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# Technological Advances: Seismic/Suspensions

Giles Hammond

Institute for Gravitational Research

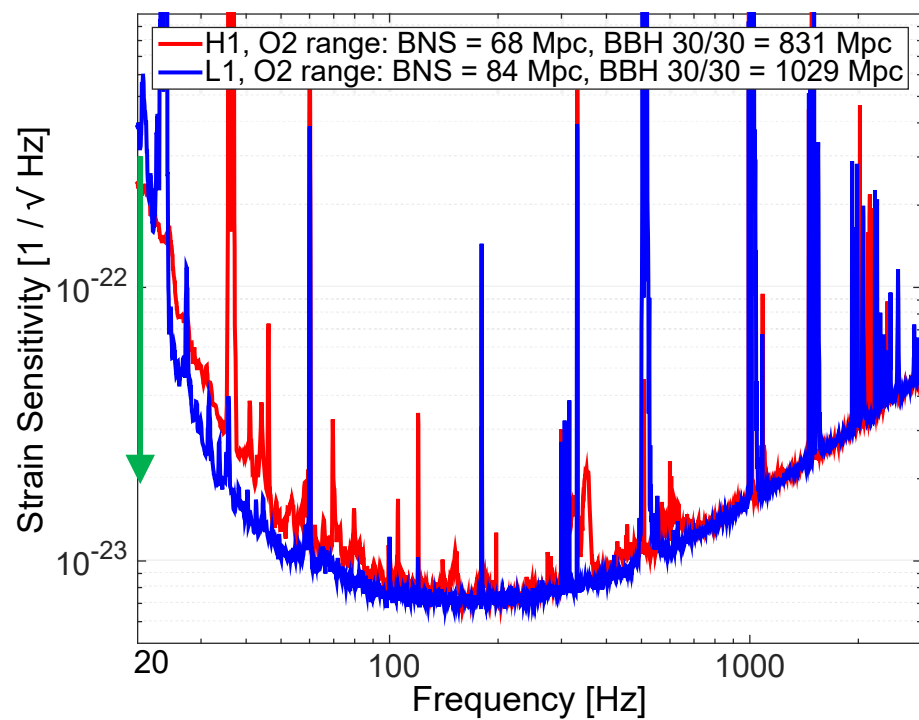
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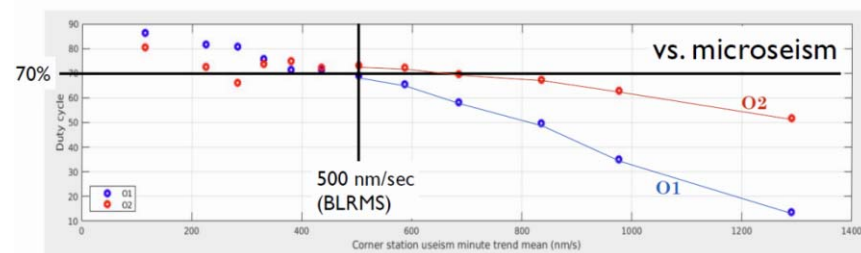
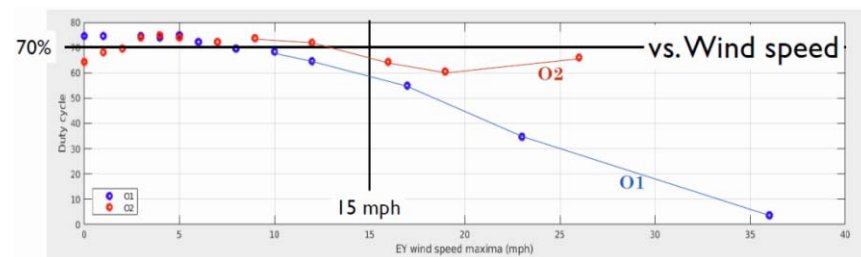
LIGO G1701267

DAWN III meeting, Syracuse NY, July 6<sup>th</sup>-7<sup>th</sup> 2017

## Instrument Noise in O2



- Good progress on commissioning
- Improved duty cycle during excess noise (wind/earthquakes) with tilt sensors



## Highlights



- Well engineered solution to install/commission robust monolithic suspensions
- Bounce/Roll mode dampers
- Collaboration with KAGRA via the ELITES programme



- To realise aLIGO design sensitivity will require coordinated effort to address low frequency noise
  - Scattered light, lots of work done during O1/O2 break at LLO. Better baffling and absorbers
  - Length-Angle decoupling (cross-coupling in suspensions/actuators)
  - Implementation/operation of seismic subtraction arrays (aLIGO/AdVIRGO)
  - Characterisation of thermal/magnetic effects with seismic/suspensions (violin modes track temperature variation)
- Switching out ETMs to operate with improved coatings
- End reaction masses with annular cutouts to reduce squeeze film damping
- Charging; seems to be a reliable method for in-situ charge measurement and mitigation



- **A+**
  - Newtonian noise subtraction using seismometer arrays: optimising the array
  - Make use of current ISI, but implement tilt sensors
  - Larger diameter beamsplitter suspension
  - Monolithic suspensions with thinner fibres (for bounce/roll/violin modes)
  - Stray light modelling & analysis
  - Laser damage thresholds & contamination control
- **3<sup>rd</sup> Generation (LIGO Voyager/Cosmic Explorer/Einstein Telescope)**

## Newtonian Noise

- Site selection (surface/underground/existing infrastructure), optimum location for multiple 3<sup>rd</sup> generation detectors
- Modelling of NN from atmospheric sources (infrasound) and structures
- Design/modification of site infrastructure
- Deployment of borehole seismometers (for 3D array)
- Low frequency torsion pendula (e.g. TOBA/TORPEDO)



- **3<sup>rd</sup> Generation cont ....**

## Seismic Noise

- Sensors with improved tilt-horizontal coupling and noise performance
- Study of upconverted noise e.g. crackling noise
- Development of new control strategies to provide best IFO performance
- Seismic platform interferometer?

## Suspensions

- Large cryogenic silicon mirrors (150-200kg) with low absorption
- Large silica mirrors (150kg-200kg)
- Cooling strategies for 123K and 20K & temperature control
- Silicon or silica springs in the final suspension stage => protective coatings
- Low temperature sensors/actuators, sensing more DOF's
- Handling/fabrication of heavy suspensions, re-engineering SUS cage
- Cooldown time, how to commission a cryogenic IFO



- With LIGO India, we need to seize the opportunity to implement the detector with A+ upgrades (researcher exchange between India and LVC)
- Build on some of the successful collaborative efforts
  - Seismic array monitoring for NN subtraction (aLIGO/AdVIRGO)
  - Collaborative activities with KAGRA (initiated under the ELITES program). Learn from the cryogenic work being undertaken in Japan
  - Need to ensure good engagement with the worldwide prototype network (Caltech 40m, AEI10m, Glasgow 10m, Gingin)
  - Opportunities for sensor development in the area of interferometric readout and accelerometers (SWG/Nikhef), and low frequency torsion pendula (KAGRA, ANU)