

# Improving seismic isolation in Advanced LIGO using a ground rotation sensor

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for

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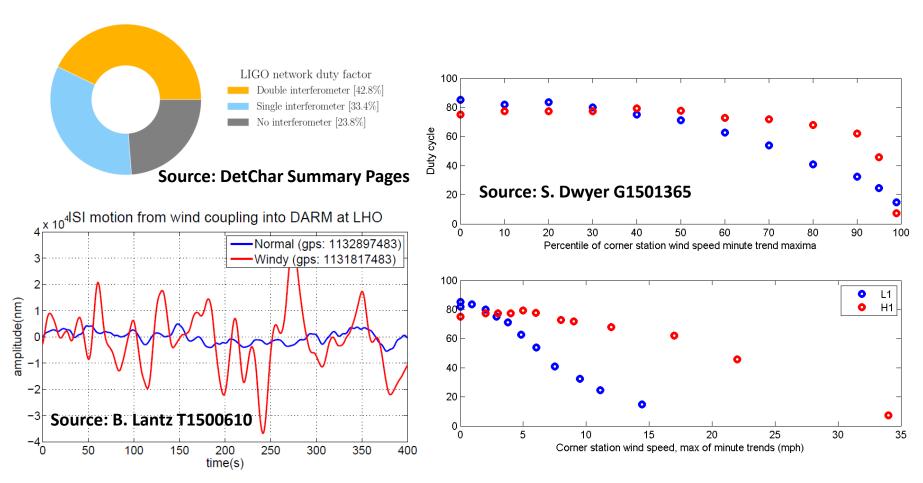
LIGO-G1600083

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- > BRS at LIGO Hanford and Tilt Data
- Next Gen/Compact BRS
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# Introduction

# Tilt-problem limits duty cycle of detectors



- Earthquakes, High Wind and High Microseism cause significant down-time of both detectors.
- All are primarily related to the problem of <u>tilt-horizontal coupling</u> producing too much ISI motion.

# Tilt versus Horizontal displacement

 Conventional seismometers and tiltmeters cannot differentiate between horizontal displacement and ground tilt.



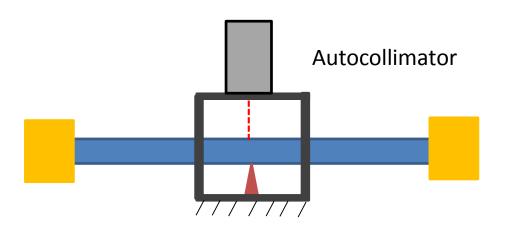
Tilt response to horizontal displacement response for <u>all</u> seismometers =  $-g/\omega^2$ 

⇒ Tilt is confused with horizontal motion at low frequencies (below ~ 0.1 Hz).

<u>Solution</u>: Inertial rotation sensors, Tilt-free seismometers or Ring-laser gyroscopes...

# Beam Rotation Sensors

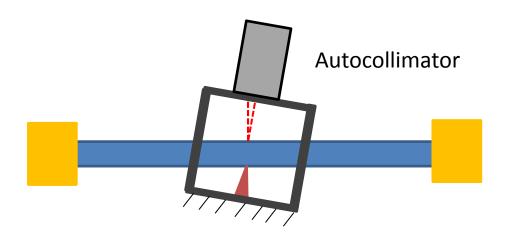
# **BRS** Concept



### Principle:

- Ground tilt is measured by measuring angle between ground and low frequency beam balance.
- Horizontal acceleration can be rejected by locating center of mass at the pivot.

