



Content:

- **1. Thermal Correction System**
 - a. Ring heater (for Giacomo Ciani)
 - **b.** Adaptive Optical Element (for Paul Fulda)
- 2. Alignment Sensing and Control
 - a. Fast beam pointing (for Daniel Amariutei and Paul Fulda)

Guido Mueller University of Florida @ LVC Meeting Nice, March 2014

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Ring Heater







aLIGO RH:

- NiCr wire around glass rod
- Thermal inhomogeneities due to
 - heat sinks
 - winding densities
- brittle (but now installed)

Requirements:

LIGO

- Astigmatism in reflection (AC side) < 3nm over ~8cm
- In transmission (RC side): Undefined



UF RH:

- Alumina coated Al
- NiCr wires clamped
- Flexible
- Compatible with current shield
- Still need final RGA scan

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Ring Heater





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Adaptive Optical Element

Useful in two places:

LIGO

- » Mode matching from IMC to PRC.
 - Must deal with high transmitted power (>50W).
 - Required optical power < $1/zR_{IMC} \sim 0.1$ D.
- » Mode matching from SRC to OMC
 - Low transmitted power.
 - Required optical power ~0.15 D [1].
- Heat load to HAM table should be <1W.
- Should not add higher-order distortions to transmitted beam (mode purity >99%).
- Vacuum compatible, reliable actuator.

[1] LIGO-G1301154 S. Ballmer





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Testing in vacuum: DBB mode purity and range measurement

- Diagnostic breadboard of aLIGO PSL design used.
- » Consists of triangular PMC cavity with PZT mirror.
- Scan cavity and observe higher-order mode peaks.
- Perform beam scan to find beam parameters.
- Actuate AOE at different levels, re-mode match and repeat scan, observe change in HOM content.
- Calculate AOE focal length.



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Adaptive Optical Element



Positive results shown at ~-0.05D (f~-20m) actuation:



Mode purity actually increased by astigmatic AOE heating.

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Adaptive Optical Element



Conclusions and next steps

- Design performed well in vacuum from mode purity and dynamic range perspective.
- Power in too much (>10W to reach ~0.5D) though.
- Implement improvements in design:
 - » Use optics with polished barrel.
 - » Use reflector to direct radiated heat to tank walls rather than bench.
 - » Try micro-etching heater material (NiCr) directly on optic barrel (alternative design).
- Run up to high-power again with recent alterations check power efficiency.
- Test for full vacuum compatibility at Caltech facility.





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Case studies:

LIGO

Thermal substrate lens

• Changes PRC mode

• Changes MM into AC from IMC

Note:

The mode matching between PRC and AC is fairly insensitive to substrate thermal lens

Assume no RH, no CO2

Try to maintain MM from front end with different power levels

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Case studies:

LIGO

Thermal substrate lens

• Changes PRC mode

• Changes MM into AC from IMC

Note:

The mode matching between PRC and AC is fairly insensitive to substrate thermal lens

 $1\mu D \sim 1W$ input power

Assume no RH, no CO2

Try to maintain MM from front end with different power levels requires differential operation

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How it works: n₀–δn 🛉 Ε, n₀+δn ↓E, n₀–δn θ_{int} B. n_0 A С В Idealized!

Fast Beam Pointing

Motivation:

LIGO

- Enable larger bandwidth in alignment servo systems
- Generate 10-mode at RF-frequency to enlarge design space for alignment sensing

 $\langle \alpha_1 \rangle$ n_o+δn • electro optic effect:

$$\delta n = \frac{n_0^3 r_{33} E}{2} = \frac{n_0^3 r_{33} V}{2d}$$

• changing the voltages will change the deflection angle:

$$\theta = \frac{2L\delta n}{h} = \frac{n_0^3 r_{33} LV}{hd}$$

Fast Beam Pointing

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Fast Beam Pointing

Testing it:

LIGO

LiNbO₃ crystal $n_0 = 2.232$ $r_{33} = 32.8 \text{ pm/V}$

• Polarization: no observed change

- the electrodes are milled in the Cu layer of PCB, with a 55 µm width gap between the two interposed rows
- thin layer of dielectric lacquer applied (dielectric strength 48 KV/mm)
- black delrin holder and spacers (dielectric strength 5.735 KV/mm)

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Using it...

Potential applications:

- 1. Fast feedback actuator for ASC
 - Possible location: from PSL table into IMC
 - Note: We would need four elements
 - two dimensions
 - two Gouy phases
- 2. Generate ASC signals for simple cavity
 - i. PDH-style:
 - Generate 10-mode
 - Reflect off cavity
 - Detect with single element PD
 - ii. Anderson-style:
 - Generate 10-mode to be resonant in cavity
 - Detect in transmission with single element PD
- 3. Generate ASC signals in coupled cavities/Advanced LIGO setup
 - Finesse modeling ... (no idea yet if that works/helps)

Fast Feedback Actuator

One-dimensional experiments:

Diagonalize WFS:

- BD1 -> QPD1
- BD2 -> QPD2

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Fast Beam Pointing

16 BD1 **Diagonalize WFS:** 14 BD2 • BD1 -> QPD1 12 **Finesse** WFS [W/rad] 9 & 01 • BD2 -> QPD2 Tuning of **QPD1** location 2 0 200 400 1200 1400 600 800 1000 1600 1800 2000 Distance from Gouy telescope [mm] WFS distance scan BD1 0.08 Qualitative agreement BD2 0.07 pretty good. 0.06 ∑ 0.05 0.04 Experiment Hope to close loops soon 0.03 0.02 0.01 500 600 700 800 200 300 400 900 1000 Distance [mm]

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Alignment Sensing

Finesse simulations for an aLIGO IMC-like linear FPC

Hope to be able to look at these ideas again since we are done with IO ...

One 10-SB on resonance in reflection allows to separate misalignment or ETM and ITM (single cavity)

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