



Status of LIGO

June 22, 2015

Daniel Sigg

LIGO Hanford Observatory

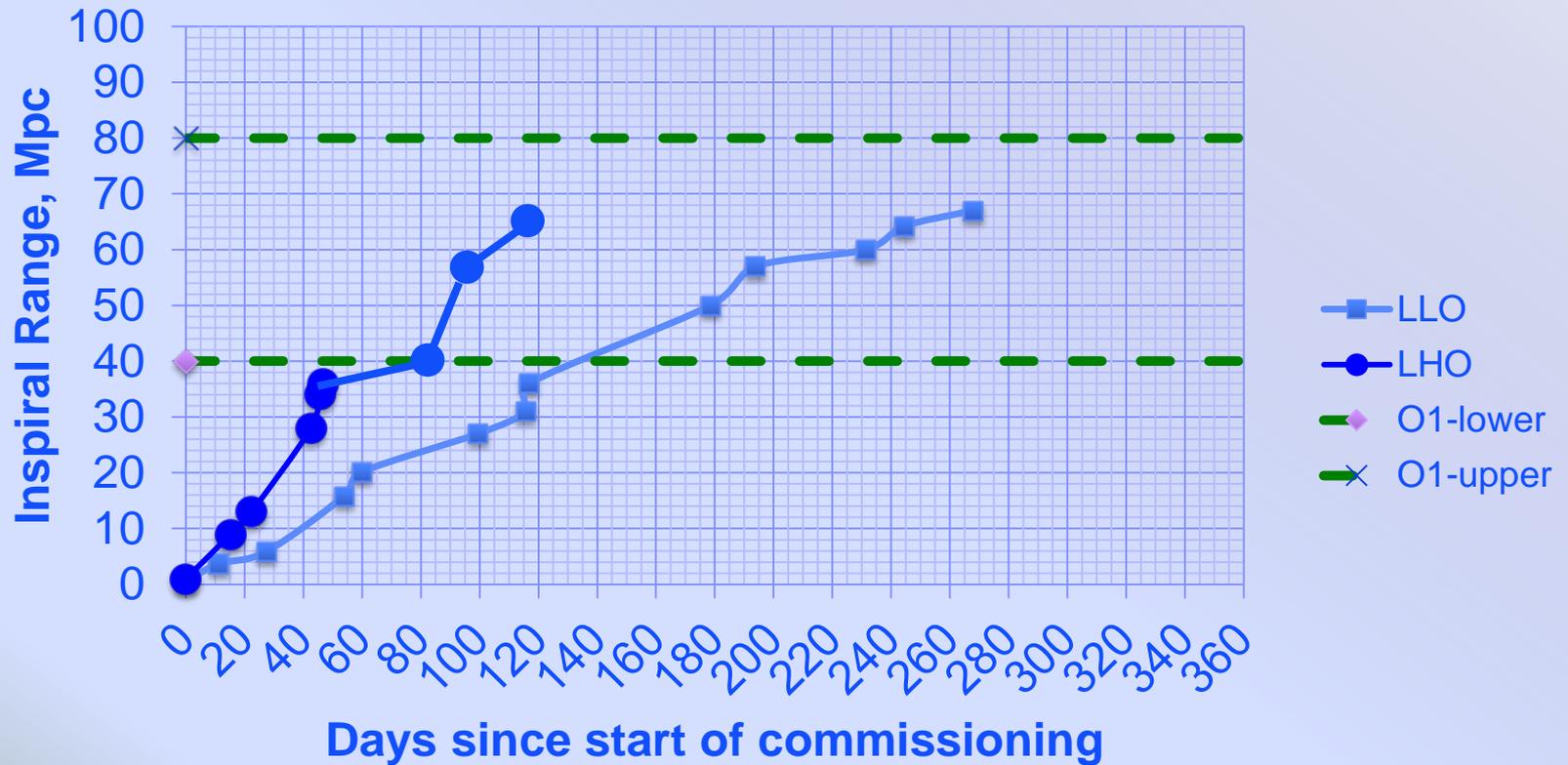
(on behalf of the LIGO Scientific Collaboration)

Highlights

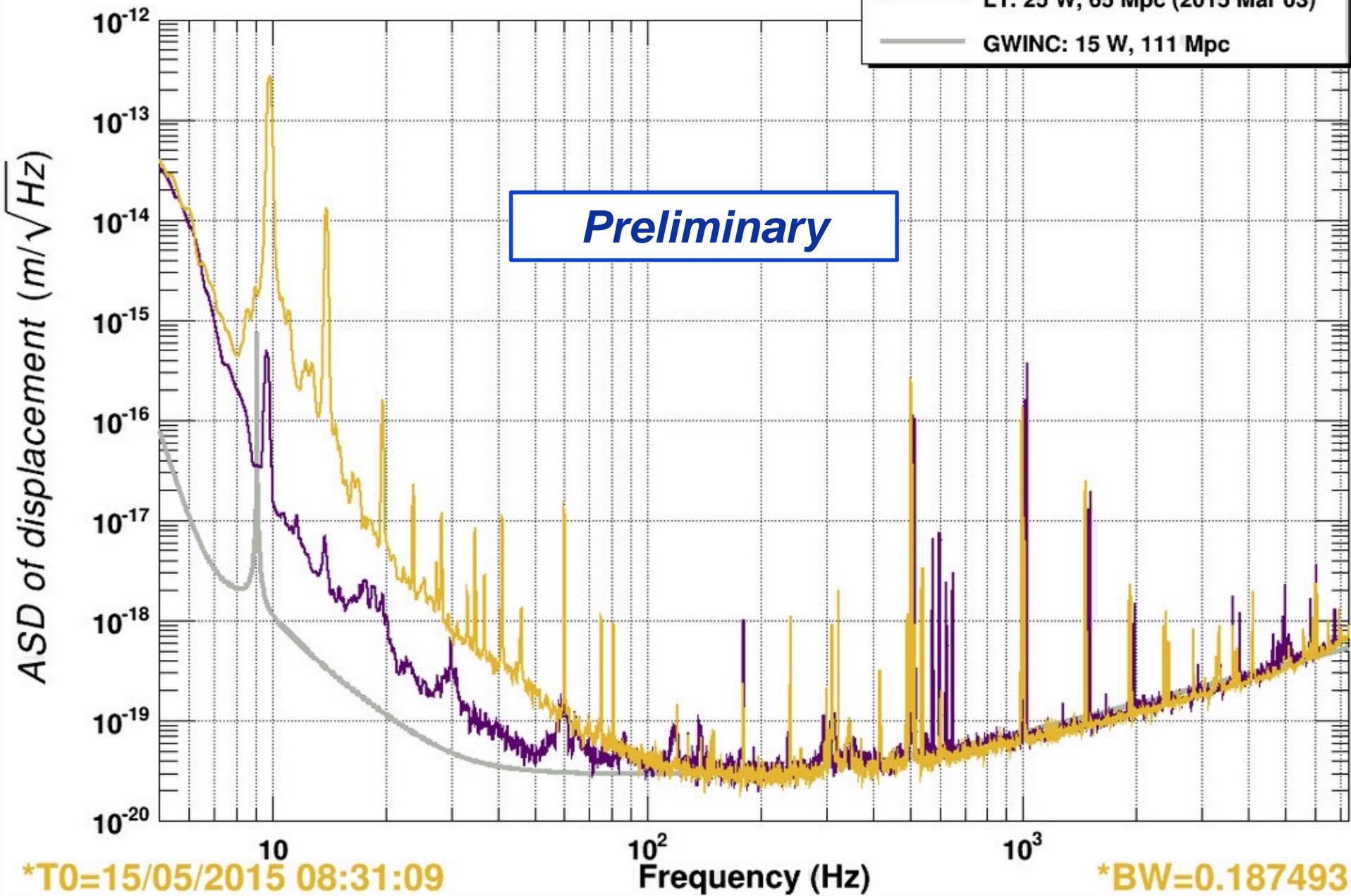
- ❑ Robust Locking Achieved
 - Livingston Observatory: First lock on May 27, 2014
 - Hanford Observatory: First lock on February 7, 2015
- ❑ Sensitivity of initial detectors surpassed quickly
 - Current inspiral range: 50 to 70 Mpc
- ❑ ER7 completed
 - 40 hours of coincident data
- ❑ Preparing for O1: September to December
- ❑ Issues:
 - High Q Resonances and Uncontrolled DOFs
 - ESD Charging Issues
 - Parametric Instabilities

