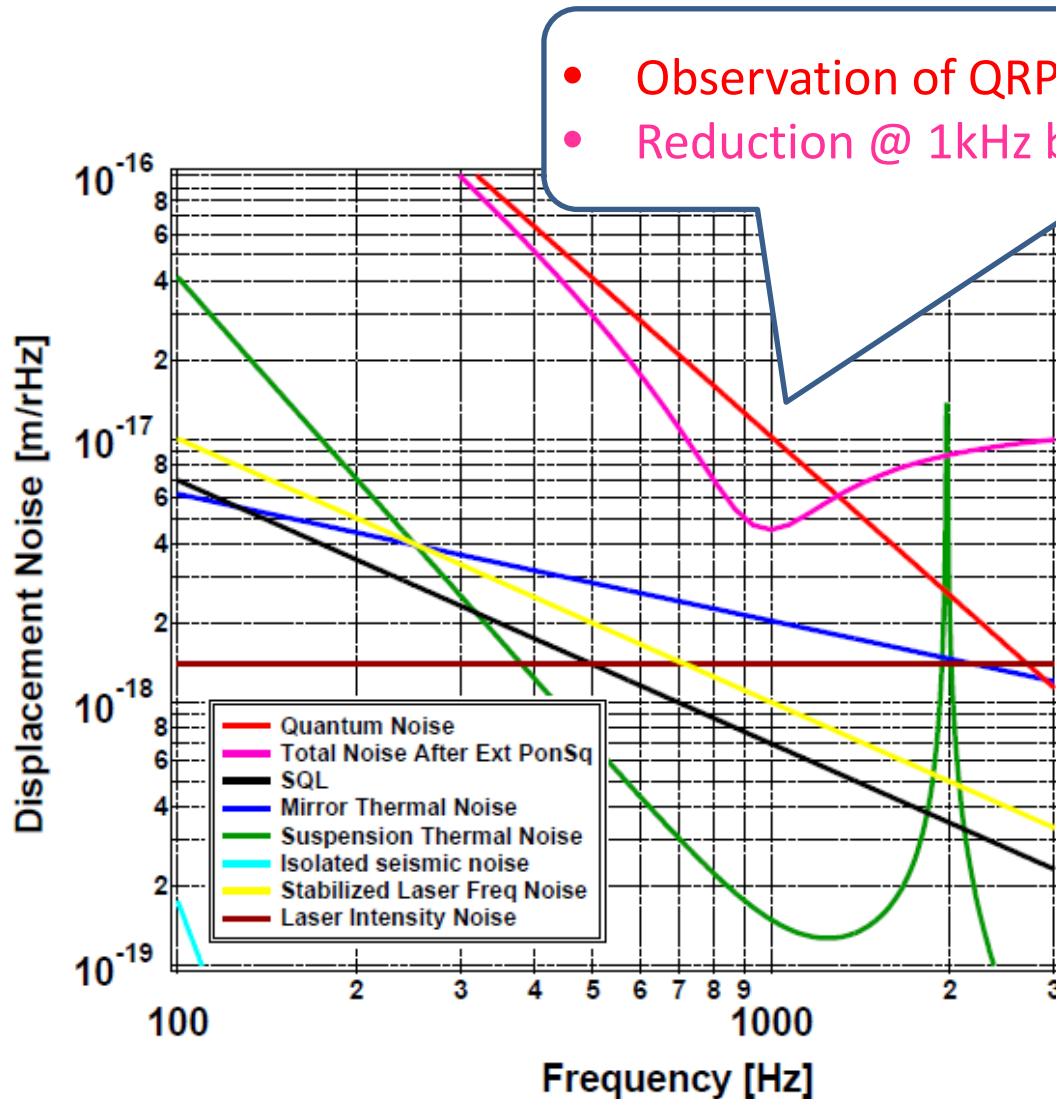


Approach

- Topology
 - Fabry-Perot Michelson interferometer
 - Homodyne detection
- Key elements
 - High Finesse cavities ($F=10000$)
 - **Tiny suspended end mirrors ($m=20\text{mg}$)**



