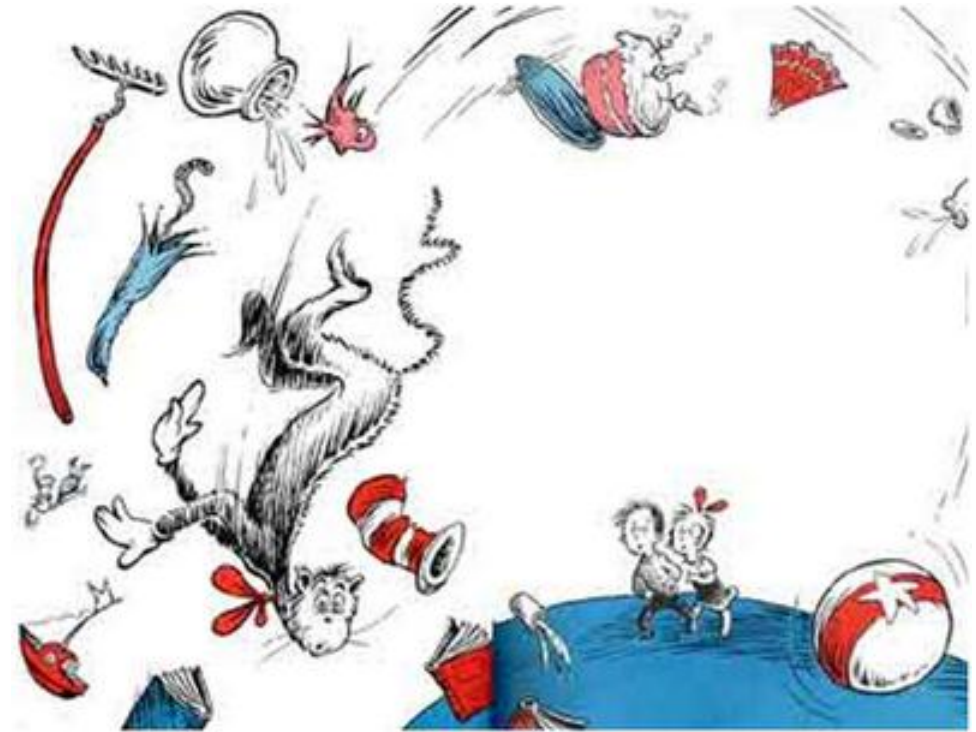




GWADW

Summary of Controls Sessions

we make it look easy, but ...



General Comments

- Controls is a new topical session for the GWADW this year
- Based on the number of participants in the controls session, we should include controls as a topical session in future GWADWs

Controls Sessions

A	System identification and modern Control
B	Optimal feedback
C	Noise feedforward and subtraction
D	Lock acquisition problems and improvements
E	Interferometer stability (Keeping the IFO Lock)
F	Next Generation Control System Architecture

- *Good sessions, good discussions, useful, but ...*
- *Perhaps too much scope – not focused enough, not enough “homework” in preparation in advance?*
- *A few too many prepared talks*
- *In future GWADWs, address next generation systems (V+, A+, Voyager, Longo,...)*

Summary D: Lock Acquisition Problems & Improvements

Acquisition overview/problems & discussions

- aLIGO (Shiela Dwyer, Keita Kawabe (LHO); Anamaria Effler (LLO))
 - Lock acquisition is deterministic for the arms (ALS) but stochastic for DRMI
 - Works most of the time; acquires in 15-30 min
 - Reasons for lock difficulty:
 - Poor initial alignment (drift)
 - Mode hopping
 - Bounce & roll modes not adequately damped sometimes (plant drifts)
 - CARM depends on RC gain
 - High winds (LHO)
 - causes DRMI lock to take more time
 - with high ETM green transmission, ALS arm lock problems
 - Simulation:
 - Optickle (and derivatives), E2E

Summary D: Lock Acquisition Problems & Improvements

Acquisition overview/problems & discussions (continued)