Progress

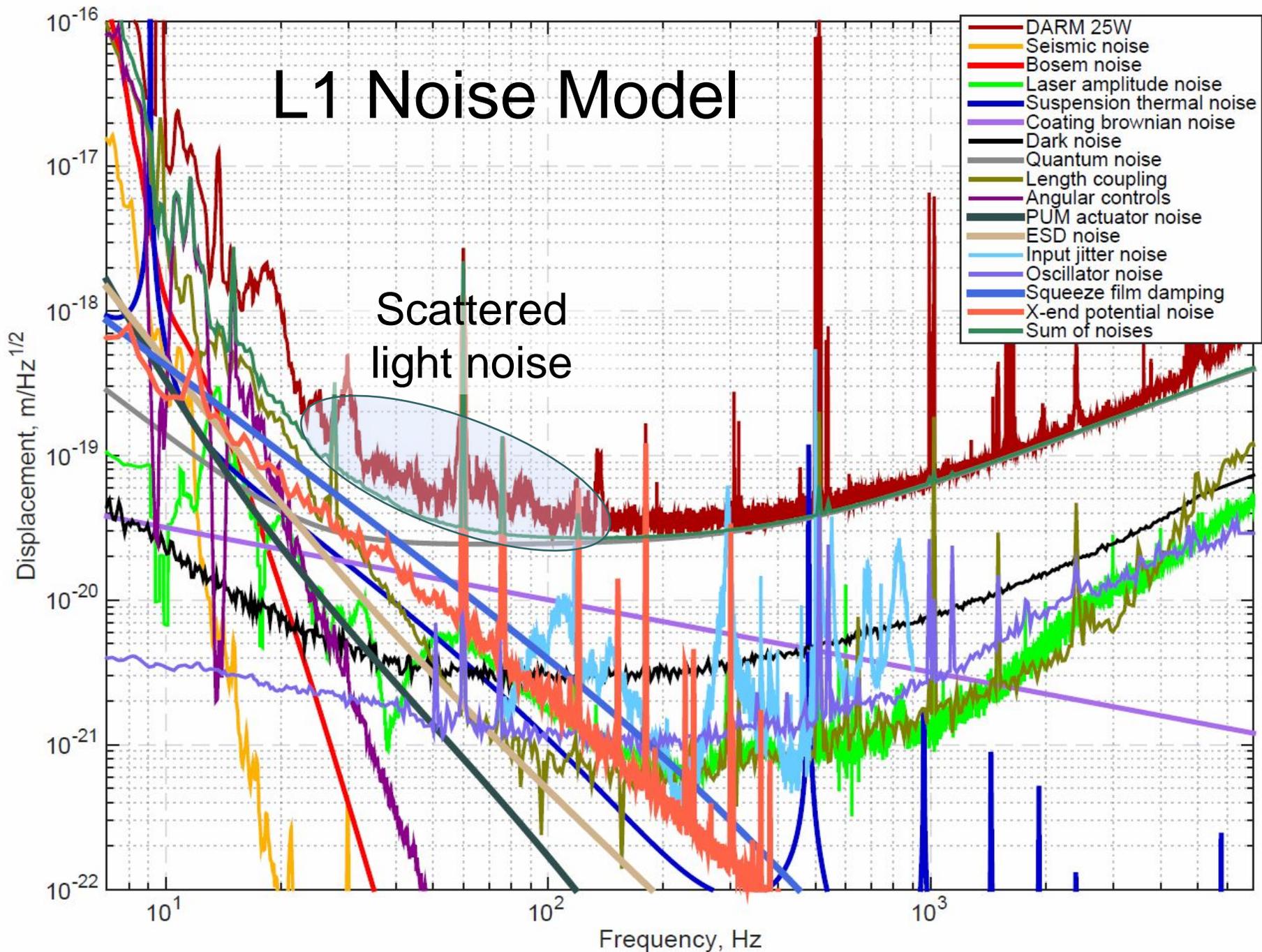
aLIGO Commissioning Progress



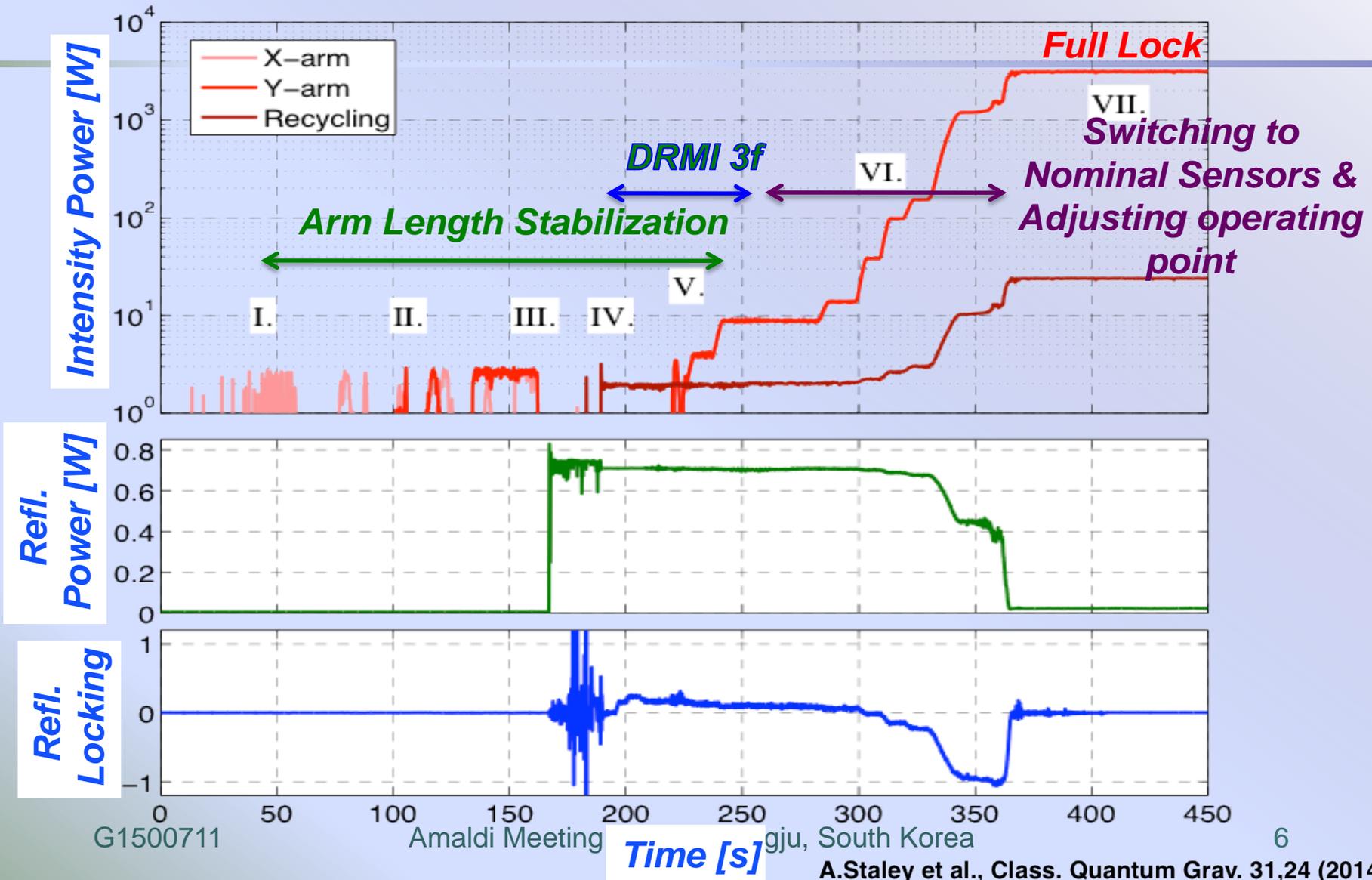
aLIGO DARM