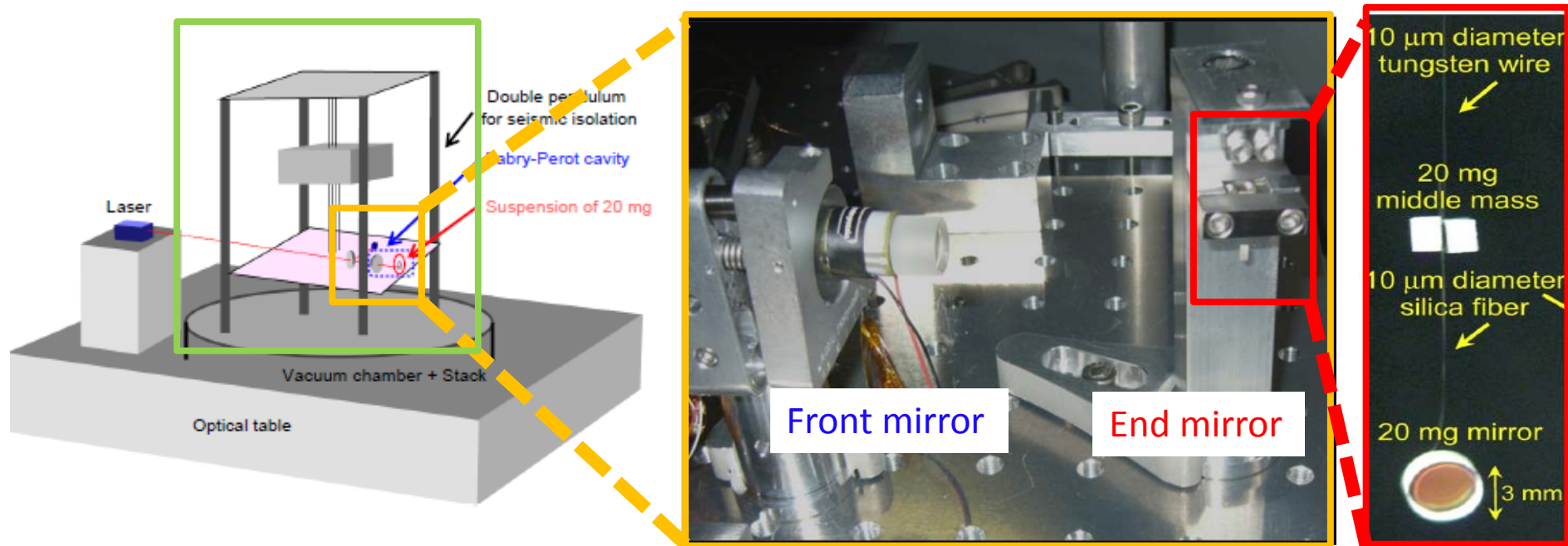
Design Sensitivity/Parameter



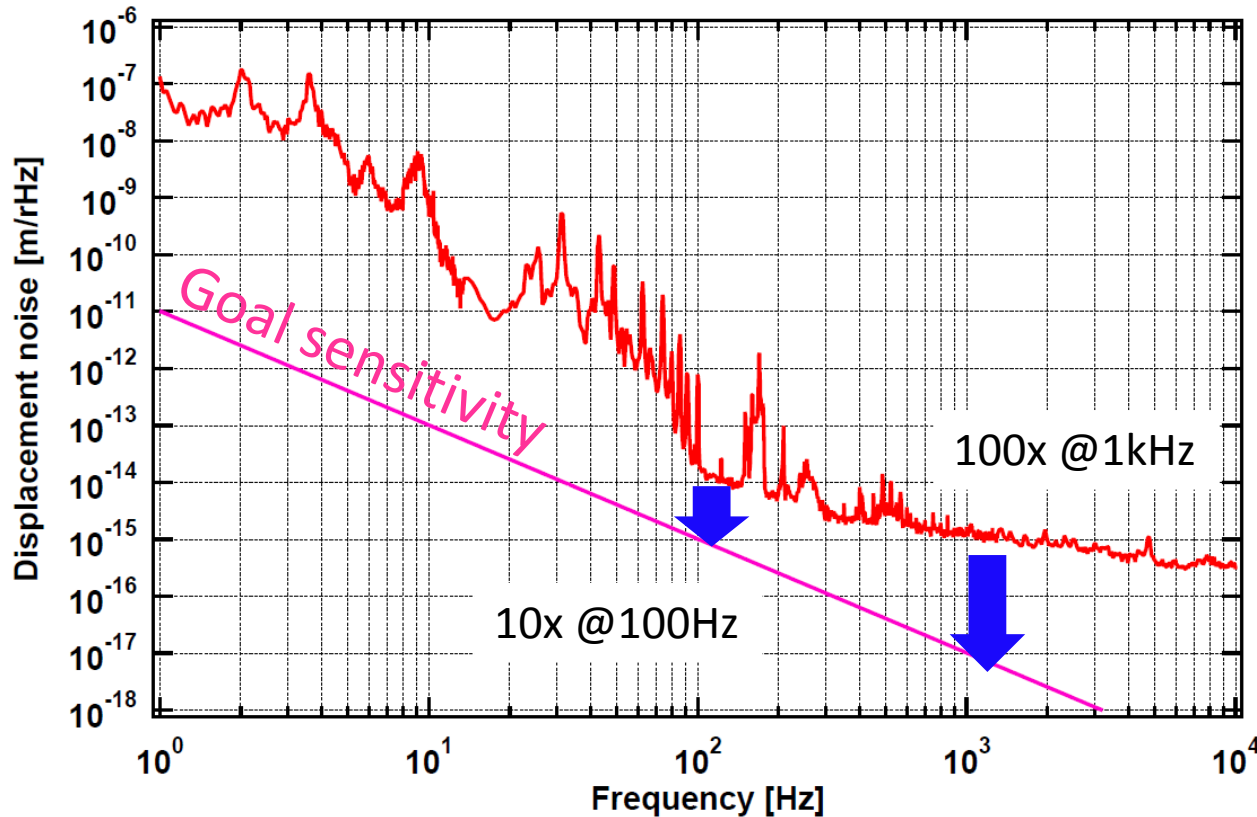
Laser power	200 mW
Injected laser power	120 mW
Finesse	10000
End mirror mass	20 mg
Diameter of end mirror	3 mm
Thickness of end mirror	1.5 mm
Front mirror mass	14 g
Reflectivity of end mirror	99.999 %
Reflectivity of front mirror	99.94 %
Optical loss	50 ppm
Beam waist of end mirror	340 μ m
Mechanical loss of substrate	10^{-5}
Mechanical loss of coating	4×10^{-4}
Length of silica fiber	1 cm
Diameter of silica fiber	10 μ m