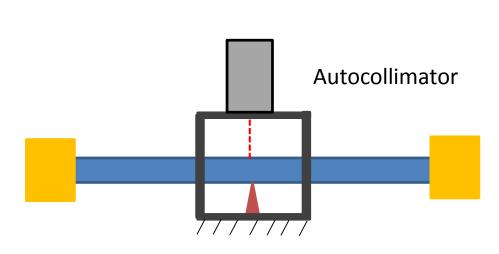
# **BRS** Concept



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#### **BRS Parameters**





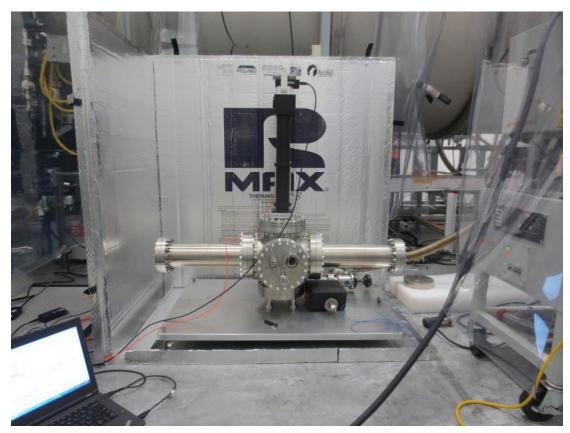
10-25 μm-thick Cu-Be Flexures

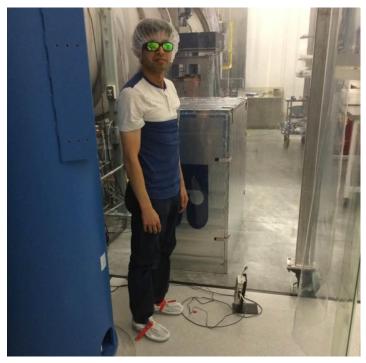
Mass	4-5 kg
Length	0.9 m
Moment of Inertia I	0.6 kg m <sup>2</sup>
Flexure stiffness	10 <sup>-3</sup> N m
Quality factor	2000-3000
Distance between pivot and center of mass	Can be tuned to less than 1 µm

Displacement rejection ratio

$$\frac{\hat{\theta}}{\hat{x}_g} \approx \frac{M\delta}{I} \approx 0.00001 [rad/m]$$