L1 Noise Model



Locking



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Amaldi Meeting

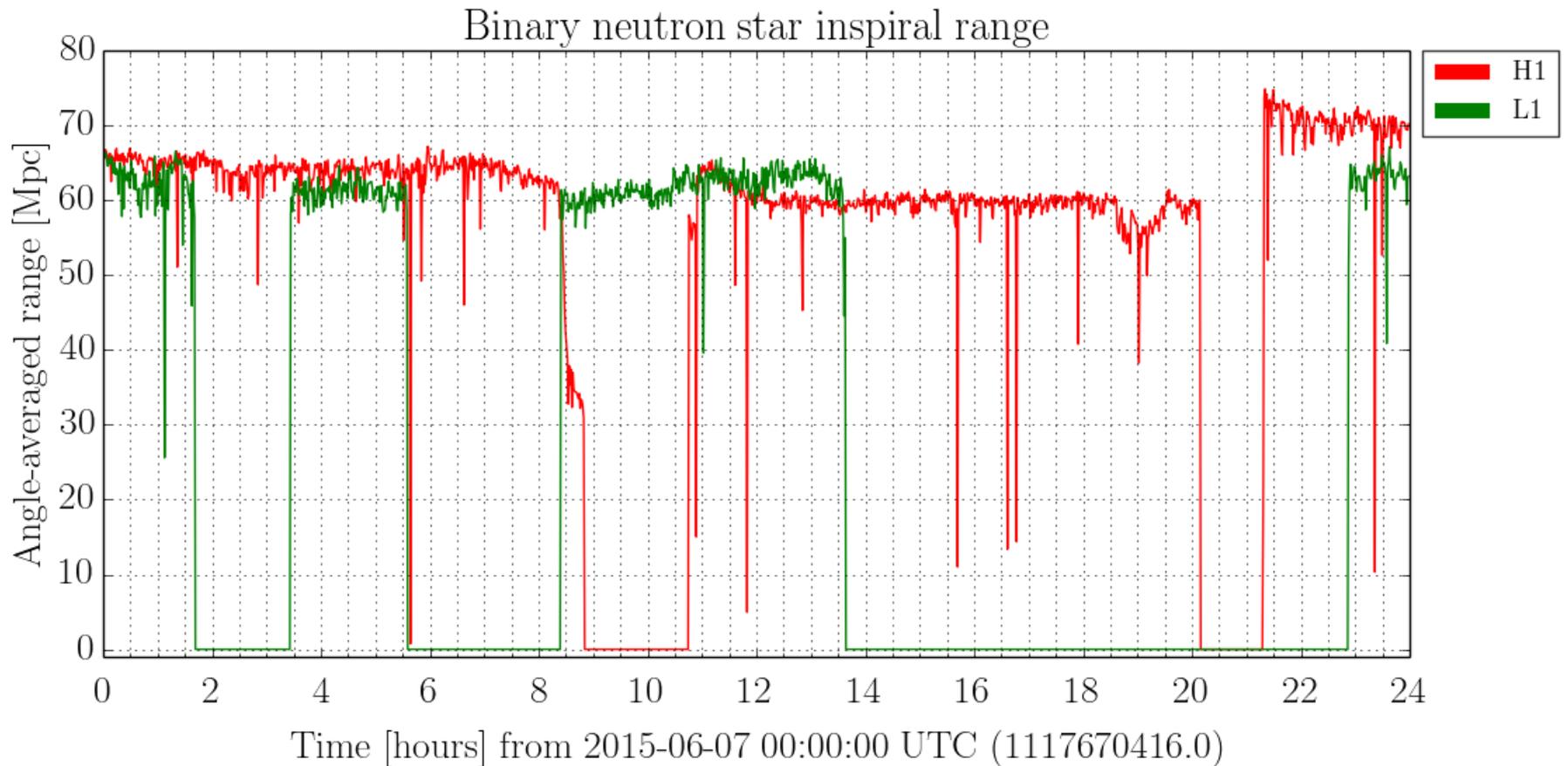
Time [s]

