

GEO600 status











CARDIFF



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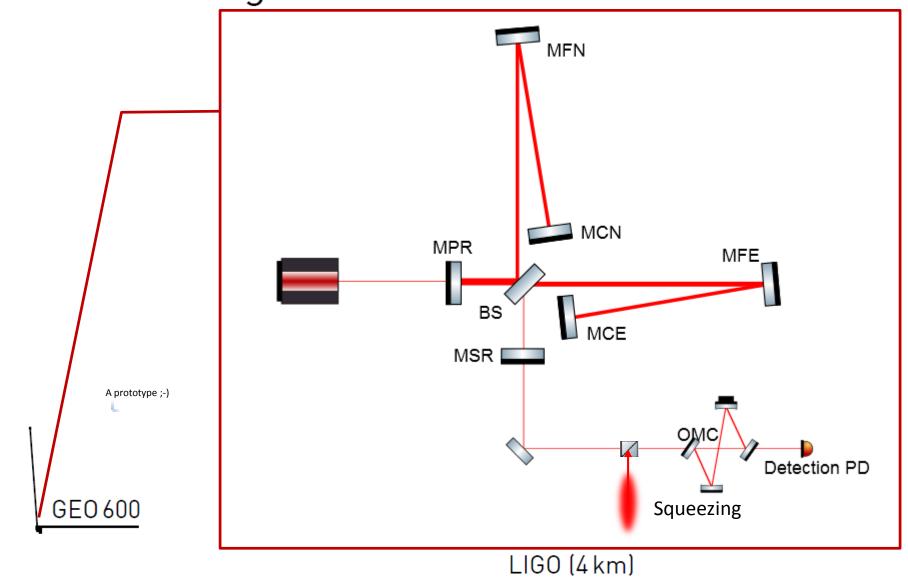
Outline

- GEO 600 & GEO-HF
- The actual status
 - Astrowatch
 - MDWS: a high bandwidth OMC AA schema
 - Squeezing: loss reduction
 - Laser power increase
- Going beyond GEO-HF?