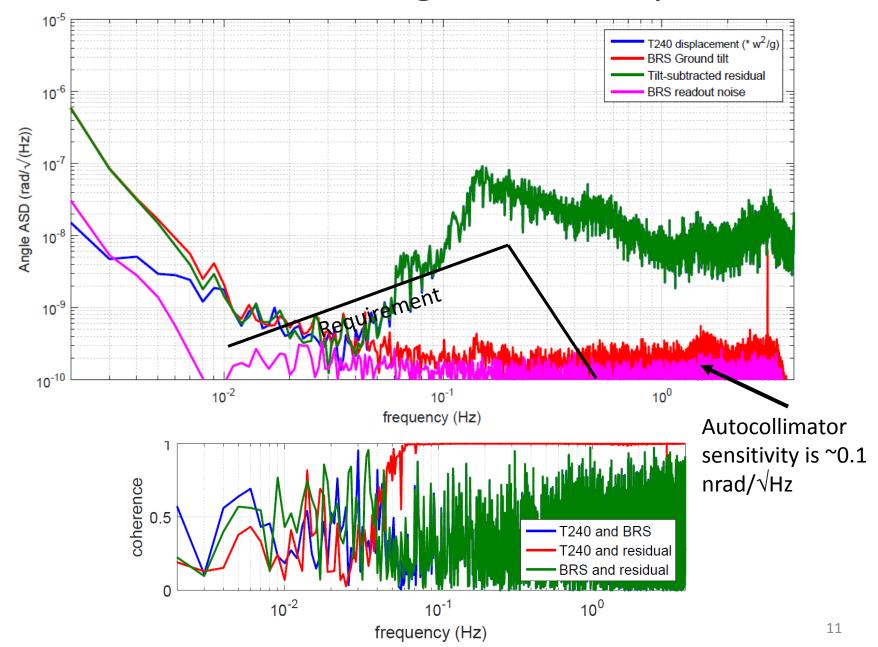
# Installation at EX VEA



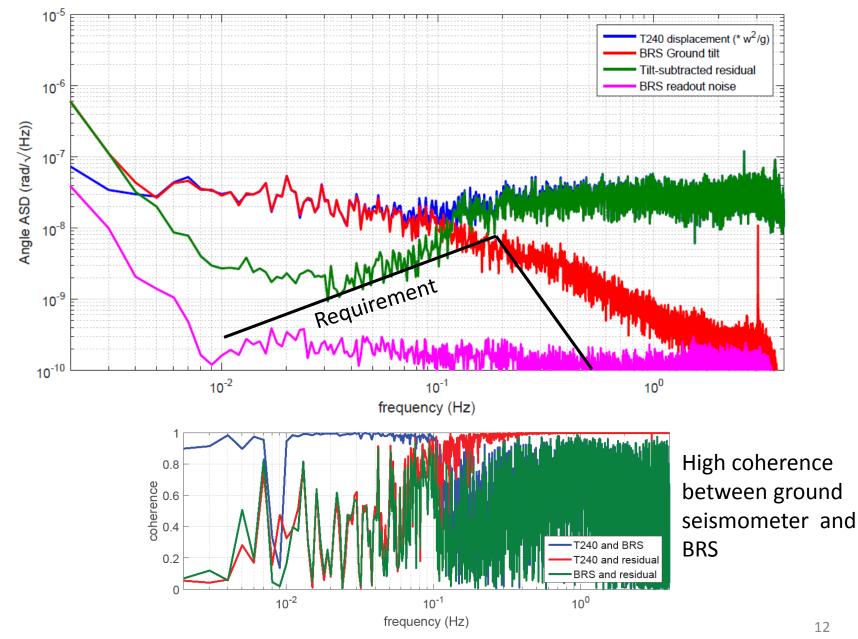


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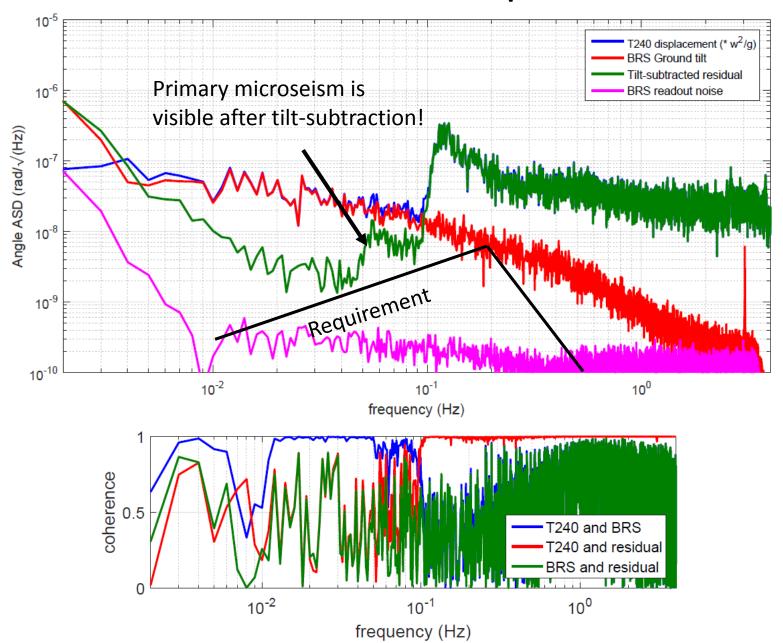
# Ground tilt during low wind speeds