gju, South Korea

7th Engineering Run

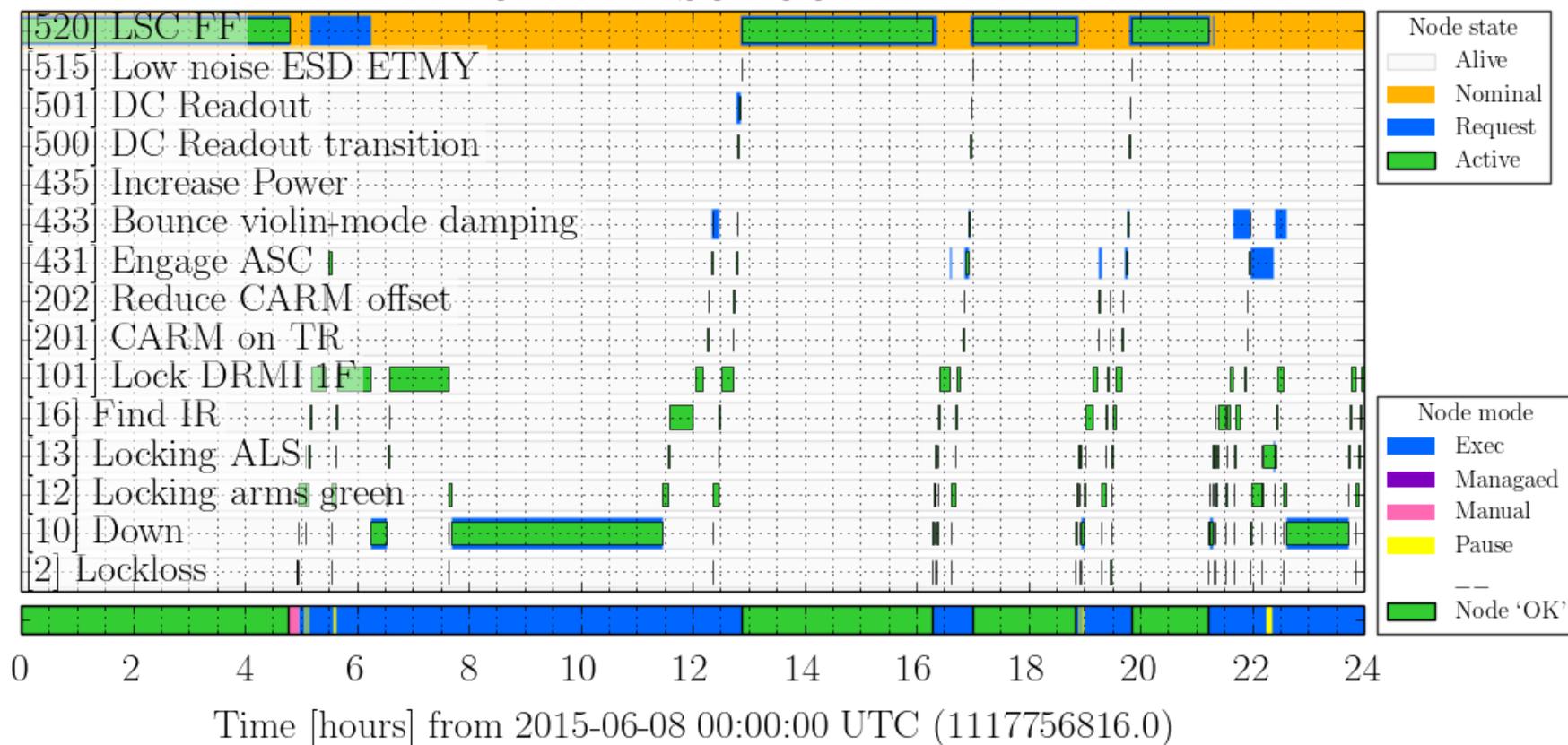
- ❑ May 16 to June 14, 2015
- ❑ First coincidence data between L1 and H1
- ❑ Test of the analysis pipeline
- ❑ H1 110 hours, L1 74 hours, 42 hours of coincidence (as of June 11)
- ❑ Successfully completed
 - Calibrated strain readout
 - Sufficient automation to be controllable by operators
 - Hardware injections
 - Blind Injections Implementation and Testing
 - Approximately 7 days of data taking

Binary Neutron Star Range (SNR 8, $1.4M_{\odot}$, sky average)

