

The LIGO logo consists of several concentric, overlapping circles of varying shades of gray, creating a ripple effect that suggests gravitational waves.

LIGO



A Preview of Future Cryogenic Suspensions for aLIGO Upgrades

Brett Shapiro

Stanford cryo people:

Brian Lantz, Tim MacDonald, Dakota
Madden-Fong (summer '13)

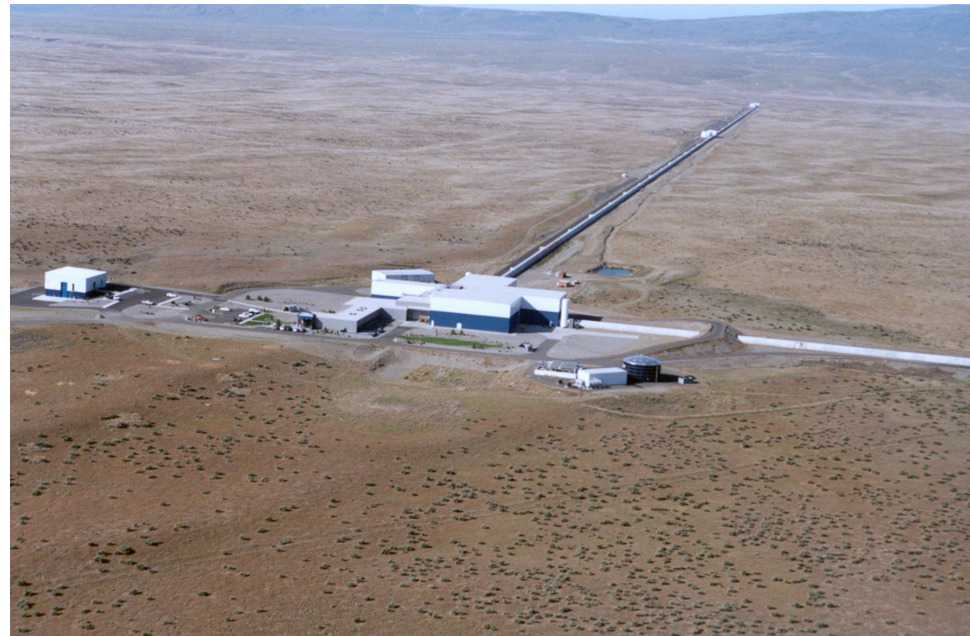
Summary

- Moving from Advanced LIGO to LIGO III
- LIGO III quad suspension design
- Steady state cooling (science mode)
 - System layout
 - Length of heat shield extending into beam tube
- Initial Cool down
 - Stanford initial cool down experiments
- Future work

Advanced LIGO Timeline



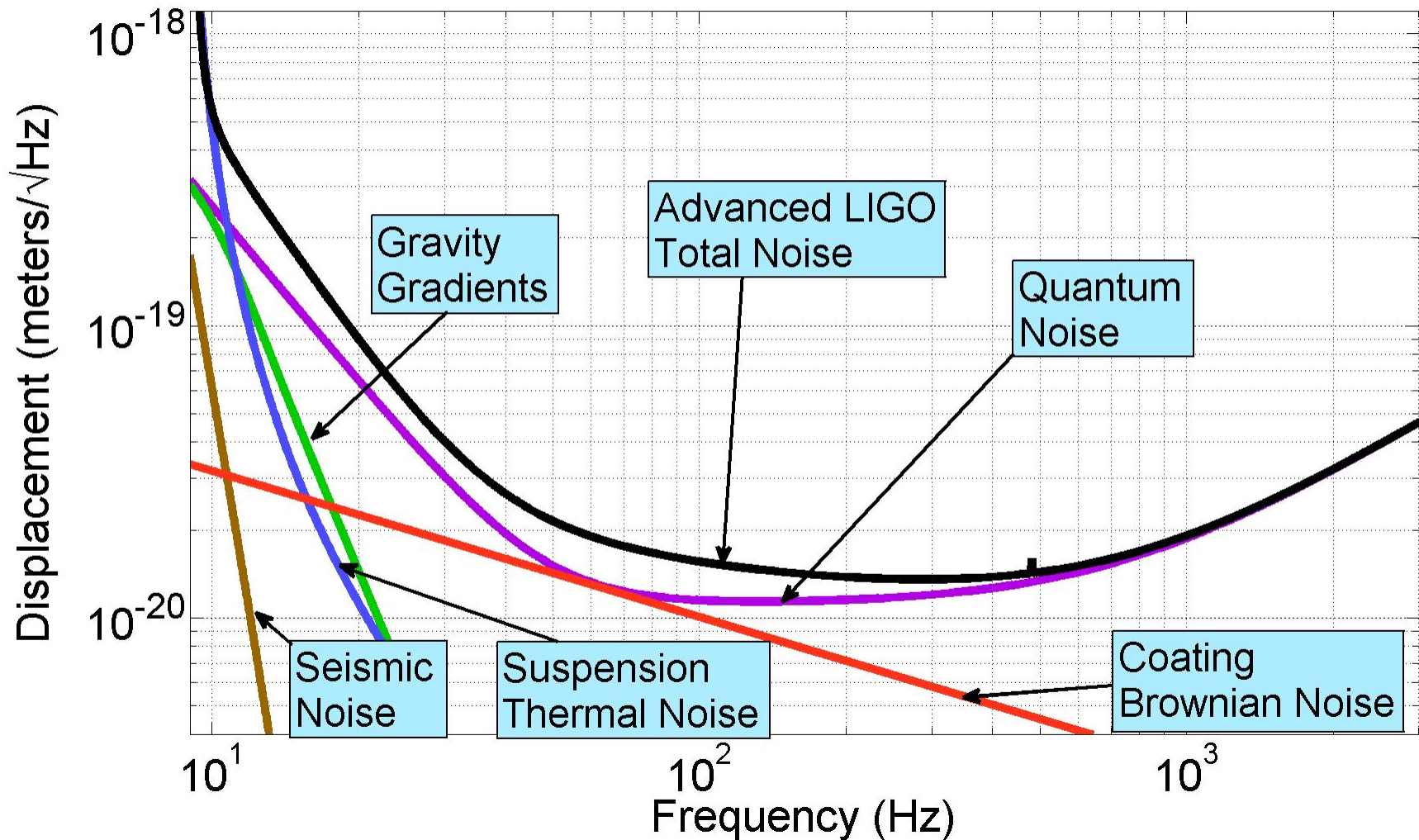
Livingston, LA



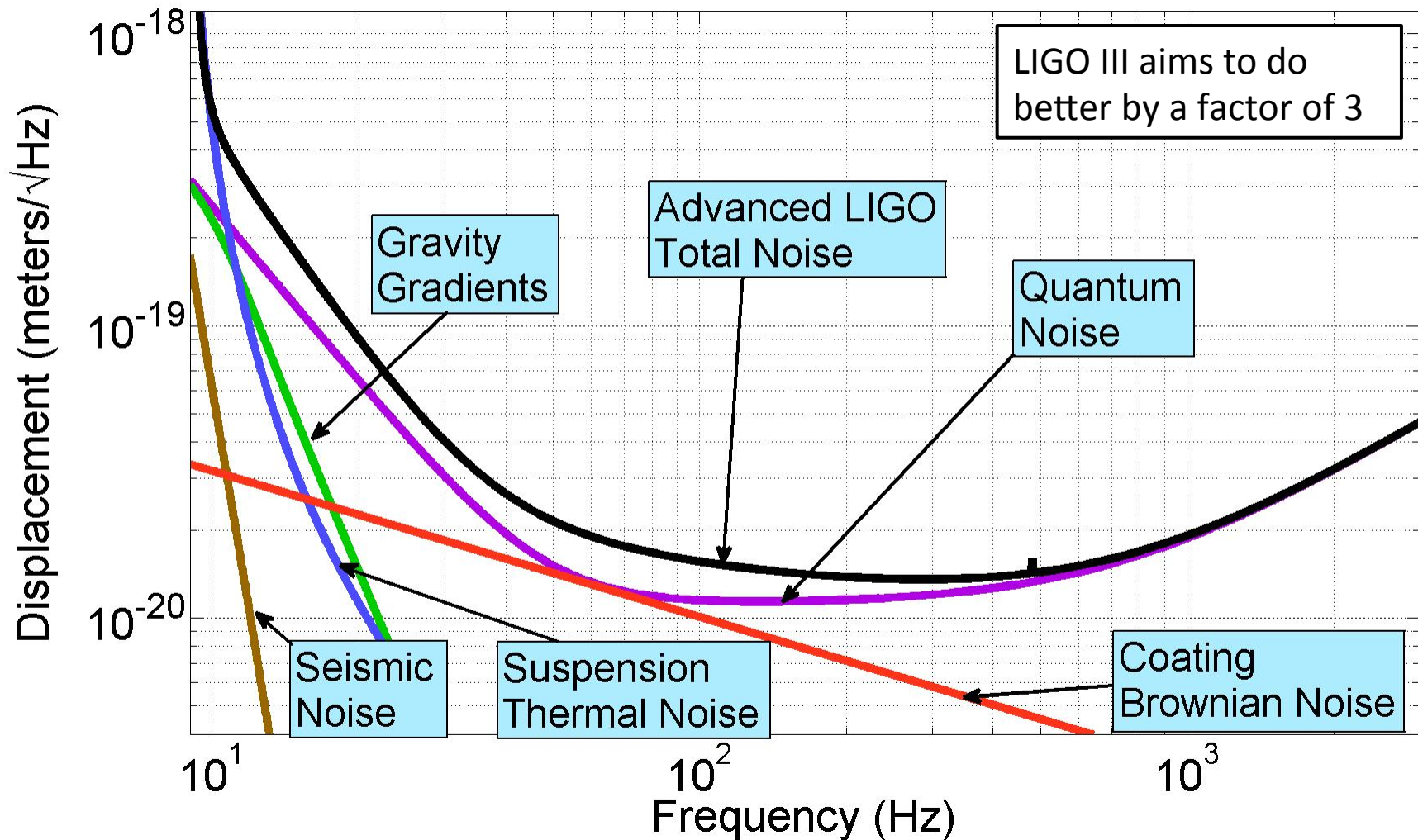
Hanford, WA



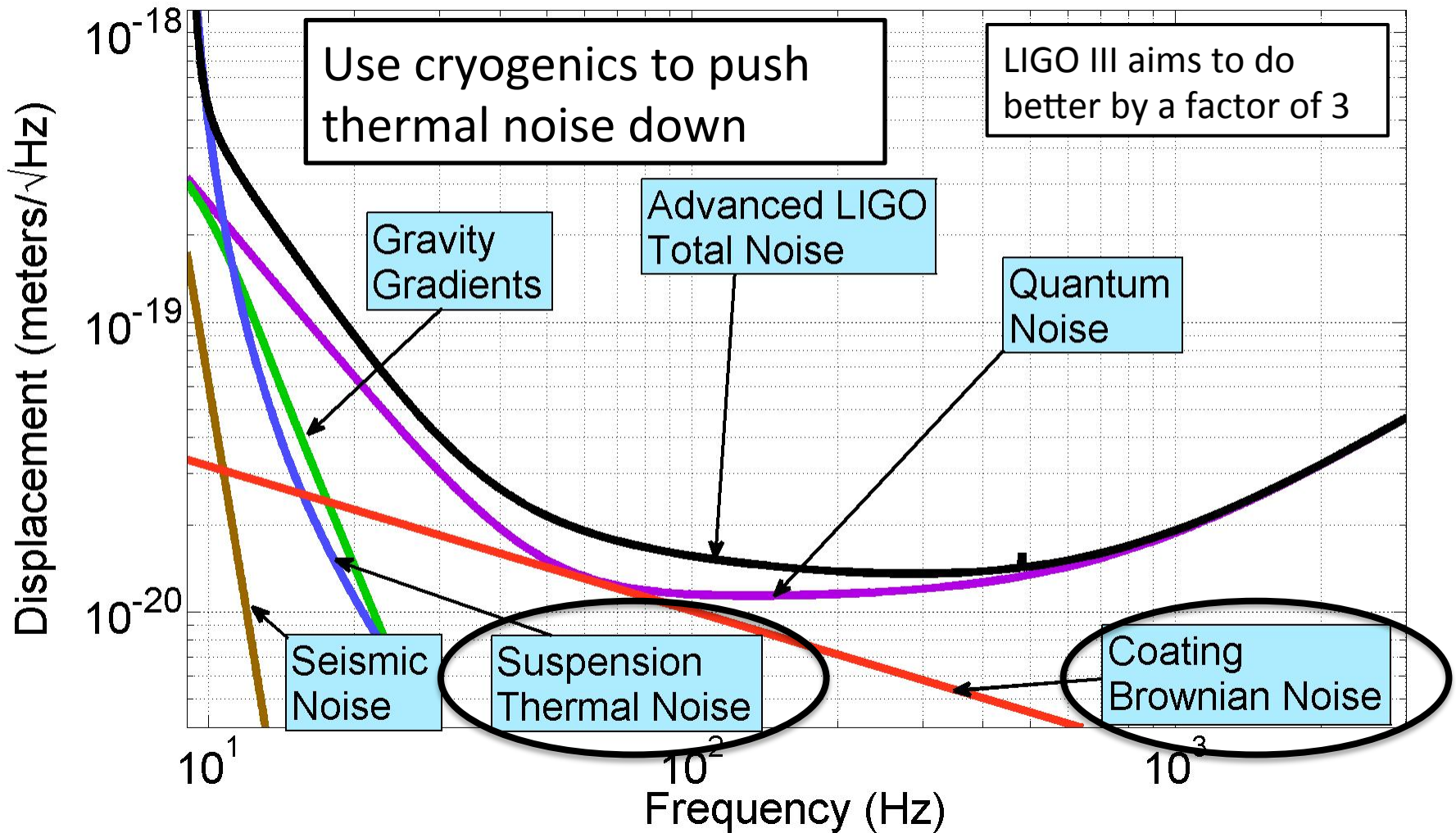
Predicted Advanced LIGO Sensitivity



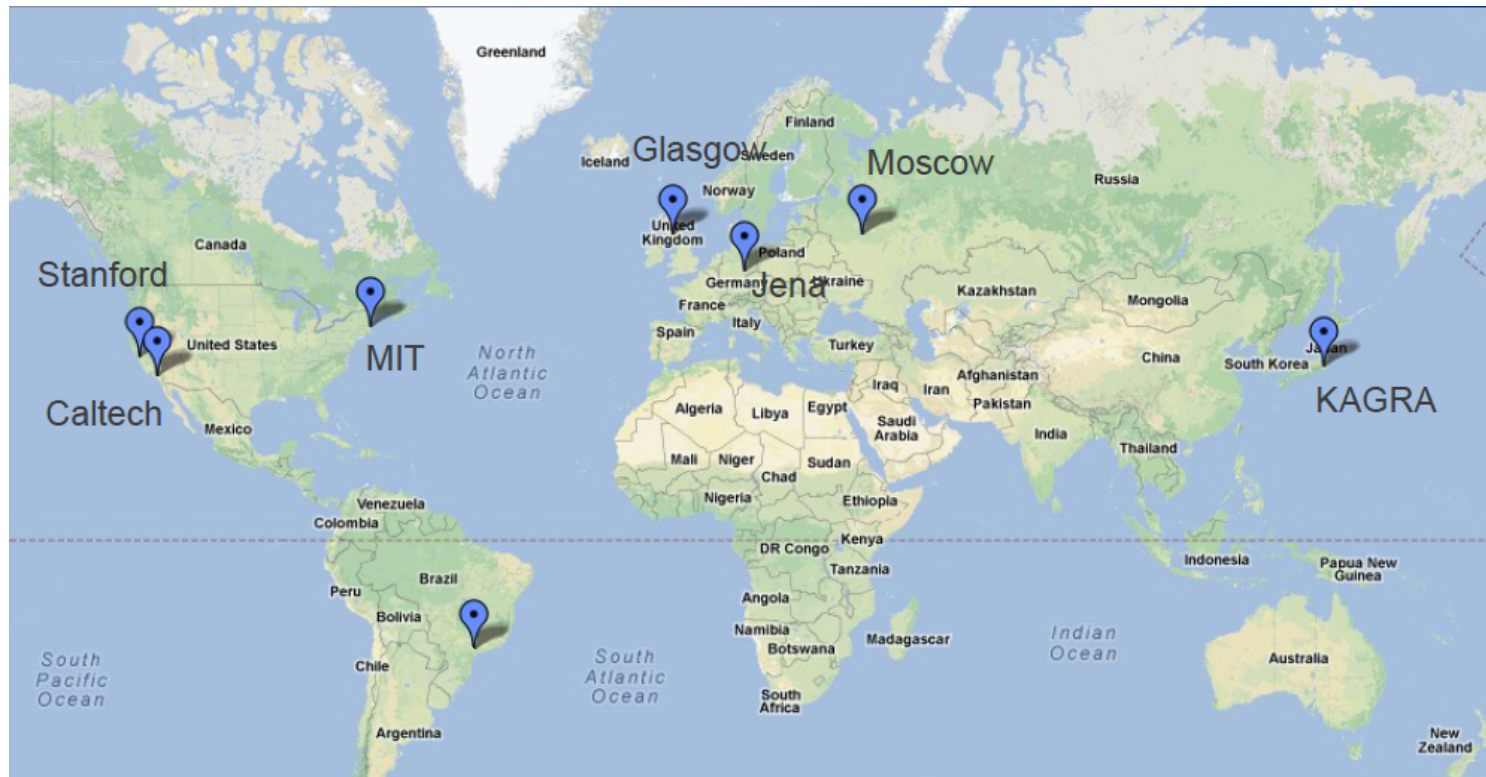
Predicted Advanced LIGO Sensitivity



Predicted Advanced LIGO Sensitivity



LIGO III cryo work distribution



- Caltech - cryogenic reference cavities; direct thermal noise measurements
- Jena/Glasgow/Moscow - mechanical loss
- MIT – high emissivity coatings
- KAGRA – 20 K sapphire suspensions
- INPE Brazil – Cryogenic multi-nested pendulum
- Stanford – optical coatings; cryogenic technology

Cyro Test Mass Problem Statement

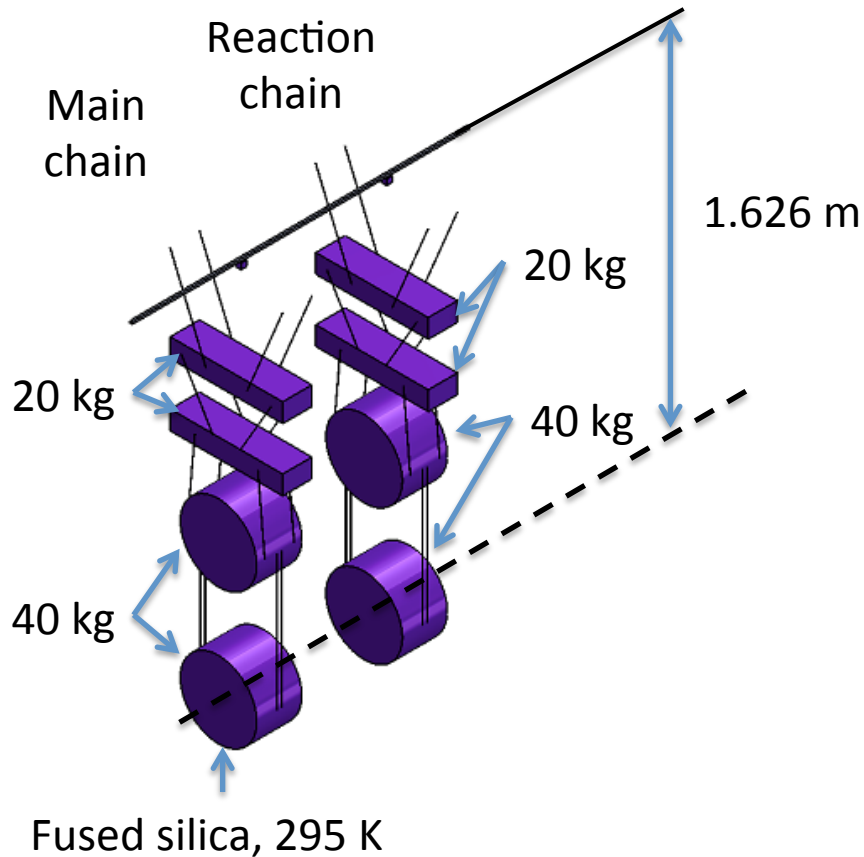
- * For LIGO III, reduce suspension and coating thermal noise by cooling the lower quad to 124 K (-149.15 C)
 - Si test masses (blue team in LIGO-T1200031)
- **2 cooling regimes:**
 - **Initial cool down - get to 124 K quickly** <- Stanford experiment
 - **Steady state - maintain 124 K once you get there**
- Include a warm-up scheme (don't forget!)
- Use the same seismic isolation platforms (ISIs, HEPIs)
 - Limit the amount of extra weight on these platforms
 - Leave the rest of the vacuum chamber warm
- Do not increase the test mass lossiness
 - Emissive coatings, heat links, thick sus fibers, optical coatings, substrate, suspension fiber bonding, etc
- Do not compromise passive seismic isolation
 - Cables, hoses, links, etc