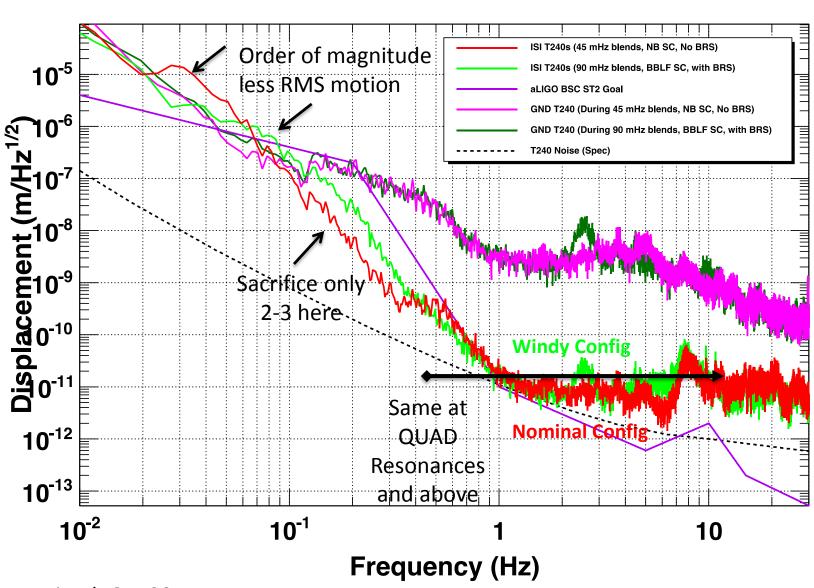
# Ground tilt during 20-30 mph winds



# Second example

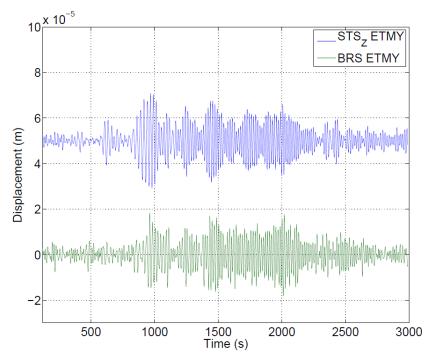


## Improving Isolation



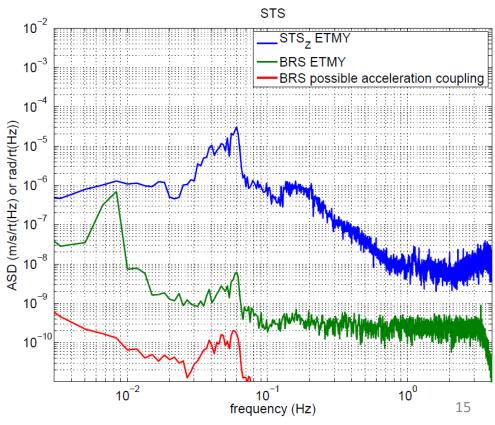
Source: J. Kissel, G1500475

# Ground tilt from M6.7 Earthquake near Australia



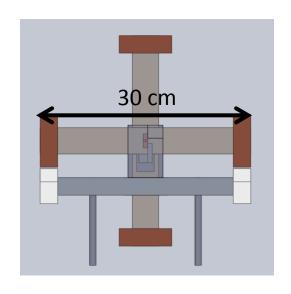
Band-passed time-series signals

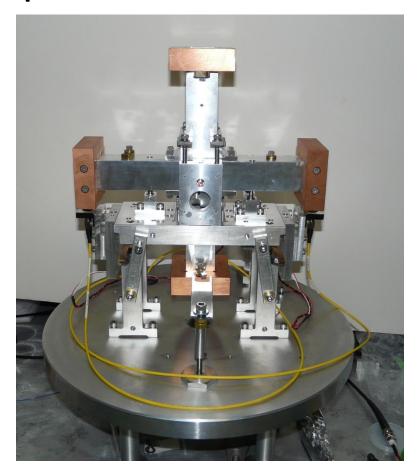
ASD of ground seismometer and BRS Data.