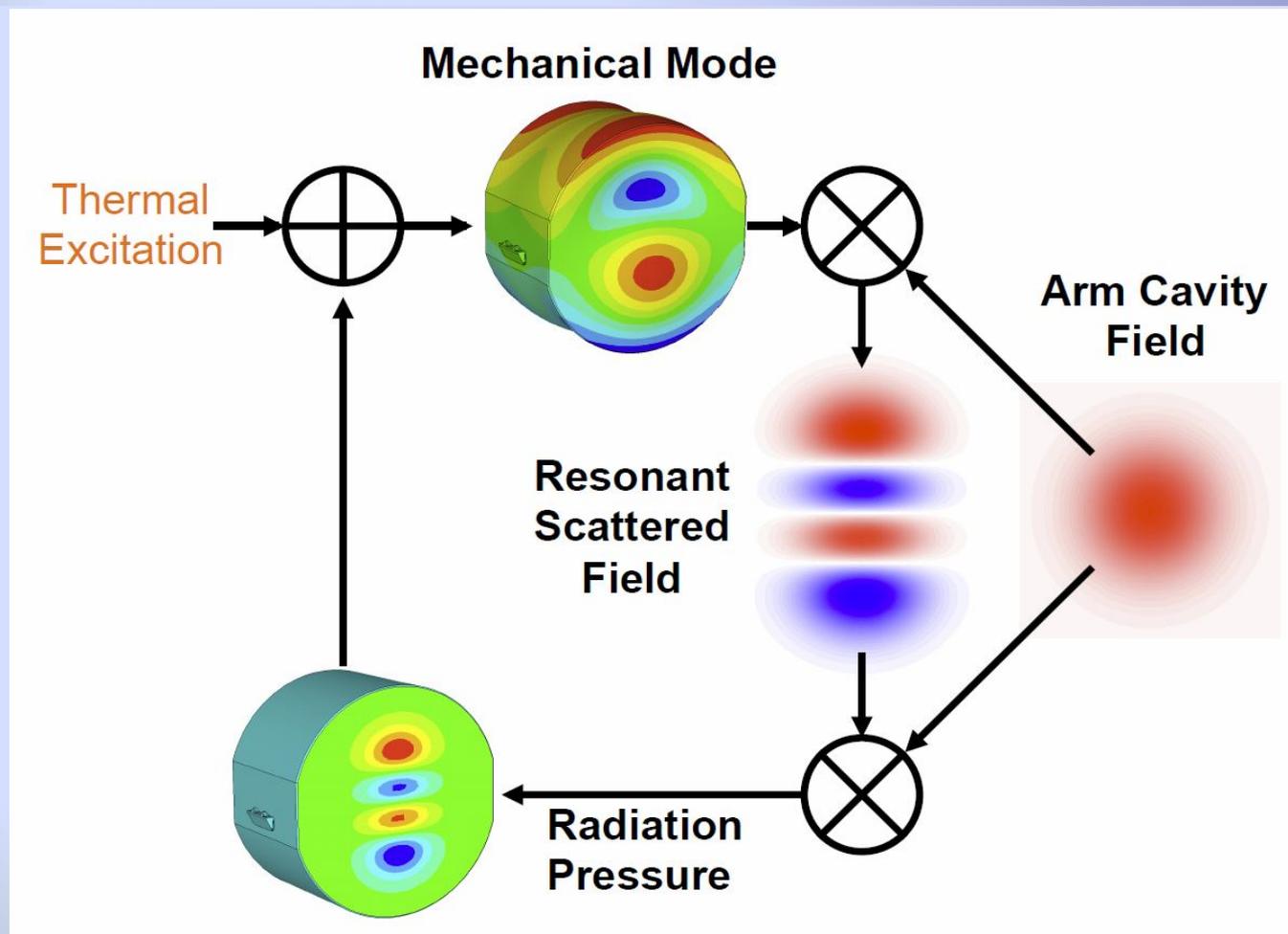


ER7 Segments (24 hours)

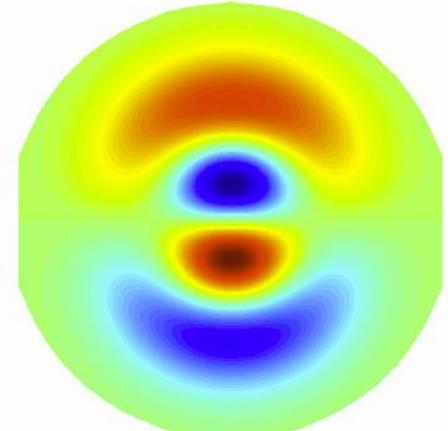
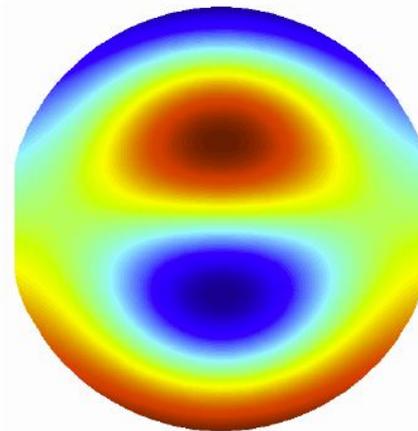
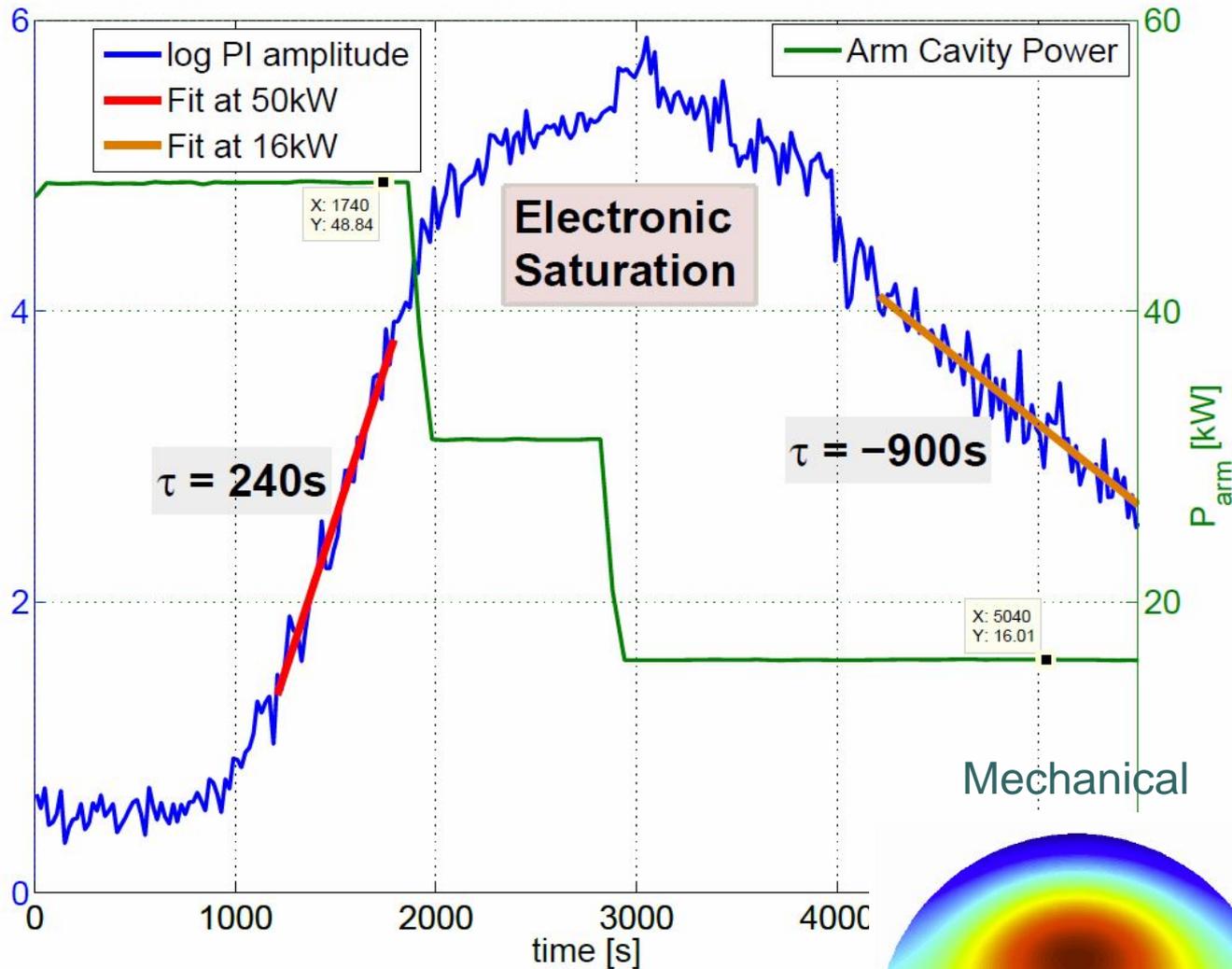
H1 Guardian ISC_LOCK state