- Virgo (Gabriele Vajente, Bas Swinkels)
 - iVirgo Thermal transients in ITM at full lock caused drift
 - solution was to wait ~10 minutes
 - Lower absorption for aVirgo
 - Planning to use same variable finesse locking technique, but don't yet have a solution for SRC
 - Added high frequency RF sidebands (132 MHz) for use in lock acquisition – less sensitive to aberrations (thermal et. al.)
 - Plan to use guided lock acquisition (ala TAMA) if needed
 - aVirgo arm cavity finesse is 450 (was 150 for iVirgo)
 - Simulation indicates that guided lock acquisition may not always be needed
 - Digital demodulation
 - Simulation:
 - E2E, Optickle, Finesse (for thermal)

Summary D: Lock Acquisition Problems & Improvements

Acquisition overview/problems & discussions (continued)

- Geo (Emil Schreiber, [Hartmut Grote])
 - “only” 3 DOFs, but PR factor is higher
 - Automated, guided, stochastic locking for each DOF in sequence
 - PRC → Michelson (with 2kHz detuned SRC via heterodyne) → SRC
 - Then OMC beacon lock acq.; switch on slow controls (alignment, power ramp up, ...)
 - ~3 – 5 min total for lock sequence
 - No (significant) lock acquisition problems
- KAGRA (Yuta Michimura)
 - Same basic approach as for aLIGO, with 2 differences:
 - Inject green ALS beam through PR2/SR2 (rather than fiber to ETMs)
 - Non-resonant SB for DRMI instead of 3f (AM, not PM)

Summary D: Lock Acquisition Problems & Improvements

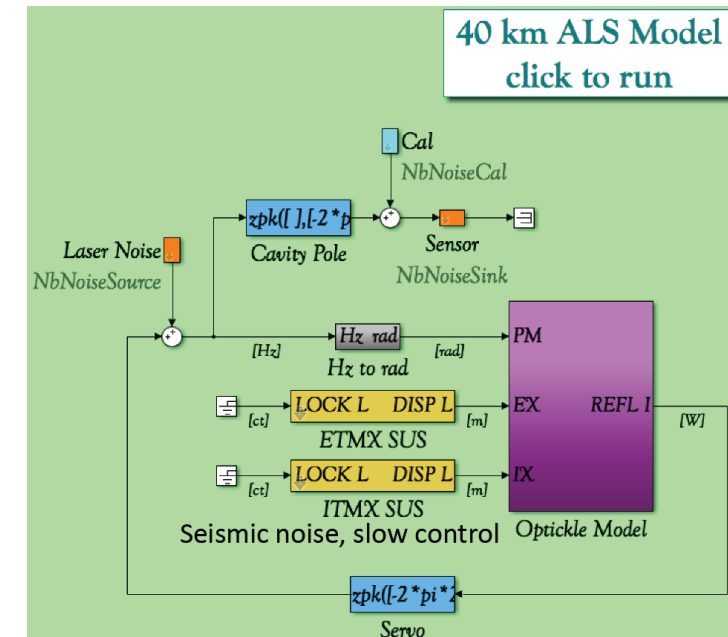
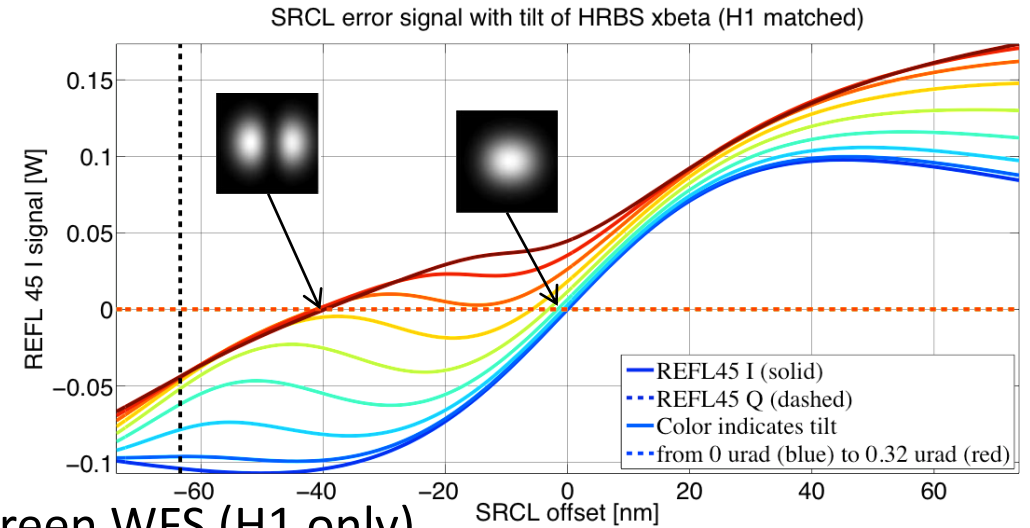
- Mode Hopping Problems (Paul Fulda)