# Next Gen/Compact BRS

# Compact-BRS

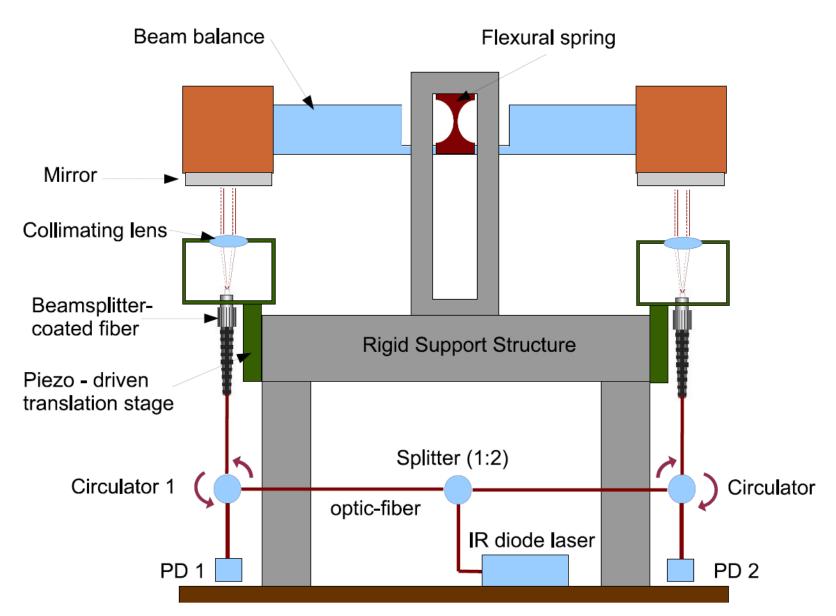




#### New features

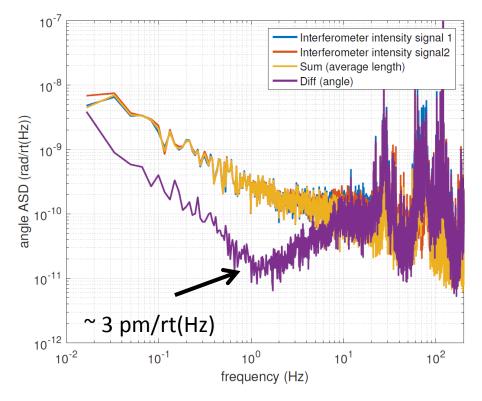
- 1. <u>Cross Shape</u> (~0 quadrupole moment) ensures first order insensitivity to gravity gradient noise.
- 2. New <u>compact</u> interferometric readout with ~10X <u>better sensitivity</u>.

#### Schematic



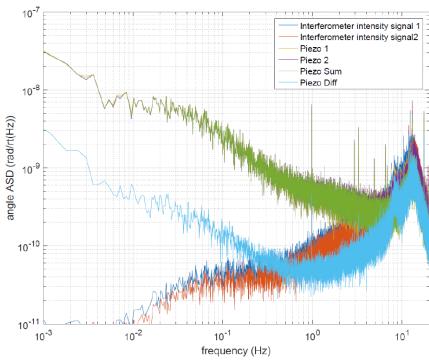
# Readout sensitivity

Individual interferometer limited by frequency noise. Differential signal (angle channel) reduced frequency noise by matched cavities.



Readout with "soft control" Piezo – stacks locked at ~ 5 mHz Tilt signal is in the PhotoDiodes

Readout with "hard control"
Piezo – stacks locked at ~ 5 Hz
Tilt Signal is in the drive



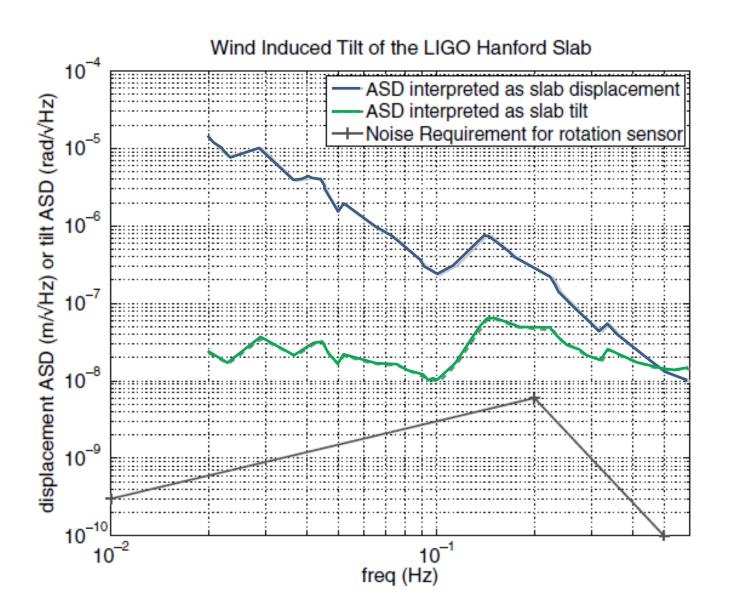
# Summary

- 1. BRS allows an unambiguous separation of tilt and horizontal acceleration between 0.01 to ~30 Hz.
- 2. BRS-1 and BRS-2 can improve total RMS differential motion of the Optical Platforms in LIGO and improve isolation, especially under windy conditions.
- 3. A new <u>compact</u> prototype (cBRS) is under development. When placed on the Optical Platforms, it can significantly boost angular isolation. It may also prove useful in Newtonian Noise measurement at 10 Hz.