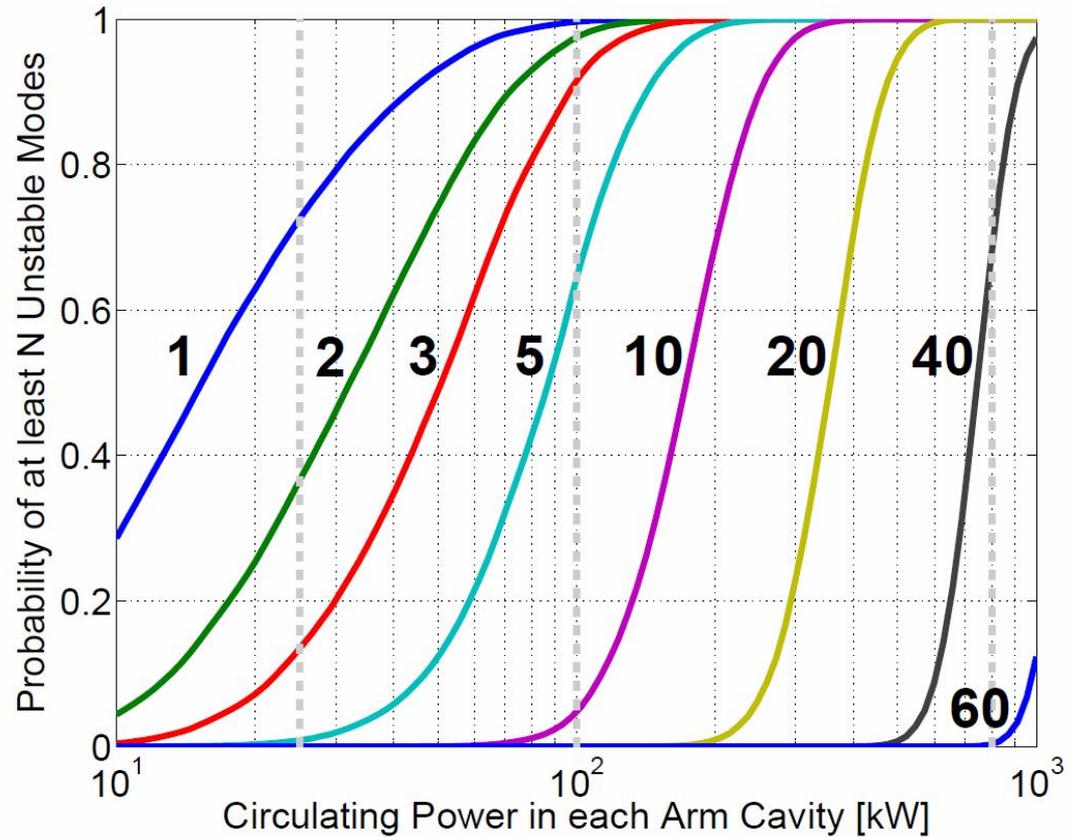
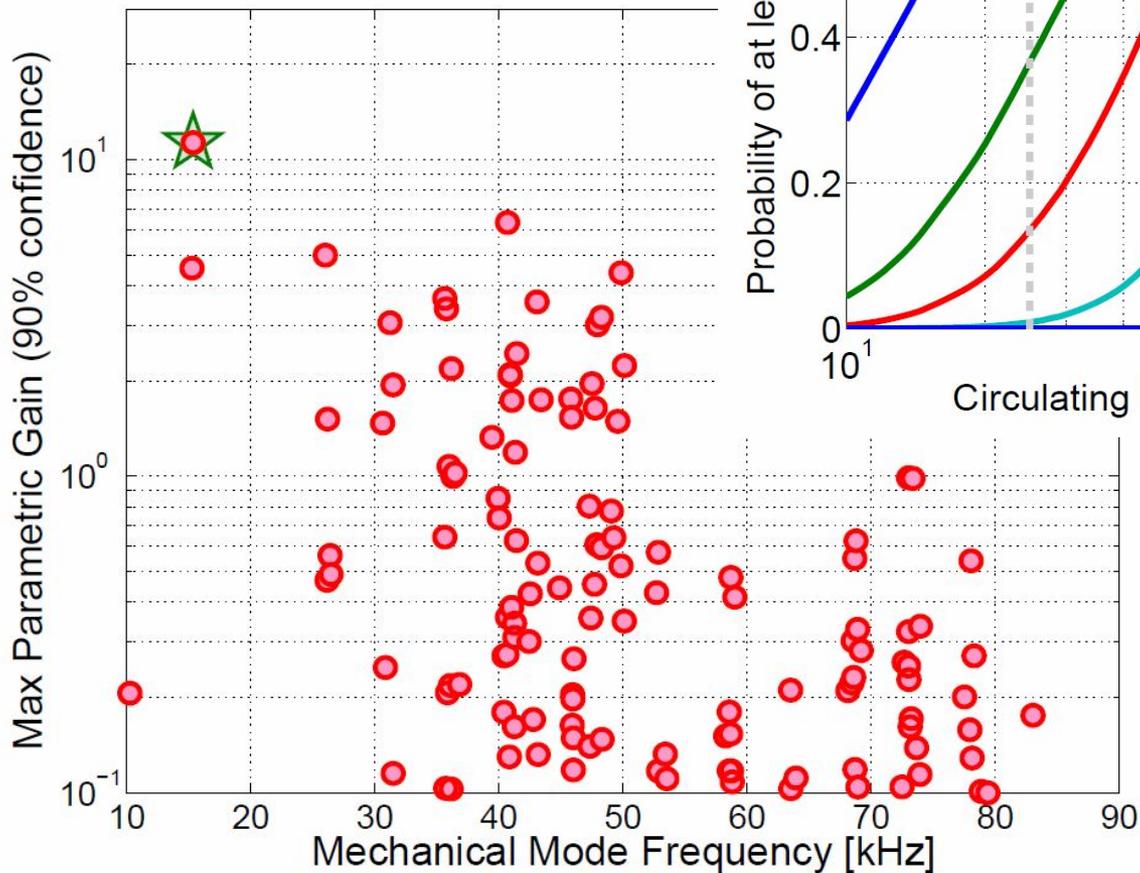
Parametric Instabilities