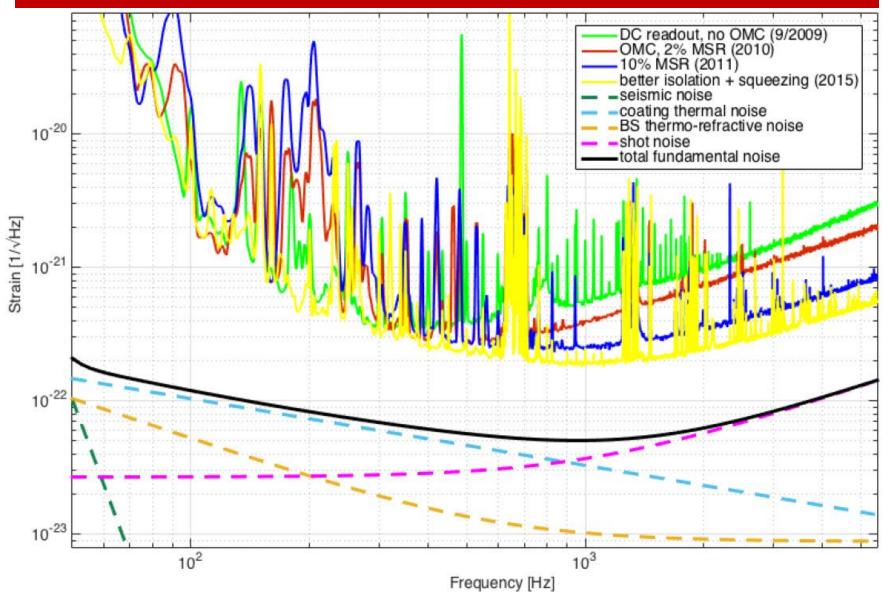


GEO 600 and GEO-HF

A *small* large-scale detector

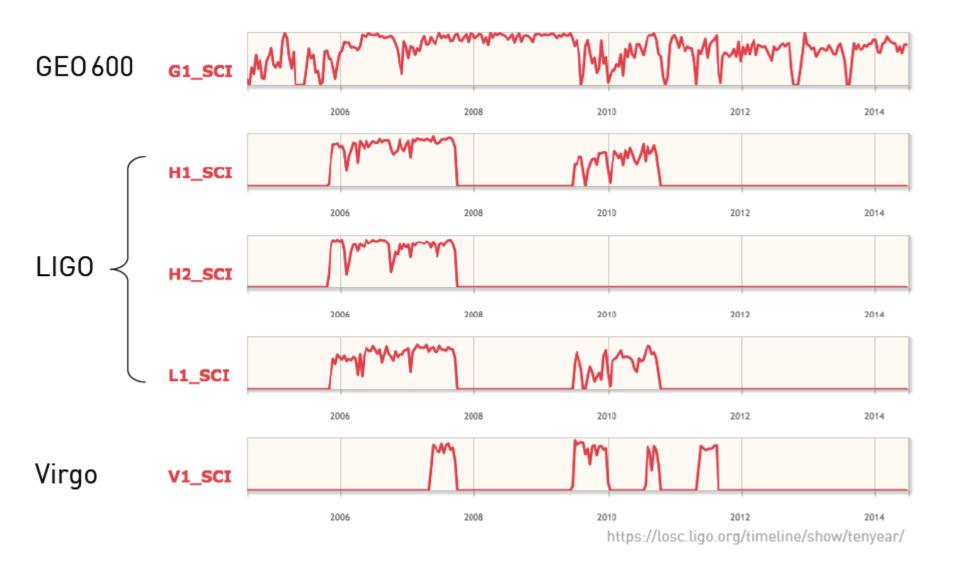


GEO 600 and GEO-HF

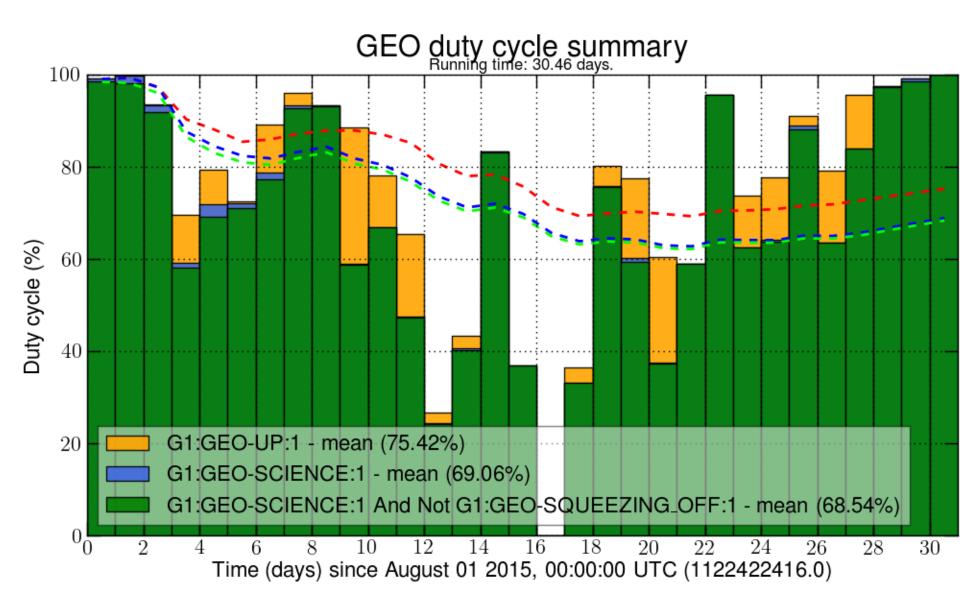


Learn more about GEO 600 and GEO-HF: **K.L. Doodley et al,** GEO 600 and the GEO-HF upgrade: successes and challenges. (P1500140-v1)