Quad Suspension Design

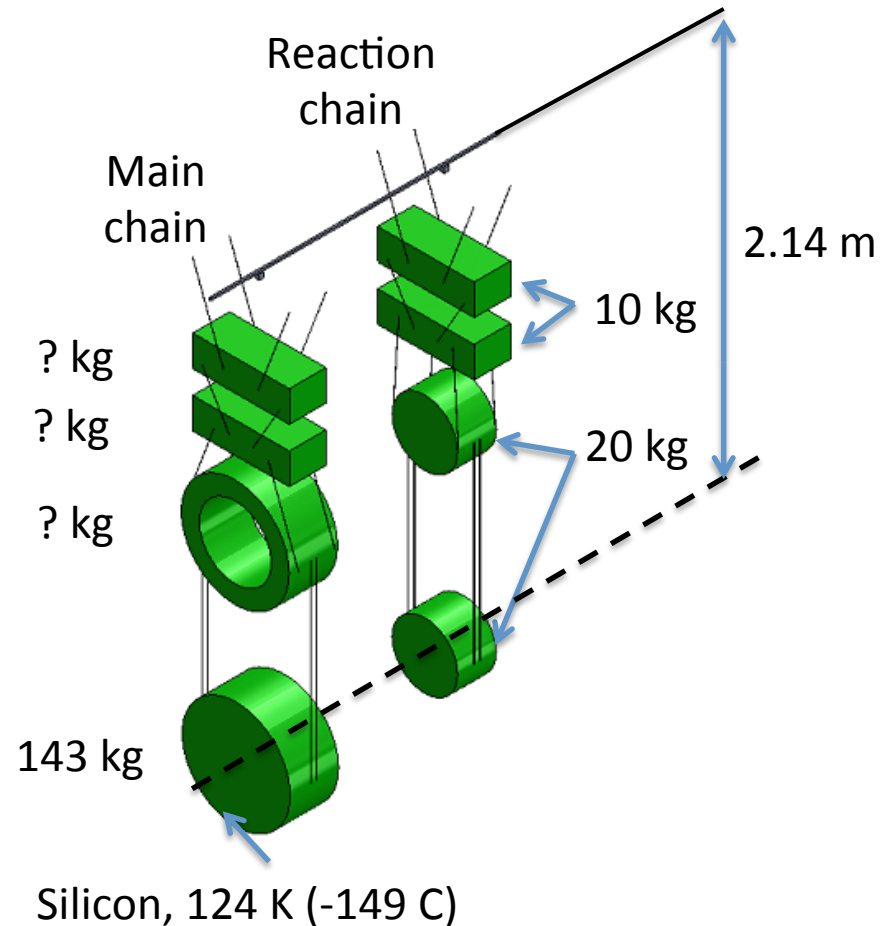
Must incorporate much larger test
mass

LIGO III Quad Suspension Design

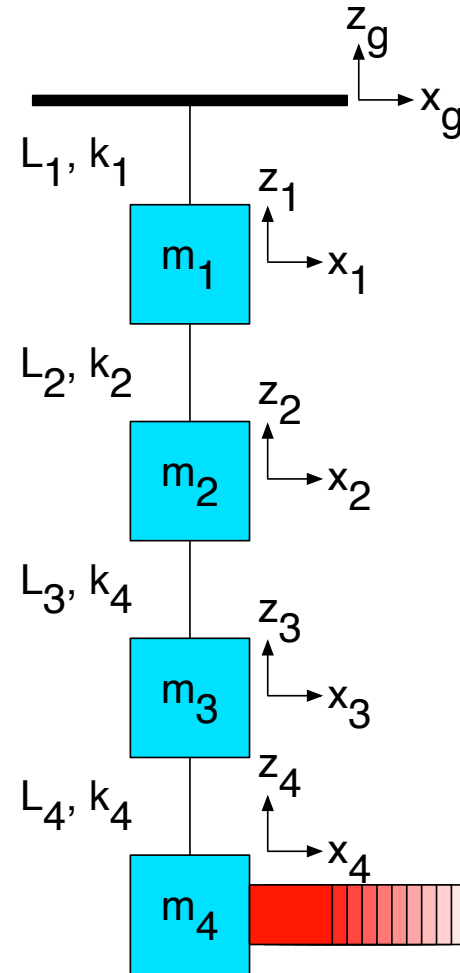
Advanced LIGO quad pendulum



Preliminary LIGO III quad pendulum

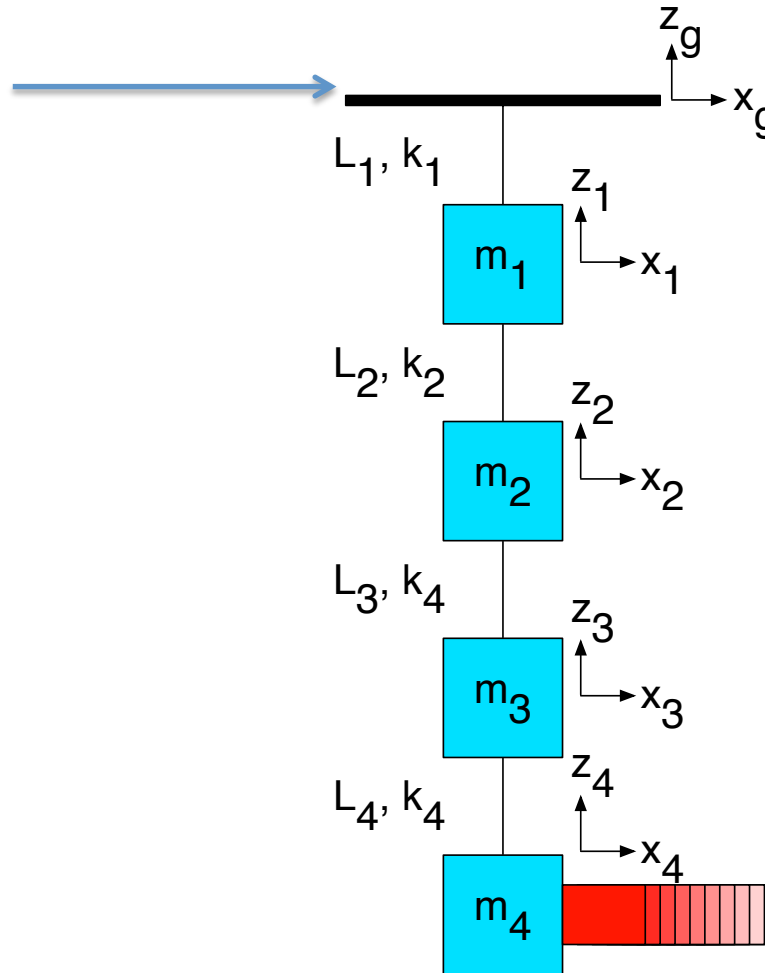


Quad Design Requirements



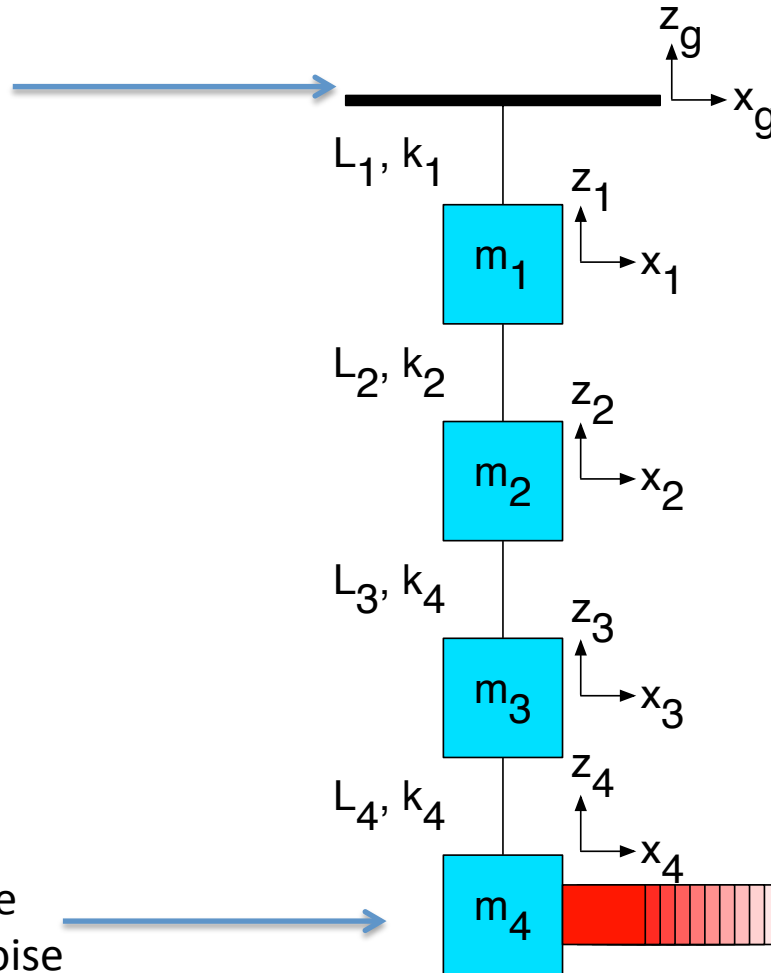
Quad Design Requirements

Payload $P \leq 270$ kg for the main chain for aLIGO BSC-ISI



Quad Design Requirements

Payload $P \leq 270$ kg for the main chain for aLIGO BSC-ISI



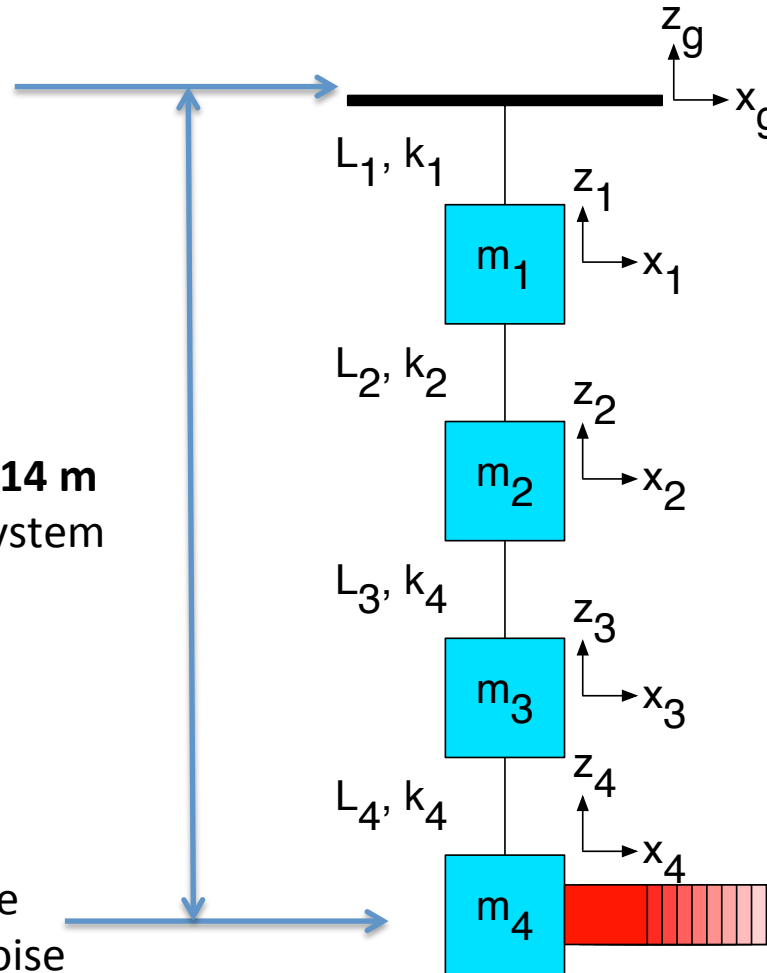
Test mass $m_4 \approx 143$ kg to minimize radiation pressure and thermal noise

Quad Design Requirements

Payload $P \leq 270$ kg for the main chain for aLIGO BSC-ISI

Total length $L \leq 2.14$ m to fit in vacuum system

Test mass $m_4 \approx 143$ kg to minimize radiation pressure and thermal noise

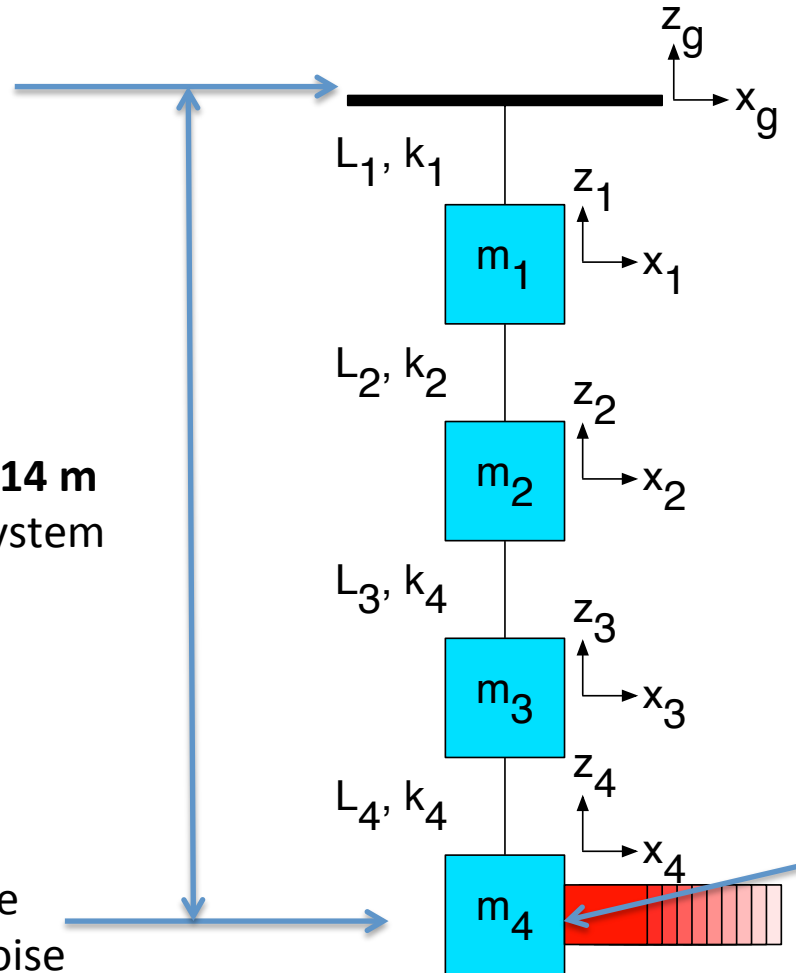


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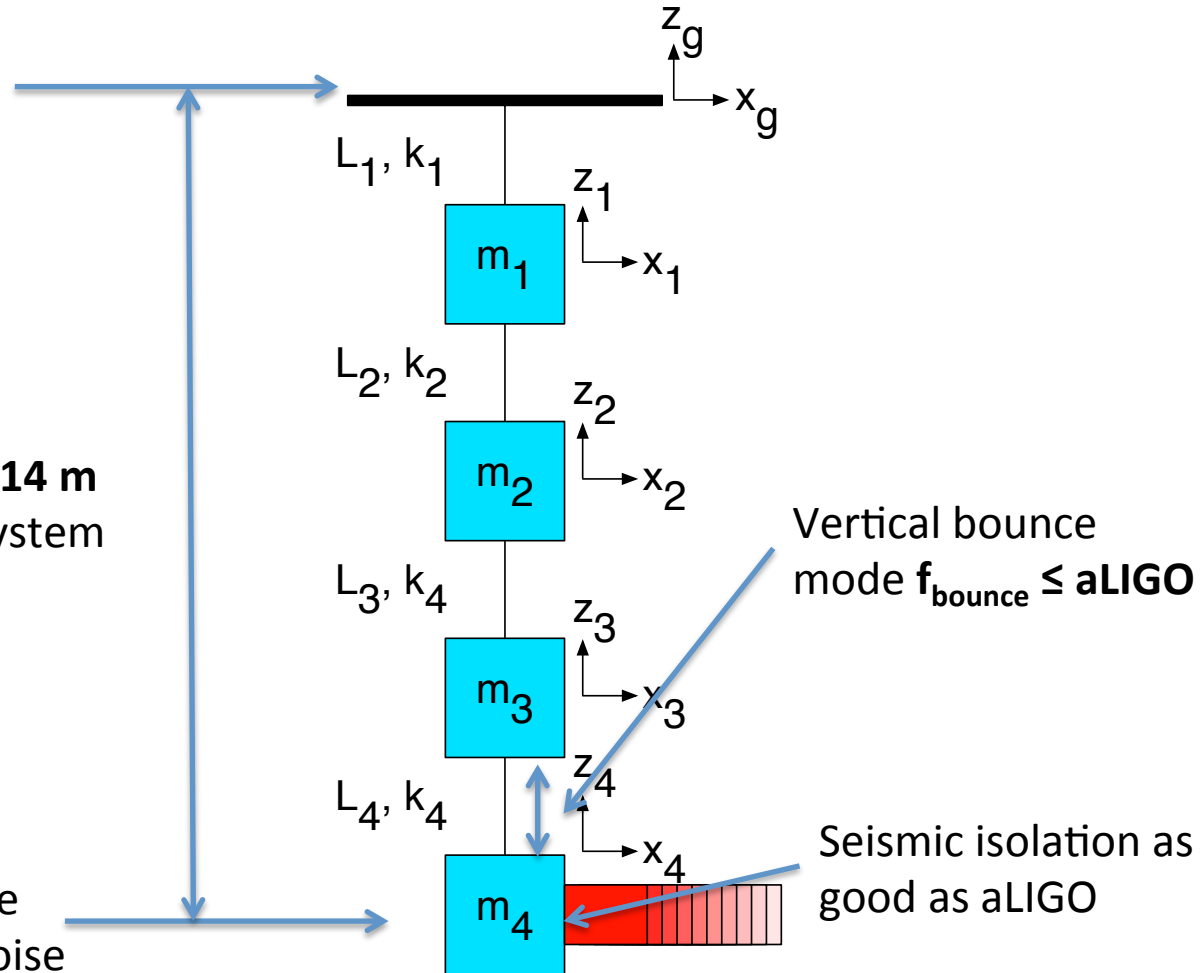
Seismic isolation as good as aLIGO

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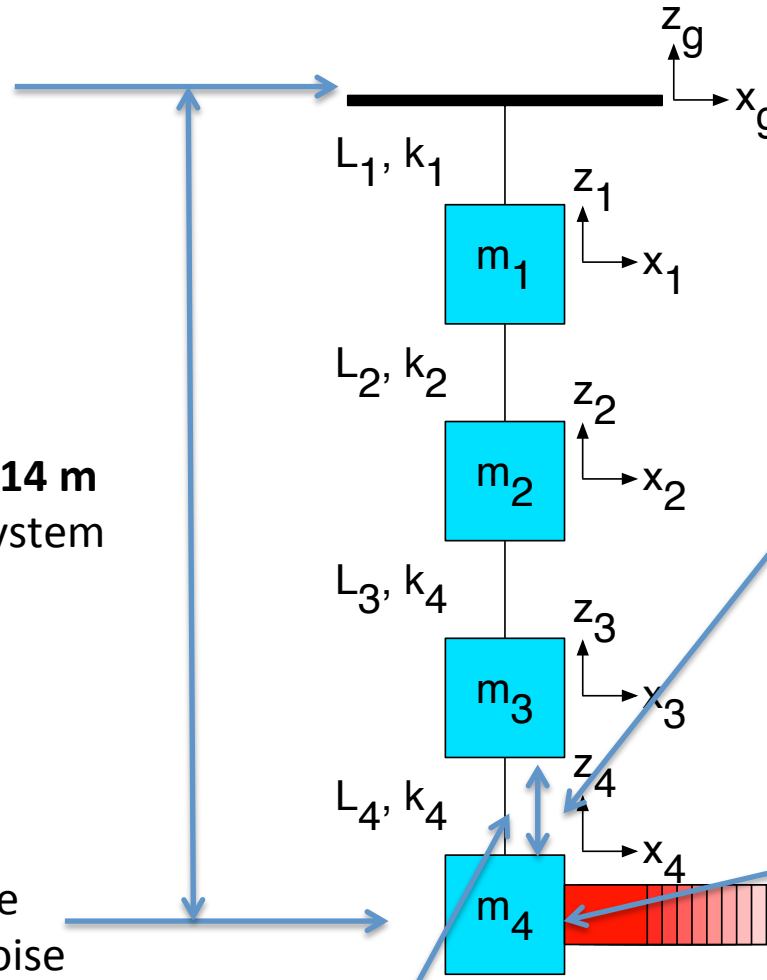


Quad Design Requirements

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Test mass $m_4 \approx 143$ kg to minimize radiation pressure and thermal noise



Vertical bounce mode $f_{\text{bounce}} \leq \text{aLIGO}$

Seismic isolation no worse than aLIGO

Si fiber stress $\sigma_4 \leq 1400$ Mpa

Very rough guess

3 Optimal Quad Designs

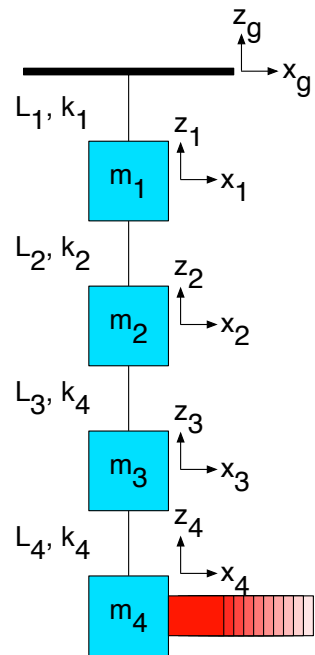
Higher
payload

Lighter
Test mass

PUM Si springs

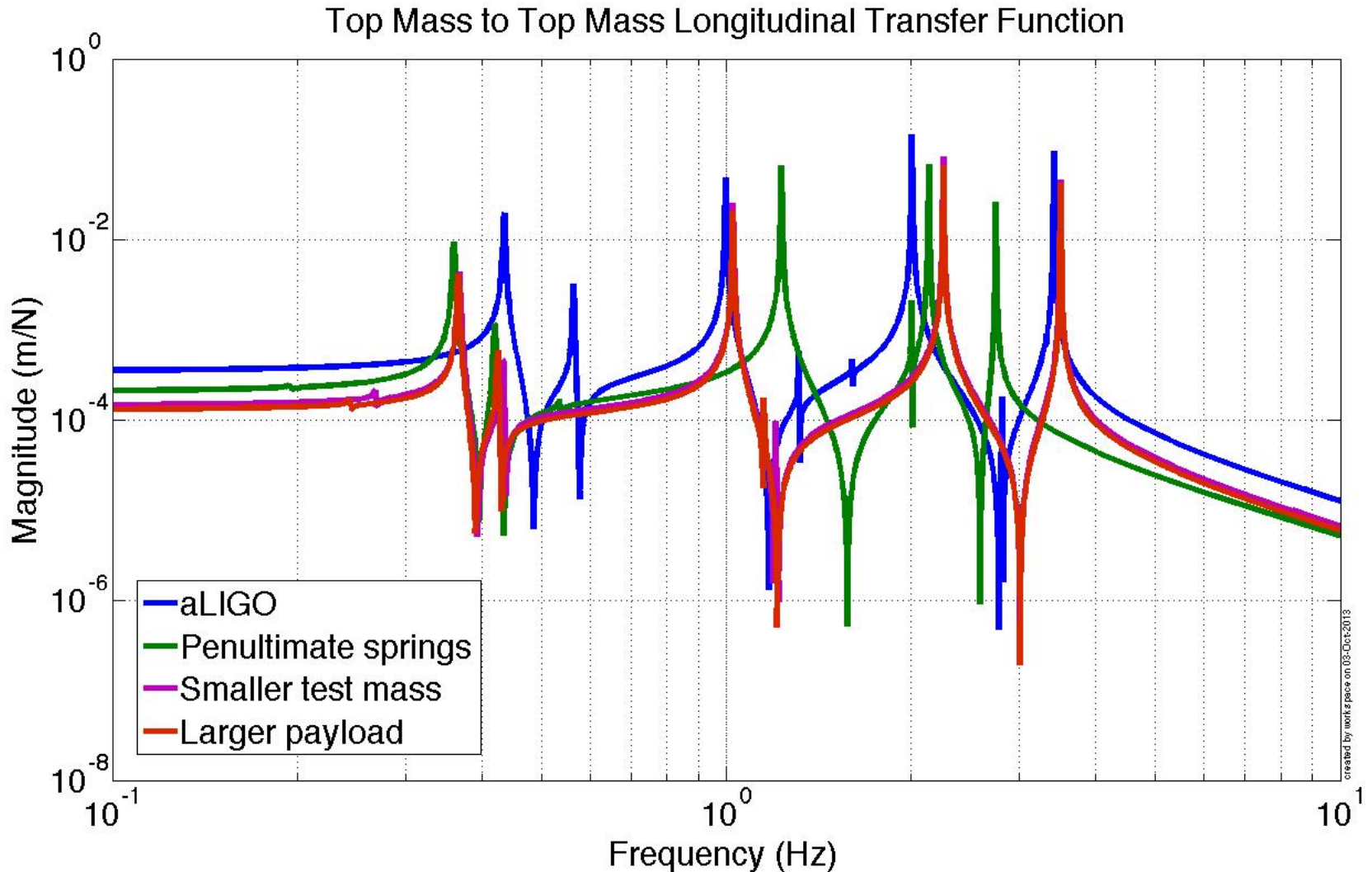
Table 3: Summary of model parameters for the three proposed modifications.

Parameters	Increased P	Decreased m_4	Penultimate Springs
P , Payload (kg)	301.9	270.0	270.0
m_1 (kg)	46.79	41.93	51.55
m_2 (kg)	39.54	35.42	41.71
m_3 (kg)	72.57	64.86	33.74
m_4 (kg)	143.0	127.8	143.0
L_1 (m)	0.372	0.372	0.535
L_2 (m)	0.372	0.372	0.535
L_3 (m)	0.372	0.372	0.535
L_4 (m)	1.025	1.025	0.535
10Hz long. isolation (m/m)	1.1×10^{-7}	1.1×10^{-7}	7.9×10^{-8}
f_{bounce} (Hz)	9.27	9.27	low, depends on springs
σ_4 , fiber stress (Mpa)	1400	1400	1400
E_4 , fiber modulus (Gpa) [6]	167.4	167.4	167.4
noise budget impact	none	slightly worse	better
relative cost	high	low	high



Note: these results are highly dependent on the allowed Si fiber stress, σ_4 . There is still much uncertainty in this value.

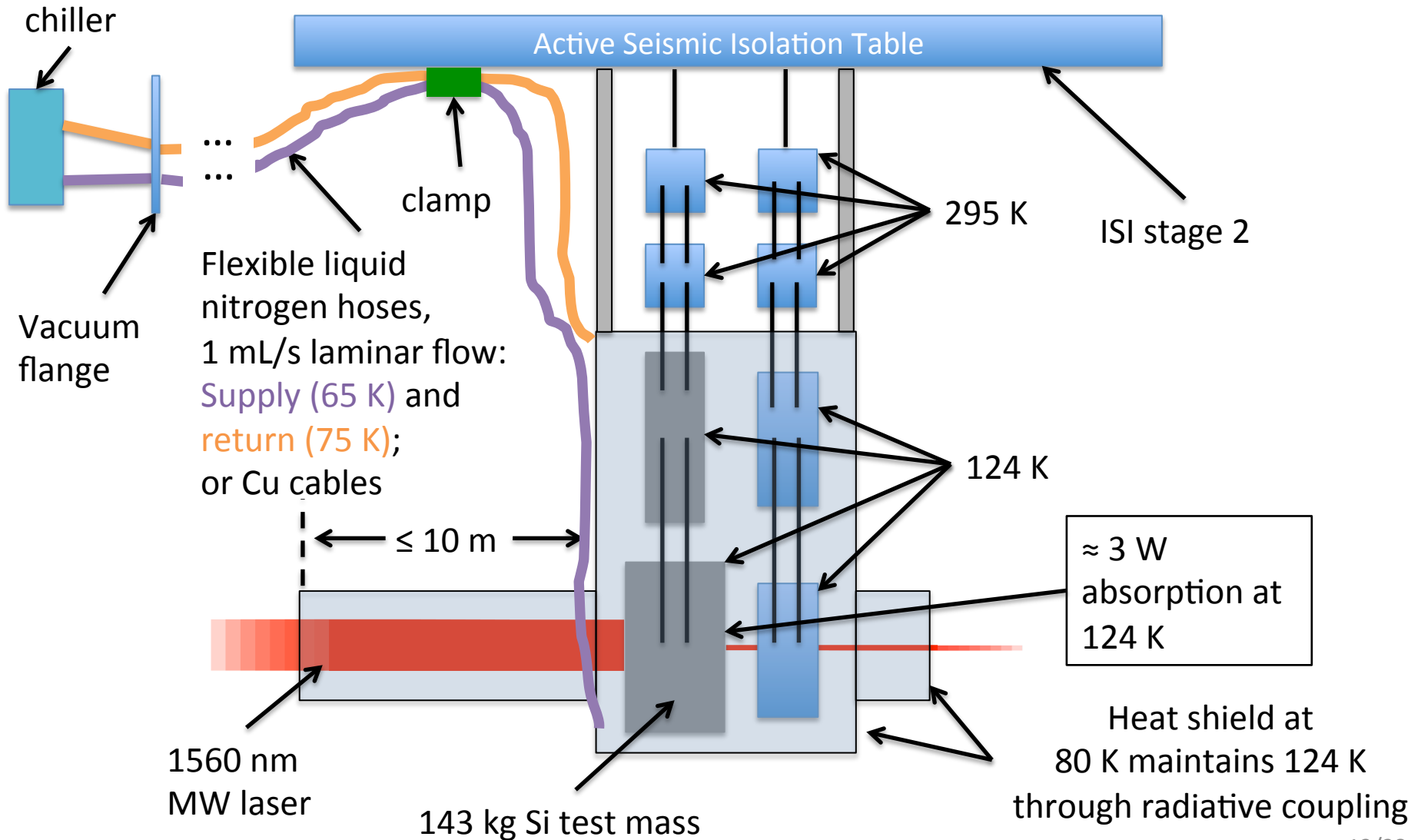
All designs permit top mass damping



Steady State Cooling

Keeping the test mass temperature
at the operating point

LIGO III Steady State Cooling



Pros and Cons of LN₂ pipe vs. Cu cable

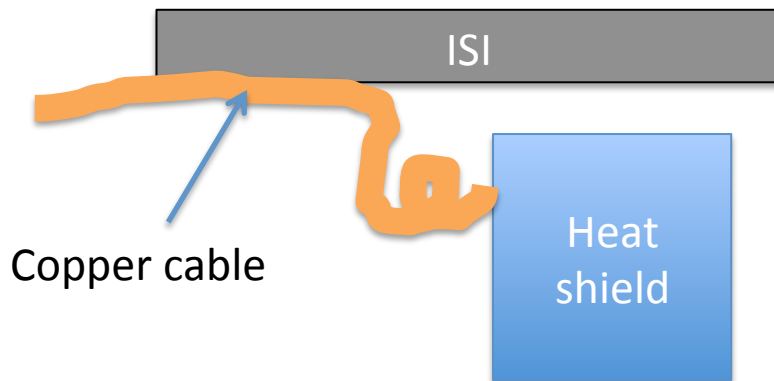
Cu cable

Pros:

- No fluid to make noise
- No LN₂ pumping mechanism
- No risk of N₂ leaks

Cons:

- Low heat transfer
- ...



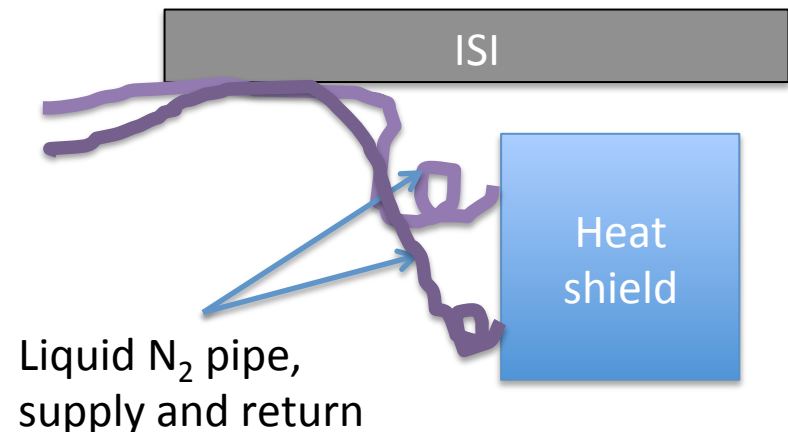
LN₂ pipe

Pros:

- High heat transfer
- No length / heat transfer trade-off
- Low weight
- ...

Cons:

- Risk of leaking
- Fluid flow could contribute noise
- ...