Initial Experimental Setup

- **Seismic isolation** for core optics is provided by a double pendulum
- **Fabry-Perot-cavity** under vacuum
- **Front mirror**: fixed mirror mounted on a Piezo actuator
- **End mirror**: 20mg mirror suspended by a 10 μ m diameter silica fiber



Achieved Sensitivity



Cavity parameter:

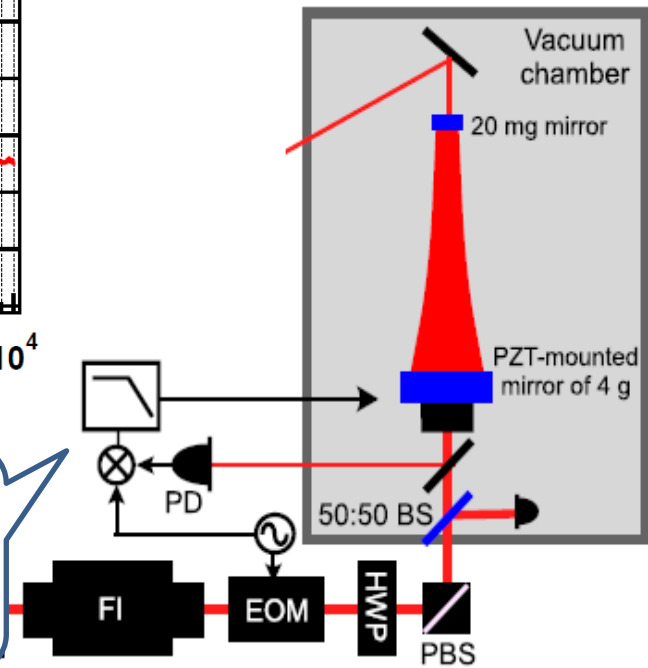
Finesse = 1400

$L = 70\text{mm}$

$\text{RoC}_1 = 2\text{m}$

$\text{RoC}_2 = \text{inf}$

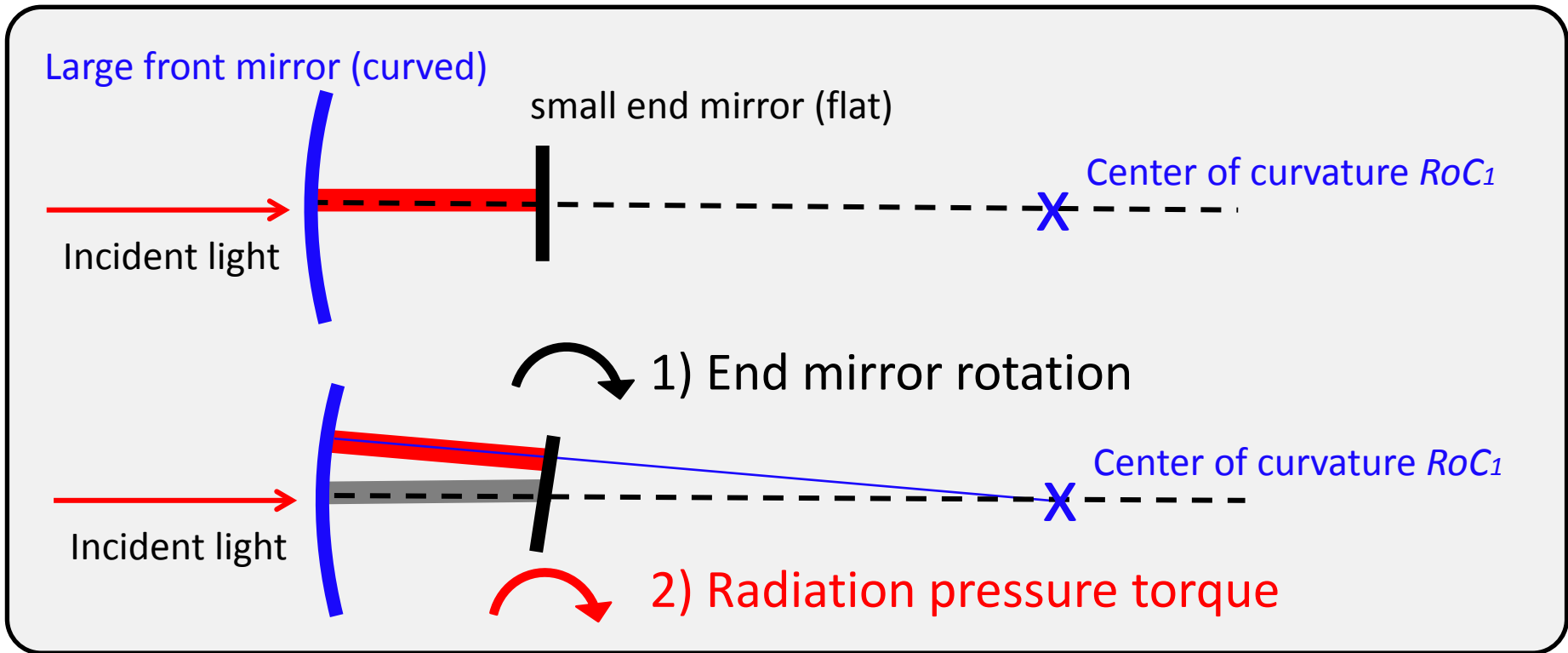
- Stabilized via PDH-scheme
- Sensitivity via calibrated feedback signal