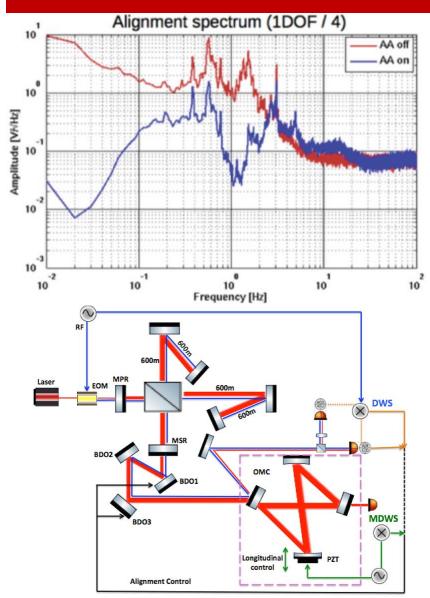
Astrowatch



Astrowatch



Modulated differential wave front sensing

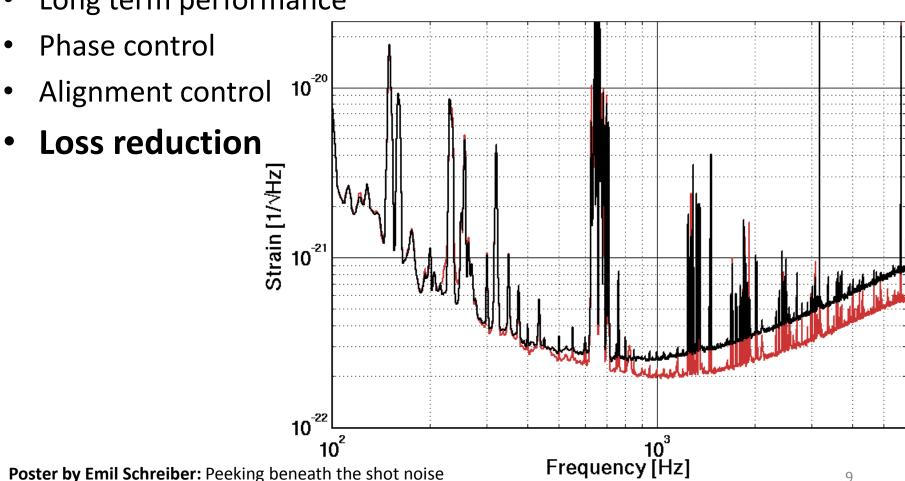


- Control bandwidth of a few Hz achieved. Goal: ~10 Hz
- Long term performance (~months) comparable to beacon dither
- uses 2f centering on sensors
- To do: decoupling of actuator DOFs, bandwidth increase and investigate benefit for operation at GEO

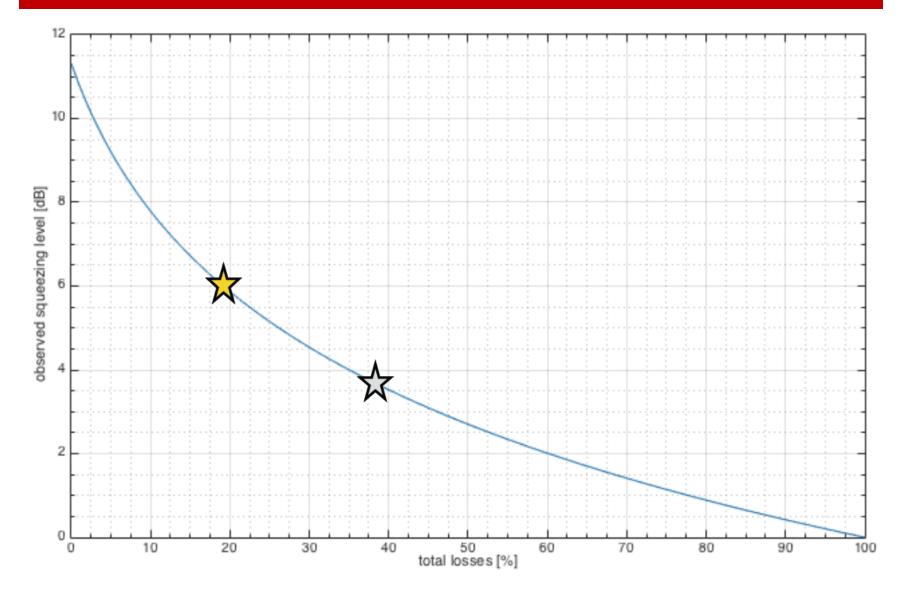
Squeezing at GEO 600

Focus in the (recent) past:

Long term performance



Squeezing at GEO 600



Squeezing: loss reduction Loss reduction has now priority: Optimizing squeezing injection path Better understand in vacuum losses Development of new low loss isolators ("Glasgow polarizer") OPO: 5% OMC and PD: TCOC **TCOb** 14% > 5%? in-air injection path: 8% > 2%?