Pros and Cons of LN₂ pipe vs. Cu cable

Cu cable

Pros:

- No fluid to make noise
- No LN₂ pumping mechanism
- No risk of N₂ leaks

Cons:

- Low heat transfer
- Cryo refrigerator must be placed near feedthrough
- High bulk: stiffness, weight, etc
 - $Q = K(A/L)\Delta T$
 1. Big L means big A
 2. Can reduce A by making cold end less than LN₂ (77 K). E.g. Cryomech's PT407 can pull 25W at 55 K.
 - Minimize stiffness by using lots of this wires, but wire dia must be > electron m.f.p.
- Thermal conductivity decreases when wire dia becomes smaller than electron m.f.p.
- Hysteresis
 - Lots of small wires sliding past each other
- High difficulty in minimizing seismic shorting
 - Minimize using:
 1. Lots of thin wires
 2. Intermediate masses along length

LN₂ pipe

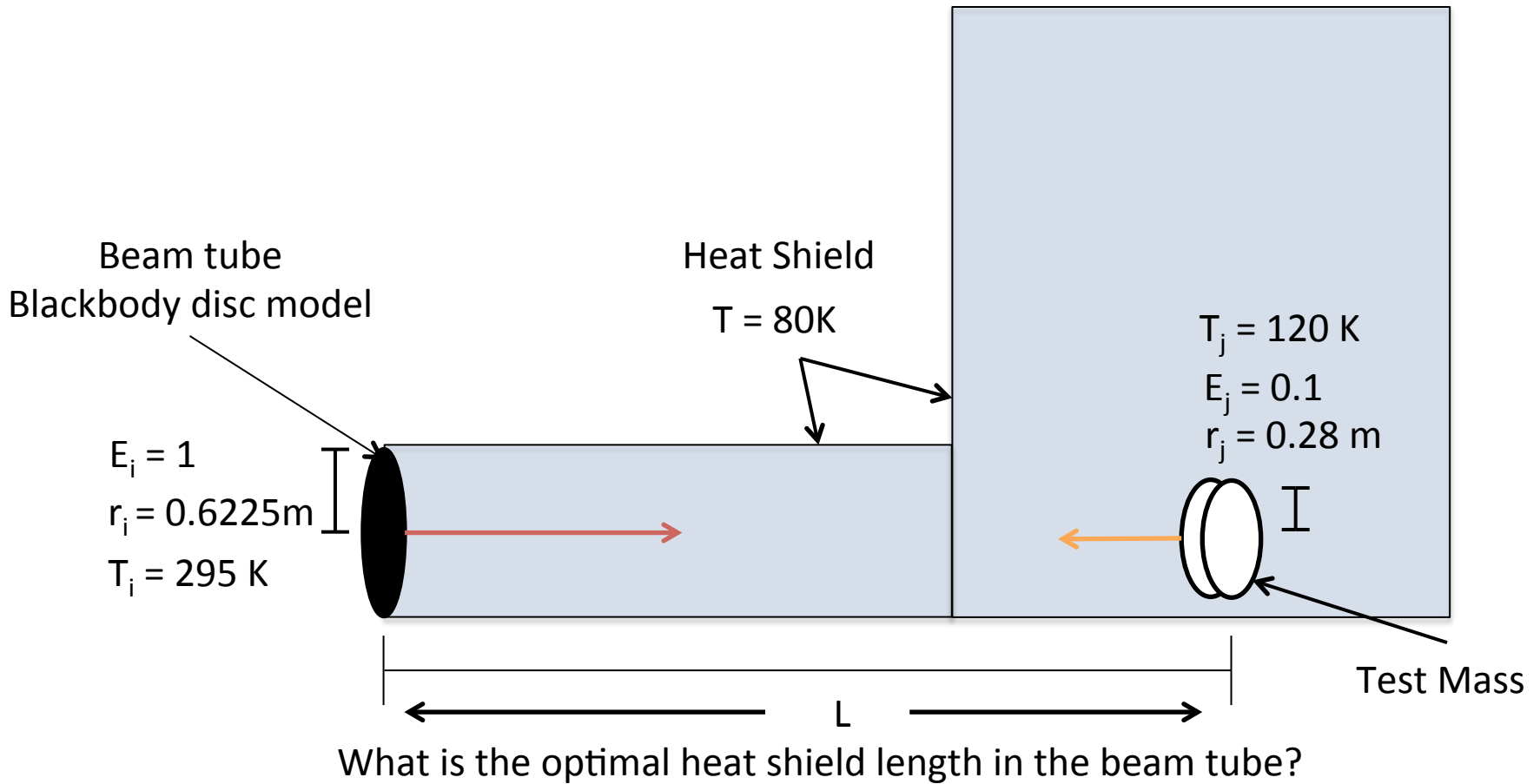
Pros:

- High heat transfer
- Low bulk
- Moderate difficulty in minimizing seismic shorting
- Length of pipe in vacuum not an issue for net heat transfer (longer pipes do require more pressure to push fluid)
- Vibrating cryorefrigerator can be placed further from vacuum feedthrough.

Cons:

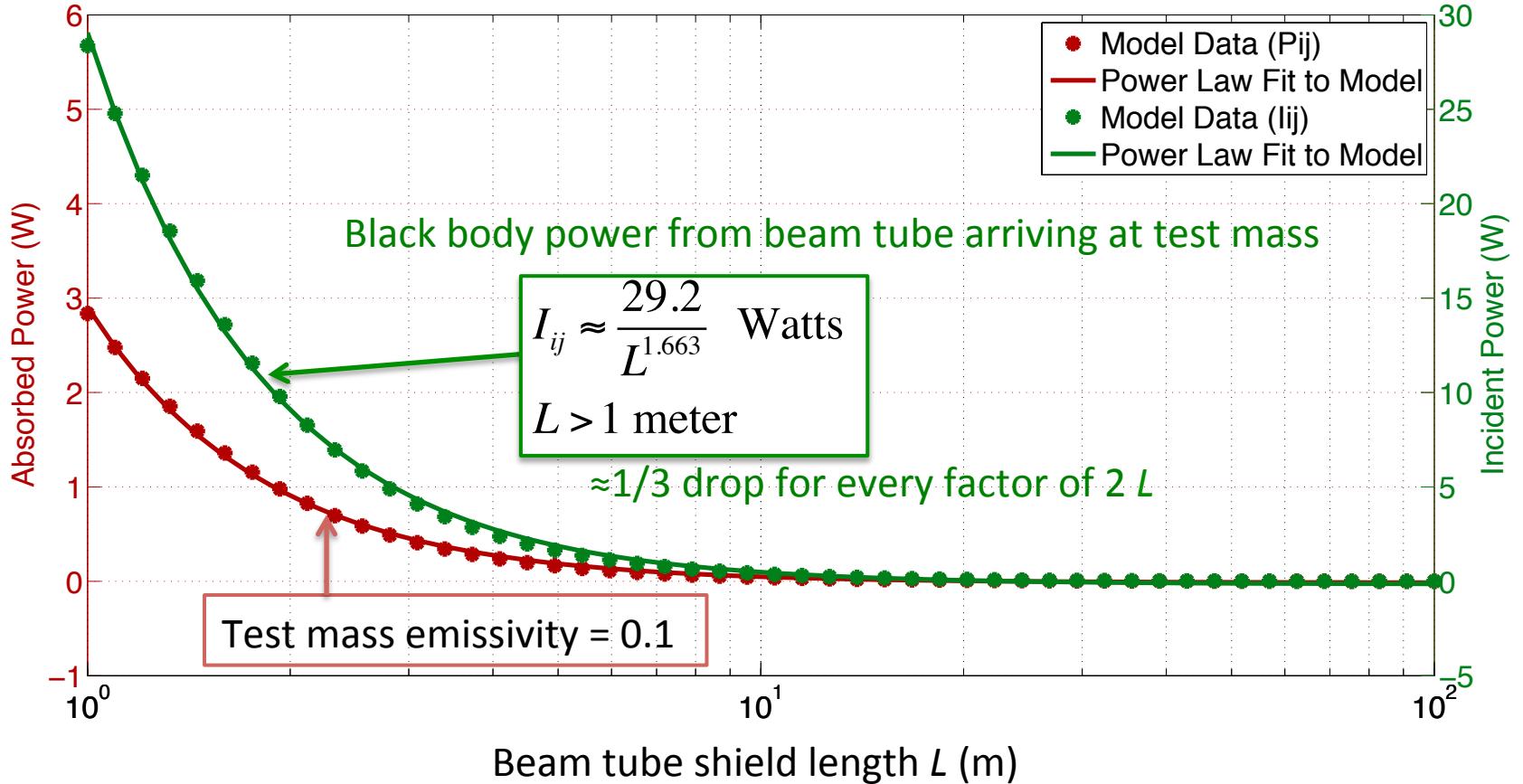
- Complex LN₂ pumping mechanism
- Risk of leaking
- Fluid flow could contribute noise
 - Minimize by:
 1. Cooling the LN₂ so it doesn't boil
 2. Ensuring laminar flow
- Pressure of flow and thermal contraction can influence table displacement
 - Minimize by locating pipes in strategic locations

Beam Tube Heat Shield Length

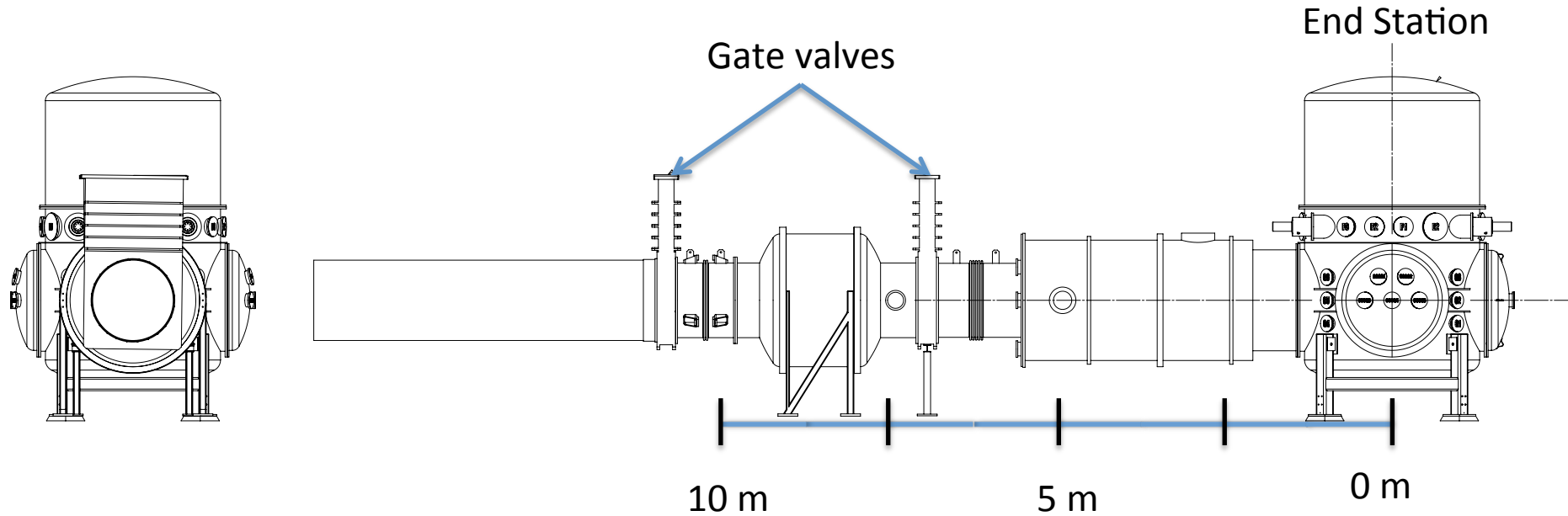


Beam Tube Heat Shield Length

Power Absorbed (P_{ij}) and Power Incident (I_{ij}) on the Test Mass Face from the Beam Tube vs Distance (L)



End Station Vacuum System



Initial Cooldown

Acquiring the operating temperature

A Lot of Heat to Remove



143 kg Test
Mass:
14 MJ to get
cold (warm)

2.7 days to 124 K
(295 K) at a
constant 60 W

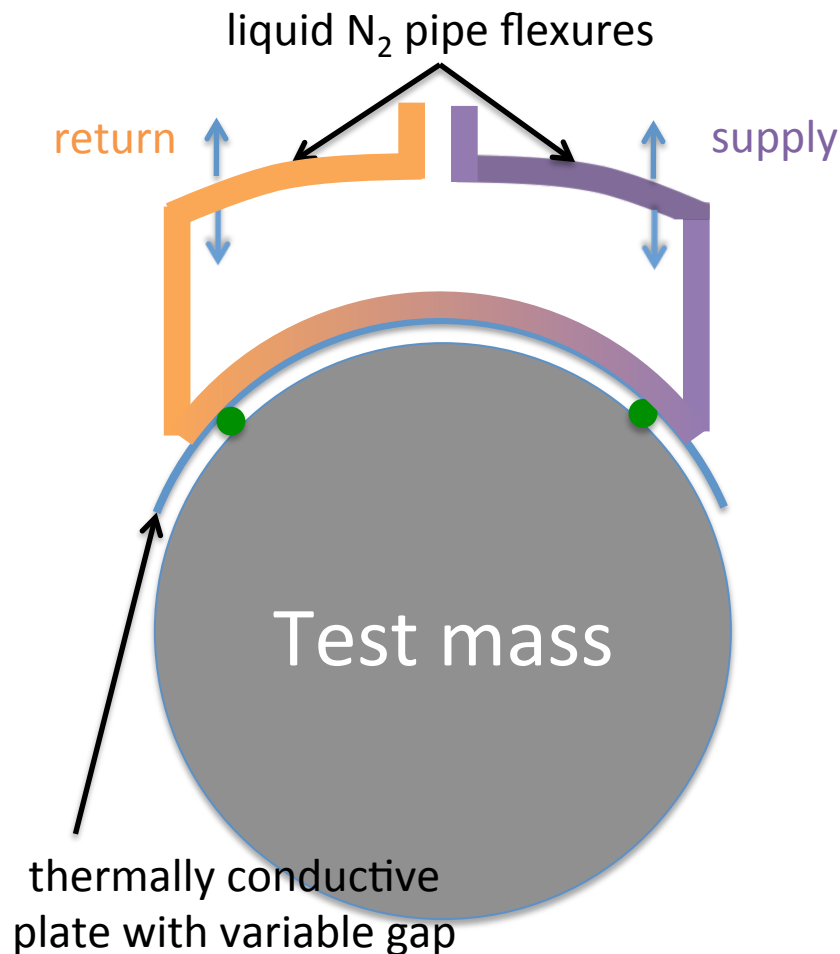
Maybe there is another
way to cool the test
mass quickly

* Radiation to the heat shield alone would take \approx 2 weeks.

Initial Cool Down Cold Link – 2 Designs

Conductive cooling, low pressure N₂ gas

Pros and Cons



Pros

- Operates in partial vacuum.
- Low heat transfer between cold and warm parts of vacuum system.
- Fine temperature control – just back away when at desired temperature
- Versatility, design permits both conductive and convective cooling.

Cons

- Requires moving parts:
 - flexible pipes
 - actuators
- Physically contacts barrel of test mass

Initial Cool Down Cold Link – 2 Designs

Pros and Cons

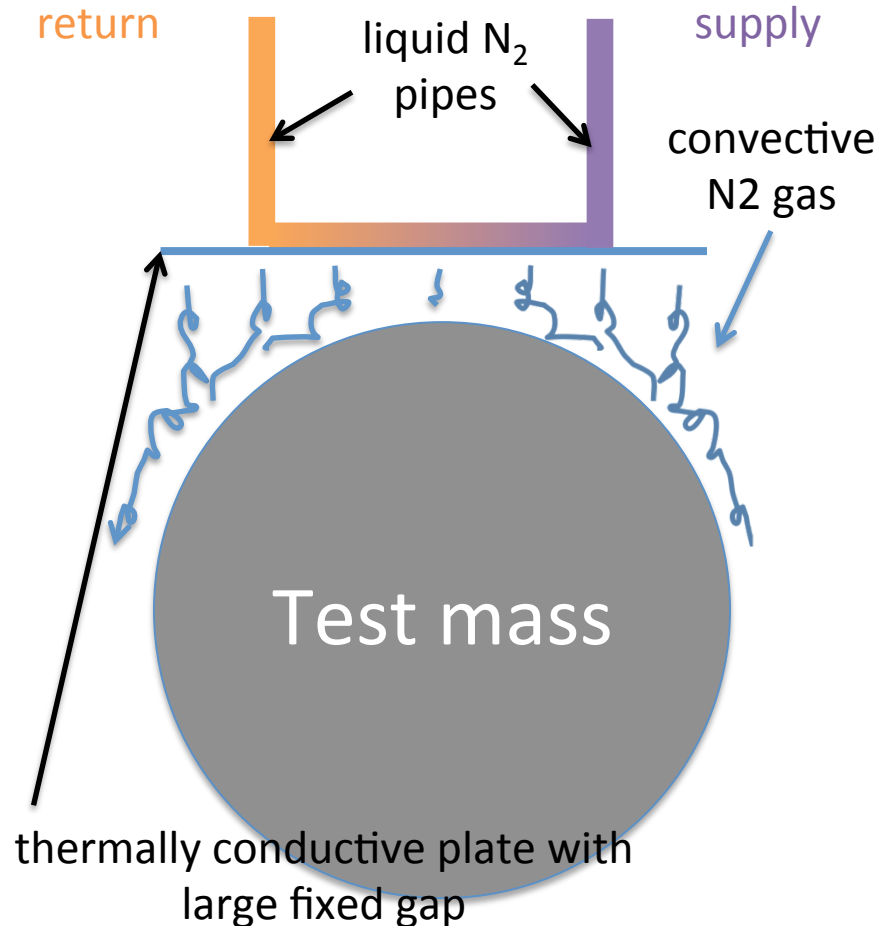
Pros

- No moving parts or actuators
- No contact with test mass
- Faster cooling than conduction

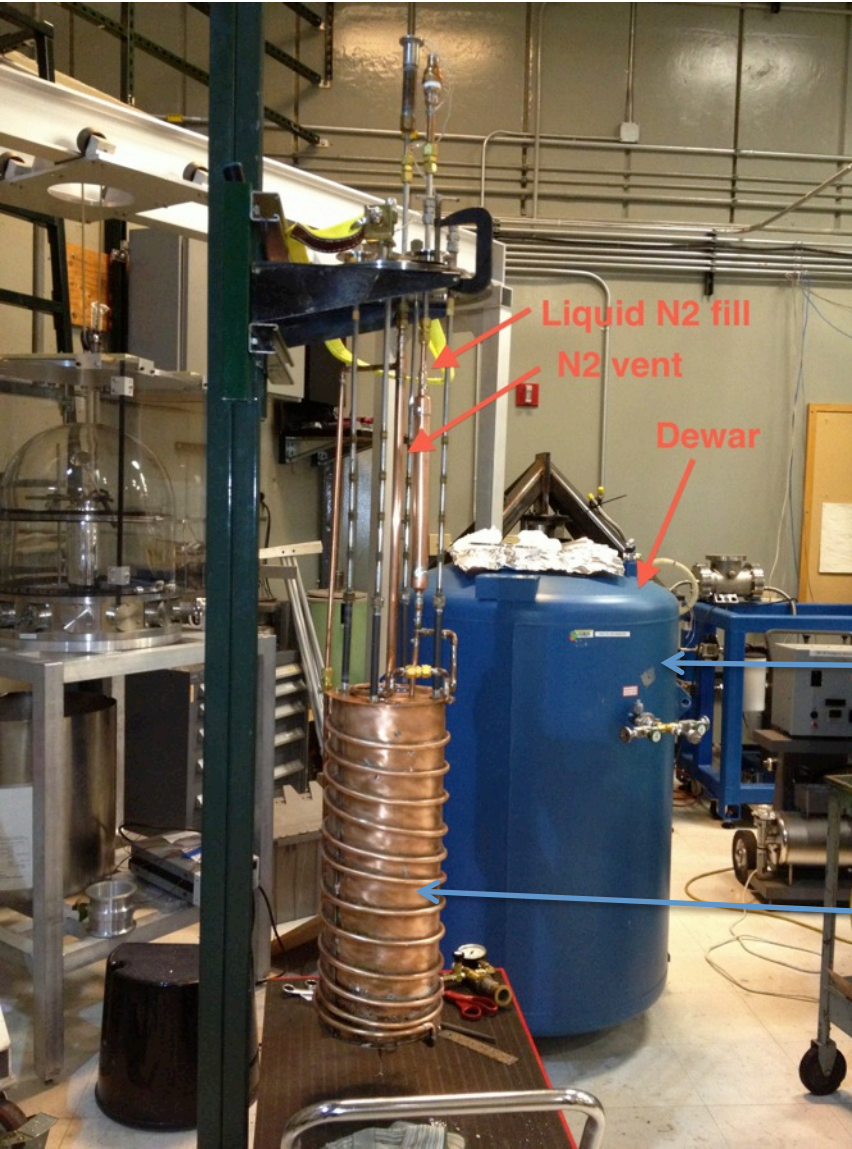
Cons

- Convection between cold and warm parts of vacuum system
- No fine temperature control – must return to vacuum to ‘turn off’ cold link.
- Does not operate under vacuum

Convective cooling, up to 1 atm N₂ gas

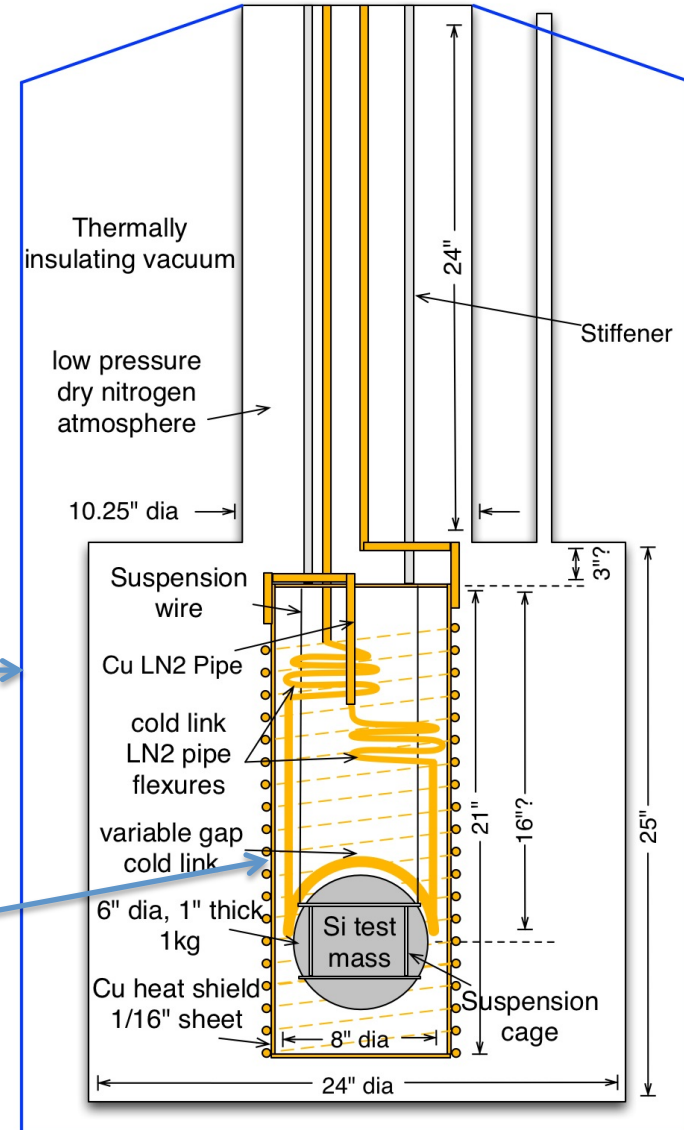


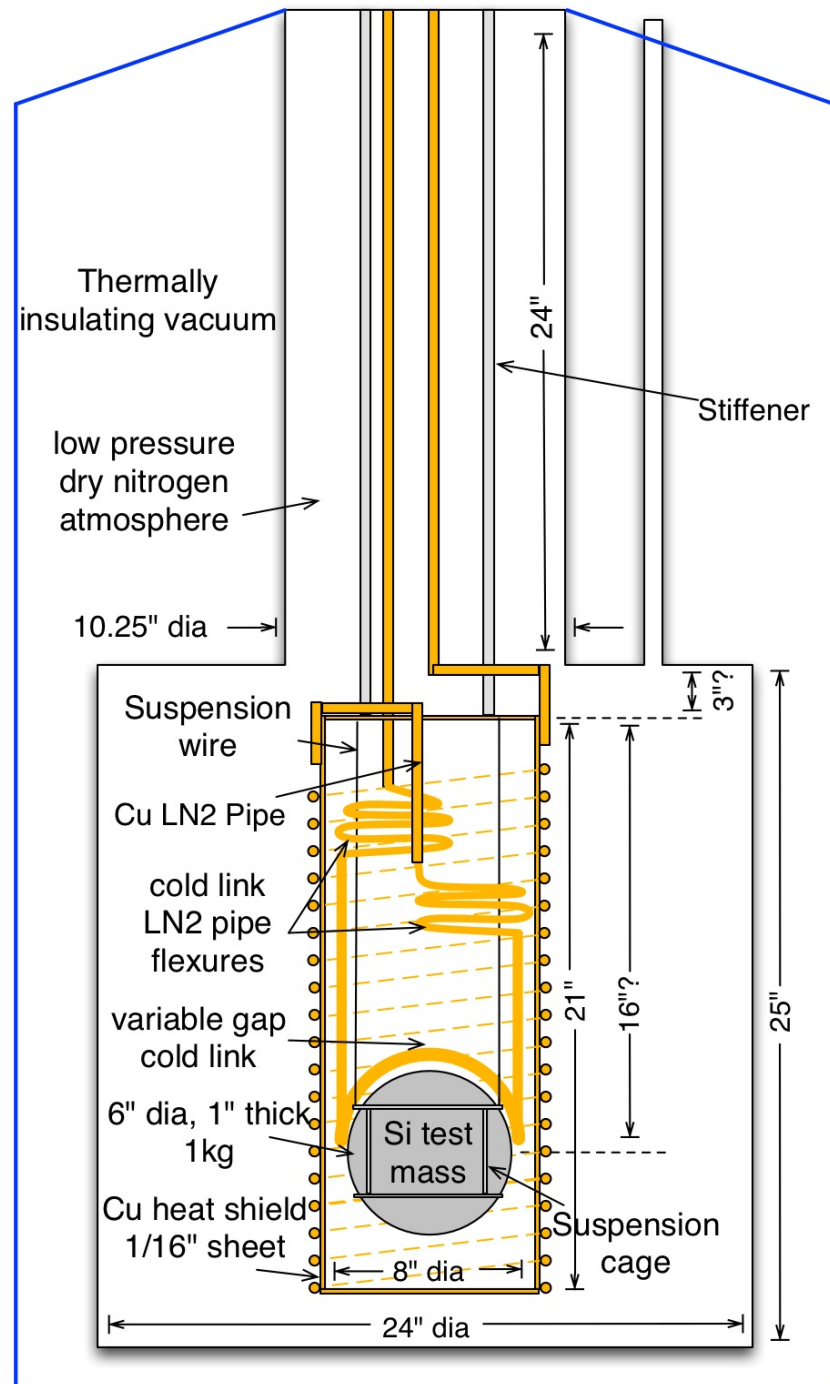
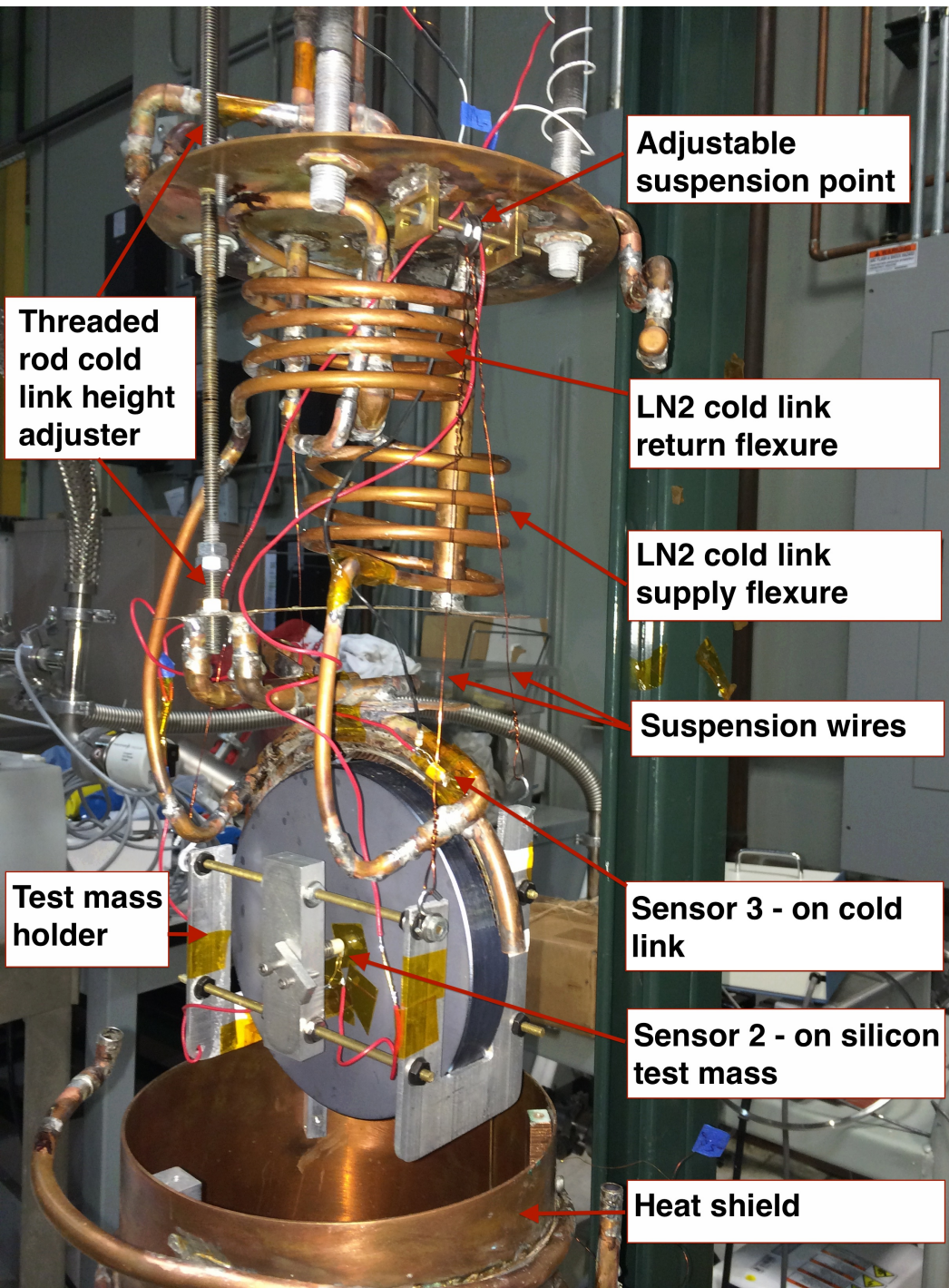
Experimental Setup



