## MIGO <br> Example and experience using SIS for aLIGO designing and commissioning

- Field Calculator using FFT + integrated analysis tool
- Field
» FP, CC with stable recycling cavity and BS
" Lock using error signal
" Telescope with arbitrary number of elements - lens, mirror, space
- Analysis tool
" Mode analysis
" Phase map to PSD, PSD to phasemap
" Flexible map manipulation
- Subtract / add zernike terms from / to phasemaps
" Hello-Vinet thermal models


## LIGO Optical configuration and source only RC + one arm

## Main interface



## A simple simulation senario

- Interferometer spec
" ITM.opt.HR_phase $=$ THERMOELASTIC( w on ITM, abs in ITM, abs in coating) + DATAFILE("ITM01.dat",-5)
- Action
" Lock => lock and calculate field
" Summary
- ETM HR

```
(in base):w=0.0619634 R=2245 z =2160.28 z0 = 427.8068 w0 = 0.01203704
(in fit ) : (wX,wY)=( 0.06195, 0.061667) R(x/y) = (2248.252, 2251.826) (x0,y0)=(0.02517 , 0 )
power / HMfrac = 228.6095 / 0.1827
(out fit ) : (wX,wY)=( 0.06195, 0.061667) R(x/y) = (-2251.263, -2247.691) (x0,y0)=(-0.02517,0 ) power /
HMfrac = 228.5973/0.1853
```

" Mode analysis

- Hermite or Laguere mode expansion with arbitrary number of modes
" Modify map and repeat for the new configuration


## LMA ETM01 coating accepting test short wavelength spiral pattern

Measured at Caltech - Z1~Z17

P-V 1nm


Using matlab to extract the spiral pattern, and use it as the phasemap in SIS

## LMA ETM01 coating accepting test short wavelength spiral pattern

- SIS analysis to understand the effect by this pattern
- Round trip loss ~6ppm $\longleftarrow$ OK
- Any other effects
" Field aberration due to this pattern
- Field in FP with this map - Field in idealistic FP ${ }^{0.6}$
- Very fine grid sizes to make sure FFT is OK
" Mode analysis if any mode could dominate
- No dominant mode for LGpm (2p+m<25) and HGmn ( $\mathrm{m}+\mathrm{n}<25$ )
" If ITM has similar pattern, can they interfere
- ITM = MAPPING
( DATAFILE("ETM01pattern.dat"), "-x","y") * 0.5

- Loss = loss by ETM + loss by ITM no additional by interference
LIGO-G1300054-v2


## LIGO <br> LMA ETM01 coating accepting test long wavelength central plateau

- Old coating system, one at a time
" The beam size on ETM is larger than that on ITM and the plateau size on ETM needs to be 20\% wider, when coating to coating variation is taken into account
- New coating using the planetary system, a pair at a time
" Higher order mode, mostly LG20, in the FP cavity is $\sim 100 \mathrm{ppm}$
- Better than old, 120ppm, and two ETMs will be "identical", but is this good enough?
- The plateau size is around the same as the old one
- Astigmatism uncertainty due to the substrate is not a major issue
- Asymmetry in the far outside is better (smaller) in the new coating
" Coupled cavity simulation
- LG20 in PRC is ~2000ppm increase by the ETM coating aberration
- LG20 in SRC shows no increase of LG20 by the mode healing
- Stable signal recycling cavity kills LG2O in SRC

Cross check by other tool is appreciated very much

## LIGO ETM01 (new using planetary) vs ETM06 (old)

Tilt, Power, Astigmatism subtracted


2013 January Commissioning Workshop

## LMA ETM01 coating accepting test long wavelength central plateau



## Why ROC(ITM) < ROC(ETM) Power loss on RM3



## Loss function not so beautiful mirror maps

$$
1-\frac{P(a \sin (2 \pi \cdot f \cdot r))}{P(\text { no aberration })}
$$




2q6m So2a