S.Sakata, GWADW (2008)

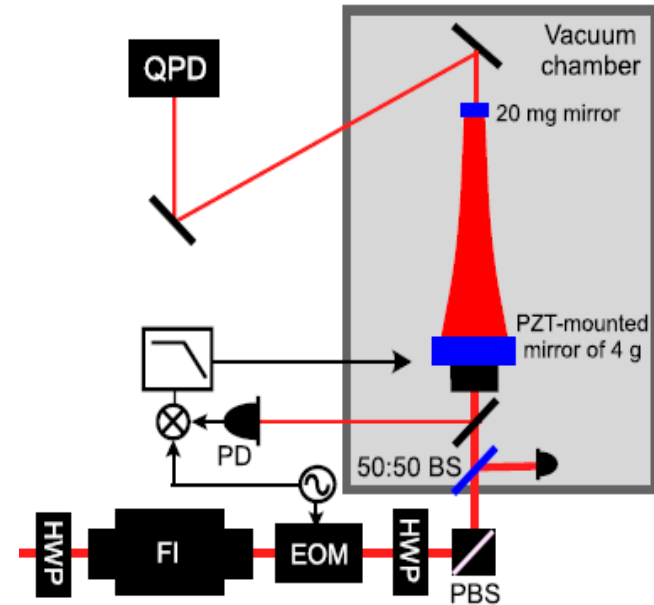
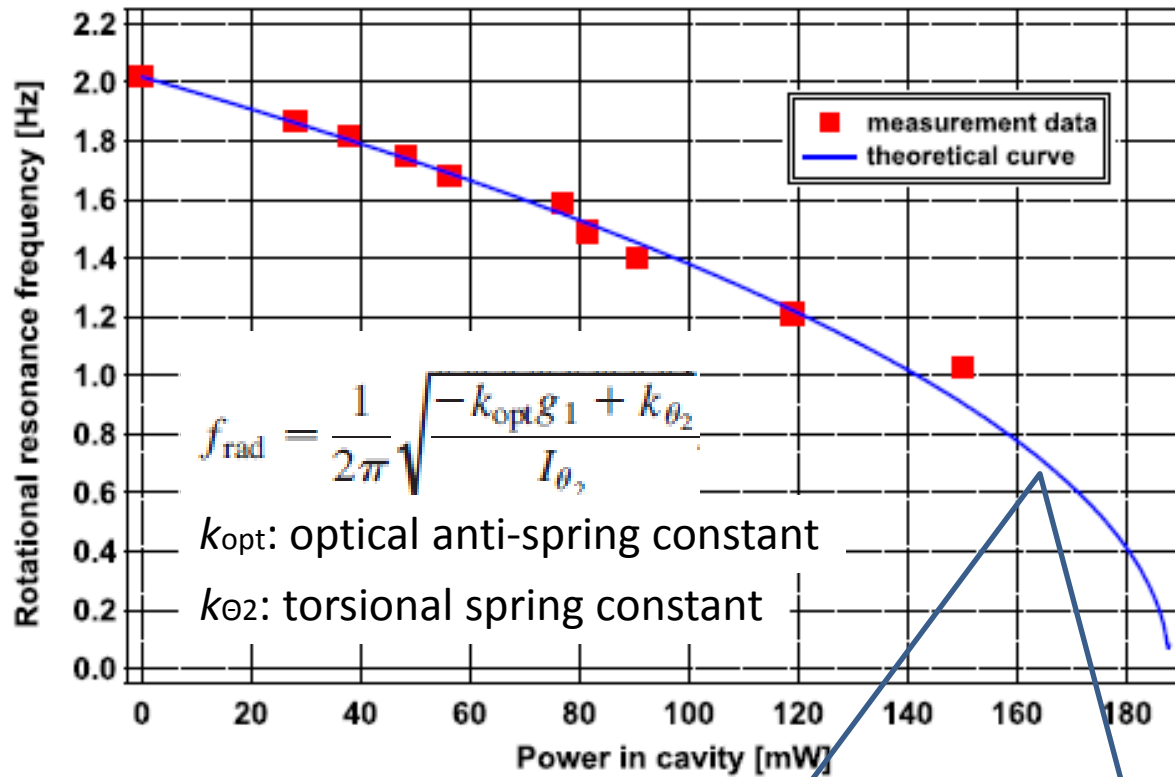
Angular Anti-Spring in Suspended Cavities