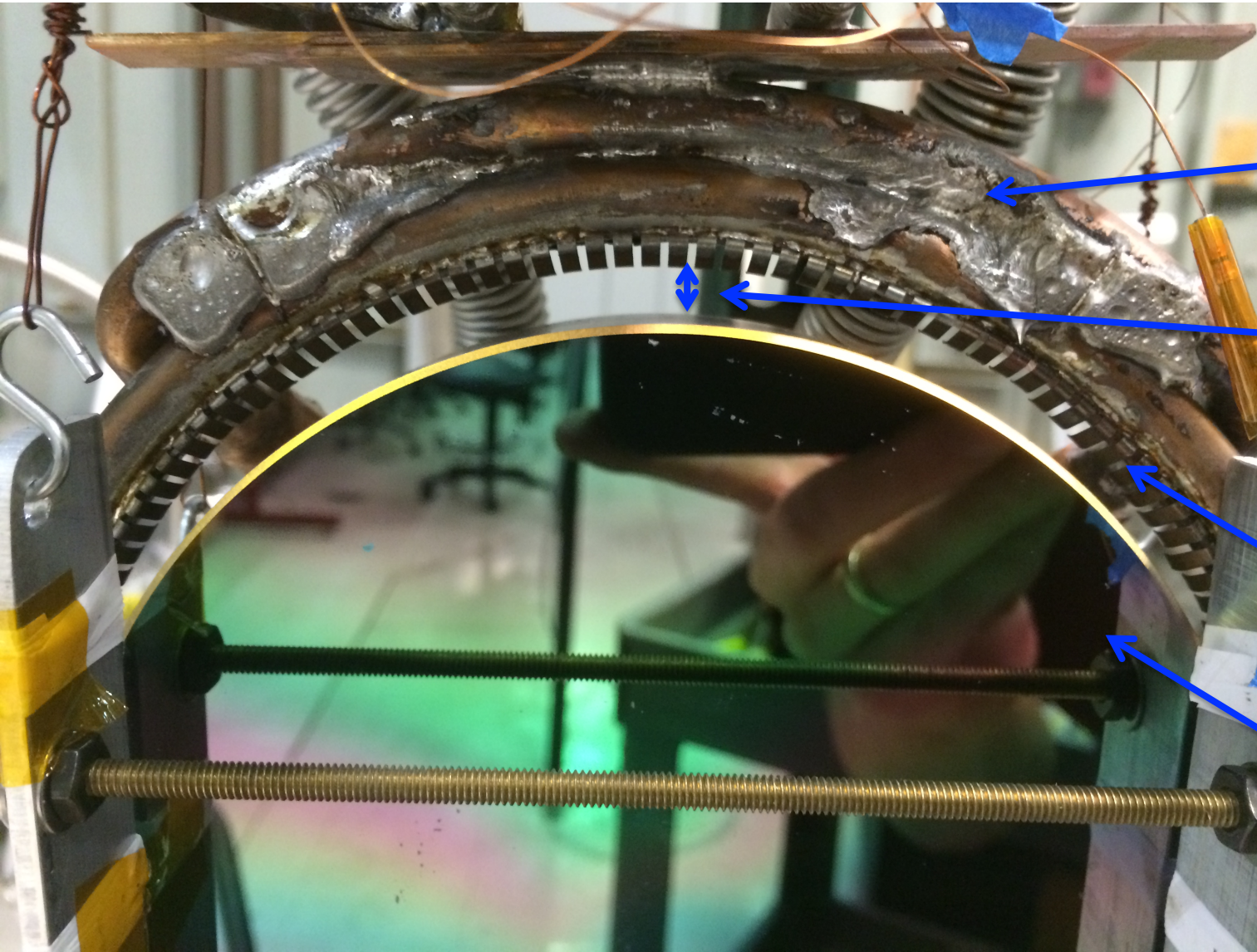
Test dewar

Heat shield





Close up of cold link



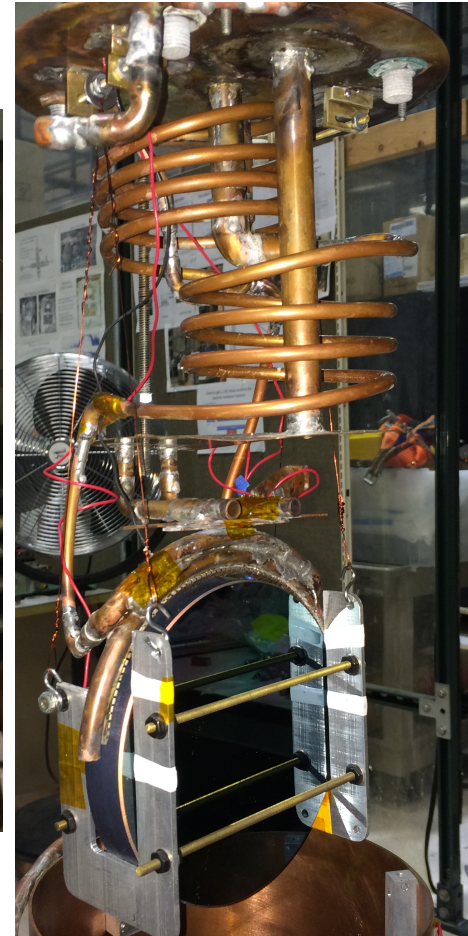
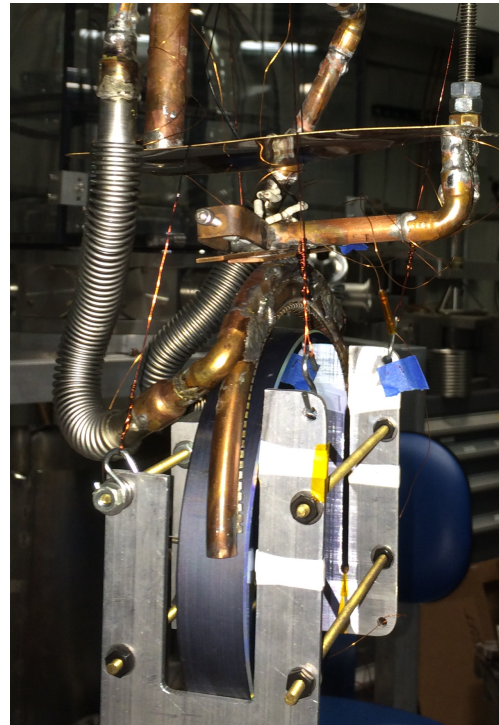
Cold link

Variable gap

Flexure 'point' contacts

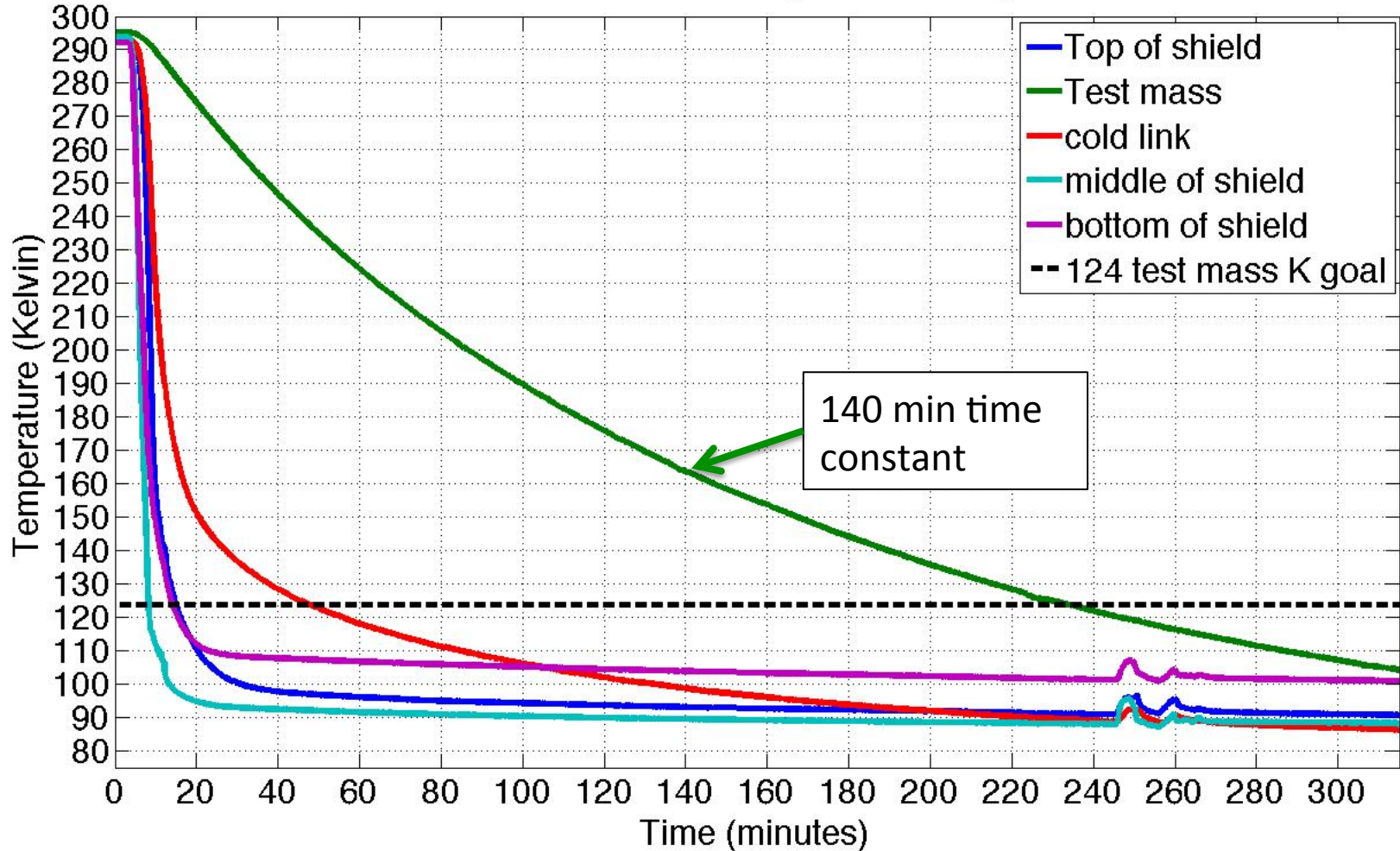
Test mass

Flexible cold link evolution



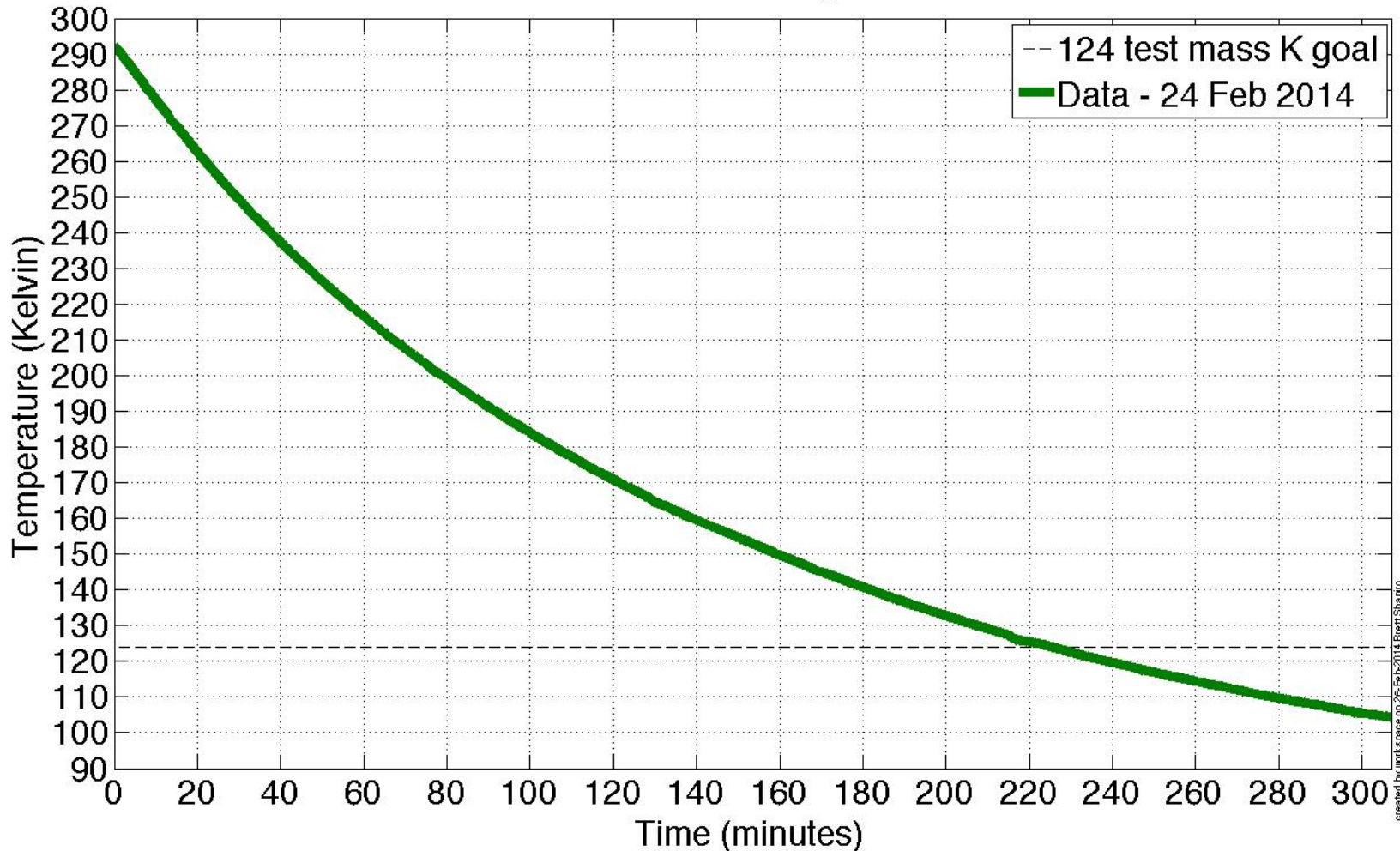
Measurement – cold link engaged

Silicon Test Mass Cooling - 24 February 2014



Test mass temperature modeling

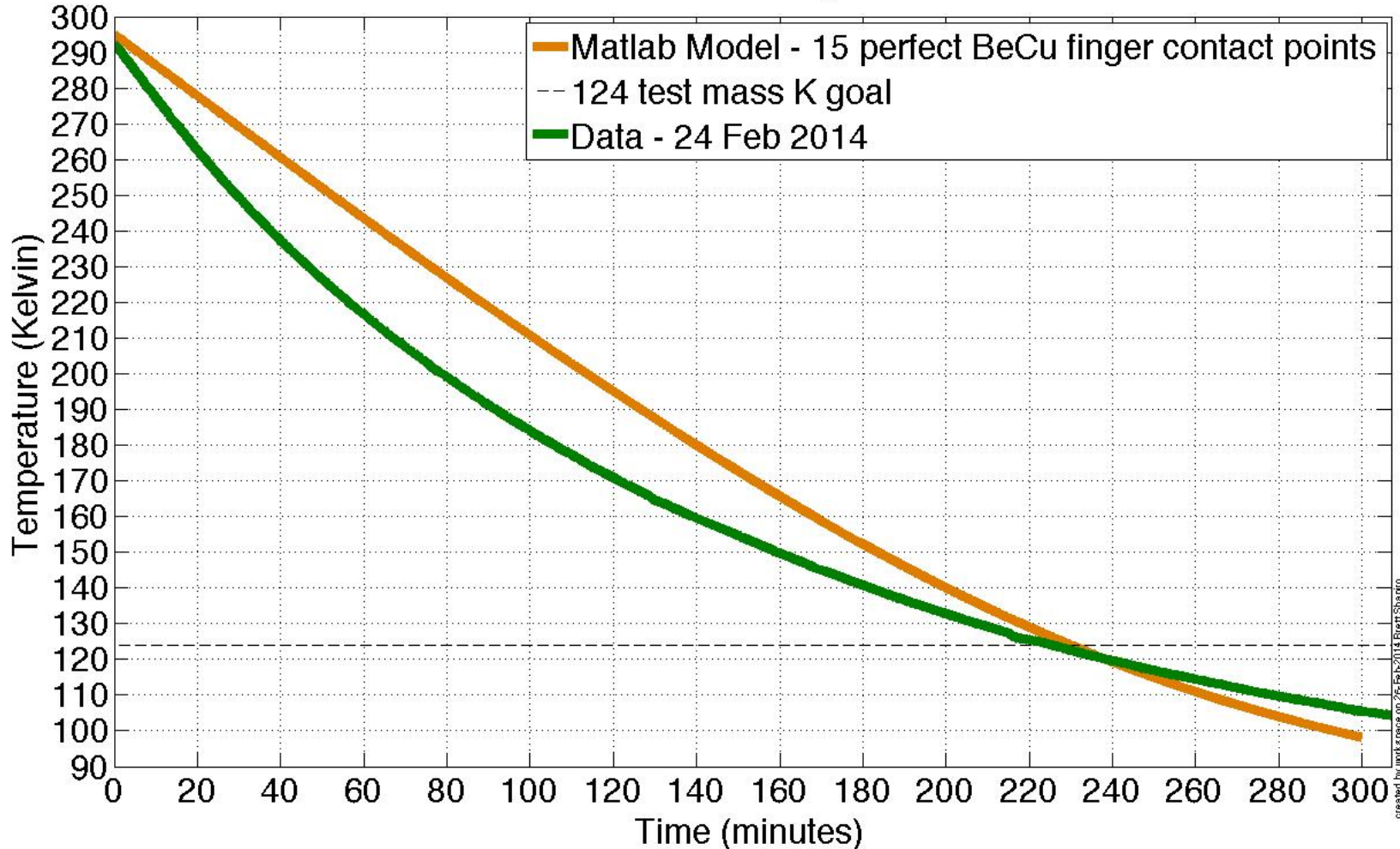
Silicon Test Mass Cooling with Cold Link



created by work space on 26-Feb-2014 Brent Shapiro

Test mass temperature modeling

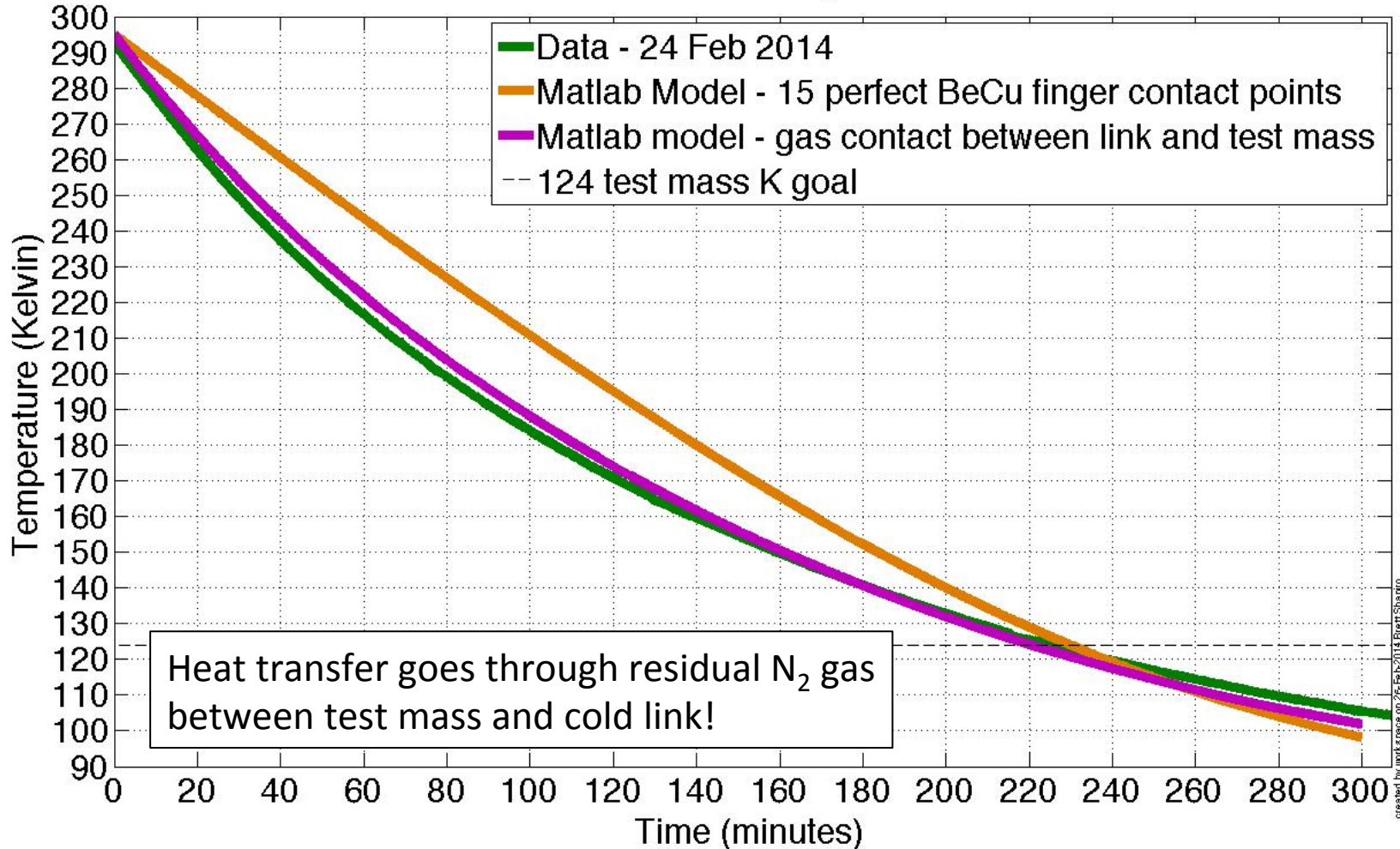
Silicon Test Mass Cooling with Cold Link



created by work@pace on 26-Feb-2014 Brent Shapiro

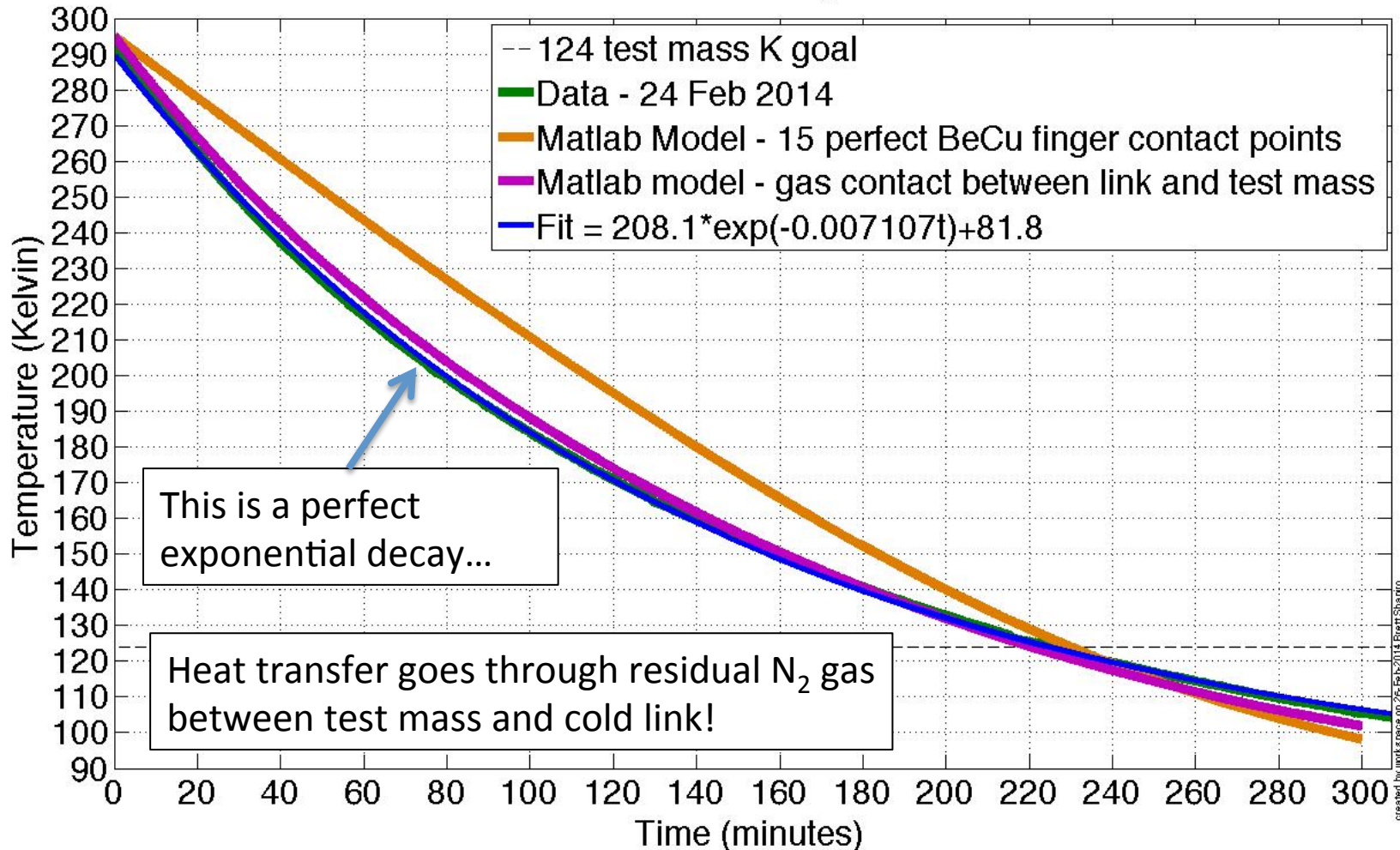
Test mass temperature modeling

Silicon Test Mass Cooling with Cold Link



Test mass temperature modeling

Silicon Test Mass Cooling with Cold Link



Exponential Temperature Decay

$$\dot{T}_{Si} + \frac{K_{CL}}{C_{Si}} T_{Si} = \frac{K_{CL}}{C_{Si}} T_{cold} \quad \longrightarrow \quad T_{Si} = T_{hot} e^{-\frac{K_{CL}t}{C_{Si}}} + T_{cold}$$