- ALS

- Causes: Low finesse; alignment
 - Solutions: improved SEI performance (tuning); ALS green WFS (H1 only)
 - Will replace ETMs with proper green transmission
 - H1 DRMI (but not L1)
 - Causes: Alignment; mode matching; Gouy phase; Cross-coupling
 - Solutions: improve alignment, increase HOM separation
 - Would be good to measure RC optics with better ROC accuracy

- Simulation for Commissioning/Controls (Chris Wipf)

- SimulinkNB – noise budget tool
 - Example application: ALS for 40 km long arm (FSR @3kHz)
 - Tool for designers and commissioners



Summary D: Lock Acquisition Problems & Improvements

- Bayesian approach for locking high finesse cavities (Manuel Marchiò, Giancarlo Cella)
 - Apply force outside of resonance (to slow down the optic) based on a dynamics model
 - State space (position, velocity) Gaussian distribution + linear dynamics → Kalman filter estimator
 - Use transmitted power to know that you are passing a TEM00 resonance
 - Plan to include PDH off-resonance signal with particle filter (sum of Gaussians) method
 - Could this approach work for DRMI, for a more deterministic lock?

D: Lock Acquisition Problems & Improvements: *goals/questions*

- From each interferometer, gather a compendium of
 - lock acquisition problems
 - Locking metrics & statistics
- Discuss approaches to address each lock acquisition problem and their relative merits and success to date
- Are the pre-lock auxiliary systems (OptLev, ALS, TCS, ...) adequate, or are improvements necessary?
- Shouldn't future interferometers be designed to observe and control all DOF (bounce, roll, violin modes)?
- Would we benefit from embedding independent sensing/actuation into the interferometers for routine diagnostic measurements & characterization?
- Which lock acquisition problems can & should be pursued on small scale research interferometers?
- SIMULATION TOOLS
- STOCHASTIC LOCKING NOT GOOD

Summary E: Interferometer Stability

Lock Loss Causes

- aLIGO (Keita Kawabe, Shiela Dwyer (LHO); Anamaria Effler (LLO))
 - We can automatically detect lock loss, but not the cause (yet)
 - Detchar has a tool for auto plotting signals of typical interest just preceding a lock loss event
 - Will make auto & publish
 - Clear need for better local sensing of poorly controlled (damped) modes: roll, bounce, violin modes
 - Not yet clear if the limited actuation authority for these modes is stable/reliable or not (i.e. do we also need better actuation?)
 - Depends on alignment & beam position
 - Bounce & roll don't change at LLO; stable with good alignment
 - Find good alignment based on RC gains (automated procedure)
 - Inspiral range creeps down (sometimes) – thermal?
 - Leading lock loss causes at LLO:
 - Parametric Instability
 - EQs every 2-3 days
 - Saturation
 - Detchar is working on ODC checks on saturation
 - DAQ system has automatic saturation detection code