Parametric Instabilities (3)



Parametric Instabilities (2)



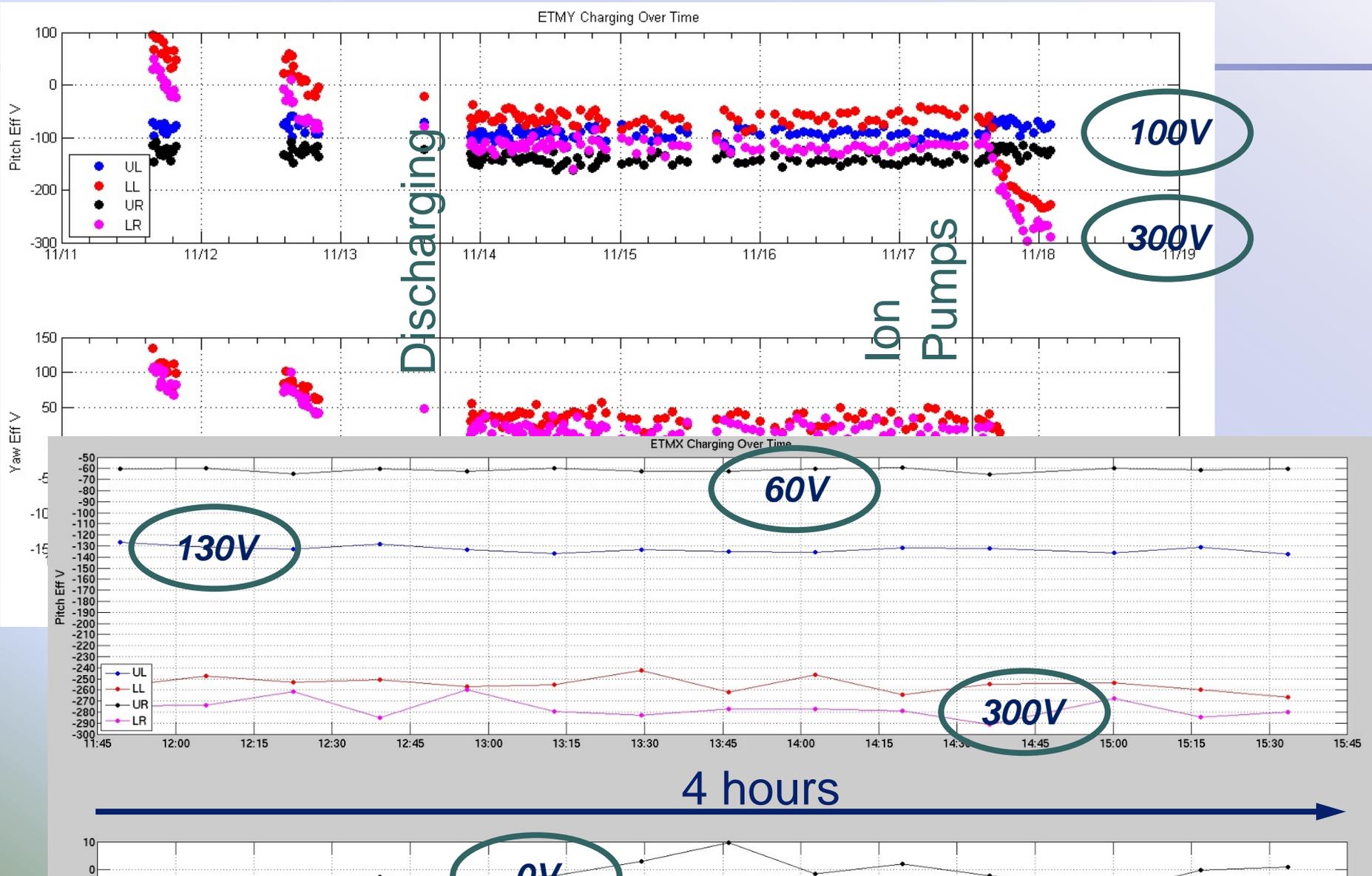
Number of Unstable Modes Depends on ROC

Thermal control of ROC can work for small N

High Power Operations

- With 25W input power
 - No significant thermal loading
 - Parametric instabilities mitigated with ETM ring heater
 - Angular instabilities under control
- At higher powers
 - Hartmann wavefront sensors to map ITMs
 - Ring heaters for ITMs and ETMs
 - CO₂ laser heating for ITM compensation plates
 - Active damping control for parametric instabilities
 - Testing a passive mode damper (shunted piezo-electric transducer)
 - Angular instabilities require more control bandwidth (induces noise)