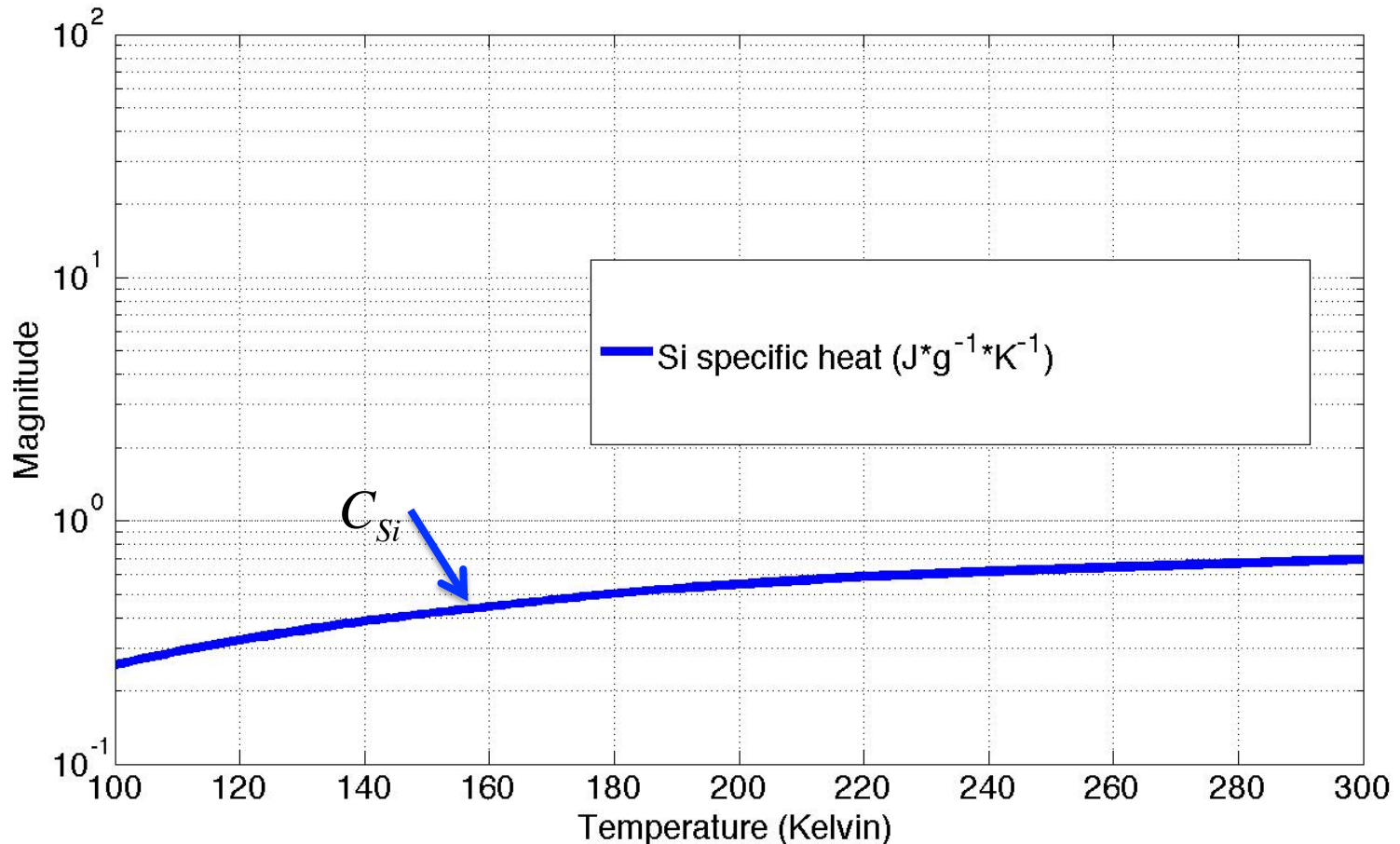
K_{CL} = thermal conductance of cold link

C_{Si} = heat capacity of silicon

- These are both functions of temperature.
- In general, the solution is not truly exponential since the time 'constant' changes.

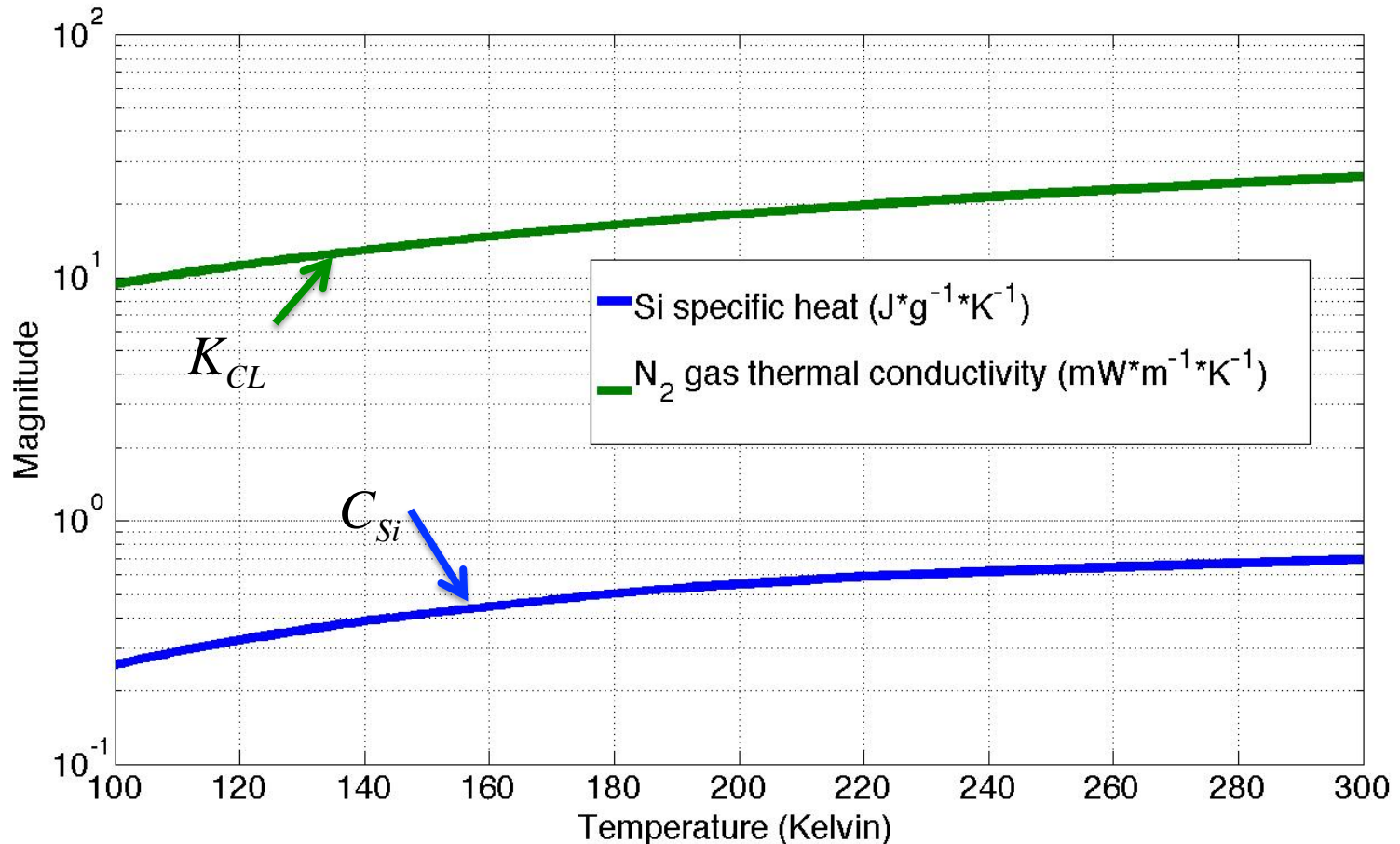
Exponential Temperature Decay

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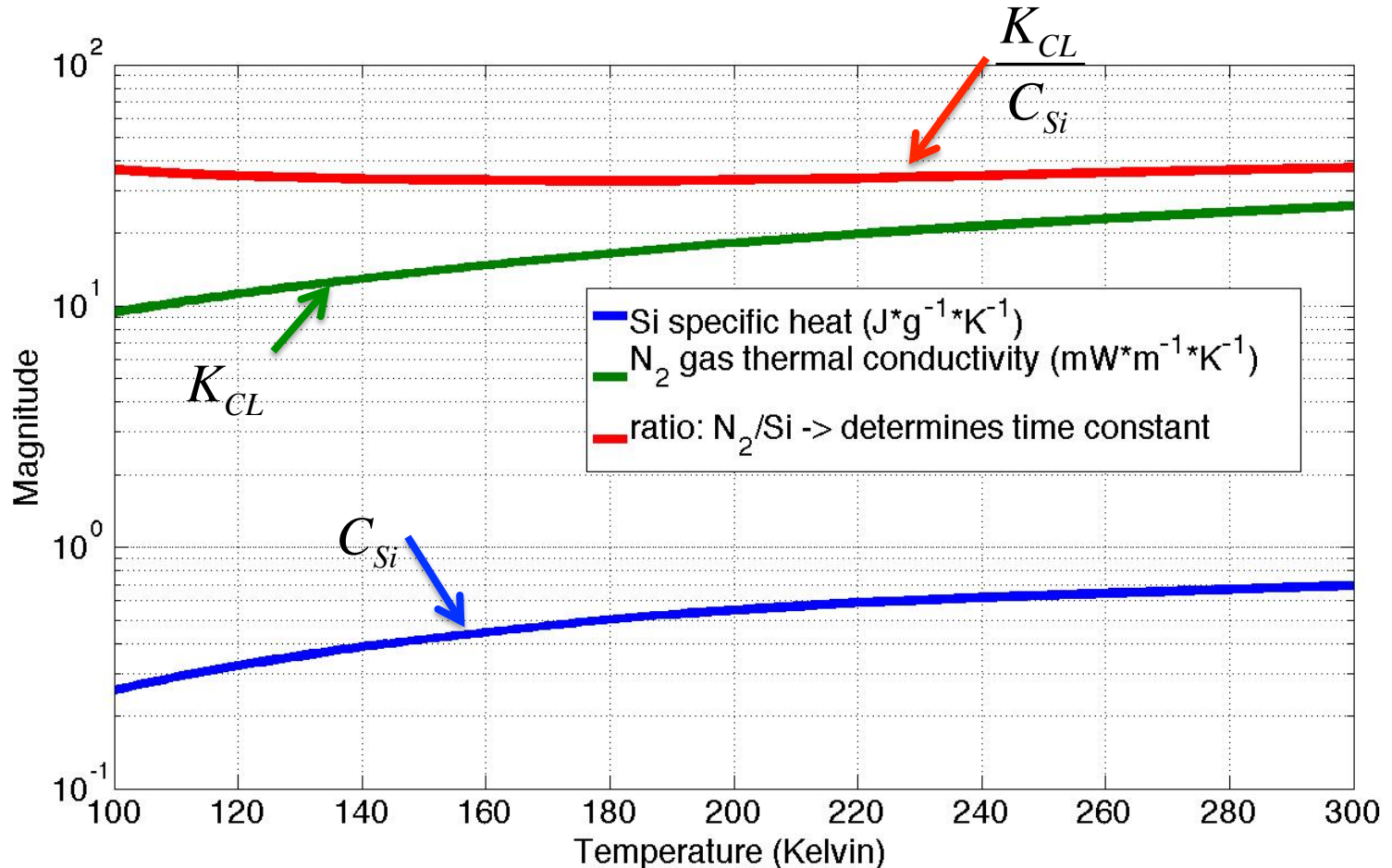
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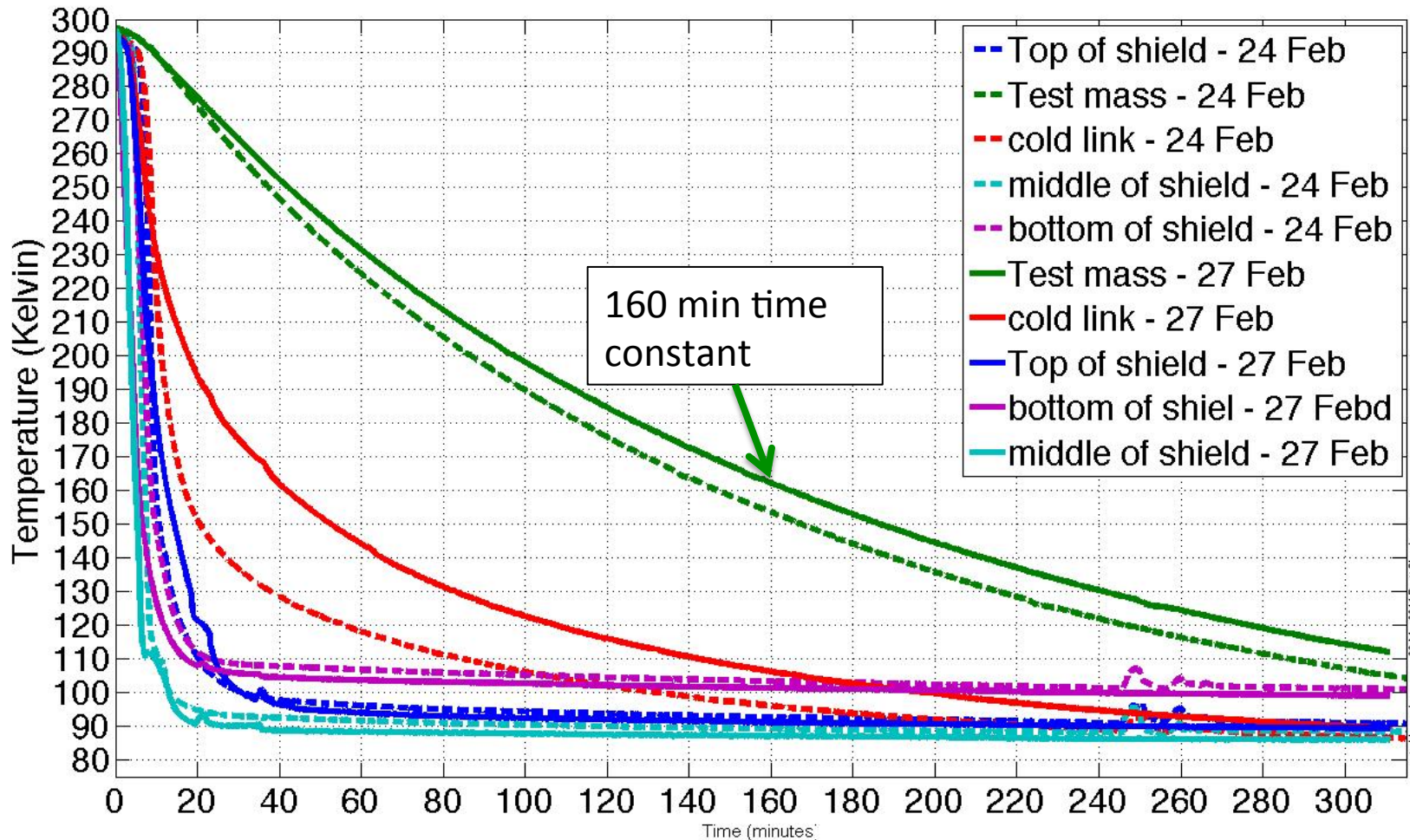
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Measurement – cold link disengaged

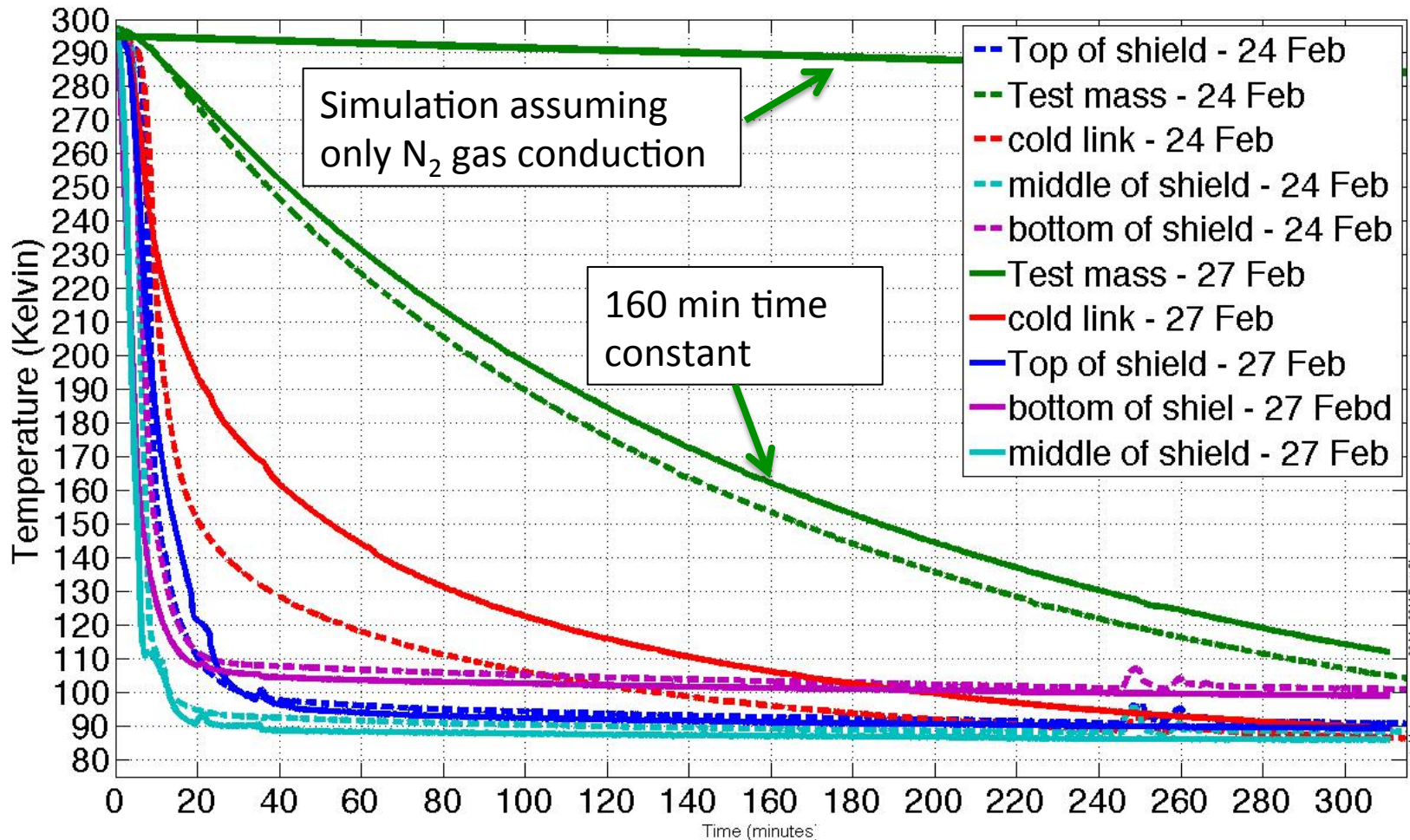
Cold link in contact - 24 Feb; Cold link backed off 0.5 inches - 27 Feb



created by workspace on 02-Mar-2014 Brett Shapiro

Measurement – cold link disengaged

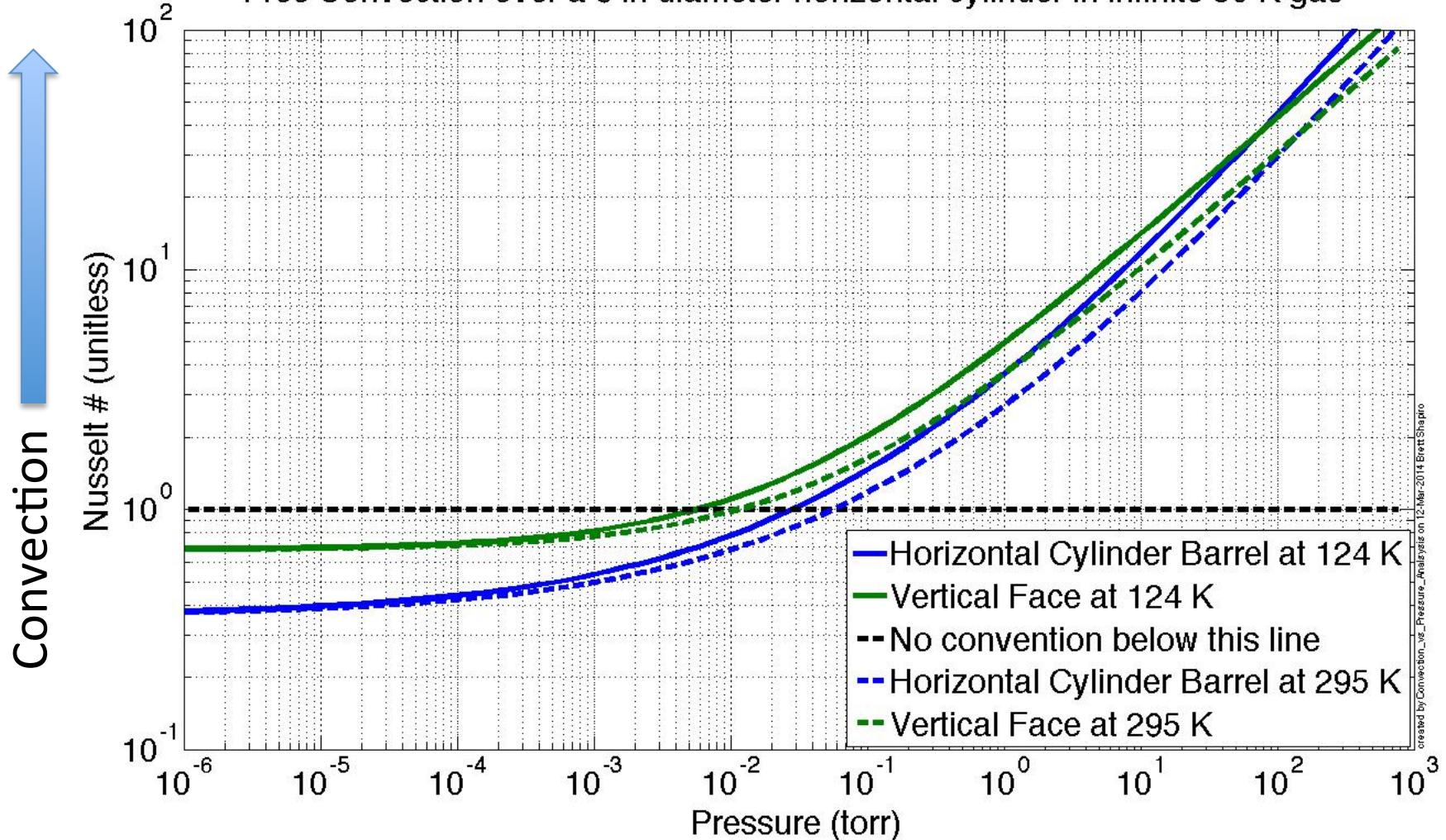
Cold link in contact - 24 Feb; Cold link backed off 0.5 inches - 27 Feb



created by workspace on 02-Mar-2014 Brett Shapiro

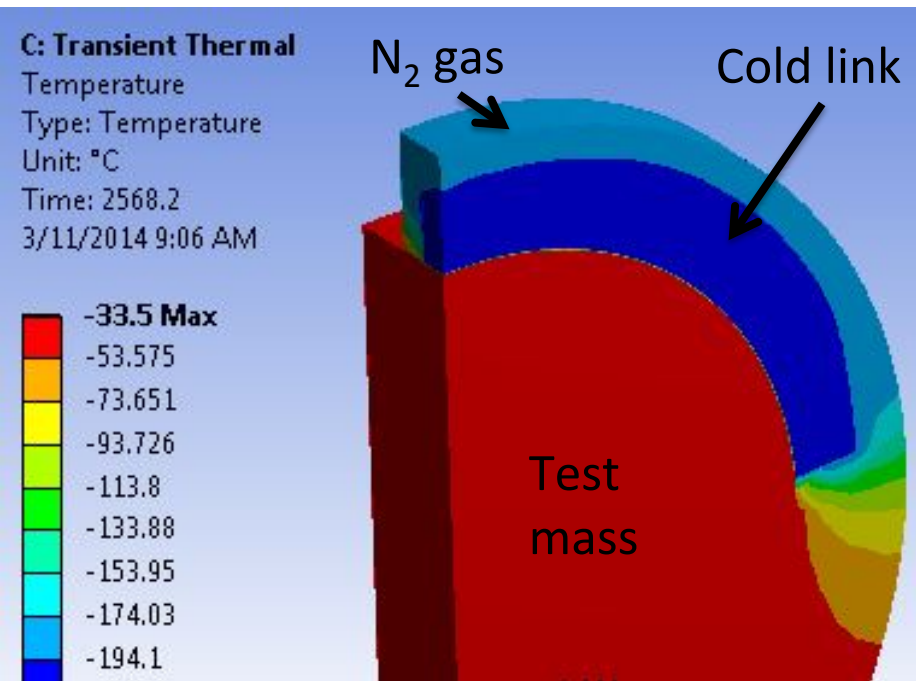
Conductive vs Convective Regimes

Free Convection over a 6 in diameter horizontal cylinder in infinite 80 K gas

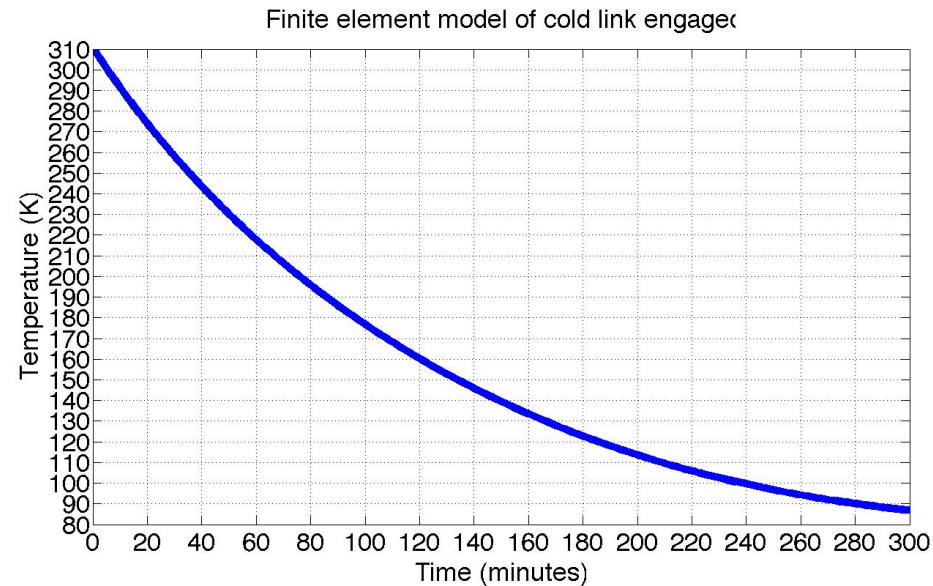


Finite Element Modeling

- Due to complexity, LIGO III designs must be verified with modeling or FEM
- Below: FEM of conduction through N₂ gas to cold link for Stanford experiment