# Thank you!

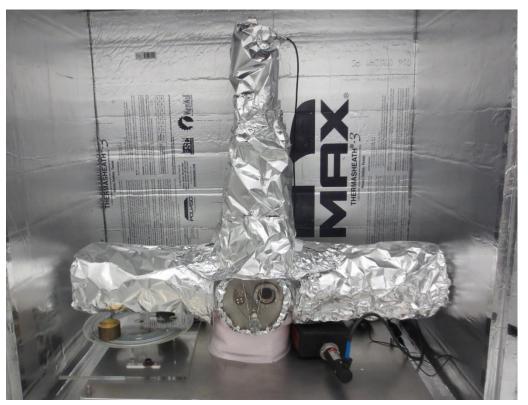
# Extra slides

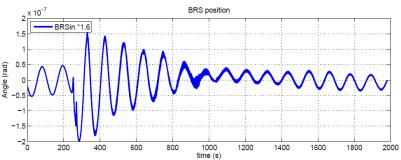
# Target sensitivity



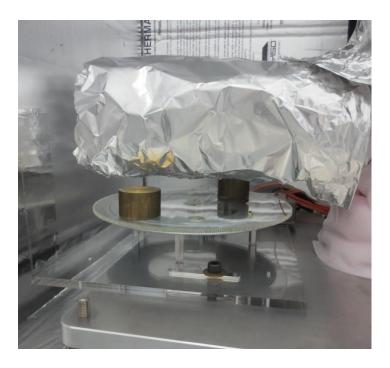
from
"Requirements
for a ground
rotation sensor
to improve
Advanced LIGO"
(B. Lantz et al.,
2009)

# Damper Installation





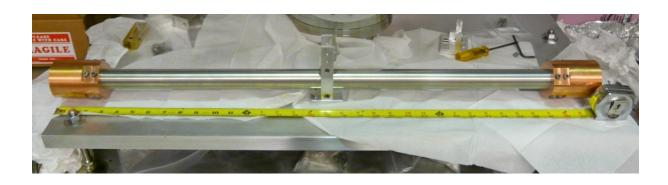
#### gravitational feedback system!



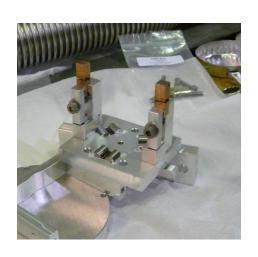
A turn-table applies a gravitational torque on the balance to actively damp the beam-balance when amplitude exceeds a threshold.

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# **BRS-2** Assembly pictures



Assembled beam and weights



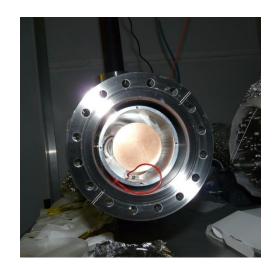




**Flexures** 

Integrated with vacuum chamber

# **BRS-2** Assembly pictures



**Capacitor plates** 



**Beckhoff modules** 



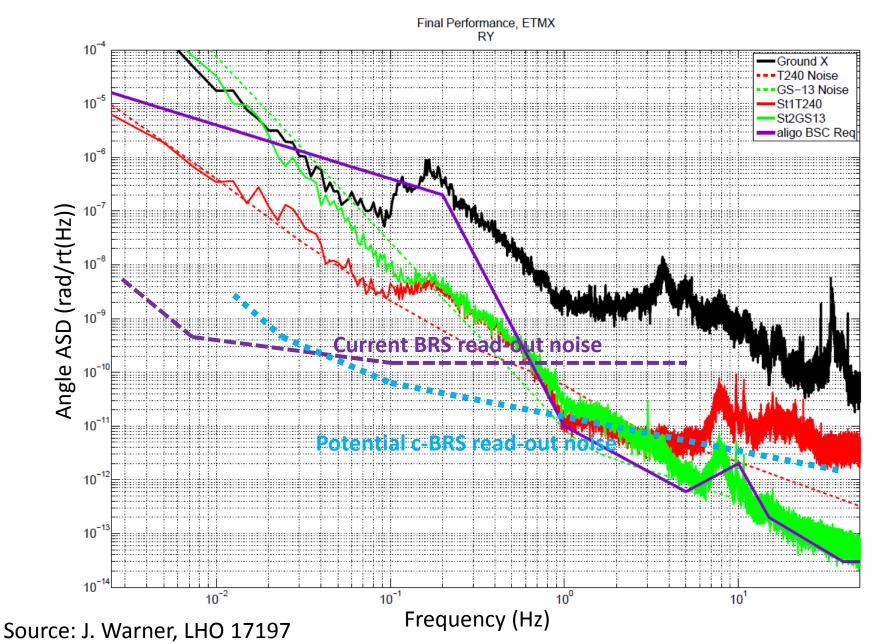


Basler CCD camera



Foam box

# Sensitivity comparison



### cBRS Tilt data