- Optical axis is shifted by mirror rotation (yaw/pitch)
 - Torque by radiation pressure
 - For high circulating power: optical torque > restoring force
 - Cavity becomes instable (angular anti-spring effect)



[see e.g. J.A.Sidles, D.Sigg, Phys. Lett. A 354, 167 (2006)]

Measurement of Angular Anti-Spring (Yaw)



- Optical anti-spring increases with laser power
- Observed via a decrease in yaw eigenfrequency
- Data was taken until near instability ($\approx 190\text{mW}$)

Cavity parameter:

Finesse = 1400

$L = 70\text{mm}$

$R_1 = 2\text{m}$

$R_2 = \text{inf}$

How to Cope the Angular Anti-Spring

Optimize the cavity geometry:

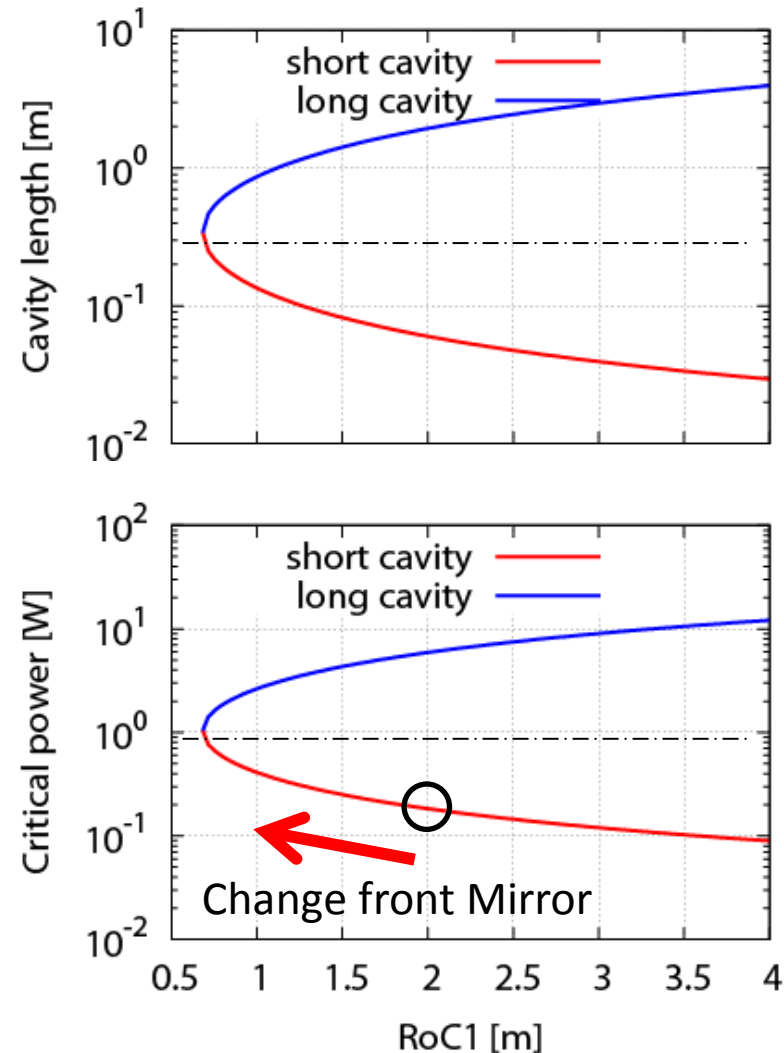
- Change cavity g-factors
- Change RoC of the front mirror

• Restriction:

- Using a flat end mirror*
- Beam radius of $340\mu\text{m}$ (coating TN)
- Reasonable cavity length $\leq 0.3\text{m}$

• Result:

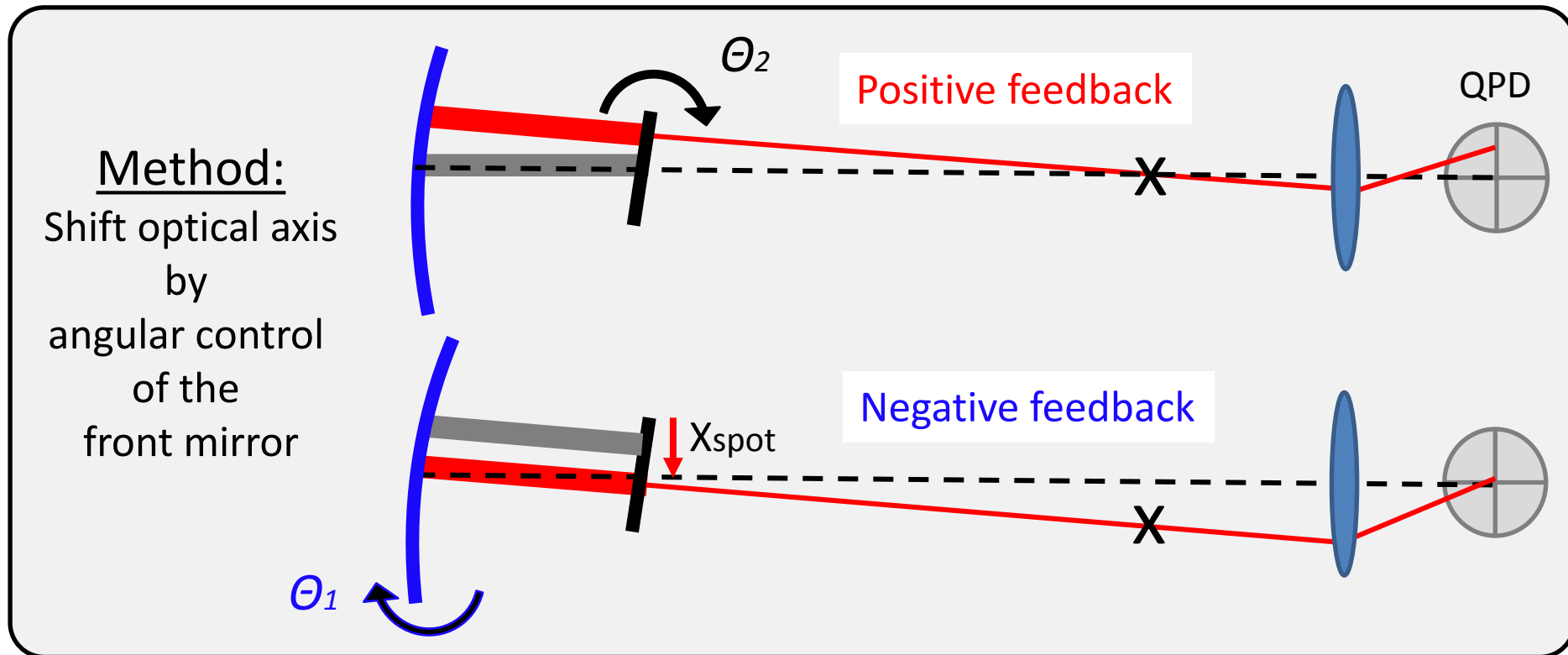
- Going to smaller RoC1
- Critical power still $< 1\text{W}$
- Circulating Power in final design $\approx 500\text{W}$
- **Angular control is required!**