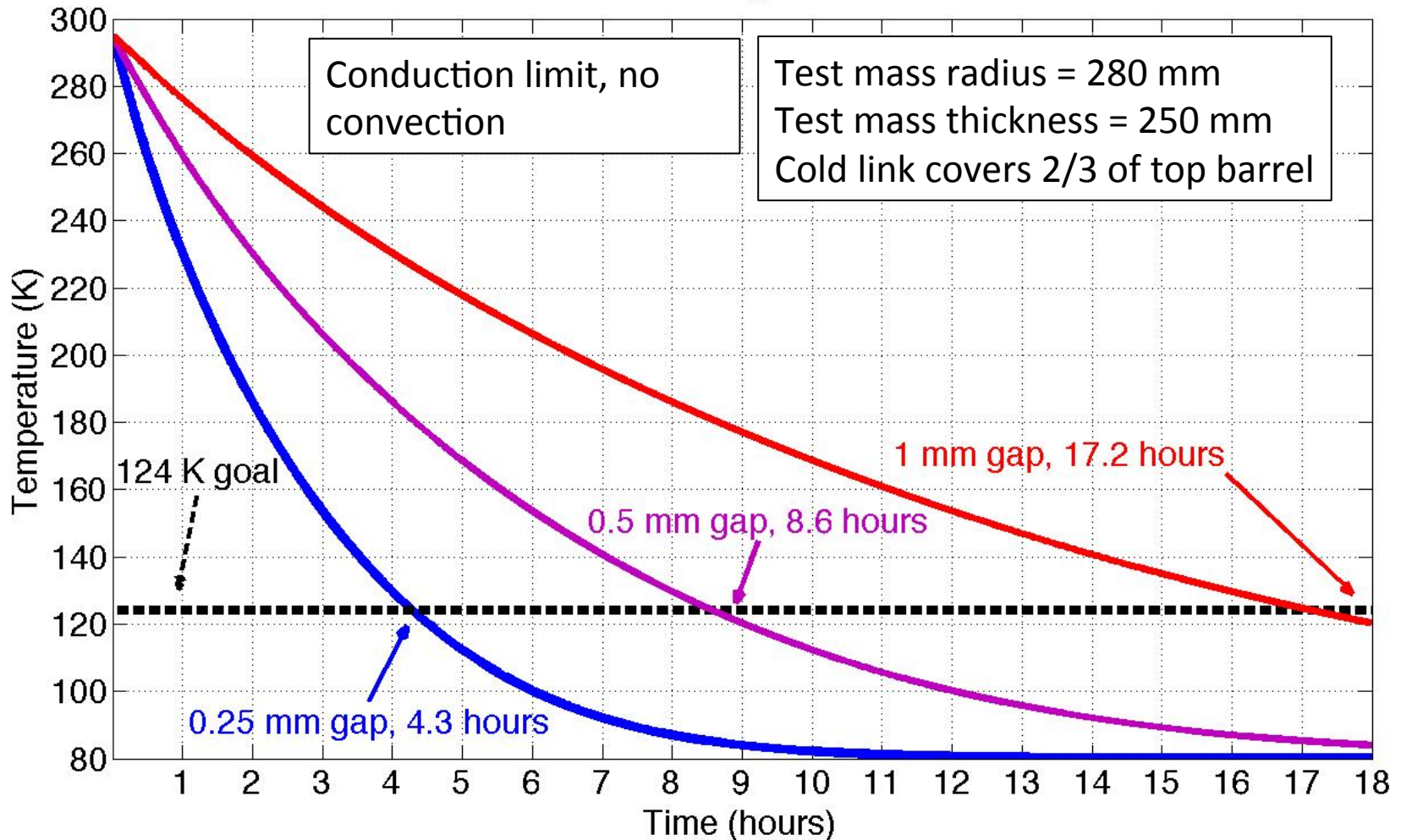
≈ 43 min into cool down



- Convective FEM is proving to require large amounts of computing power

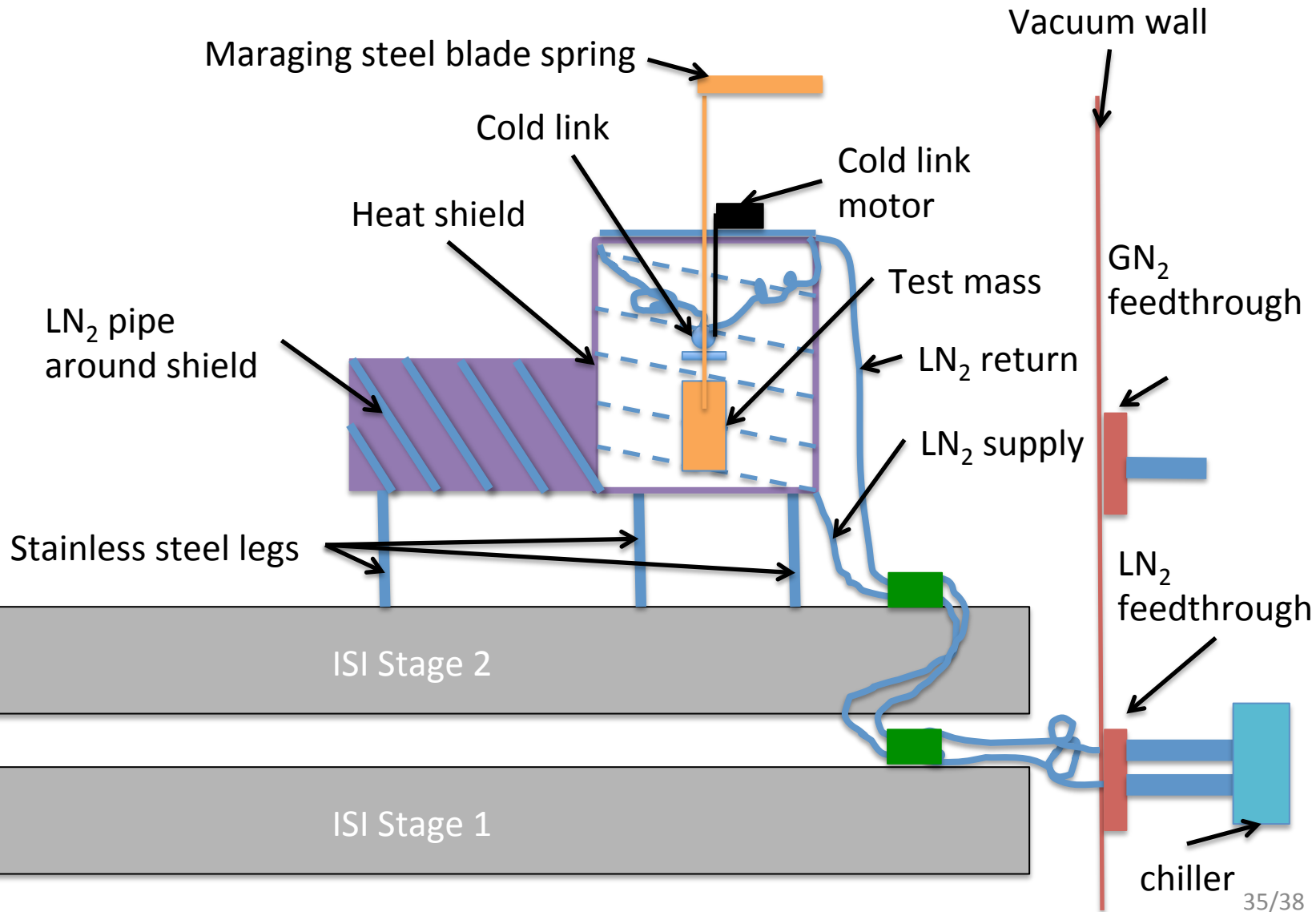
Cold link on a LIGO III test mass

Cold link on a 143 kg silicon test mass



Stanford Next Gen Experiment

Stanford Next Gen Experiment Layout



LIGO

Future work

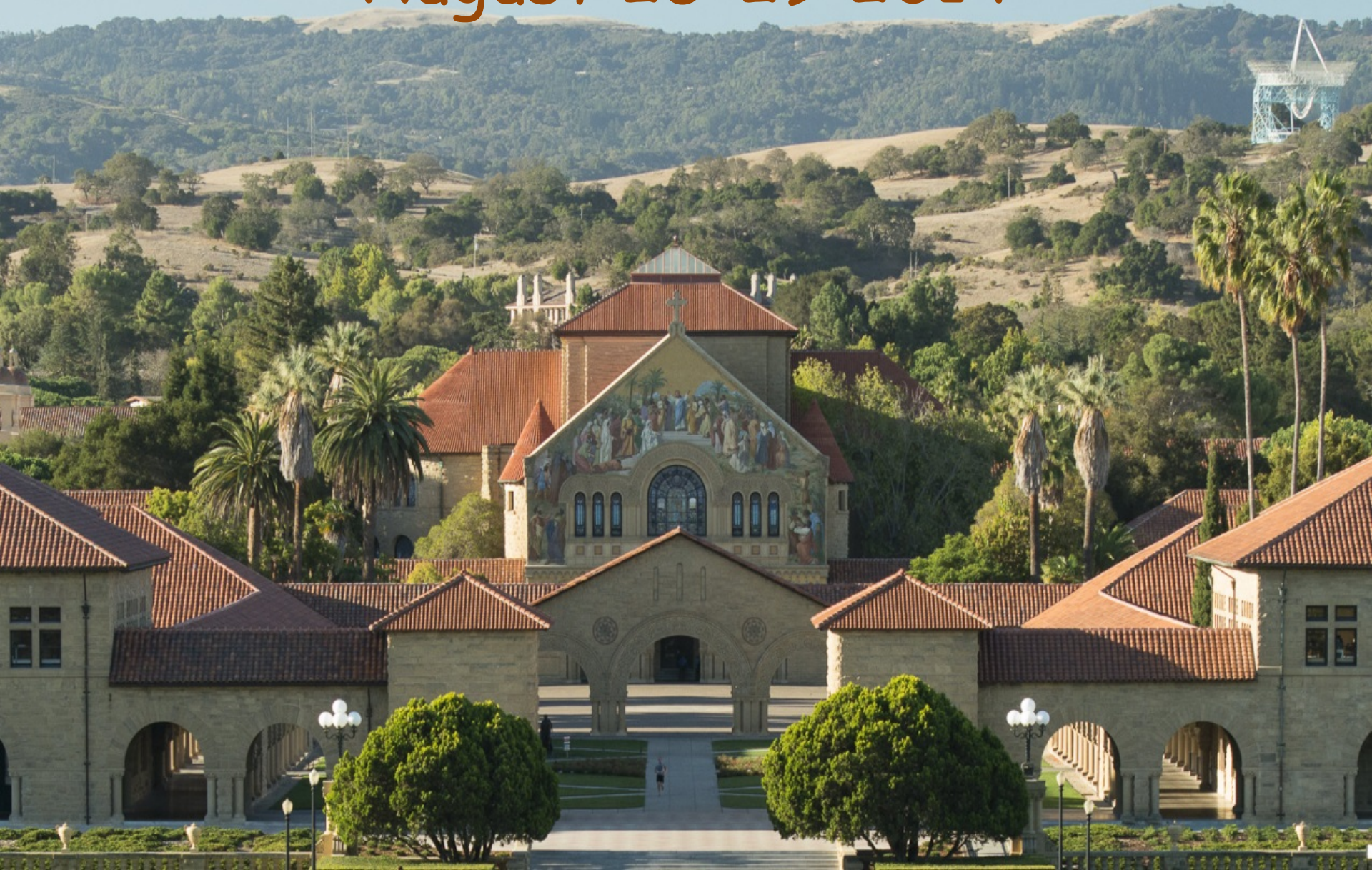


Next generation experiment using the Stanford ETF (experimental test facility)

- More realistic LIGO setup
- Measure temperature drifts on LIGO hardware, e.g. blade springs
- Measure seismic noise of nitrogen delivery and/or copper cables
- Test heat shield design
 - Black coatings
 - baffles
- Test a variety of cooling techniques
- System integration: how to make all this stuff work together
- Implement in stages
 - Cables/hoses first – test seismic noise
 - Heat shield and suspended optic
 - Install cryogenic refrigerator
 - Cavity?
 - Anything we haven't thought of yet

LVC STANFORD

August 25-29 2014



LVC STANFORD

August 25-29 2014

Accommodations

Registration – webpage coming soon

Carothers Hall

- (30) Single Rooms: \$77.20 per person/night
- (85) Double Rooms: \$56.70 per person/night

Benefits:

- 5 minute walk to conference.
- Outdoor facilities available:
- Telephone in each
- Comfortable and clean facility.
- Onsite management.
- Free laundry facilities.
- Linens provided
- Quiet areas including business center/lounge.
- Crothers Hall is exclusive to LVC.
- Free shuttle service to shopping/restaurants

Alternate accommodations (hotel)
\$180-350/night

\$250 Student rate
\$500 Non-student rate

Includes:

- Breakfast and lunch
- Banquet dinner
- Conference rooms
- AV and AV Technicians
 - Poster session

Does not include:

- Parking passes (\$10/day)
- Transportation to/from airport
 - Poster printing

Contact Claudette Earl
cearl@stanford.edu

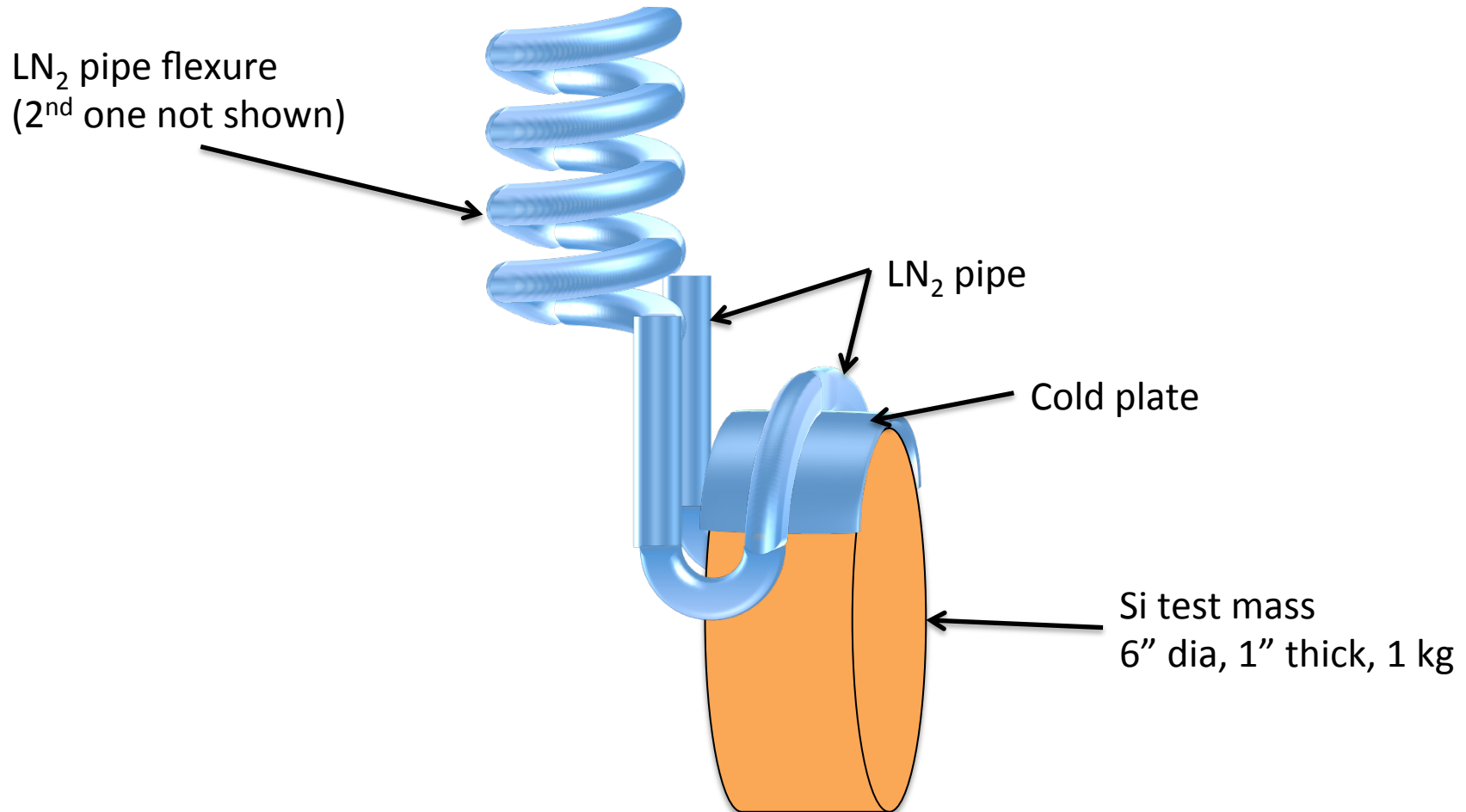
The LIGO logo consists of several concentric, slightly irregular circles in a light gray color, resembling ripples or gravitational waves, positioned in the top-left corner of the slide.

LIGO



Backups

Test Mass in Next Gen Prototype

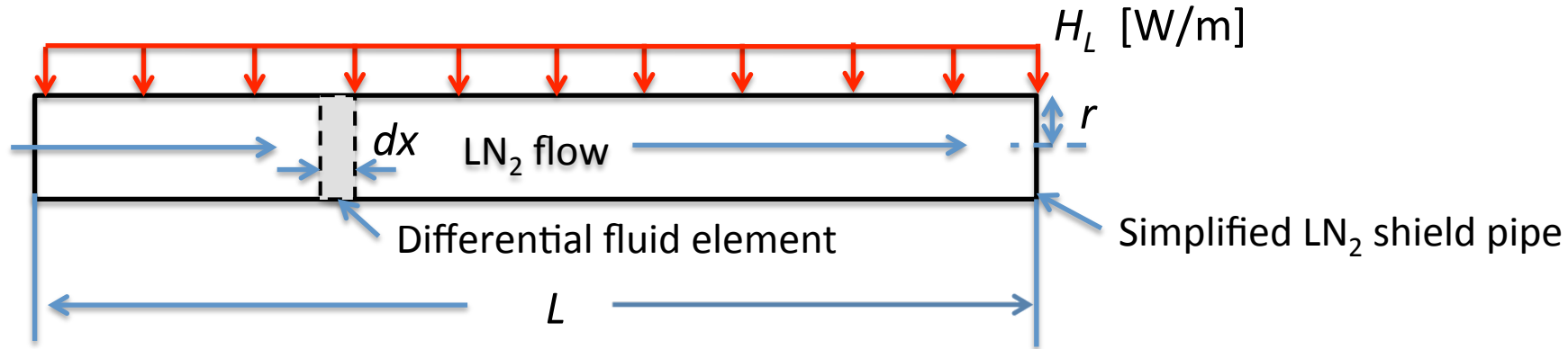


Experimentation lessons learned

- Air dominates most heat flow across contacts
- Cold links should have distributed contacts
- Solder is not leak tight against high pressure cryogenic fluids – welding is probably best
- Cryogenic fluid should have 1 flow path
- Send fluid from bottom up
- Use fatter pipes to minimize fluid pressure
- Minimize the number of materials in the plumbing – joining and contraction issues
- Leave room for differential contraction
- Silicon diode temperature sensors are great

Steady State LN₂ flow rate

What flow rate do we need to maintain the shield at about ≈ 80 K?



Liquid nitrogen parameters

$$C_p = 2041 \text{ J/kg} \cdot \text{K} \text{ at boiling}$$

$$\rho = 806.1 \text{ kg/m}^3$$

Q_m = mass flow rate in kg/s

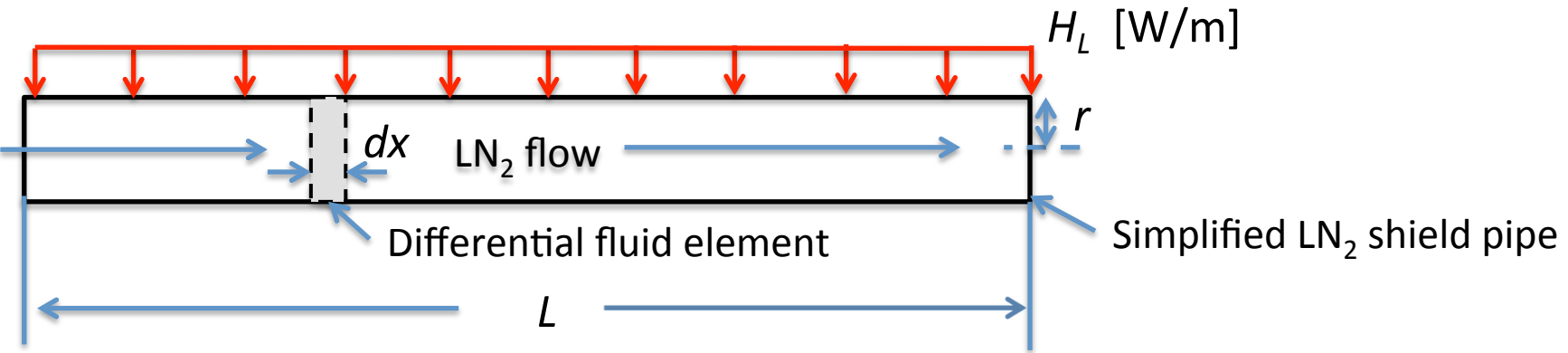
$$Q_v = Q_m / \rho = \text{volumetric flow rate in m}^3 / \text{s}$$

$$1 \text{ m}^3 = 1000 \text{ liter}$$

Heat load in shield

$$P = 10 \text{ W}$$

$$H_L = P / L \text{ [W/m]}$$



Velocity of fluid element for a given flow rate

$$v = \frac{Q_v}{\pi r^2} = \frac{Q_m}{\rho \pi r^2} \text{ [m/s]}$$

Time it takes fluid element to travel pipe

$$t = L / v = \frac{L \rho \pi r^2}{Q_m} \text{ [s]}$$

Mass of fluid disk-element

$$dm = \rho \pi r^2 dx \text{ [kg]}$$

Heat capacity of fluid element

$$C_p dm = C_p \rho \pi r^2 dx \text{ [J/K]}$$

Power into fluid element

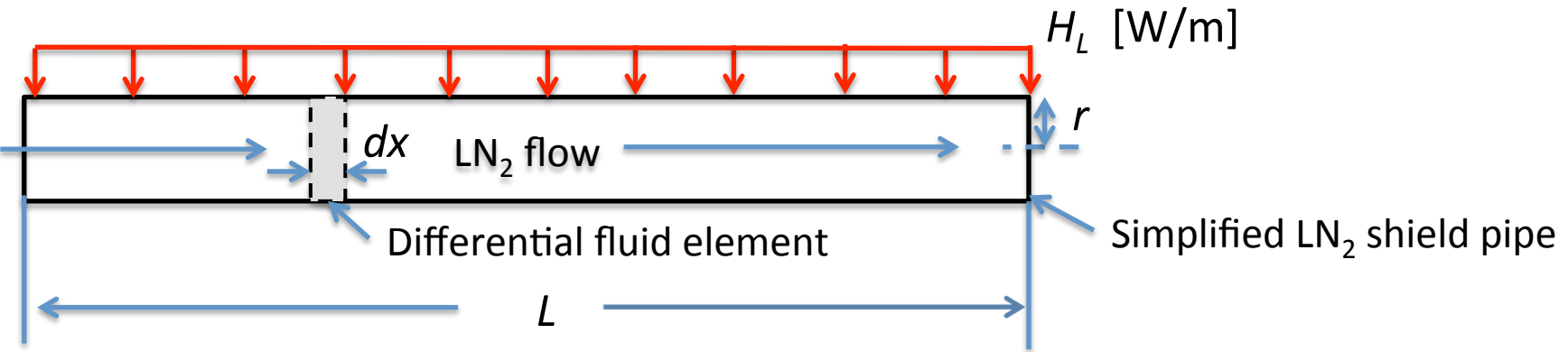
$$H_L dx = \frac{P}{L} dx \text{ [W]}$$

Total heat energy into fluid element
for time in pipe

$$H_L t dx = \frac{P \rho \pi r^2}{Q_m} dx \text{ [J]}$$

Temperature rise in fluid

$$\Delta T = \frac{H_L t dx}{C_p dm} = \frac{P}{\rho C_p Q_v} \text{ [K]}$$



Volumetric flow rate

$$Q_v = \frac{H_L t dx}{C_p dm} = \frac{10^6 P}{\rho C_p \Delta T} \text{ [ml/s]}$$

if we allow for

$$\Delta T = 10 \text{ K}, P = 10 \text{ W}$$

then,

$$Q_v = \frac{10^6 (10)}{806.1(2041)10} = 0.61 \text{ ml/s}$$

Scaling laws of LN₂ pipes vs. Cu cables

Cu cable

$$H = K_{th} \frac{N\pi r^2}{L} \Delta T$$

$$K_L = \frac{NEr^4}{nD^3(1+\nu)}$$

LN₂ pipe

$$H = \rho C_{P, LN_2} Q_v \Delta T$$

$$Re = \frac{2\rho Q_v}{\mu N \pi r_i} < 2000$$

$$K_L = \frac{NE(r_o^4 - r_i^4)}{nD^3(1+\nu)}$$

$$\Delta P = \frac{8\mu L Q_v}{N \pi r^4} < 1 \text{ atm}$$

In general,
 $K_{\text{any axis}} \propto \frac{r^4}{L^3}$

H = heat flow [W]

Re = Reynolds #

K_L = Longitudinal coil spring stiffness [N/m]

P = pipe pressure [Pa]

r = conductor radius

ρ = LN₂ density [kg/m³]

C_{P, LN_2} = specific heat of LN₂ [J/kg·K]

Q_v = LN₂ flow rate [m³/s]

K_{th} = thermal conductivity

L = pipe or conductor length [m]

T = temperature [K]

E = Young's modulus [Pa]

ν = Poisson ratio

μ = LN₂ viscosity [Pa·s]

D = coil spring diameter [m]

n = number of spring turns

N = number of conductors

Scaling laws cont.