Summary E: Interferometer Stability Lock Loss Causes (continued)

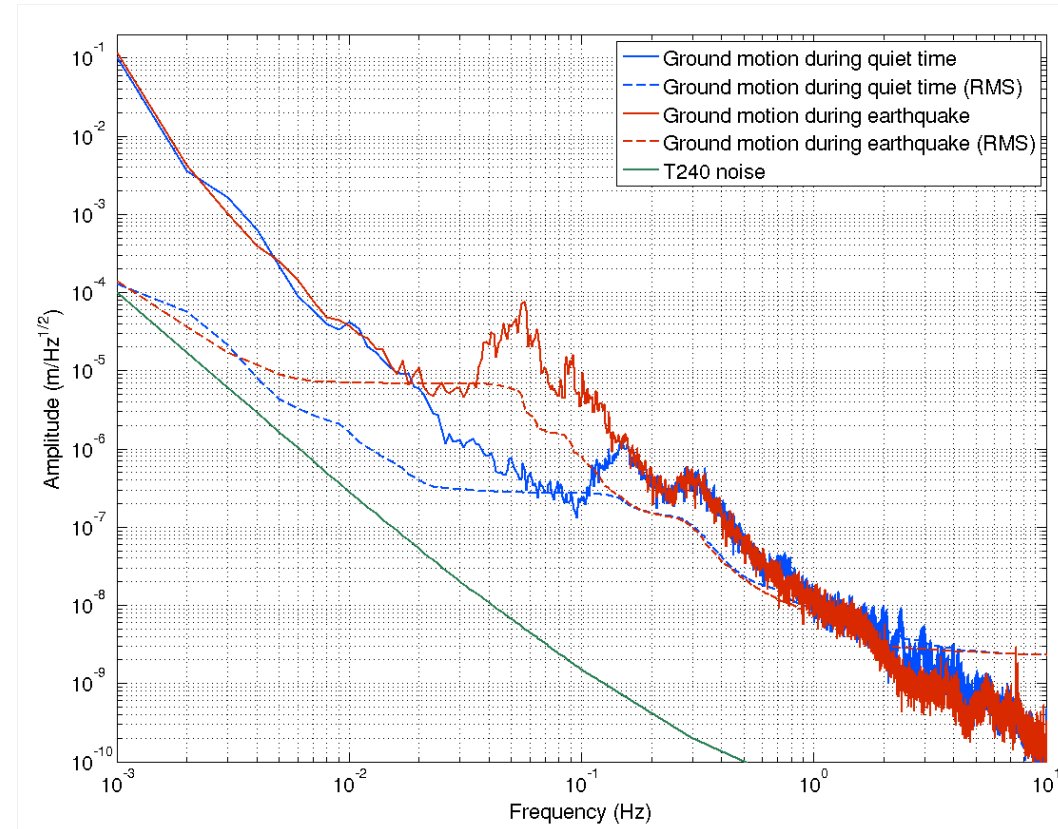
- Virgo (Bas Swinkels)
 - Reviewed VSR1 – 3 lock/unlock statistics
 - Improved caused lock loss events from 1-2/week (mag 6, Indonesia) to 1 per 2-3 weeks (mag 7, Indonesia)
 - VSR3: 90% locked, 80% science, 143 hr longest lock
 - Maintenance period (typ 4 hr/week) was largest contributor
 - Instrument/hardware reliability comparable to maintenance
- Geo (Emil Schreiber)
 - Auto lock loss detection (& plotting) and re-acquire

Summary E: Interferometer Stability

- Ground Tilt (Krishna Venkateswara)
 - Wind induced gnd tilt is significant for LHO (> 20 mph, 15% time)
 - Optic yaw response in high wind
 - Might be bilinear coupling with mis-centered beam?
 - Could be mis-balanced coil drivers
 - Investigate ISI to SUS cross-coupling, esp. yaw
 - Suggestion to simulate wind (high amplitude motion in ISI) to discover coupling path
 - LIGO pursuing a 2nd Beam Rotation Sensor (BRS) for LHO testing
- TCS (Alastair Heptonstall)
 - Plan to maintain ITM/CP thermal lens with CO2 laser when main beam losses lock – Operator sets thermal lens power with Guardian
 - Need to separately tune SRC & PRC (Virgo does this already for PRC) – LIGO considering the addition of a “ring heater” (or similar) to SR3 & PR3