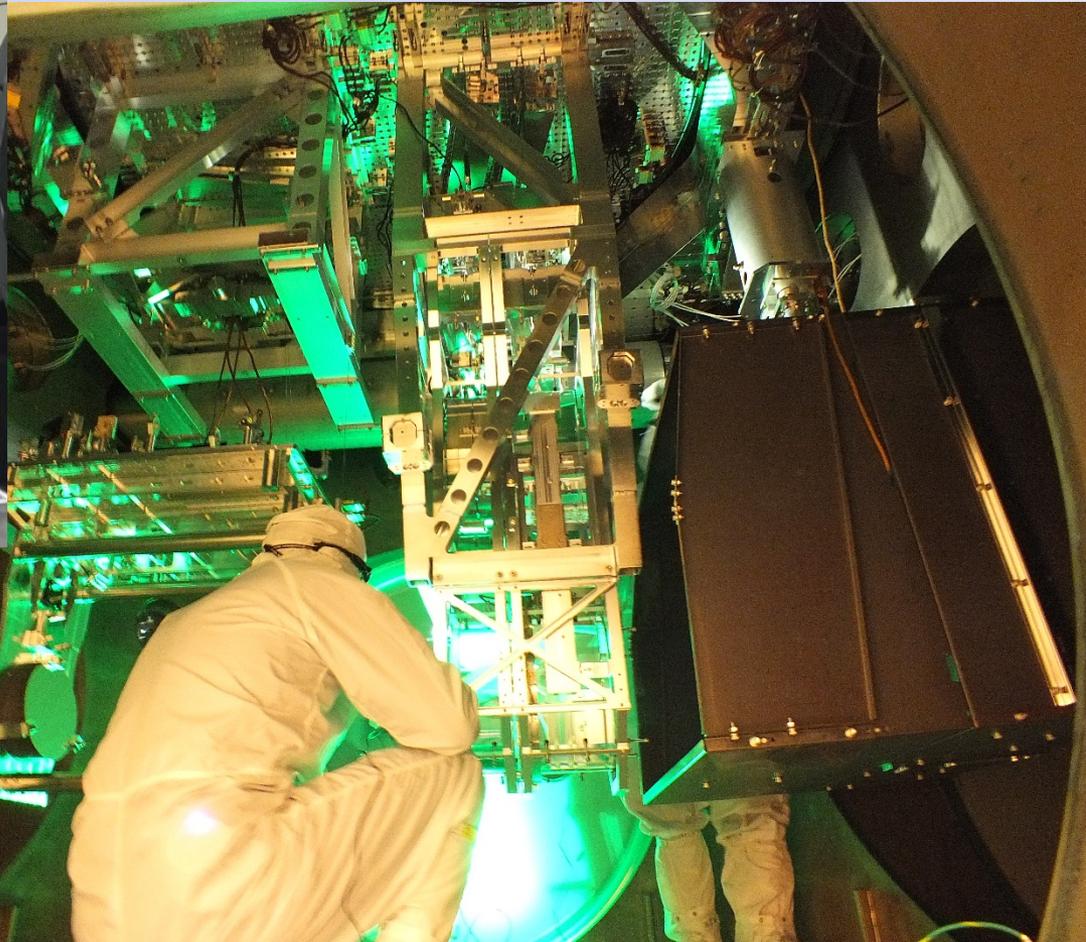
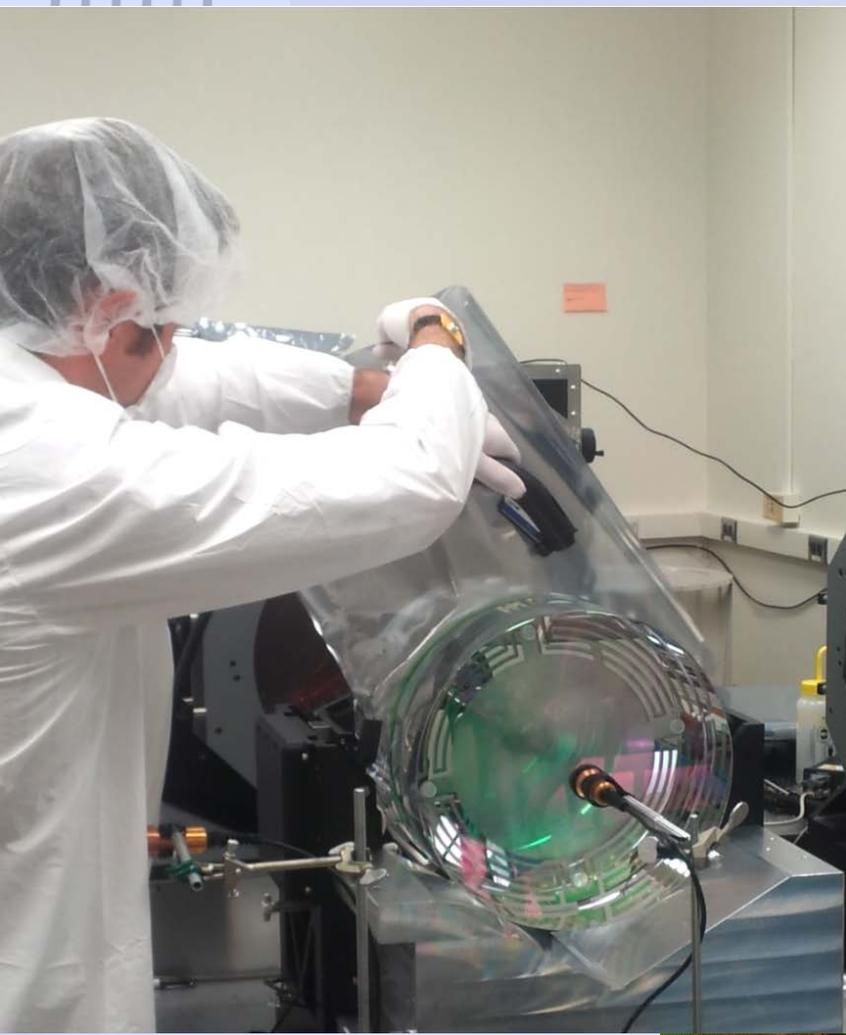
Electro Static Drive



Electro Static Drive (2)

- ❑ Why so much charge? Can be 100s of Volts
 - First contact leaves a charge behind
 - Charge variation appears to be caused by ultraviolet light emitted by ion pumps mounted at the test mass chambers (recharging)
- ❑ Additional Noise and Uncertainty
 - Field lines must terminate somewhere (e.g. ring heater)
 - Drive strength depends on charge, adds a linear term
 - Charge is not uniform and varying
- ❑ In-vacuum ion injecting did not fix the problem
 - Hard to reach the back surface
- ❑ In-air, in-situ neutralization procedure
 - Test mass stays discharged (without an ion pump)

Test Mass Discharging



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Summary

- ❑ Sensitivity of initial detectors surpassed quickly
- ❑ Robust locking achieved
- ❑ Parametric instabilities are real; need to be dealt with
- ❑ Test mass charge is a problem
 - Discharging effort underway
- ❑ Looking forward to the first observing run
- ❑ What's left?
 - High power operations: thermal control, PI, alignment instability
 - Low frequency noise hunting & controls optimization
 - Make the seismic system work during bad weather
 - Damping of high Q modes