Cu cable

$$r = \sqrt{\frac{H}{\Delta T} \frac{L}{N\pi K_{th}}}$$

LN₂ pipe

$$Q_v = \frac{H}{\Delta T \rho C_{P, LN_2}}$$

$$\text{Re}_{\max} = \frac{2}{\mu N \pi r_i C_{P, LN_2}} \frac{H}{\Delta T} = 1200$$

$$r_i [N] = \frac{2}{\text{Re}_{\max} \mu N \pi C_{P, LN_2}} \frac{H}{\Delta T}$$

In general,

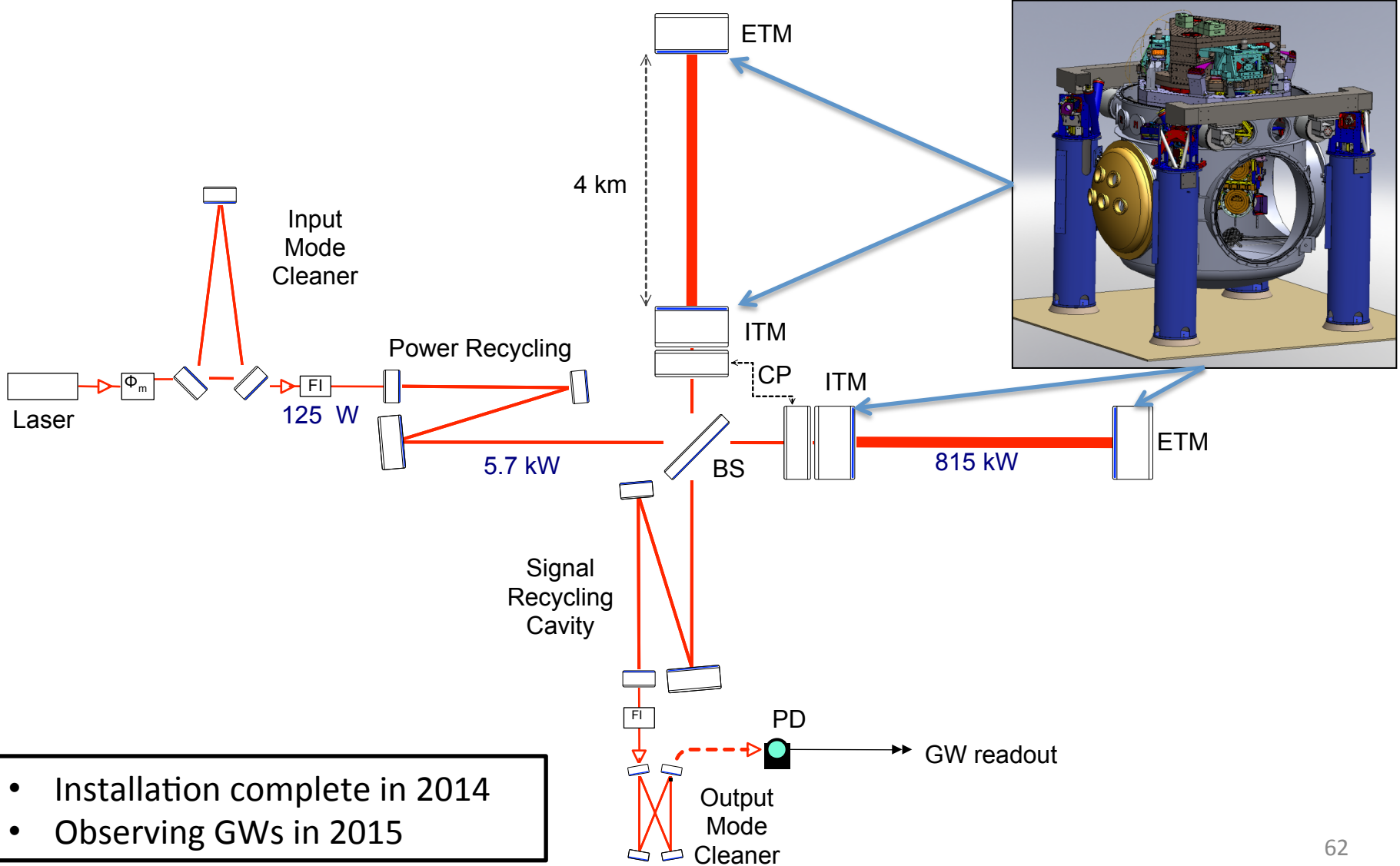
$$K_{\text{any axis}} \propto \frac{r^4}{L^3}$$

$$K_L = \frac{NEr_i^4 \left[\left(1 + \frac{t}{r_i [N]} \right)^4 - 1 \right]}{nD^3(1+\nu)}$$

$$K_L = \frac{E}{nD^3(1+\nu)N} \left(\frac{H}{\Delta T} \frac{L}{\pi K_{th}} \right)^2$$

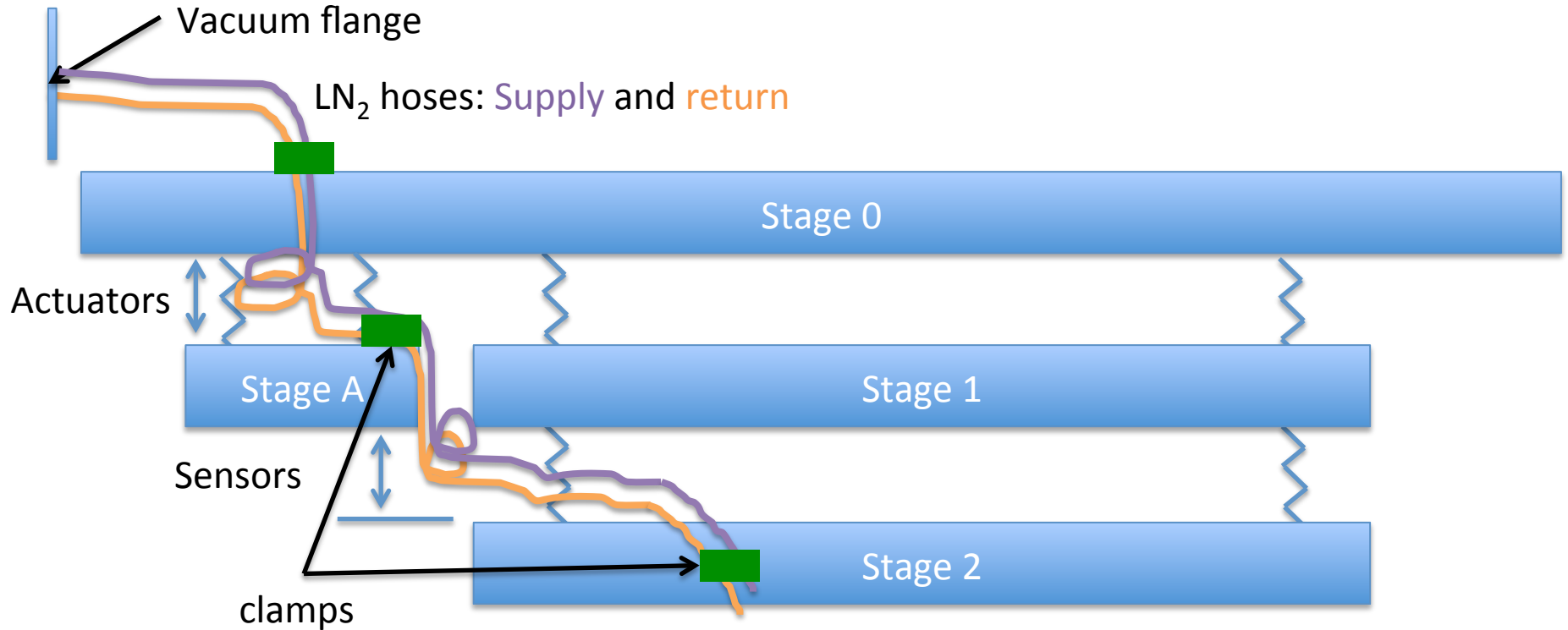
$$K_L = \frac{E \left[\left(1 + \frac{t}{r_i [N]} \right)^4 - 1 \right]}{nN^3 D^3 (1+\nu)} \left(\frac{2}{\text{Re}_{\max} \mu \pi C_{P, LN_2}} \frac{H}{\Delta T} \right)^4$$

Advanced LIGO Layout



- Installation complete in 2014
- Observing GWs in 2015

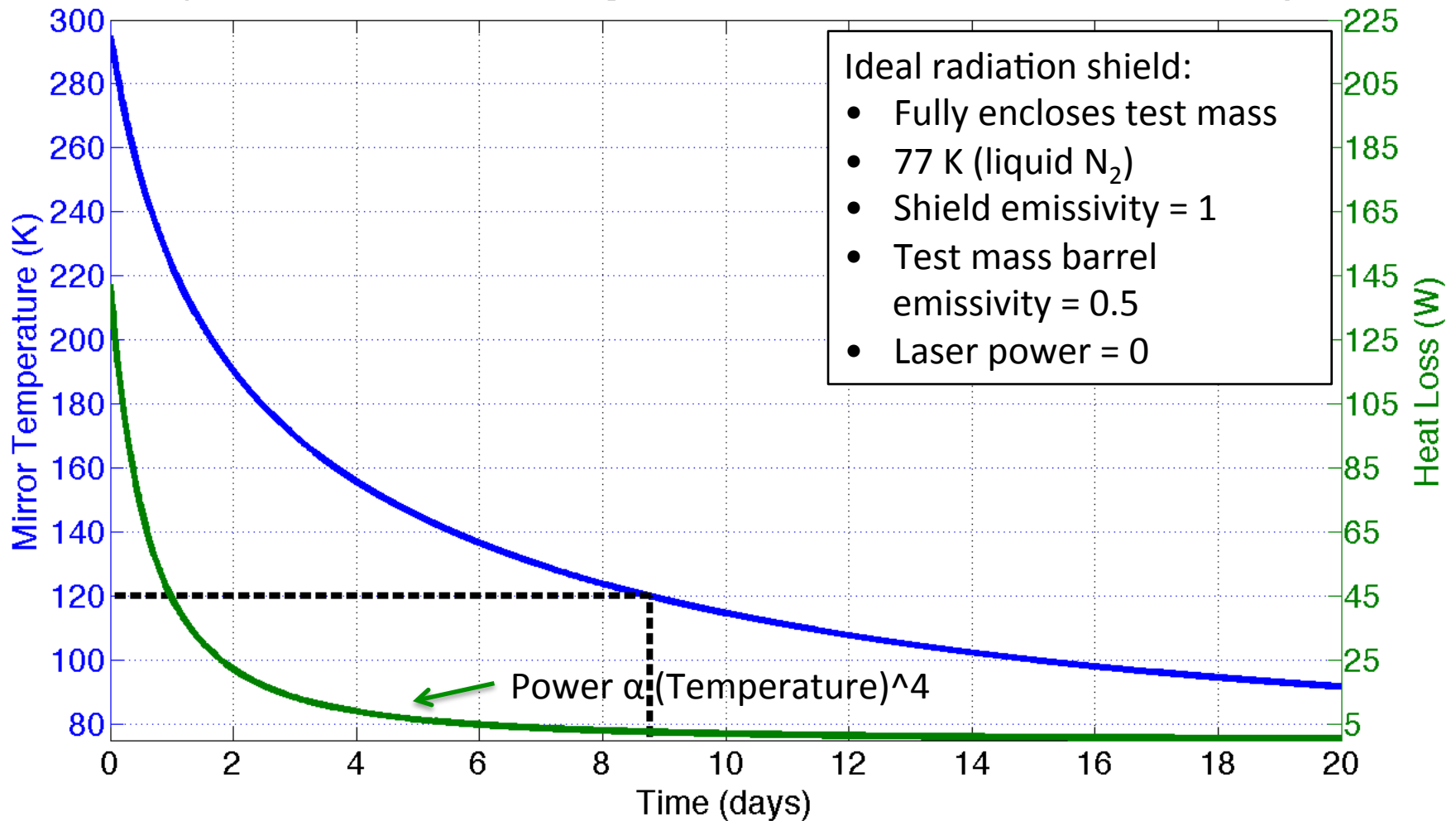
How to get a LN₂ Hose to ST2



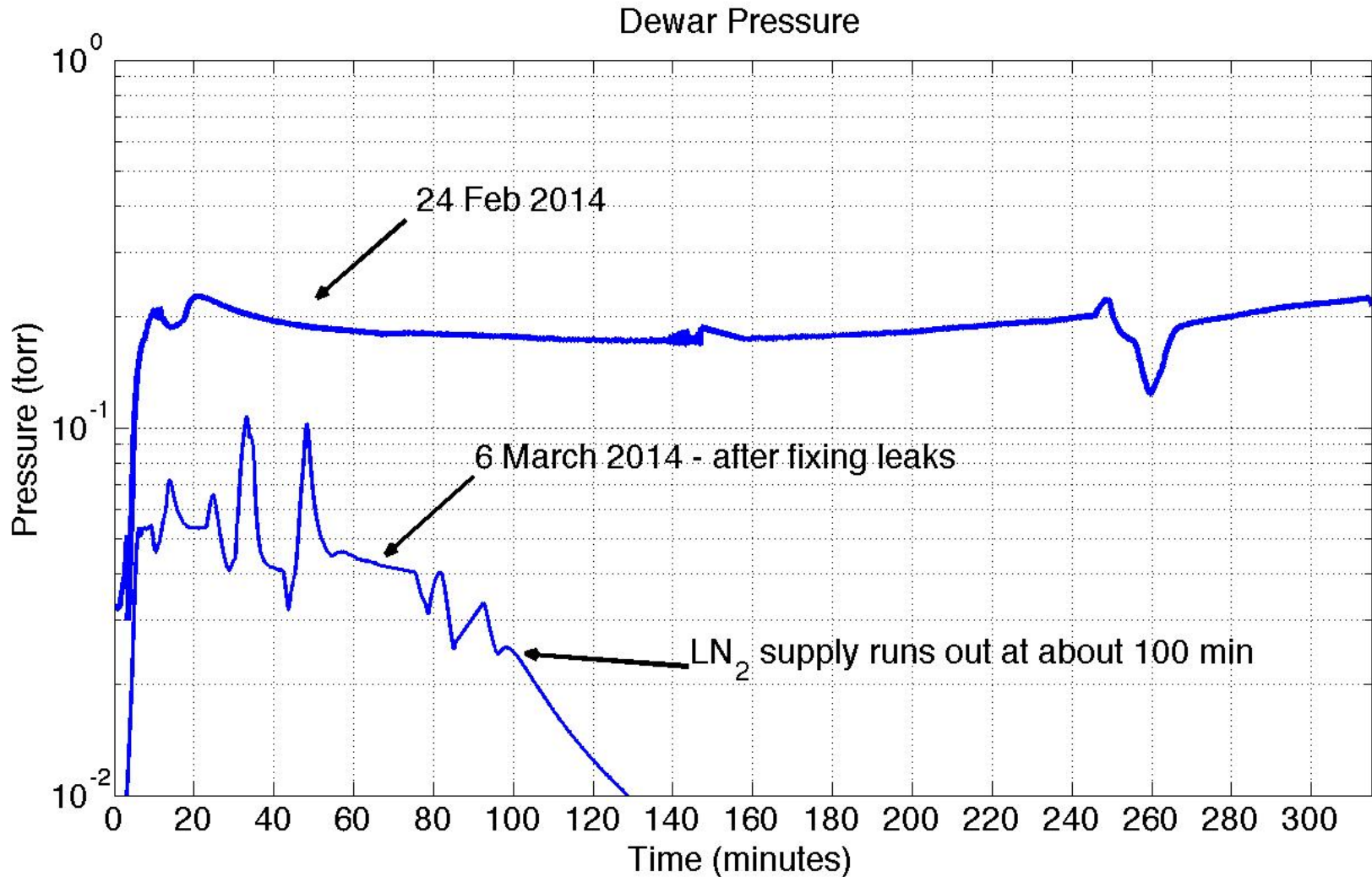
Extra stage, A, in parallel with stage 1 carries hose. Stage A is actuated to follow stage 2 so the hose does not short seismic isolation. Stage A sensor noise is set by the stage 2 isolation requirement (so it follows stage 2 and not the sensor noise).

Test Mass Radiation Simulation

Temperature evolution of a 143 kg silicon test mass. Time to 120 K = 8.7653 days.

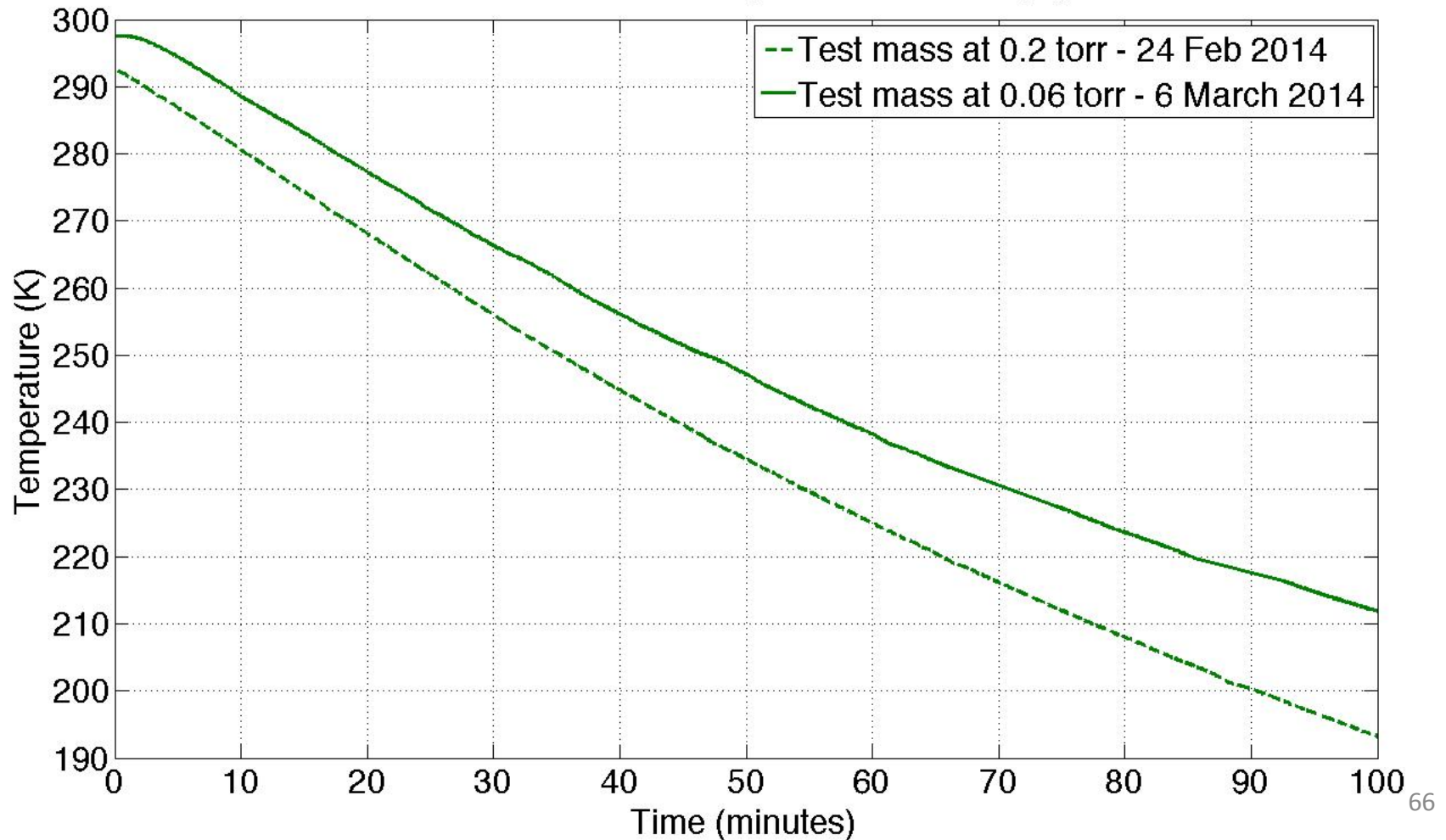


Dewar pressure during measurements

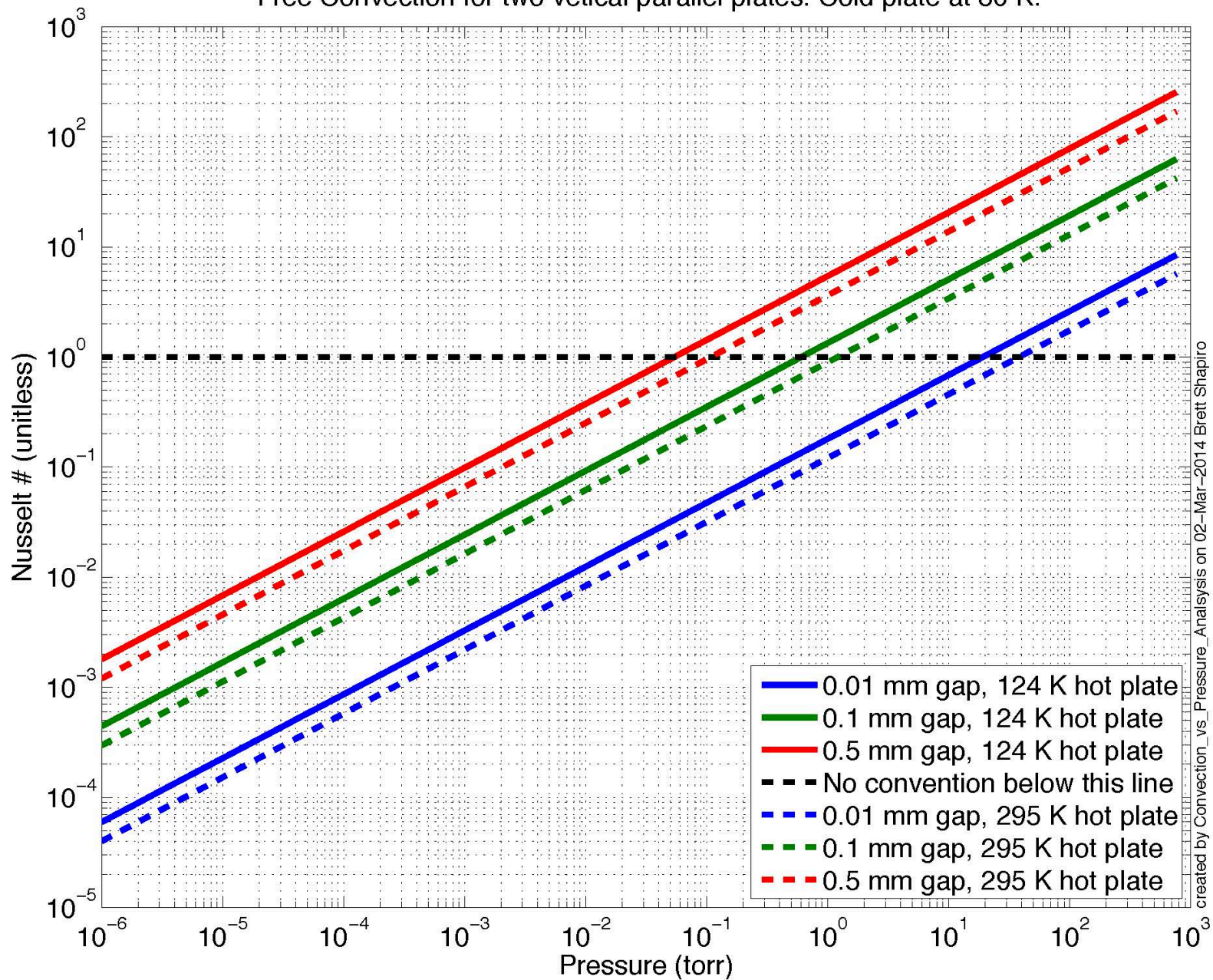


Effect of pressure on test mass temp

Silicon Test Mass Cooling - Cold link disengaged



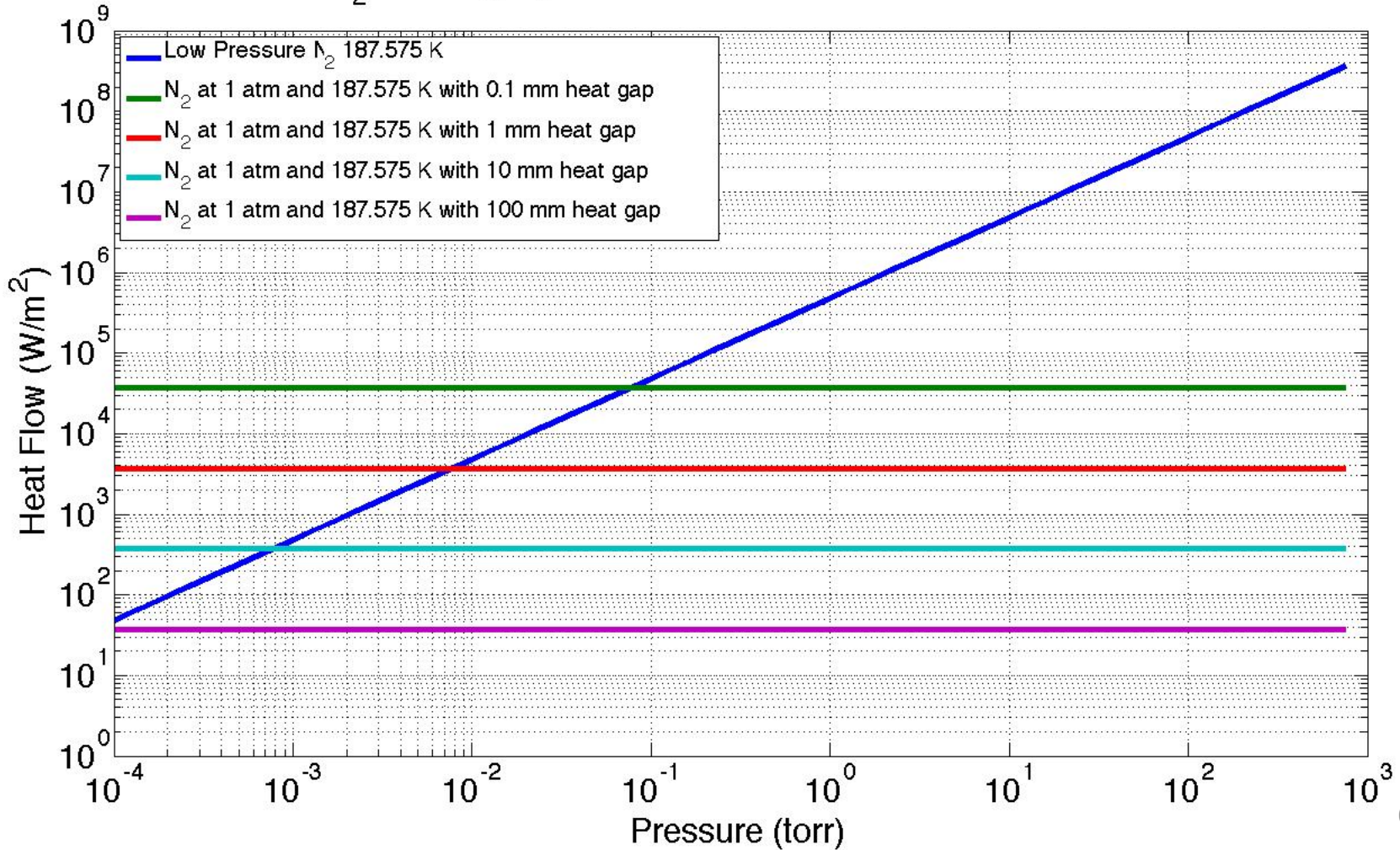
Free Convection for two vertical parallel plates. Cold plate at 80 K.