* A curved 20mg mirror is advantageous, but not available so far

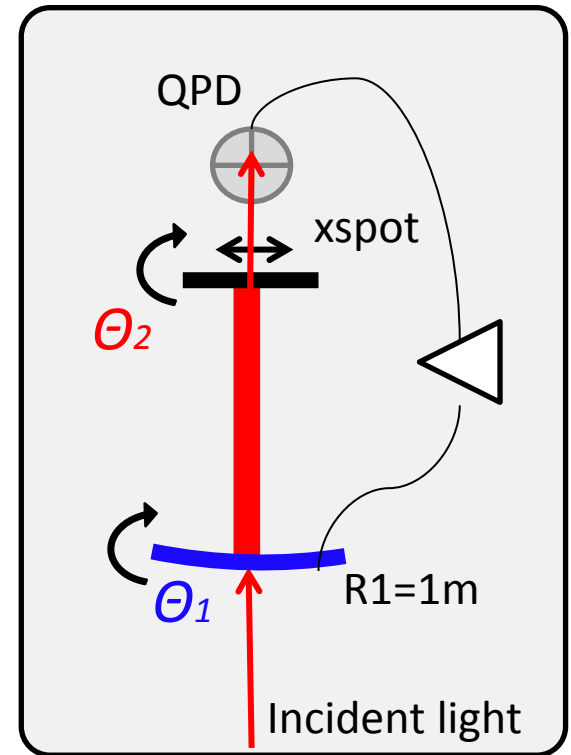
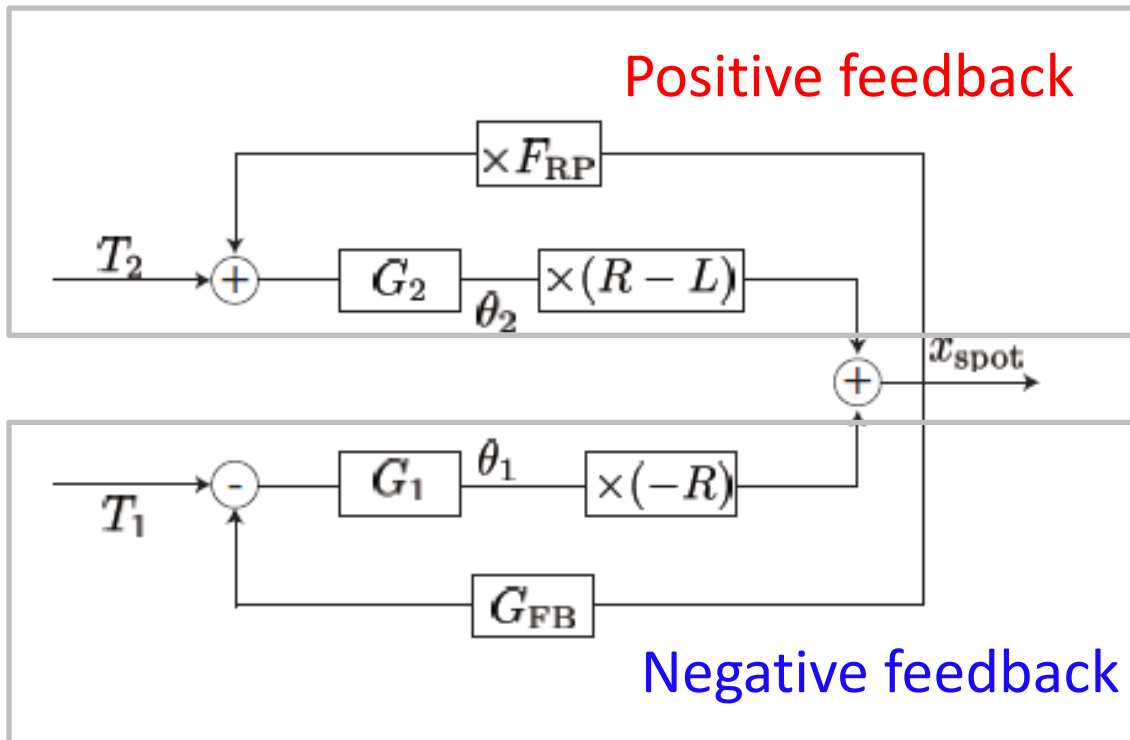
Angular Control – Idea

- Difficult to apply an external force to the small mirror directly
 - Idea: Control the small end mirror via radiation pressure



- Error signal is derived via a quadrant diode (QPD)

Angular Control - Model



T_i : torque, G_i : pendulum transfer function

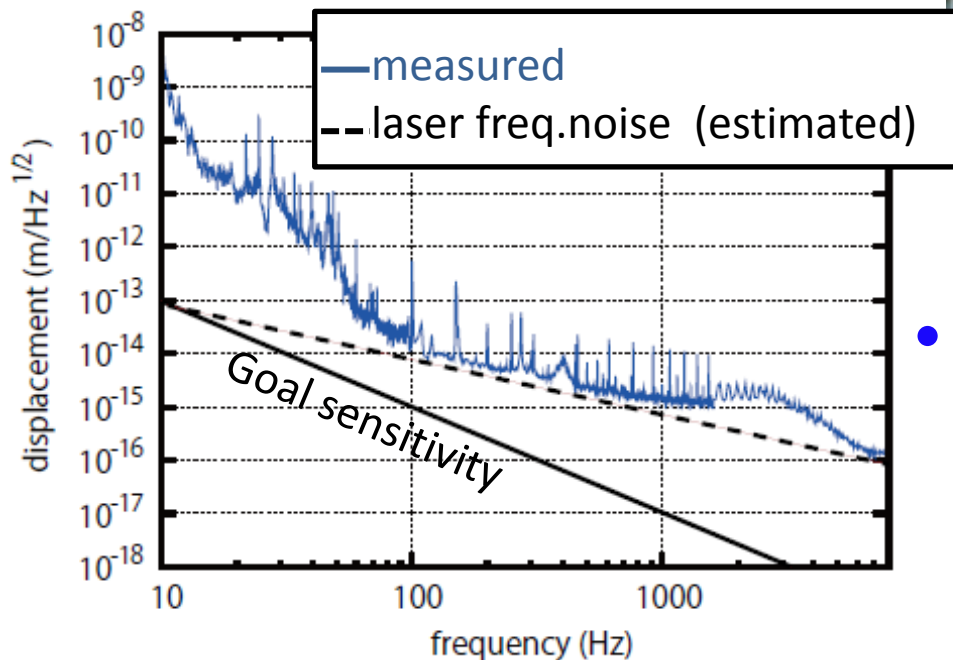
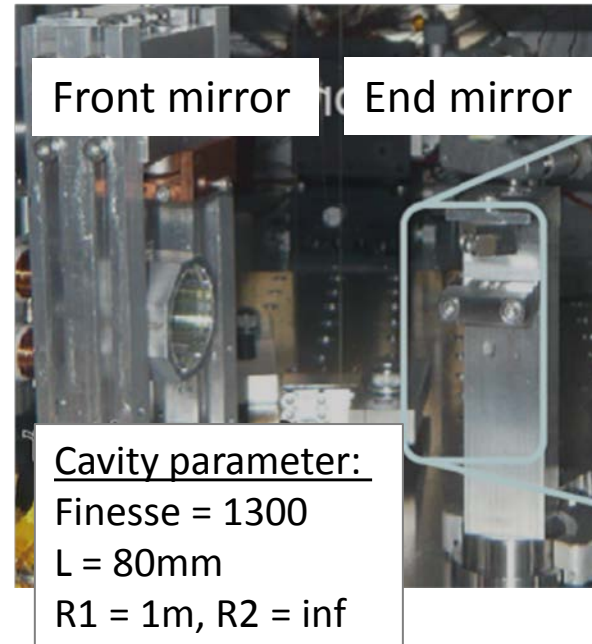
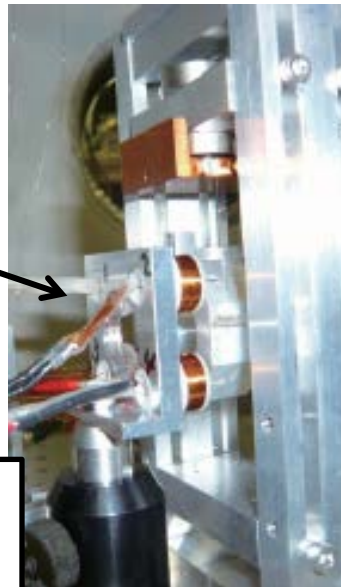
- **Model for angular control to suppress x_{spot} (for final design):**
 - G_{FB} : can provide sufficient gain at low frequencies (zero/pole @ 200/500Hz)
 - Need to suppress electronic noise around 1kHz (yaw to longitudinal coupling)