Summary E: Interferometer Stability

- Earthquakes (Sebastien Biscans, et. al.)
 - LIGO has a few unlocks per week
 - (maybe same order of magnitude for Virgo)
 - Tilt is not well known for EQs
 - Displacement ASD high < 0.1 Hz
 - LIGO ISI platform (apparently) tilts a lot in response
 - ISI platform trips on S-wave or surface waves , approx. 30 min typ from EQ p-wave \rightarrow time to prepare
 - With USGS network at least a few min warning for “all” EQs
 - EQ Monitor (Jan Harms, Michael Coughlin) can predict to amplitude factor of 4 & 1-2 sec arrival
 - Reduce digital gain & increase analog gain
 - What to do to prepare & ride-out a mild EQ, e.g. higher blend freq.
 - Each EQ is different; requires testing. LASTI?



E: Interferometer stability (maintaining lock)

goals/questions

- Gather a compendium of lock loss event causes
 - Do we have detector infrastructure for automatic detection, logging, and calculation of statistical metrics?
- Discuss approaches to address each lock loss cause and their relative merits and success to date
- Are the lock loss prevention auxiliary systems adequate, or are improvements necessary?
- Would we benefit from embedding independent sensing/actuation into the interferometers for environment sensing?
- Which lock loss problems can & should be pursued on small scale research interferometers?

Summary F: Next Generation Control System Architecture

- How can we improve controls infrastructures to facilitate development of "modern" or "advanced" controls?
- Rolf Bork and Bas Swinkels gave brief overviews of the current aLIGO and aVirgo controls infrastructure
- Leverage commercial off-the shelf (LIGO: server-class computers, PCIe I/O, Beckhoff, RT linux patch, EPICs, etc.)

vs

- flexibility and control of custom (Virgo: DSPs, TOLM, ALP, etc.)
 - Both can work.
 - Virgo hopes to move from ALP to Tango. Should LIGO move from EPICs to Tango to leverage the entire LVC community?
 - No group conclusions; not really all the right people/expertise



Summary F: Next Generation Control System Architecture

- We then heard from people working on controls technology about what they would like to see:
 - Increase availability to modern/advanced algorithms that are ready to move beyond R&D (e.g. drop in "RCG part"):
 - State Space/LQR
 - Kalman filters
 - LQG adaptive algorithms (prototyped at 40m Lab)
 - Adaptation rate $<$ loop rate
 - Leverage multiple processors to offload adaptive calculations from front-ends, or
 - multi-rate front-end calculations
 - Increase accessibility, e.g. reduce hardware requirements, to facilitate development of prototypes, simulations, etc.
 - Modularity and flexibility in the infrastructure is the key to facilitating future controls development
 - Build a set of canonical problems (e.g. SUS angular control) and matching algorithms (e.g. Kalman filter) to develop a super-set of control system requirements
 - Couple the control system design tool (RCG for LIGO) directly to a simulation (e.g. SimPlant prototype at 40m Lab)

Final Comments

- How can we as a community share and sustain a growing expertise in the application of advanced/modern controls to GW interferometer systems?
 - LVC wiki pages for Controls
 - Get a critical mass
 - Share examples, code, experiences
 - Regular LVC (virtual) meetings on Controls
 - Perhaps an “official” LVC working group on controls?
 - Detchar “gold star” vetted tools wiki page
- In future GWADWs:
 - Address next generation systems (V+, A+, Voyager, Longo, ...)
 - Consider truly advanced techniques, with adequate prep (NN, deep learning, ...)