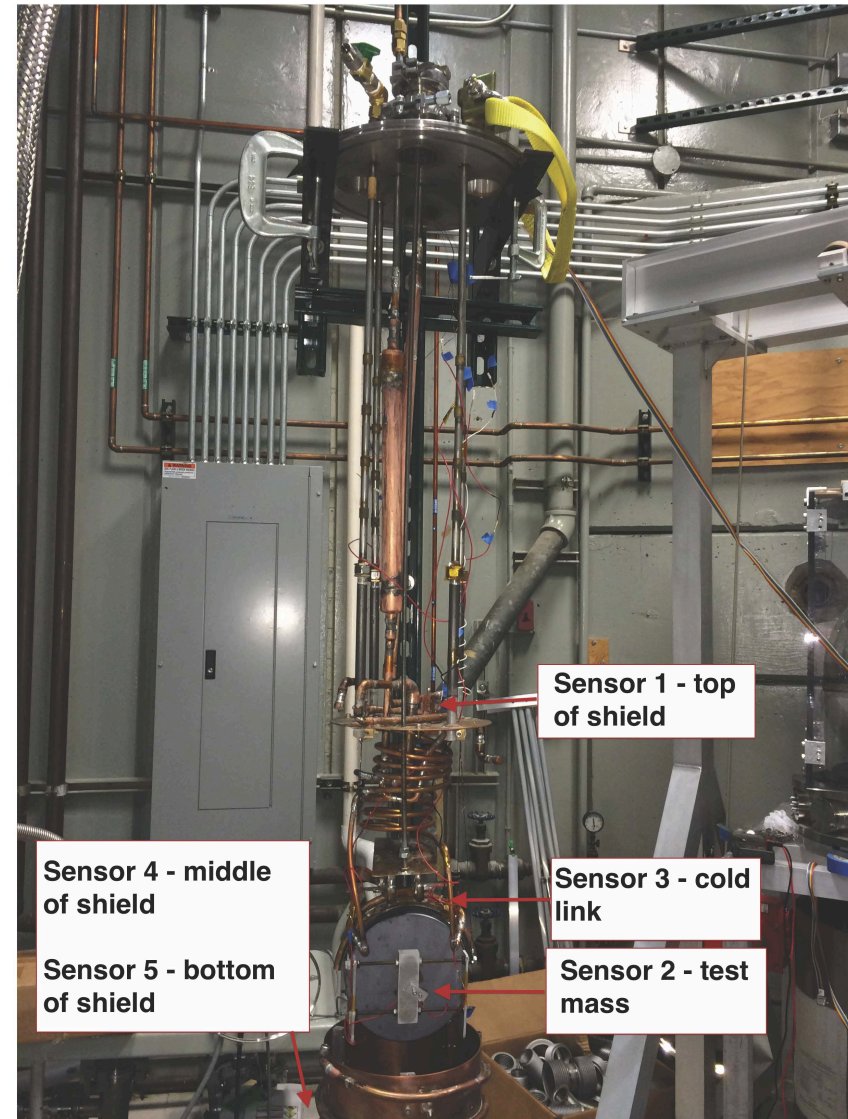
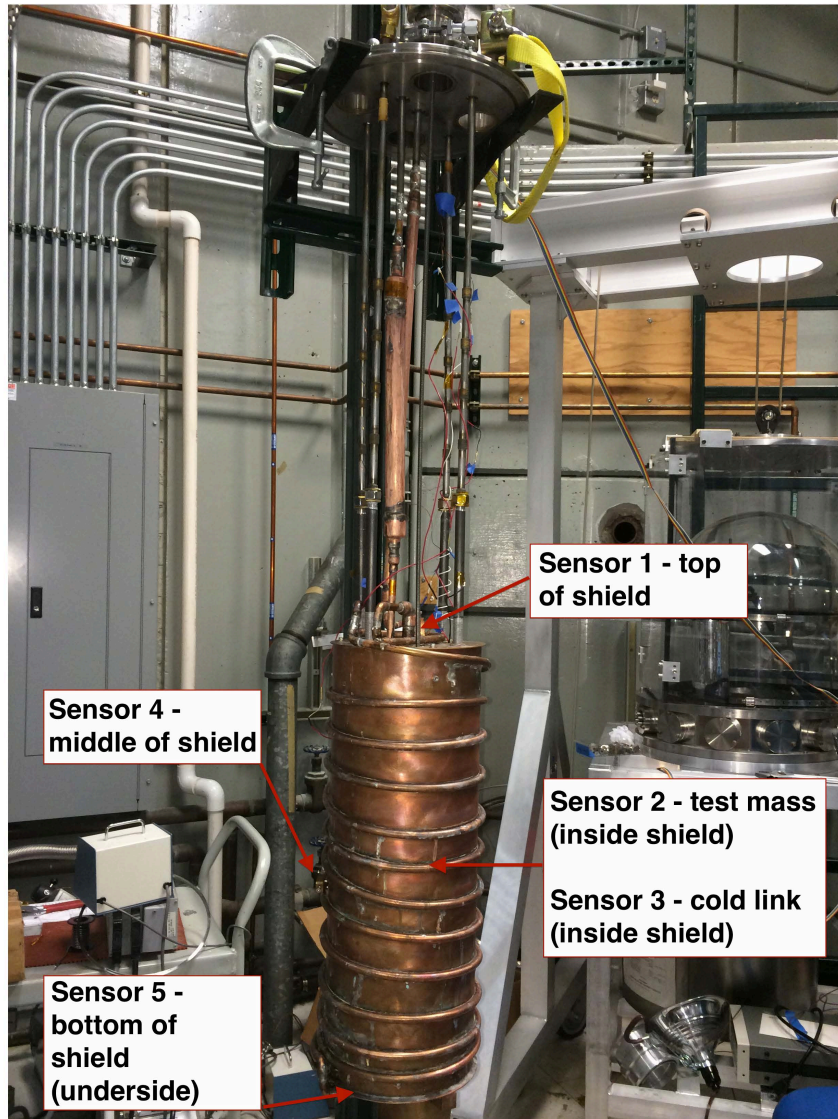
created by Convection_vs_Pressure_Analysis on 02-Mar-2014 Brett Shapiro

N₂ gas therm. cond. vs pressure

Heat flow with N₂ exchange gas from a 295.15 K test mass to a 80 K heat sink.



Temperature Sensor Locations



Test Mass Temperature Equations

$$\dot{Q}_{Si} = K_{CL} \Delta T$$

$$\dot{T}_{Si} = \frac{\dot{Q}_{Si}}{C_{Si}}$$

$$\dot{T}_{Si} = \frac{K_{CL}}{C_{Si}} (T_{cold} - T_{Si})$$

$$\dot{T}_{Si} + \frac{K_{CL}}{C_{Si}} T_{Si} = \frac{K_{CL}}{C_{Si}} T_{cold}$$

$$T_{Si} = T_{hot} e^{-\frac{K_{CL} t}{C_{Si}}} + T_{cold}$$

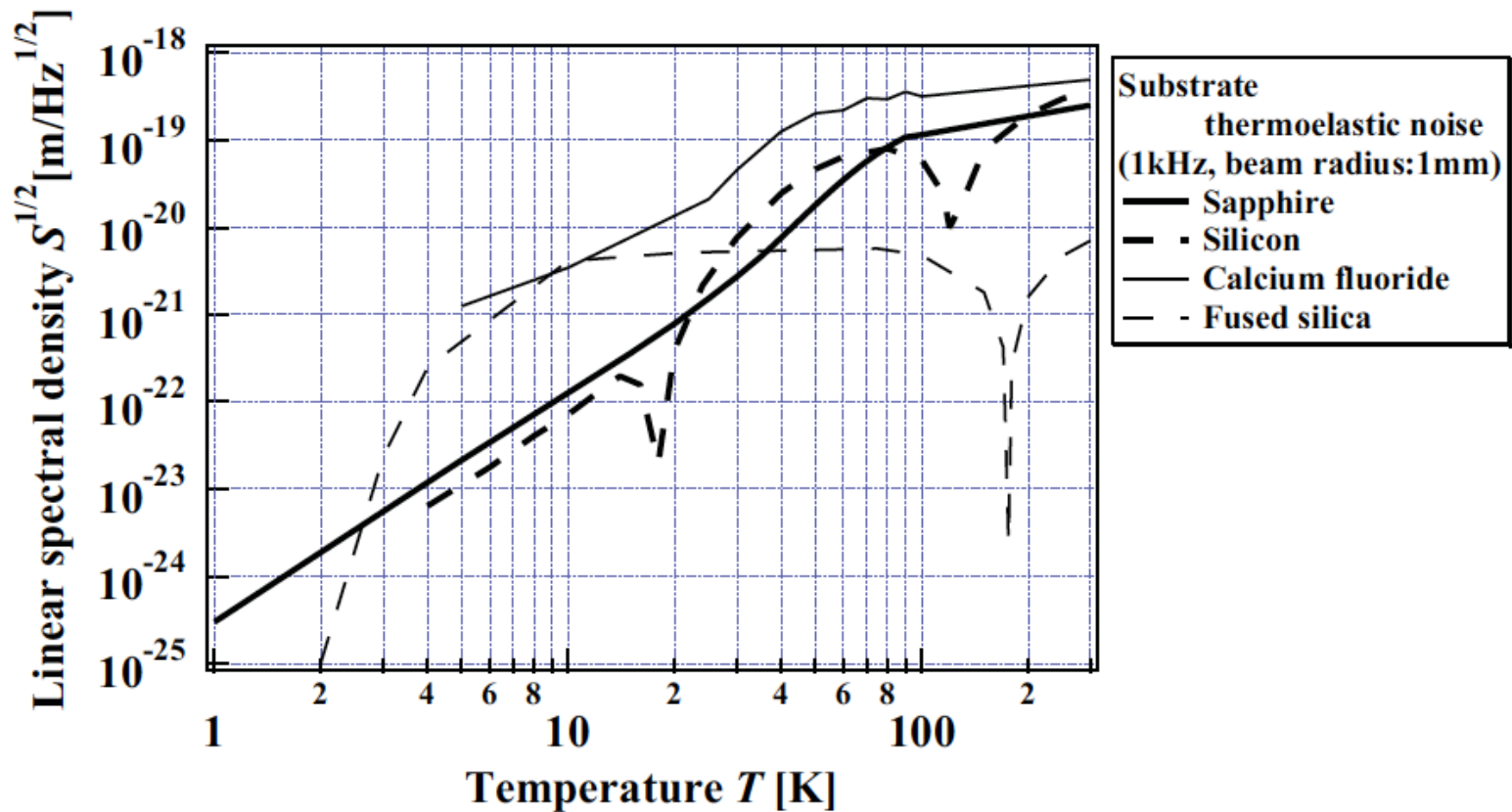
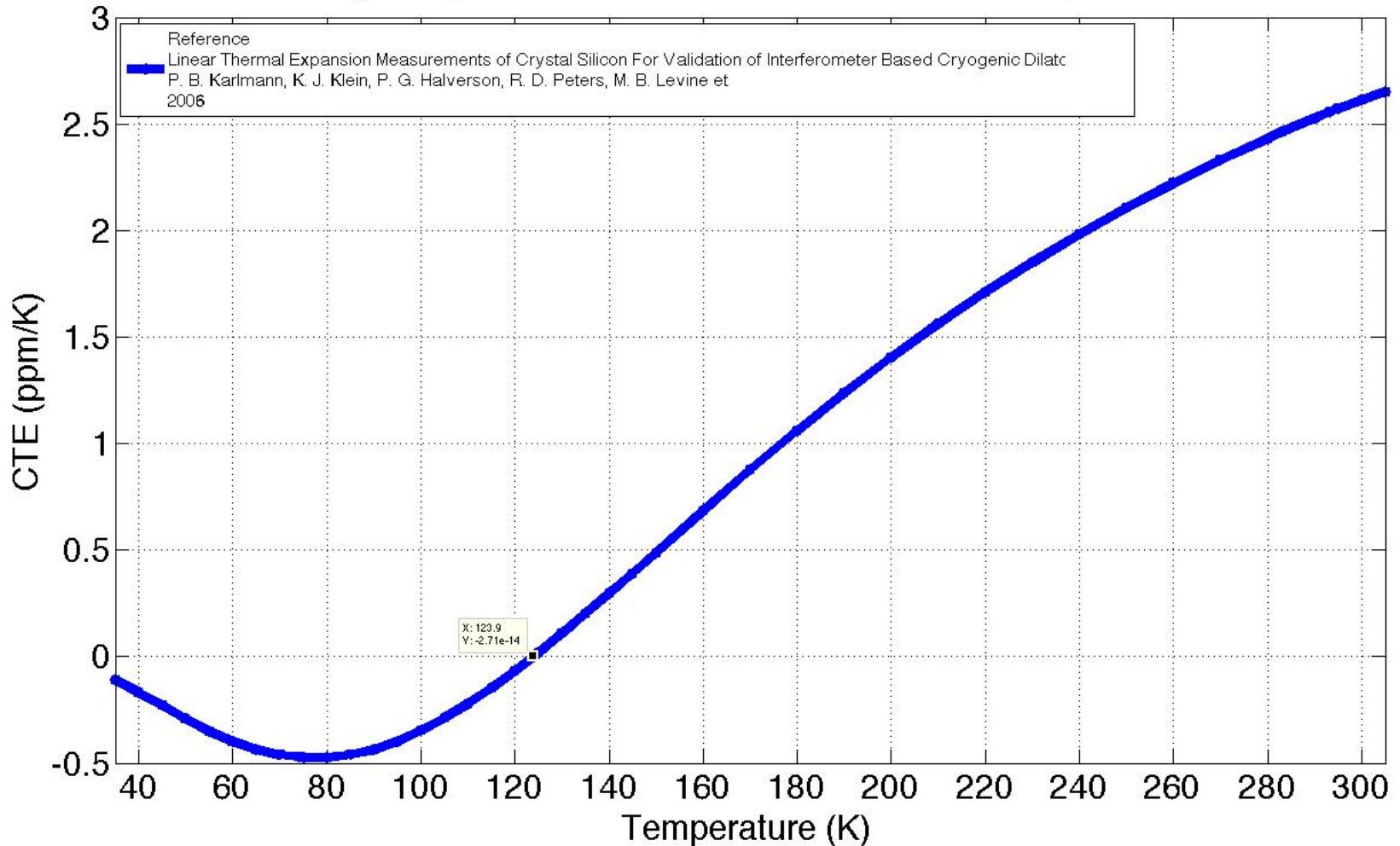


Figure 7.4 Temperature dependence of substrate thermoelastic noise. Frequency f is 1 kHz and beam radius w_0 is 1 mm.

ref: Harry, Bodiya, Desalvo. Optical Coatings and Thermal Noise in Precision Measurement. 2012. pg 113.

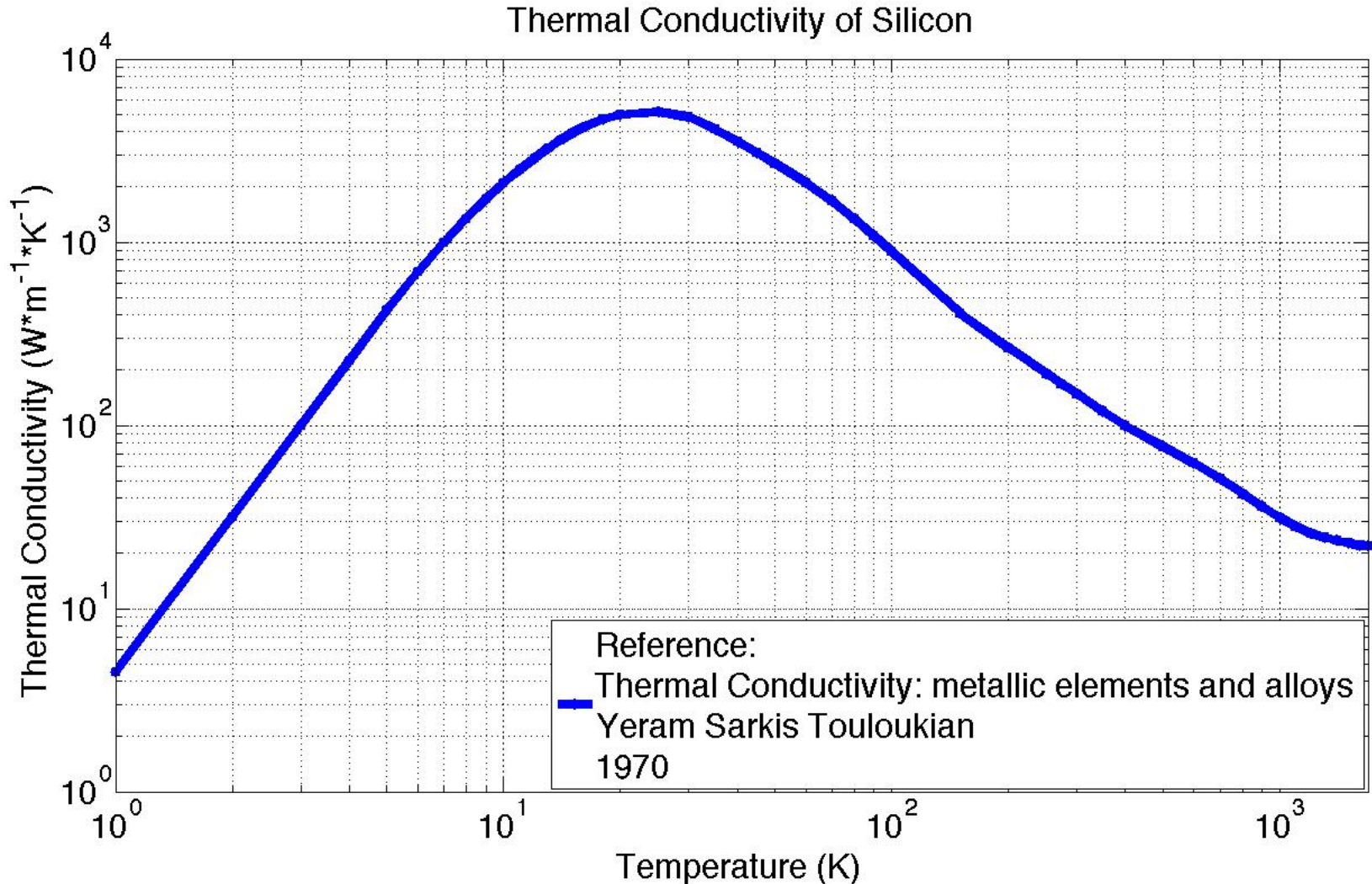
Si CTE vs Temperature

Single Crystalline Silicon Coefficient of Thermal Expansion

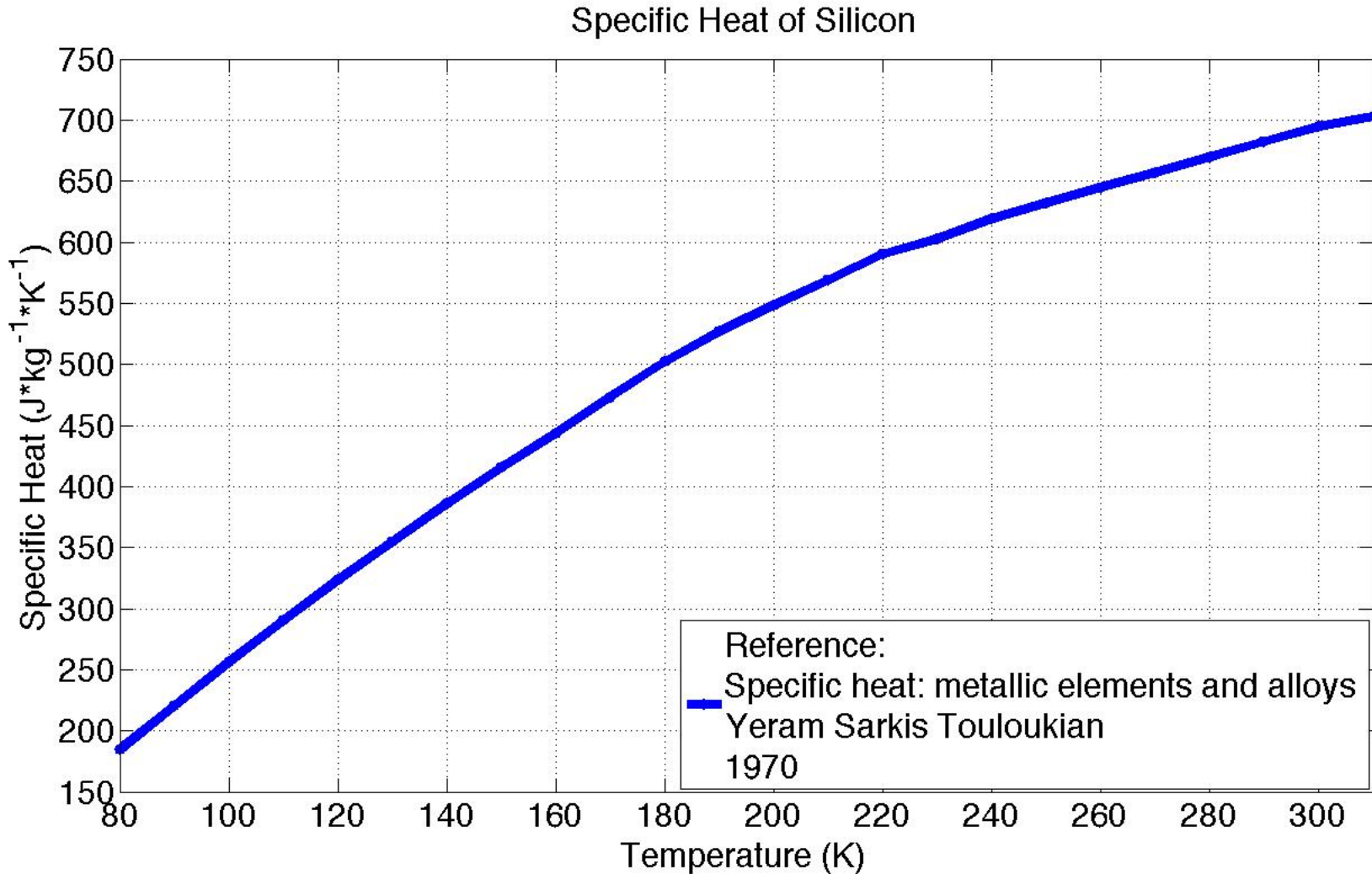


Thermoelastic component of thermal noise goes to zero with CTE.

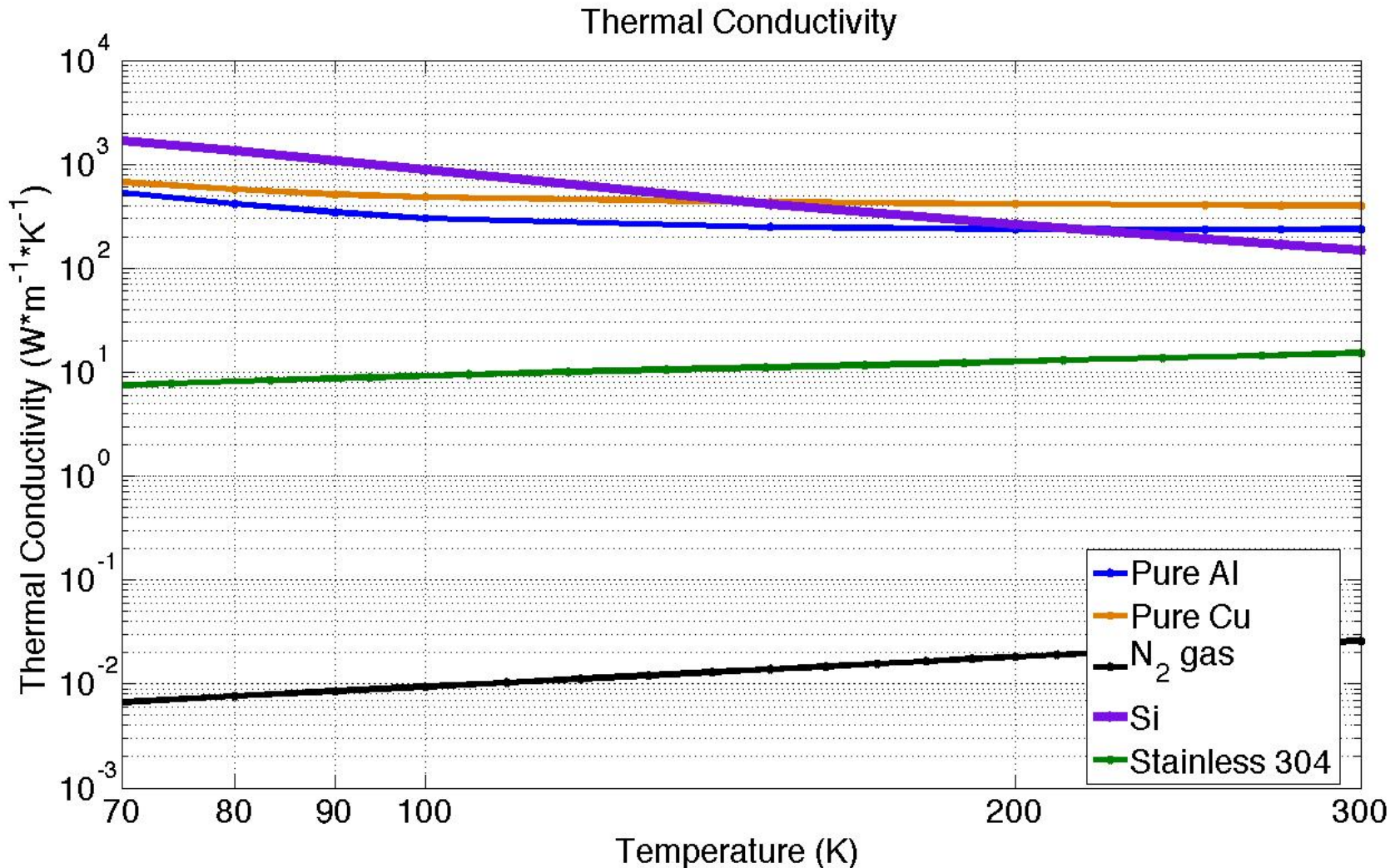
Si Thermal Conductivity vs Temp.



Si Specific Heat vs Temperature



Thermal Conductivity of Materials



References: Stainless: NIST, CRYOGENICS TECHNOLOGIES GROUP, <http://cryogenics.nist.gov/>

Al, Cu, Si: Thermal Conductivity: metallic elements and alloys, Touloukian 1970

N₂ gas: Thermal conductivity: nonmetallic liquids and gases, Touloukian 1970

Other Problems To Solve

- Flexibility if liquid N₂ hoses or Cu cables
- Temperature/height control of blade springs
- Test mass temperature control
- Test mass temperature tolerance
- How to measure temperature?
 - Measure acoustic modes – Young's modulus is temp. dependent
 - Measure test mass diameter – combined with CTE data gives temperature
 - Infrared camera
- Emissivity of optical coatings
- Lossiness of emissive coatings
- Good emissivity estimates/measurements of Si?
- Power absorption in Si and Si coatings (ppm, W, etc)?
- How noisy is flowing laminar liquid nitrogen: seismic, Newtonian?
- Optical coating thermal noise at 124 K
- How to actuate the test mass – is the ESD out?
- Can we put viewports in the heat shield?