S. Sakata, internal document (2008)

T.Mori, Ph.D. thesis, The University of Tokyo (2012)

Results with a Fully Suspended Cavity

- Suspended front mirror
- coil-magnet actuators
- increase the actuation range
- **stable longitudinal lock (low power)**
- sensitivity comparable to prev. setup



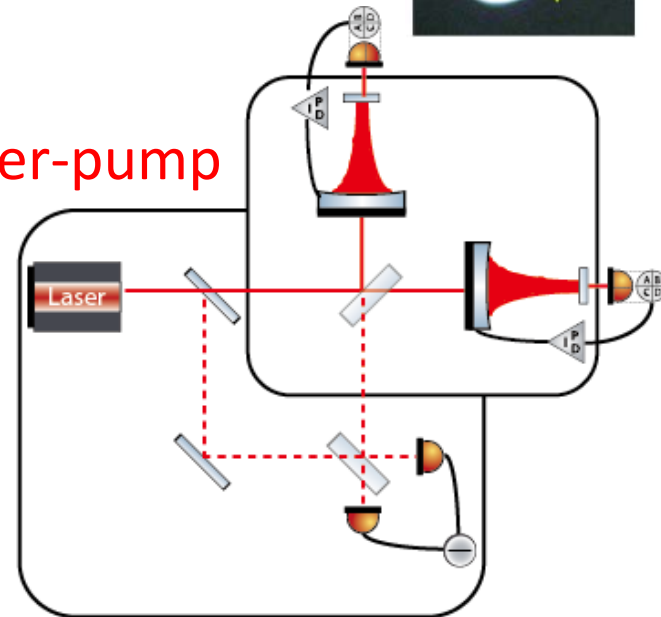
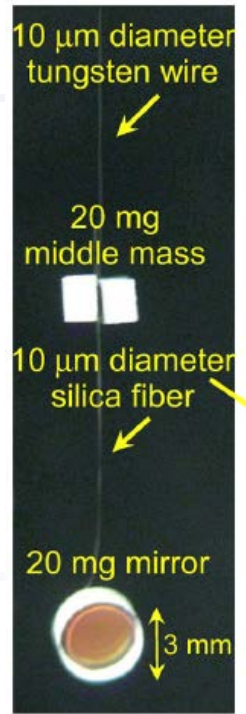
- Angular control via front mirror
 - control circuit was installed
 - **stabilization of angular anti-spring via RP not demonstrated yet**

Summary

- Goal: Observation and Reduction of QRPN
 - Key element: **20mg mirror suspended by a 10 μ m silica fiber**
- Fully suspended cavity locked stable (at low power)
 - Model/setup for angular control via radiation pressure

Outlook

- Demonstrate **angular control** (with preliminary setup)
 - Investigate feasibility for the final design
- Upgrade the vacuum system by an **ion-getter-pump**
 - More stable working conditions
- Setting up the **full interferometer**
 - Common noise rejection



Thank you



Arrival at ICRR, March 2012