Poster by Emil Schreiber: Peeking beneath the shot noise

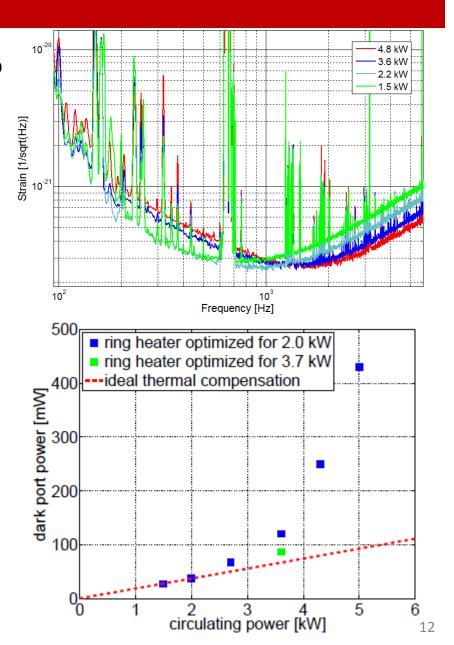
in-vacuum injection path: 13% > 4%?

BDO1

Squeezing Bench

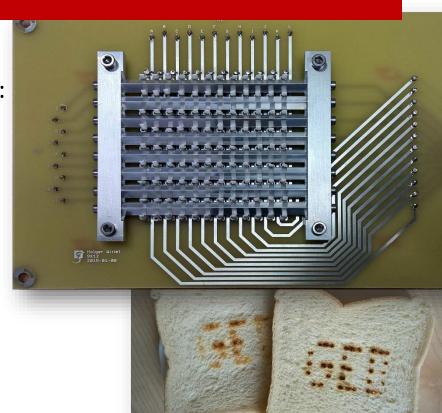
Laser power increase

- Operation with low power, i.e. 2 to 3kW impinging on the BS
- Main reason: power dependent noise at medium frequencies & thermal compensation
- Increase (again) effort to commission GEO for high power operation
- thermal compensation of the BS thermal lens:
 - At one folding mirror
 - At the beam splitter



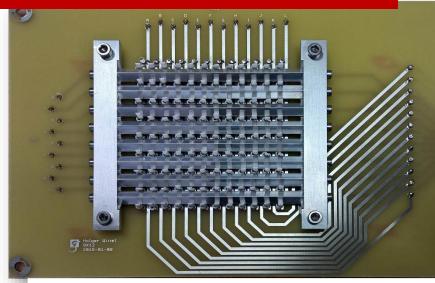
TCS at the BS

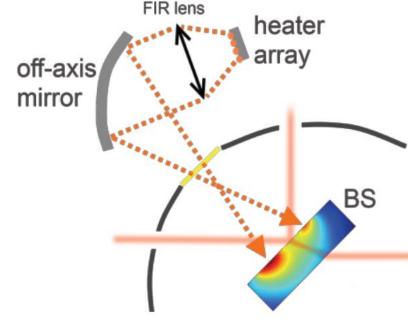
- Matrix heater: 9x12 small heaters, projected to BS surface and can act on:
 - thermally induced HOMs
 - "cold" HOMs, caused by mirror imperfections
 - Heat the BS mitigate cold locking



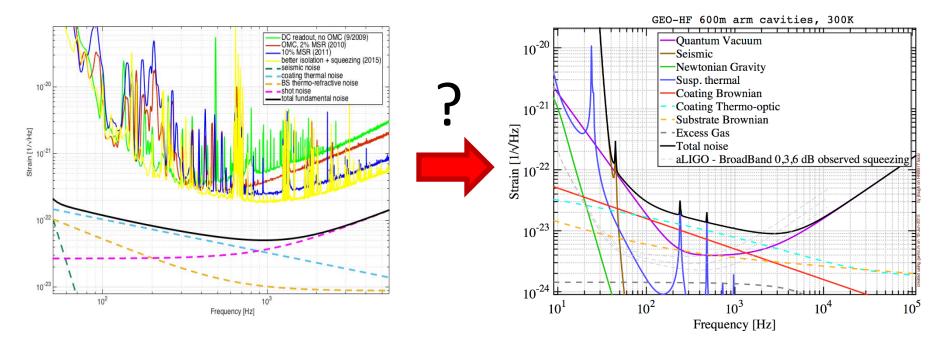
TCS at the BS

- Matrix heater: 9x12 small heaters, projected to BS surface and can act on:
 - thermally induced HOMs
 - "cold" HOMs, caused by mirror imperfections
 - Heat the BS mitigate cold locking
- Optimization of heating pattern ongoing
- Many strategies possible: (Simulationdriven guessing, using phase images, optimization algorithms)
- Preliminary results:
 - ~30% reduction of thermal HOMs
 - ~30% reduction of cold HOMs





Beyond GEO-HF?

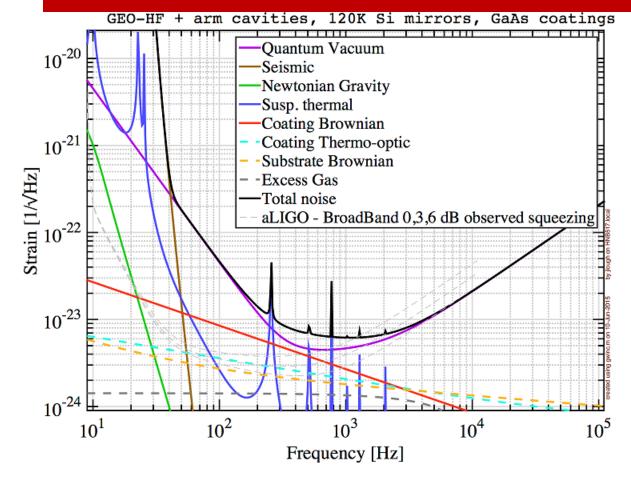


1. Options for GEO 600 as a data-taking instrument

- 2. Options for GEO 600 as a non-data-taking instrument
 - Large prototype
 - Fundamental Physics aspects like vacuum QED
 - ...

Option 2. needs to be discussed

Going beyond GEO-HF: matching AdvLIGO at HF?

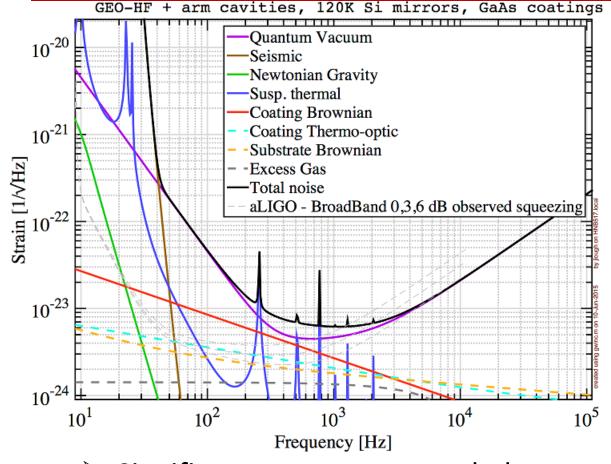


How to:

- Arm cavities
- Laser powers ~MW
 in the arms
- 6dB squeezing
- Cryogenic?
- Si mirrors? (1550nm)
- AlGaAs coatings

• ...

Going beyond GEO-HF: matching AdvLIGO at HF?



How to:

- Arm cavities
- Laser powers ~MW in the arms
- 6dB squeezing
- Cryogenic?
- Si mirrors? (1550nm)
- AlGaAs coatings
- ..
- → Challenging but possible!

- Significant man power needed
- Not compatible with our other commitments
- Continue with discussing non-data-taking options
- Until O2: Astrowatch + instrument science

Summary

- Options for a large scale upgrade of GEO 600 to match AdvLIGO design sensitivity above a few kHz was carefully analyzed. It seems possible but not compatible with our other commitments. (Astrowatch, contribute to 2nd and 3rd generation...)
- Continue Astrowatch with high duty cycle until O2
- Carry out a solid instrument science program (such as):
 - Squeezing, in particular loss reduction
 - Control like MDWS as new OMC AA system
 - Laser power increase/thermal compensation
 - Continue discussion of a program which makes full use of the existing infrastructure!

